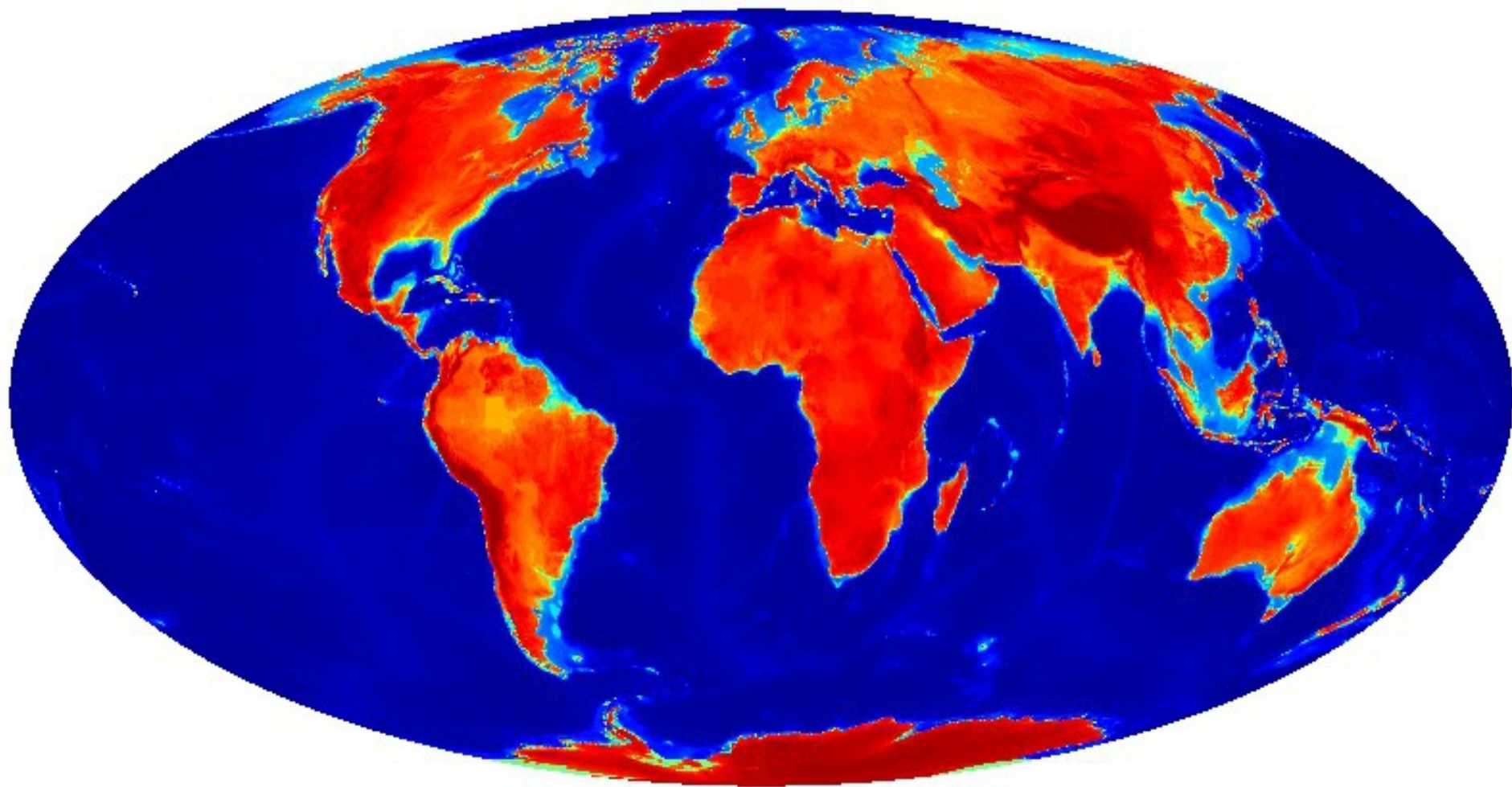


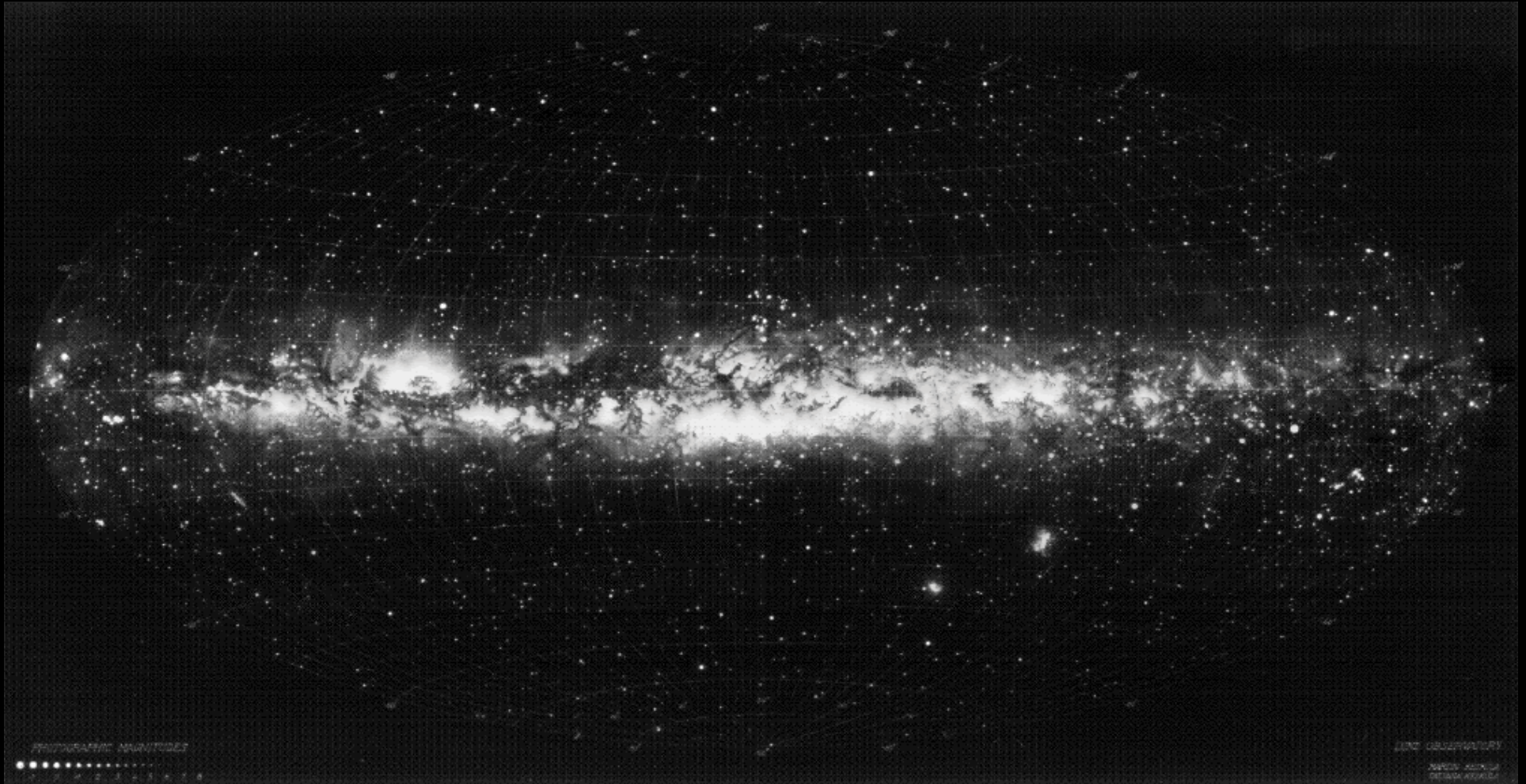
# All from Nothing: the structuring of our Universe

*Simon White*  
*Max Planck Institute for Astrophysics*





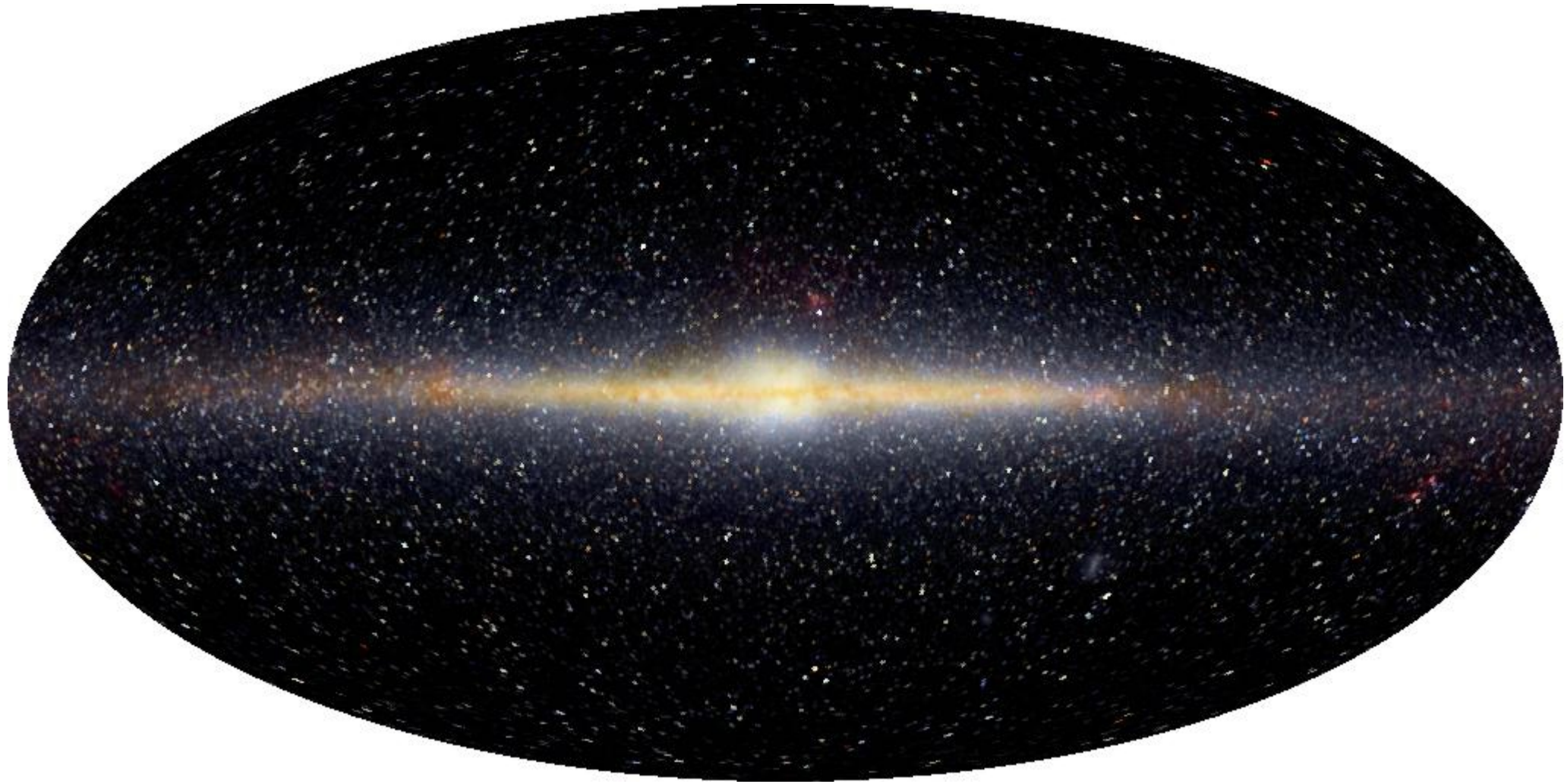
# Star map of the whole sky



..out to 10,000 light-years



# COBE's infrared map of the whole sky



..out to 30,000 light-years



# The Andromeda Nebula: our biggest neighbour



..out to 2,000,000 light-years



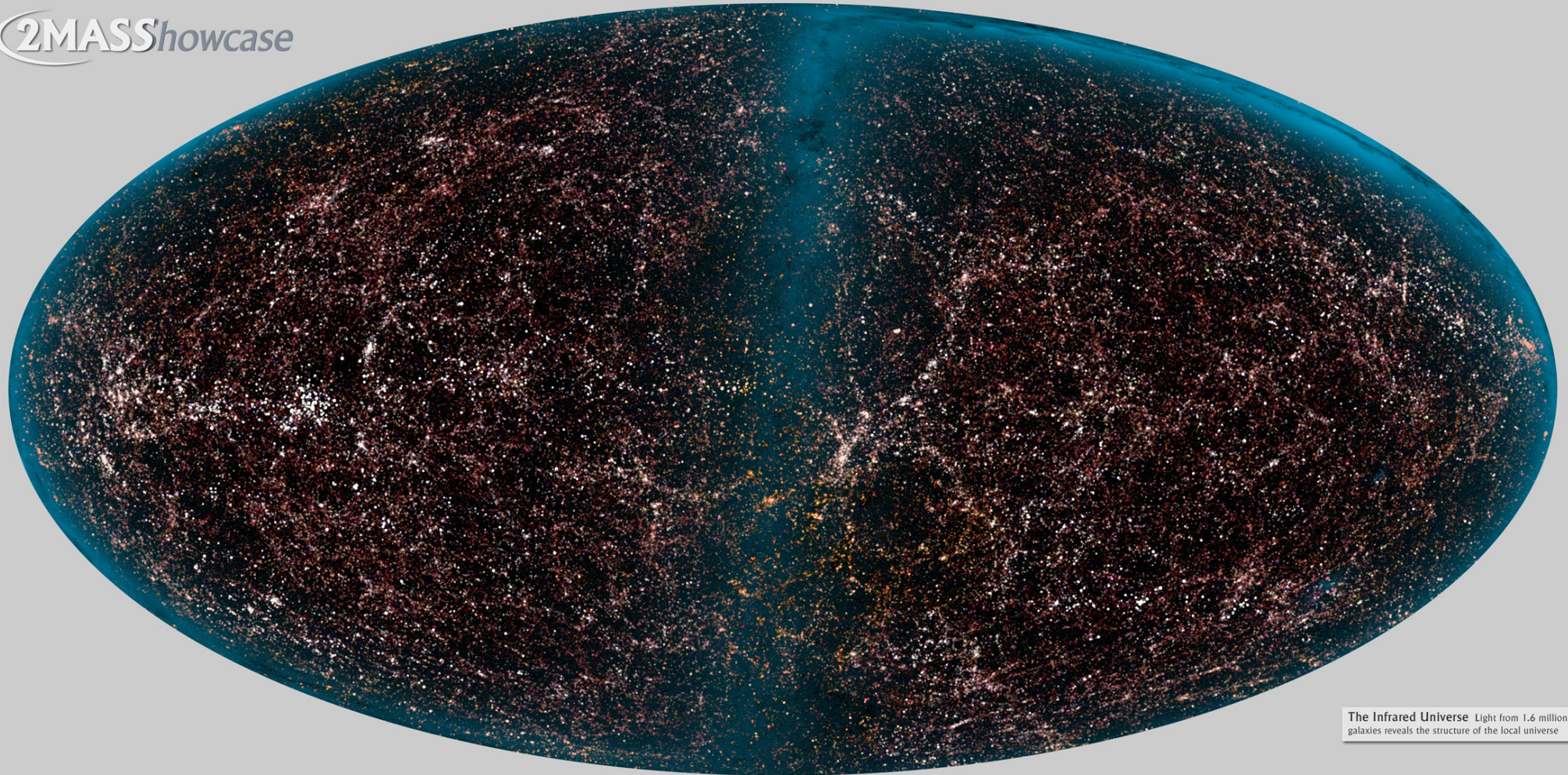
**NGC 4414 -- another  
Galaxy like our own**





# A galaxy map of the whole sky

2MASS*showcase*



**The Infrared Universe** Light from 1.6 million galaxies reveals the structure of the local universe

Two Micron All Sky Survey Image Mosaic: Infrared Processing and Analysis Center/Caltech & University of Massachusetts

..out 1,000,000,000 light-years





The deepest  
optical image  
ever made

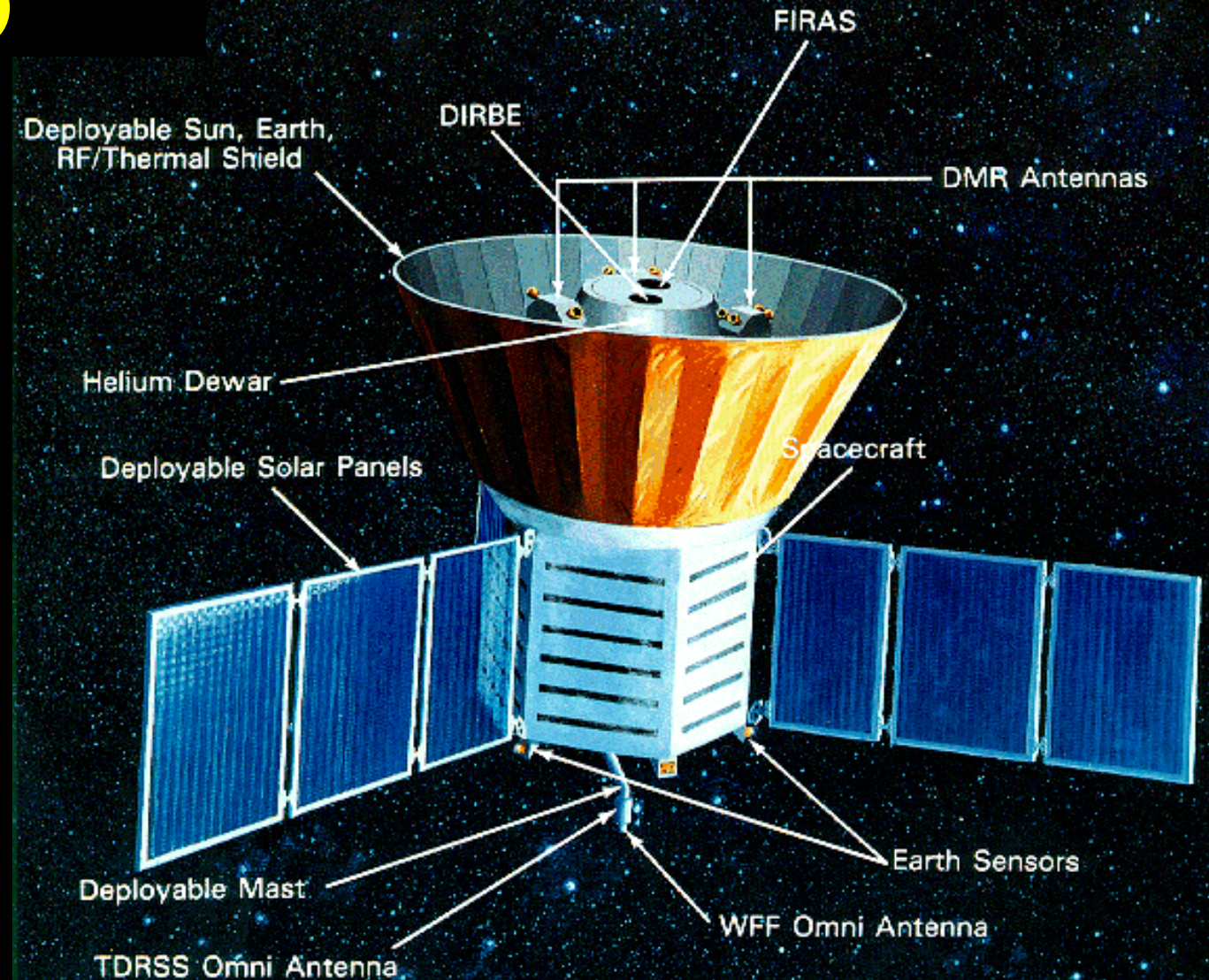
A 300 hour  
exposure with  
the Hubble  
Space Telescope

..out to more than 30,000,000,000 light-years



# The COBE Satellite (1989 - 1993)

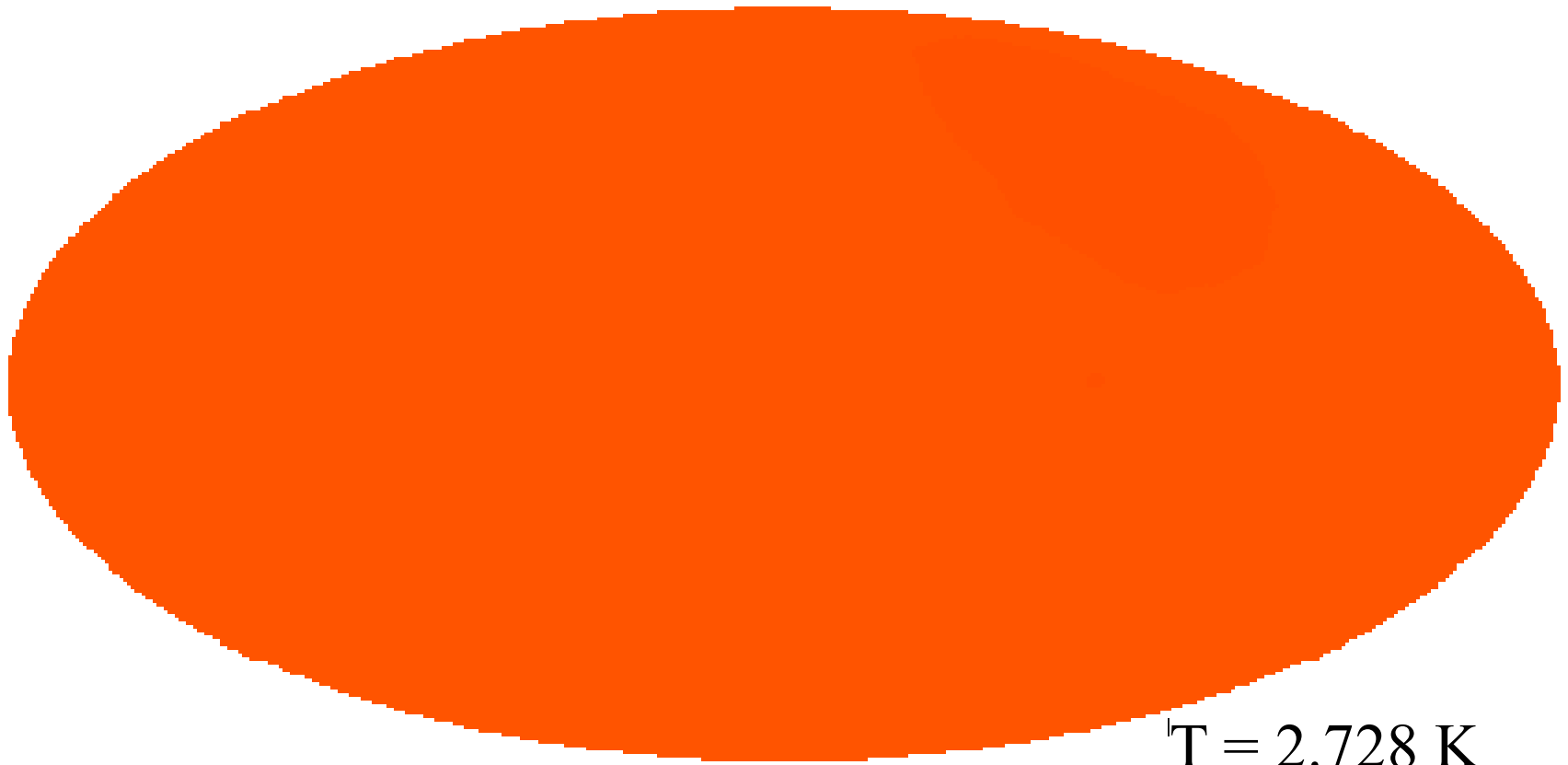
- Two Instruments mapped the whole sky at microwave and infra-red wavelengths
- One instrument took a precise spectrum of the sky at microwave wavelengths



Nobel Prize 2006



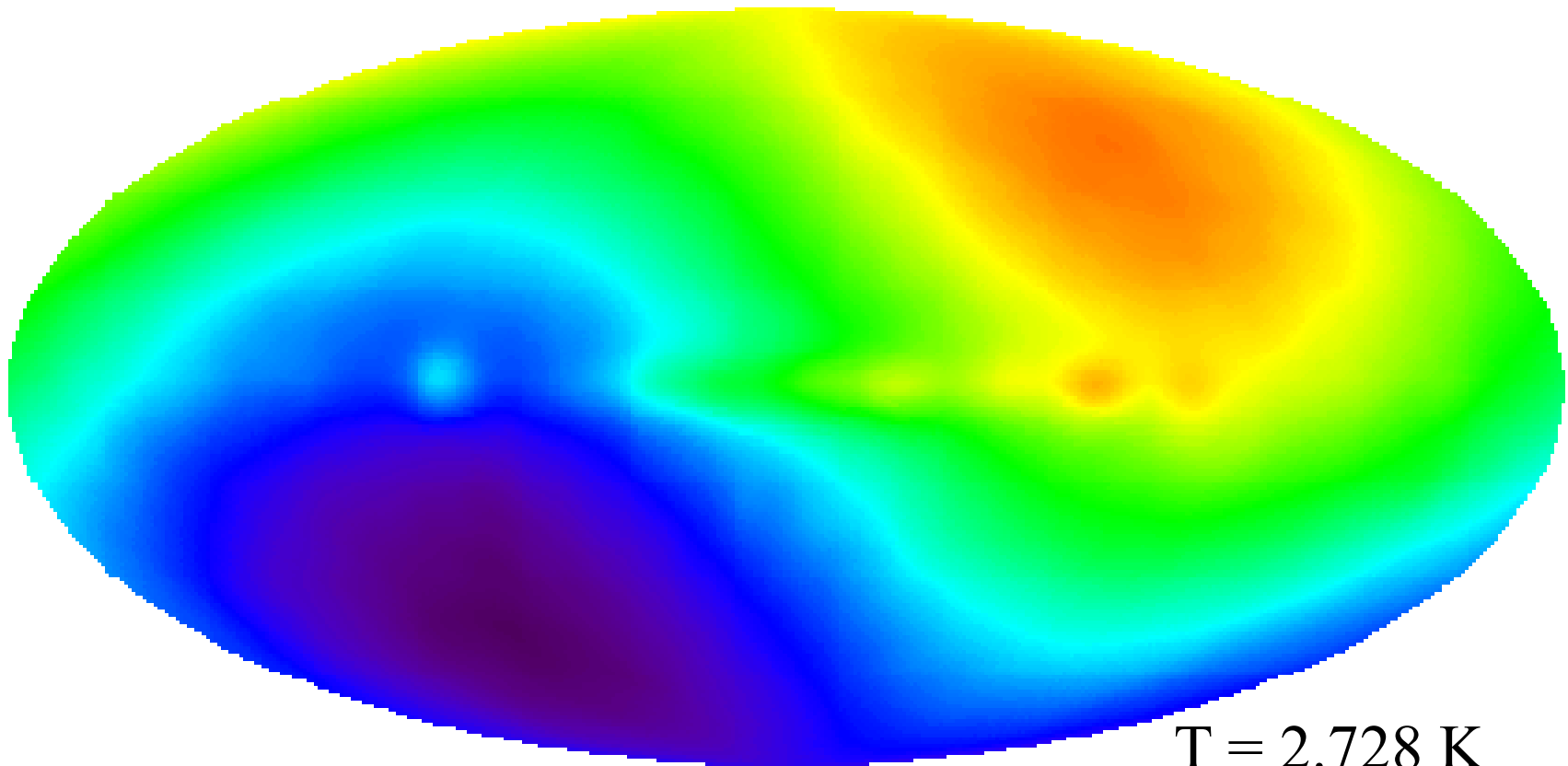
# COBE's temperature map of the whole sky



$T = 2.728 \text{ K}$   
 $\Delta T = 0.1 \text{ K}$



# COBE's temperature map of the whole sky

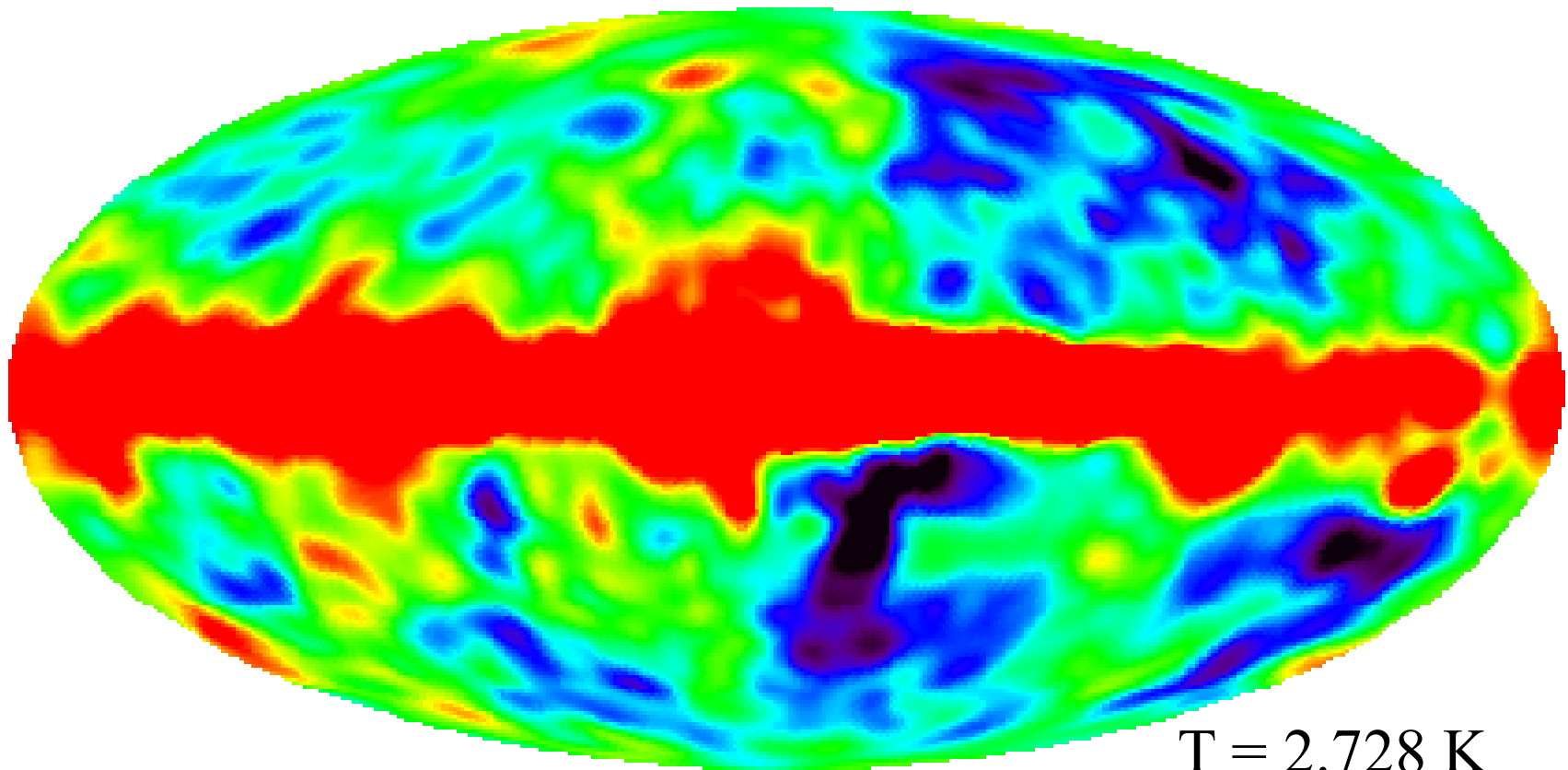


$$T = 2.728 \text{ K}$$

$$\Delta T = 0.0034 \text{ K}$$



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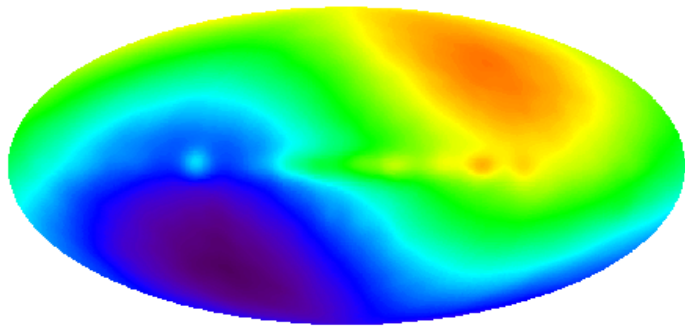


$$T = 2.728 \text{ K}$$

$$\Delta T = 0.00002 \text{ K}$$

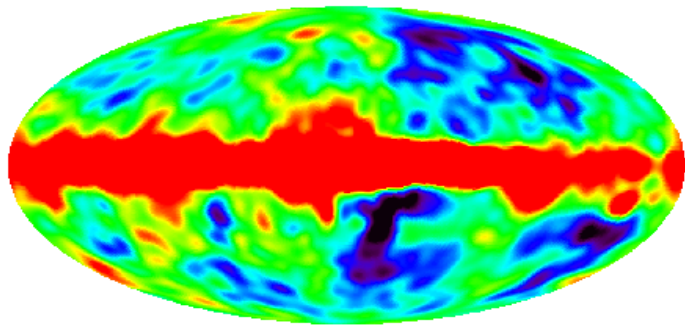


# Structure in the COBE Map

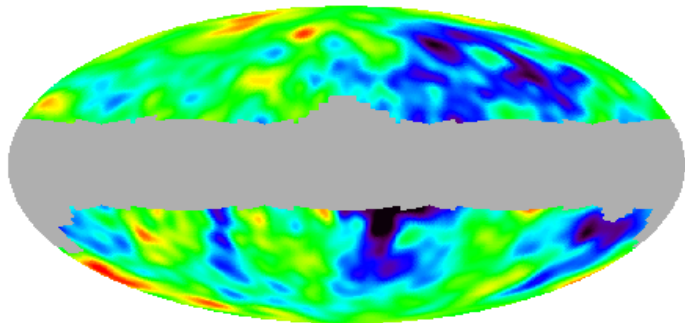


- One side of the sky is cold, the other hot  
our Galaxy's motion through the cosmos

→  $V_{\text{Milchstrasse}} = 600 \text{ km/s}$



- Radiation from dust and gas in our own Galaxy



- Structure in the microwave background itself



# Structure in the microwave background

Where is the structure?



# Structure in the microwave background

Where is the structure?

In cosmic clouds at the far edge of the visible Universe



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What are we seeing?



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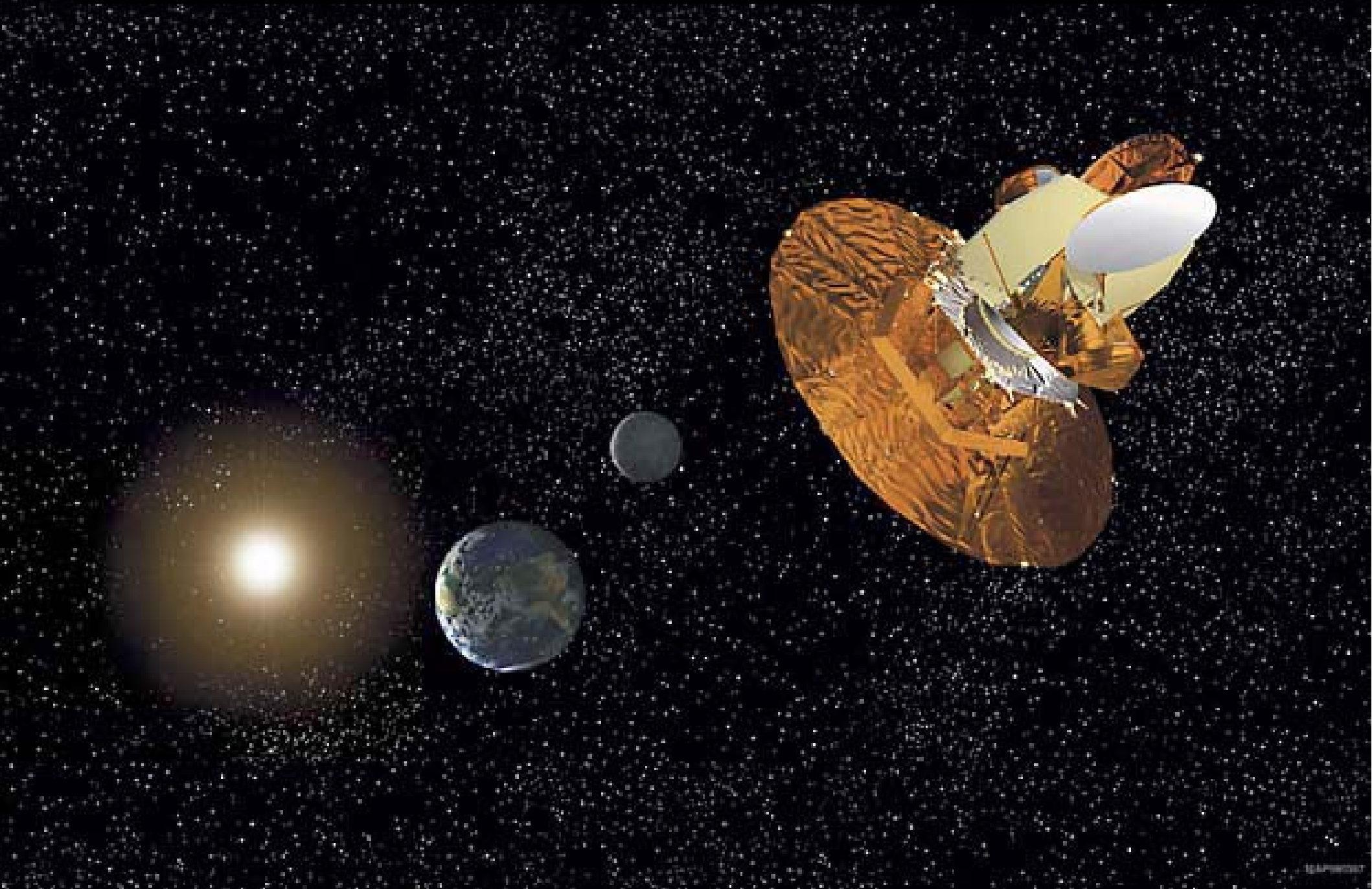
When was this structure created?

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What has this structure become?

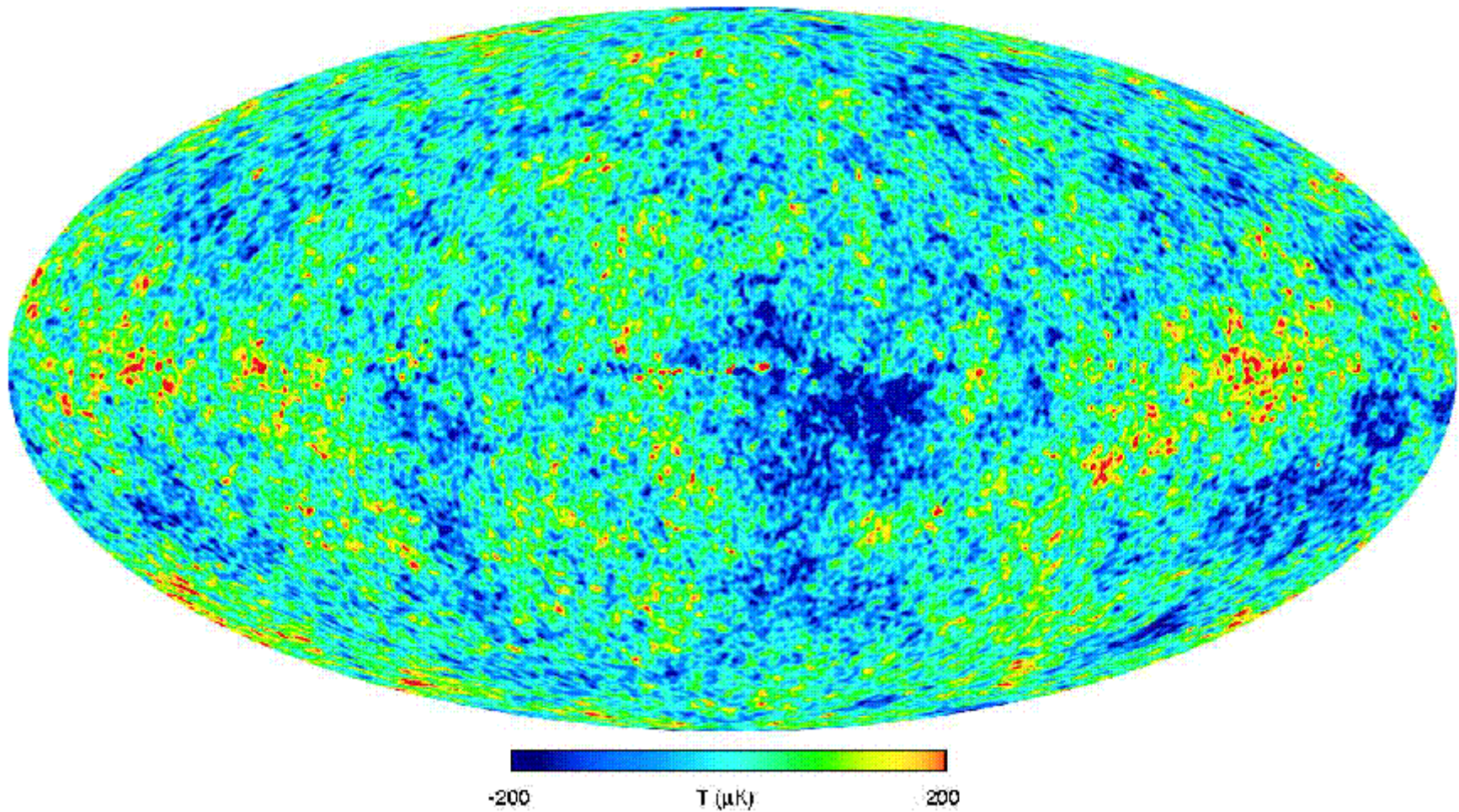
Everything we see around us (galaxies, stars, planets, people...)

# The *WMAP* satellite at Lagrange point L2





# *WMAP's* map of the whole sky



Bennett et al 2003

# What do we learn from these structures?

The pattern of structure is influenced by three things:

- the geometry of the Universe

- the content of the Universe

- the process which created the structure



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  - ordinary atomic (baryonic) matter
  - Dark (non-baryonic) Matter
  - Dark Energy
- the process which created the structure

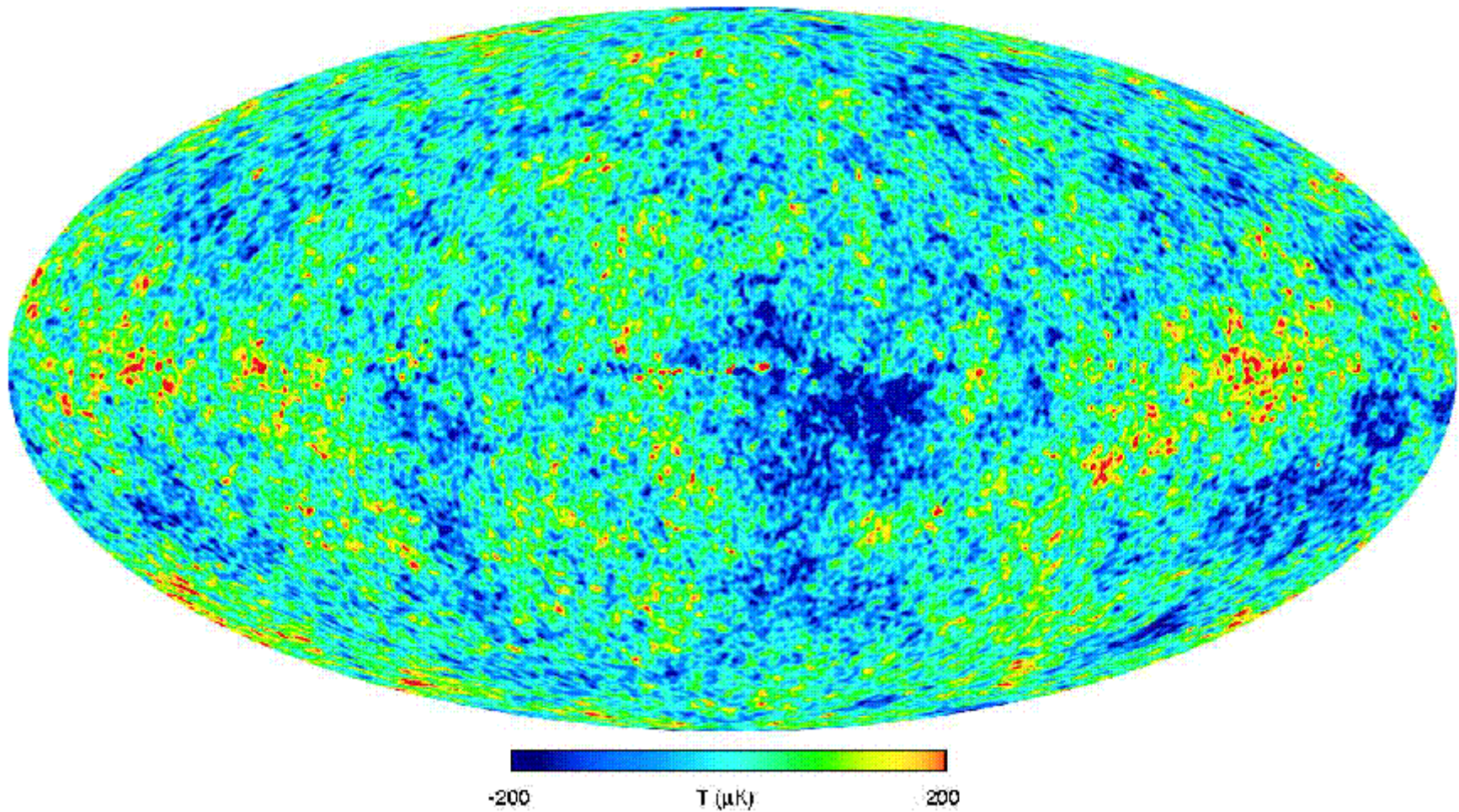
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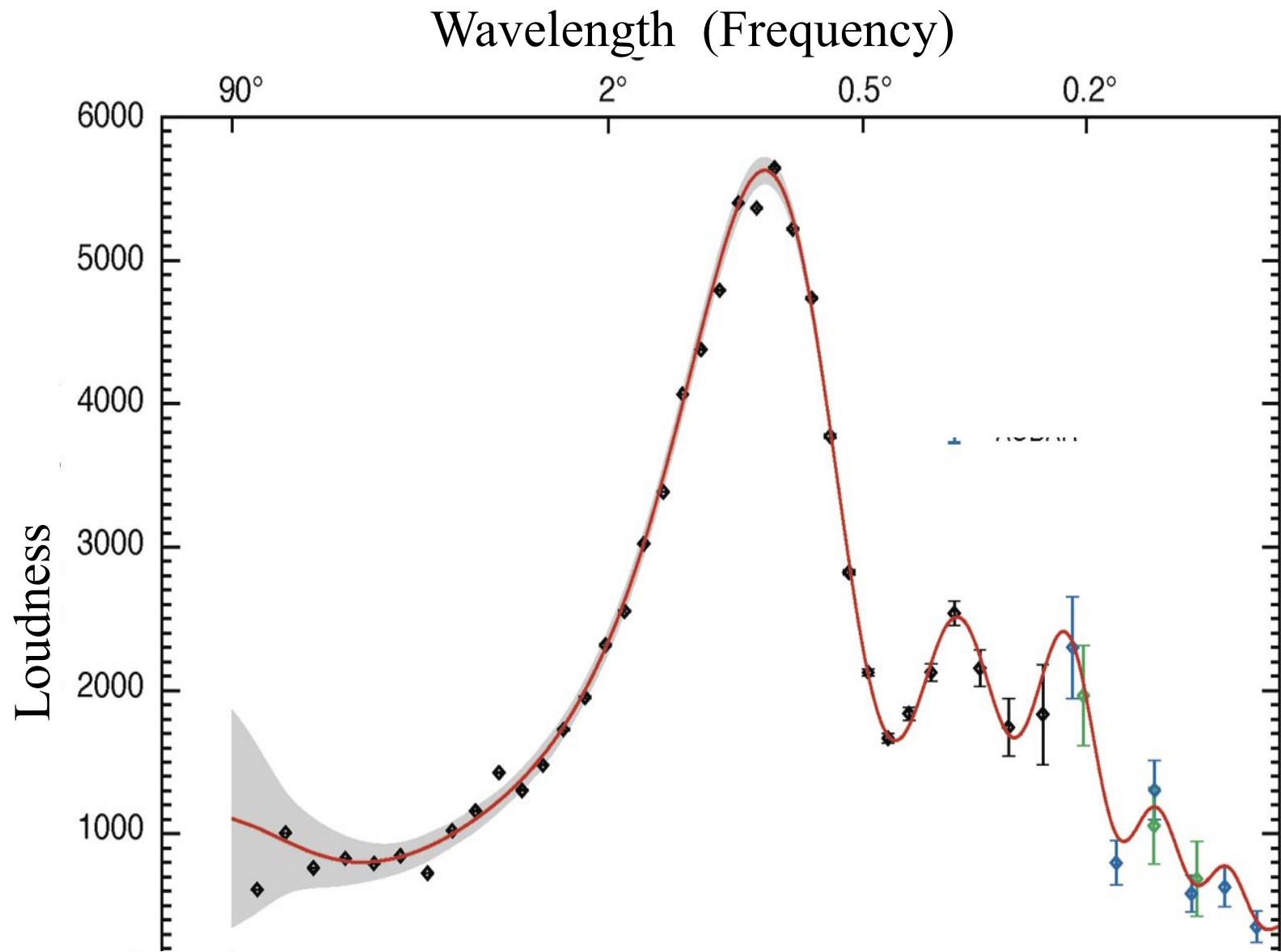
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  - Dark Energy
- the process which created the structure
  - The earliest instants of creation



# *WMAP's* map of the whole sky



Bennett et al 2003



The harmonic content of sound waves in the cosmic clouds:  
*WMAP* measurements compared to a theoretical prediction

# What did we learn from *WMAP*?

- Our Universe is flat – its geometry follows Euclid's laws



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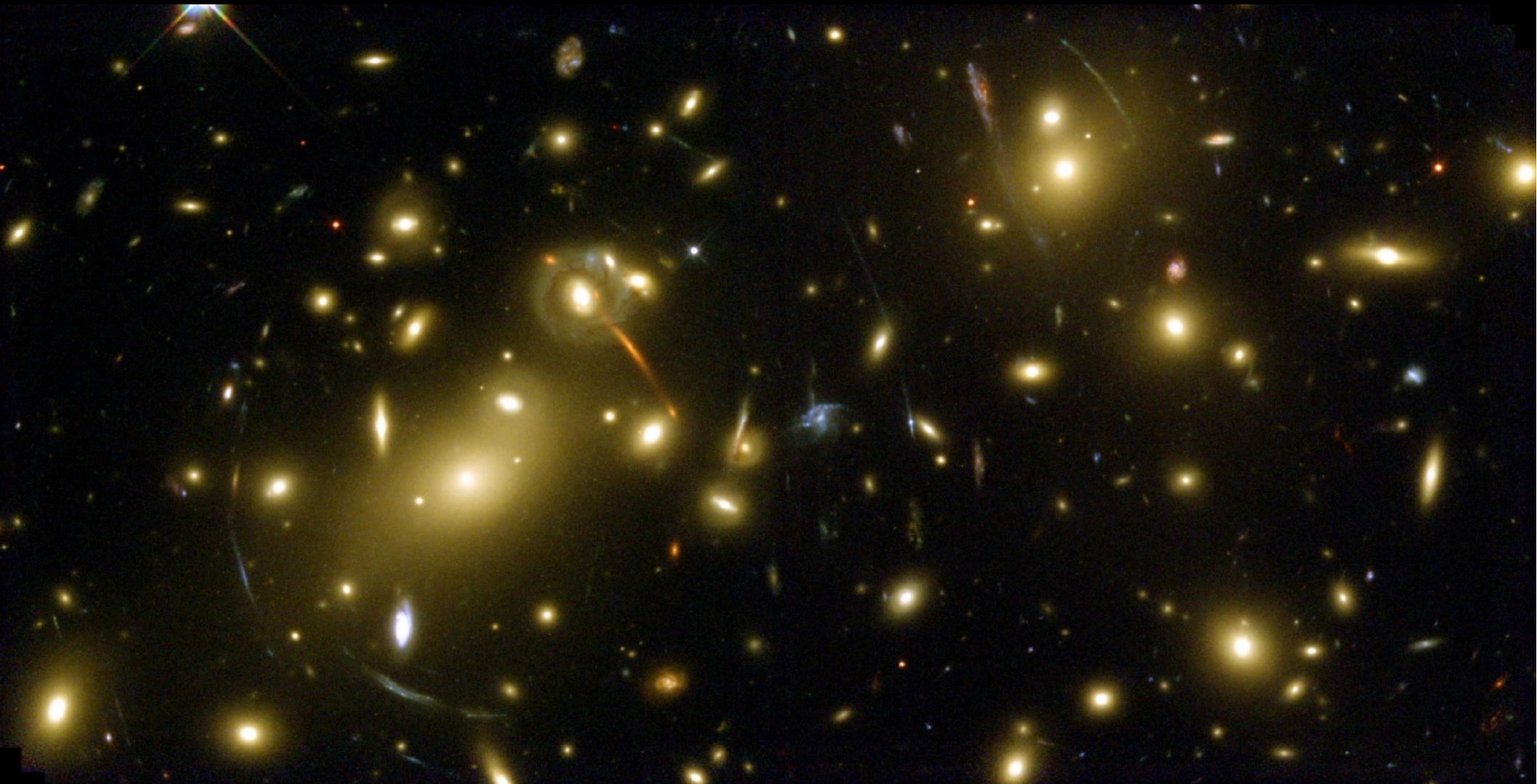
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(already “seen” through gravitational lensing effects)

# A galaxy cluster acting as a gravitational lens

Abell 2218

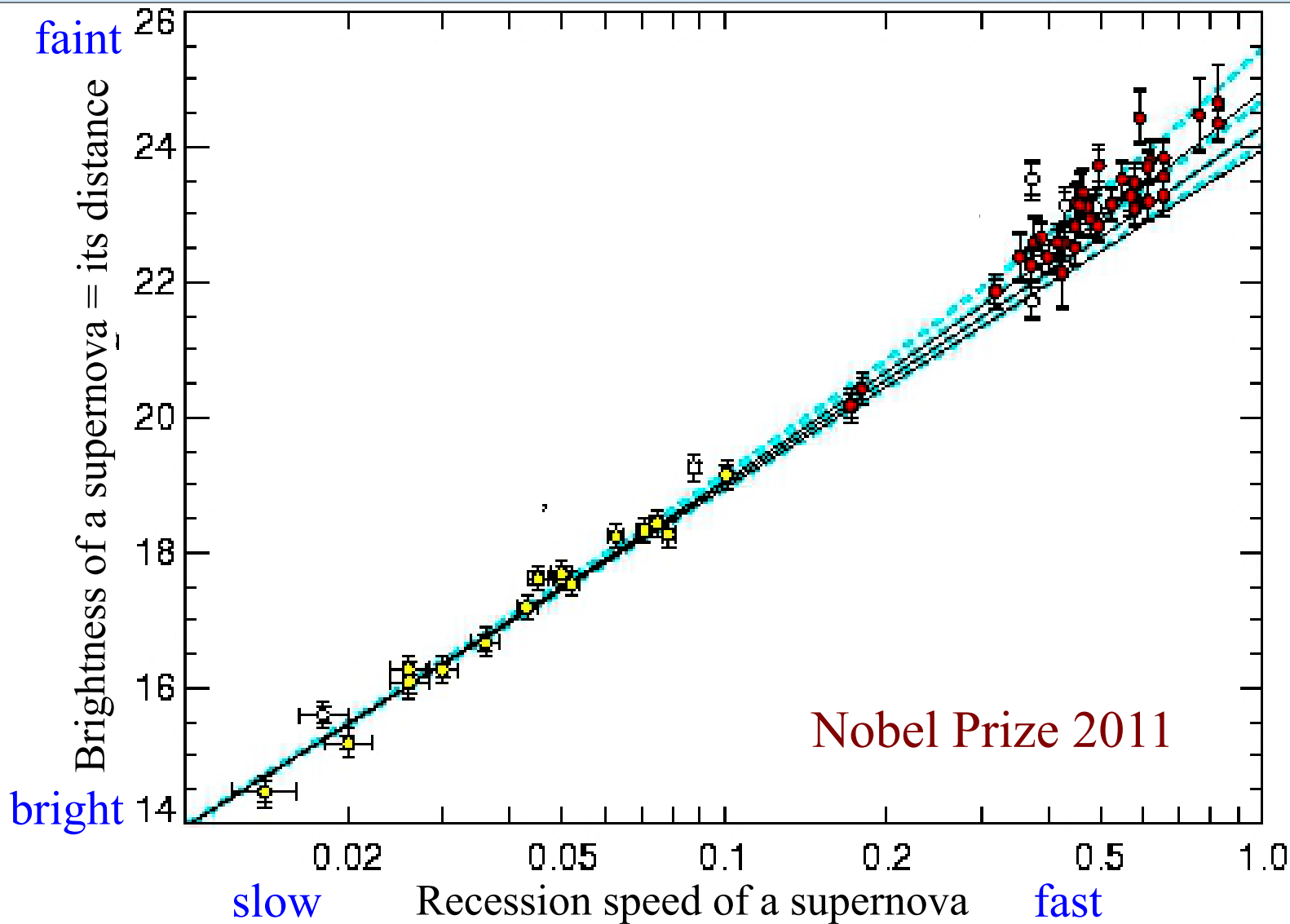




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# Dark Energy drives cosmic acceleration



Distant supernovae, seen when the Universe was younger, are fainter than expected —→ the Universe expands **faster** today than in the past

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Everything formed from the Vacuum!

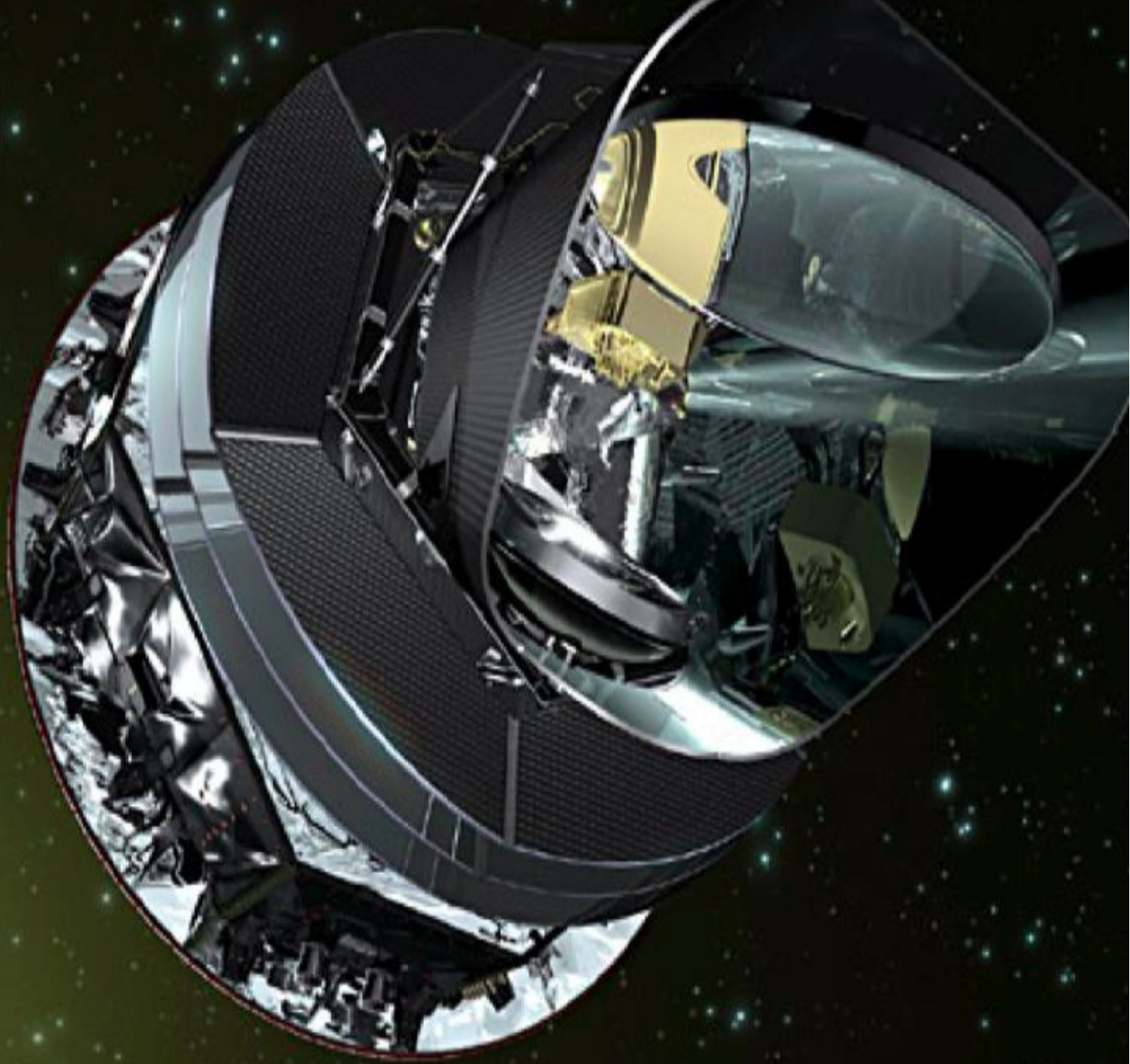
# Launch of the *Planck* satellite



Kourou, French Guyana: May 14, 2009



# *Planck* at L2





# The nine *Planck* maps

30 GHz

44 GHz

70 GHz

100 GHz

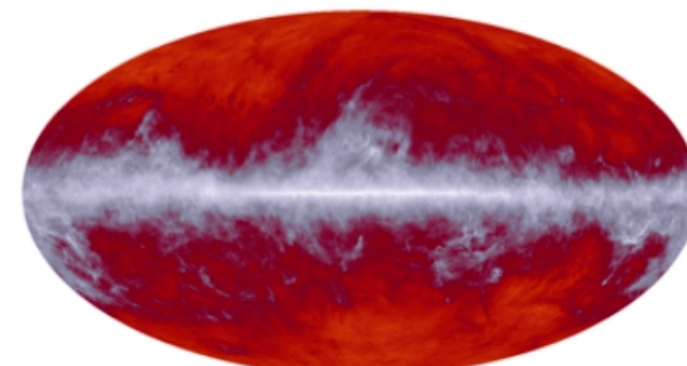
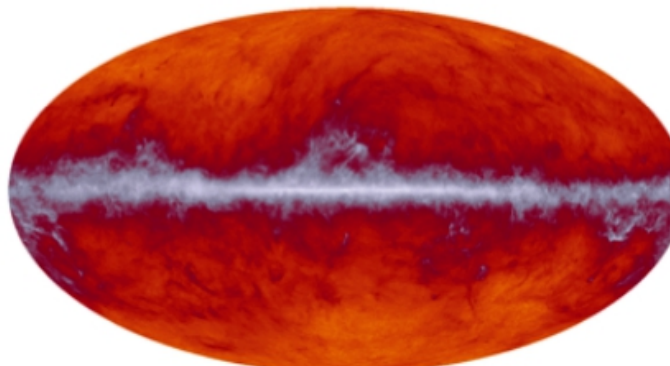
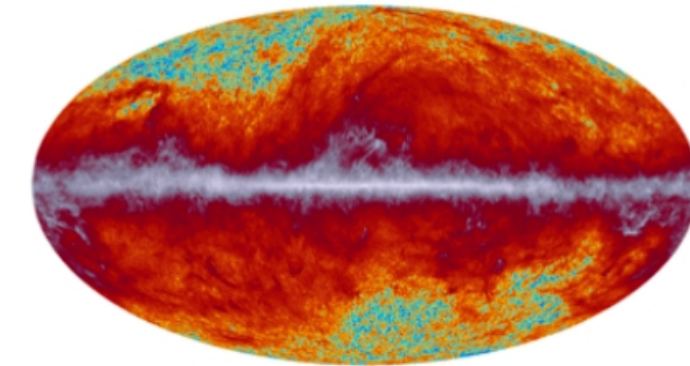
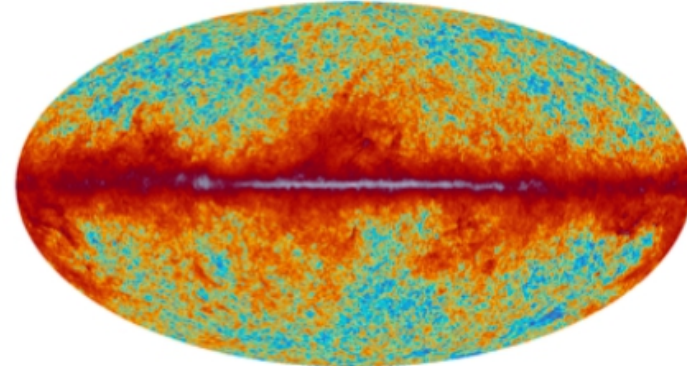
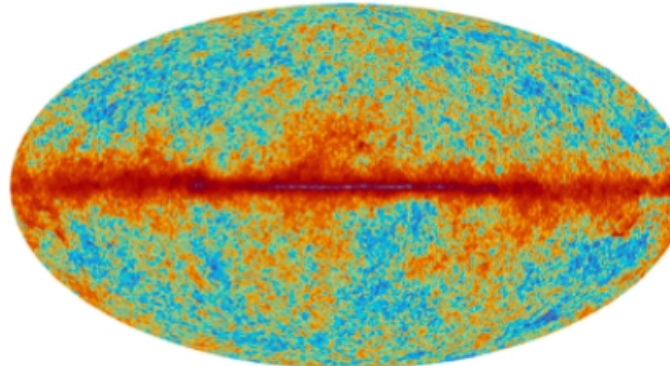
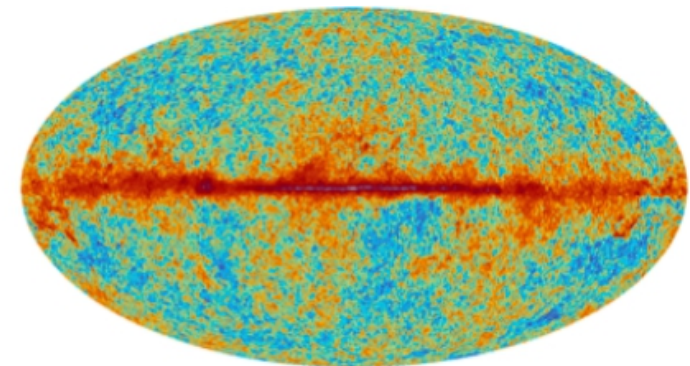
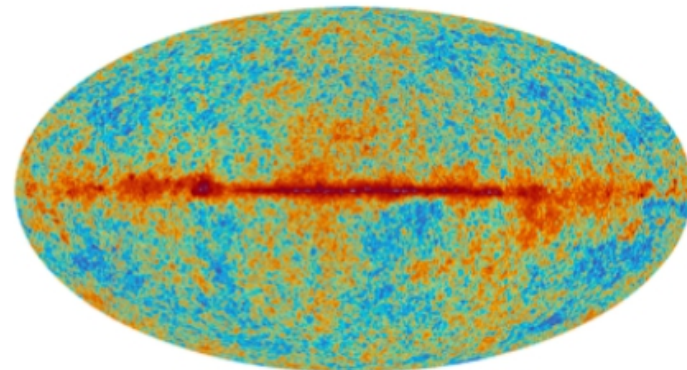
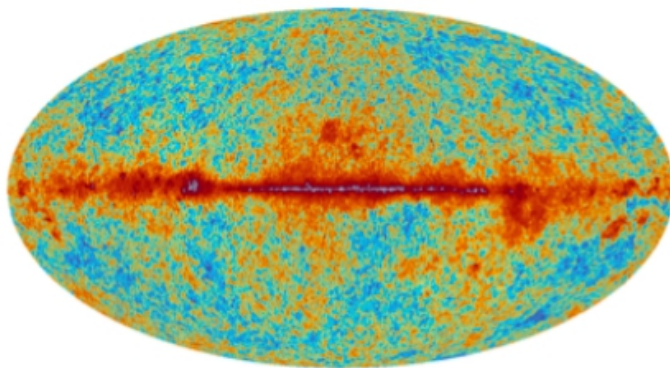
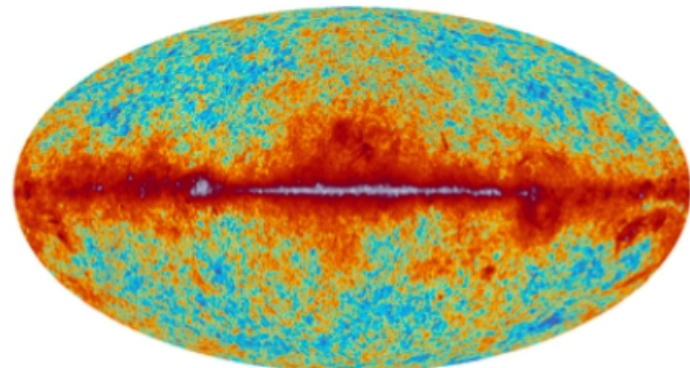
143 GHz

217 GHz

353 GHz

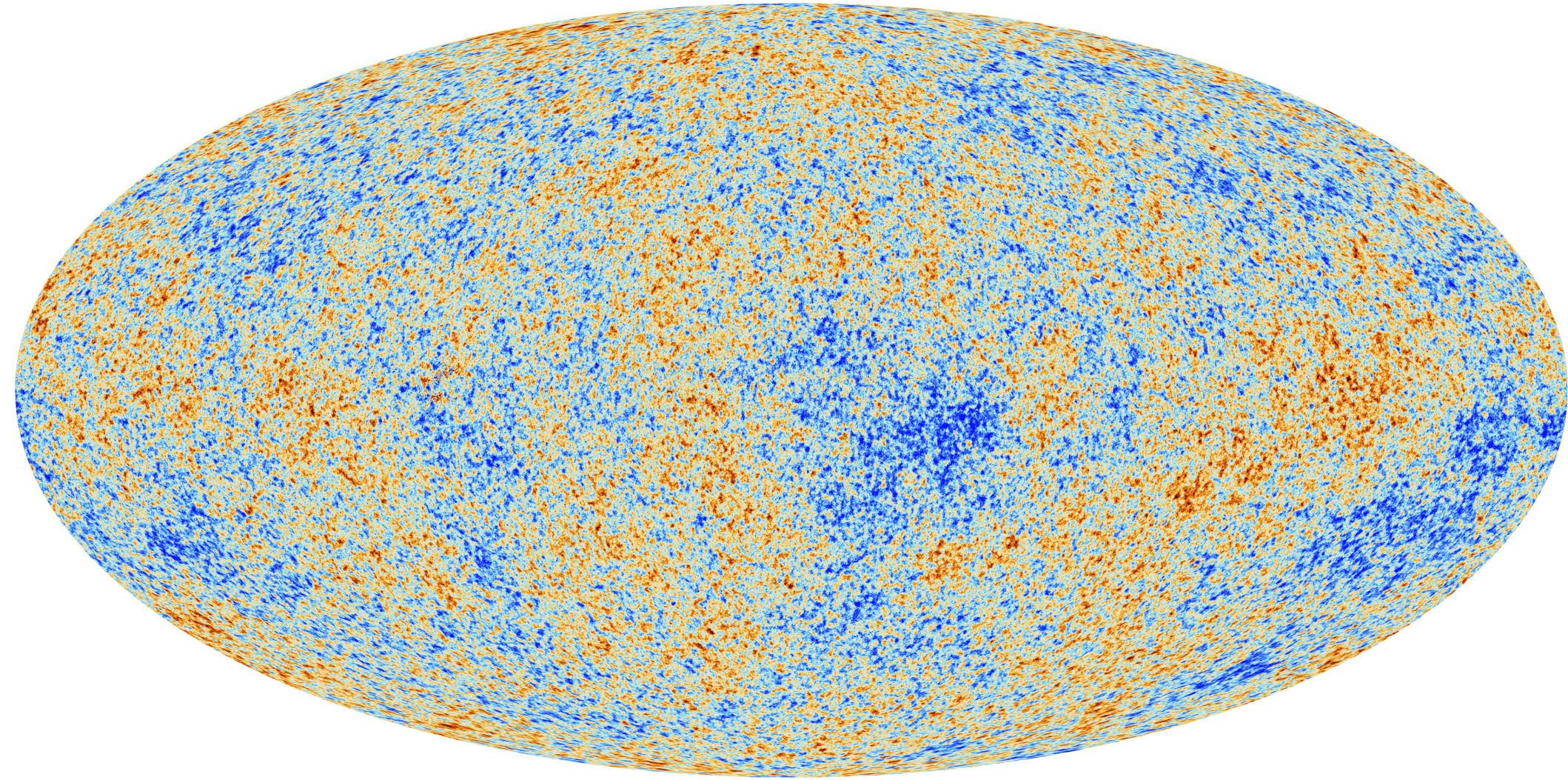
545 GHz

857 GHz





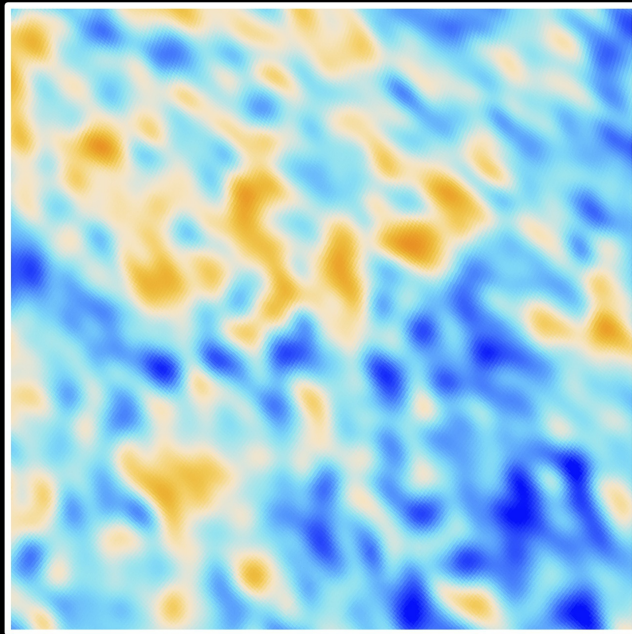
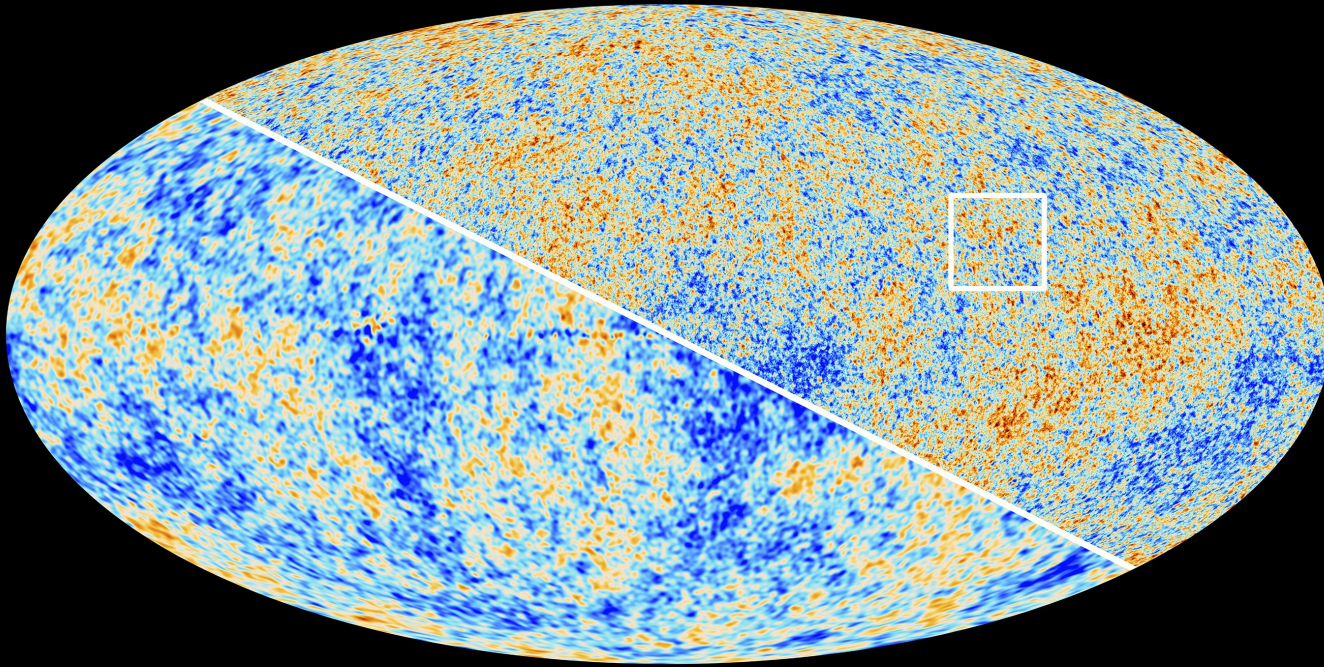
# The *Planck* map of the microwave background



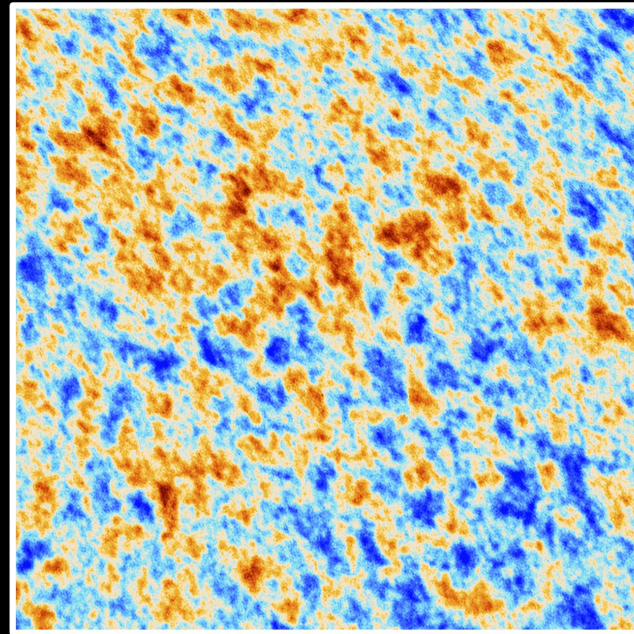
An image of the boundary of the observable Universe



# *Planck and WMAP in comparison*



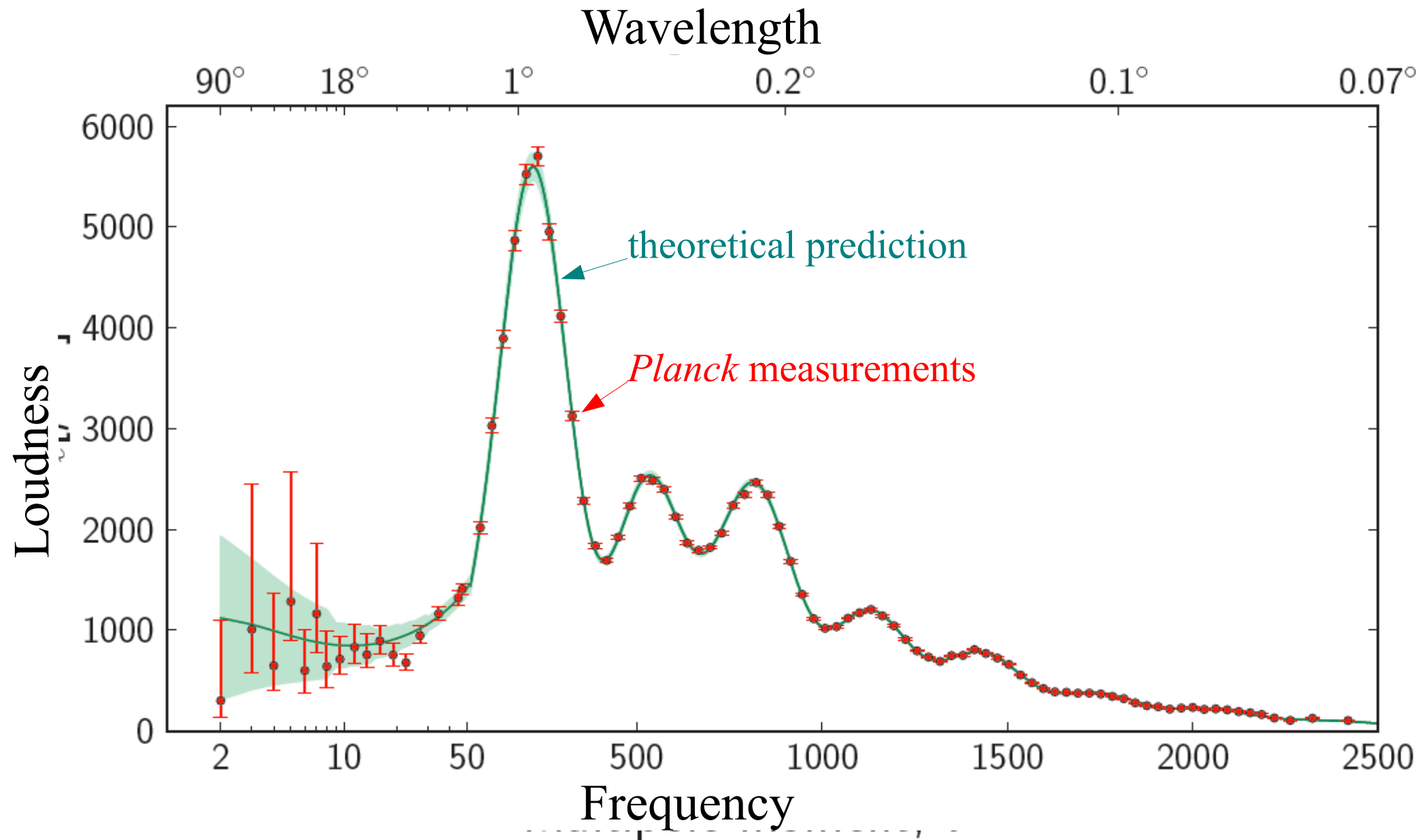
*WMAP*

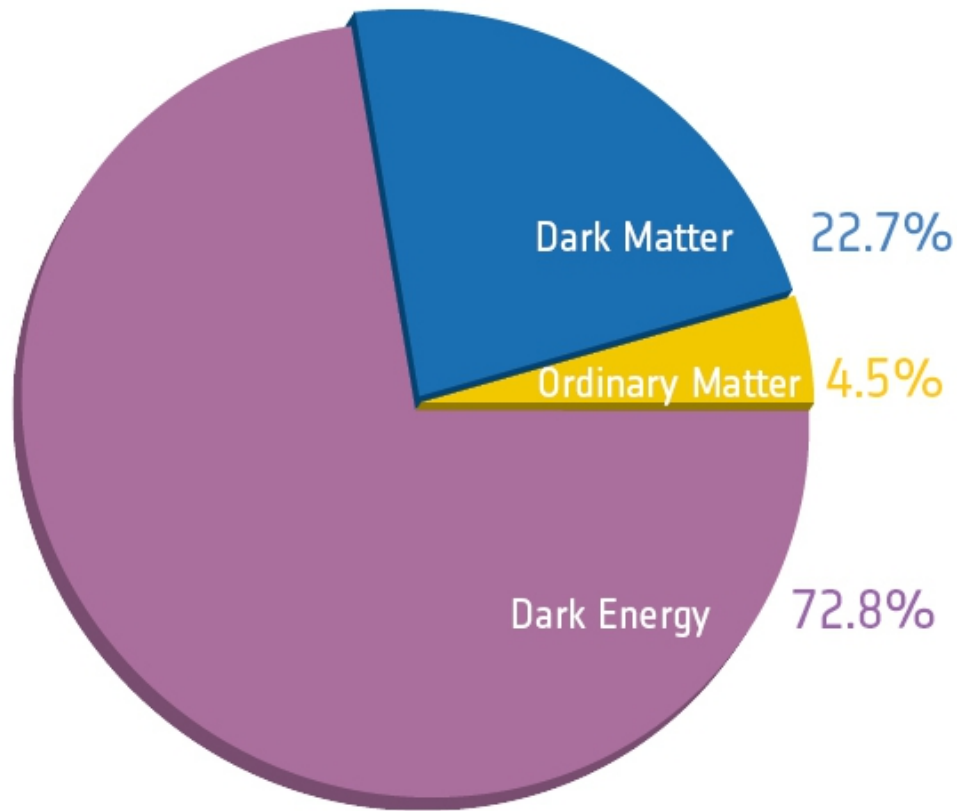


*Planck*

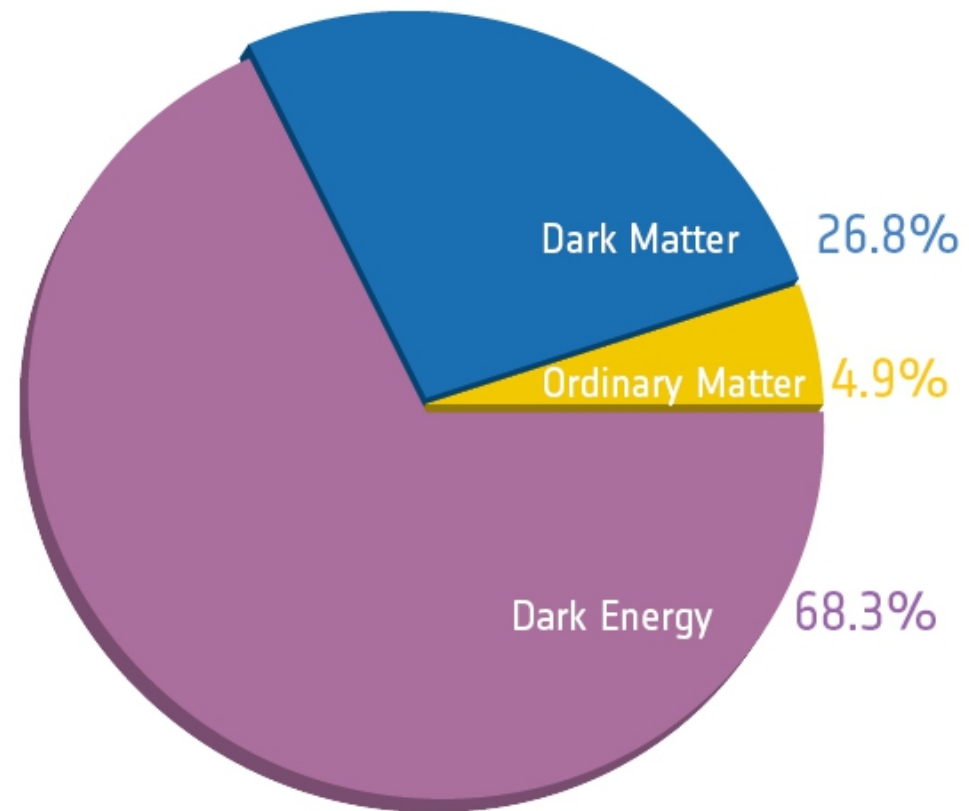


# Sound content of the cosmic clouds according to *Planck*





before *Planck*



after *Planck*

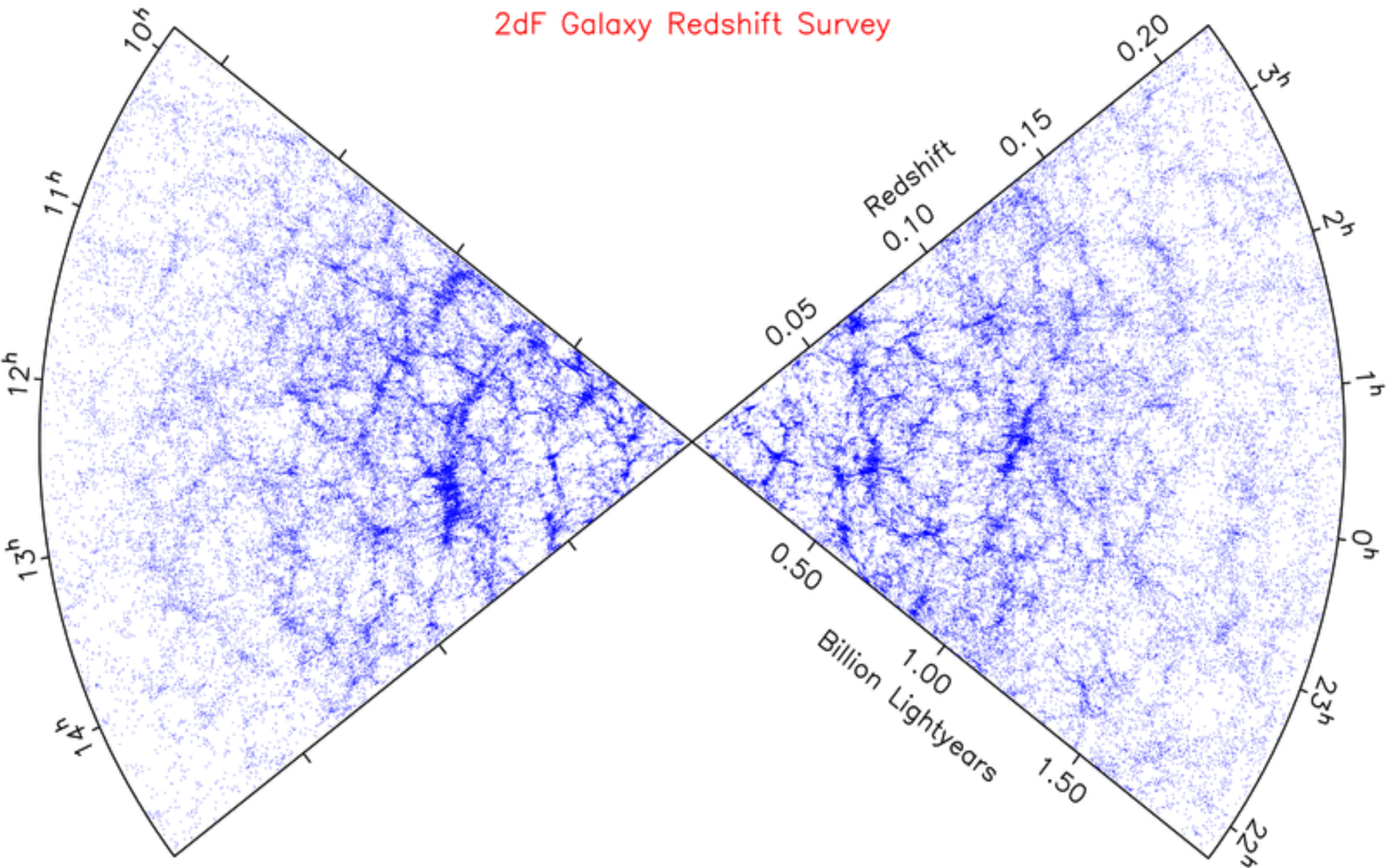
The Universe is still flat – now to better than 0.5% precision

Its expansion rate is 7% slower than previously thought, so its age has increased by 80,000,000 years!

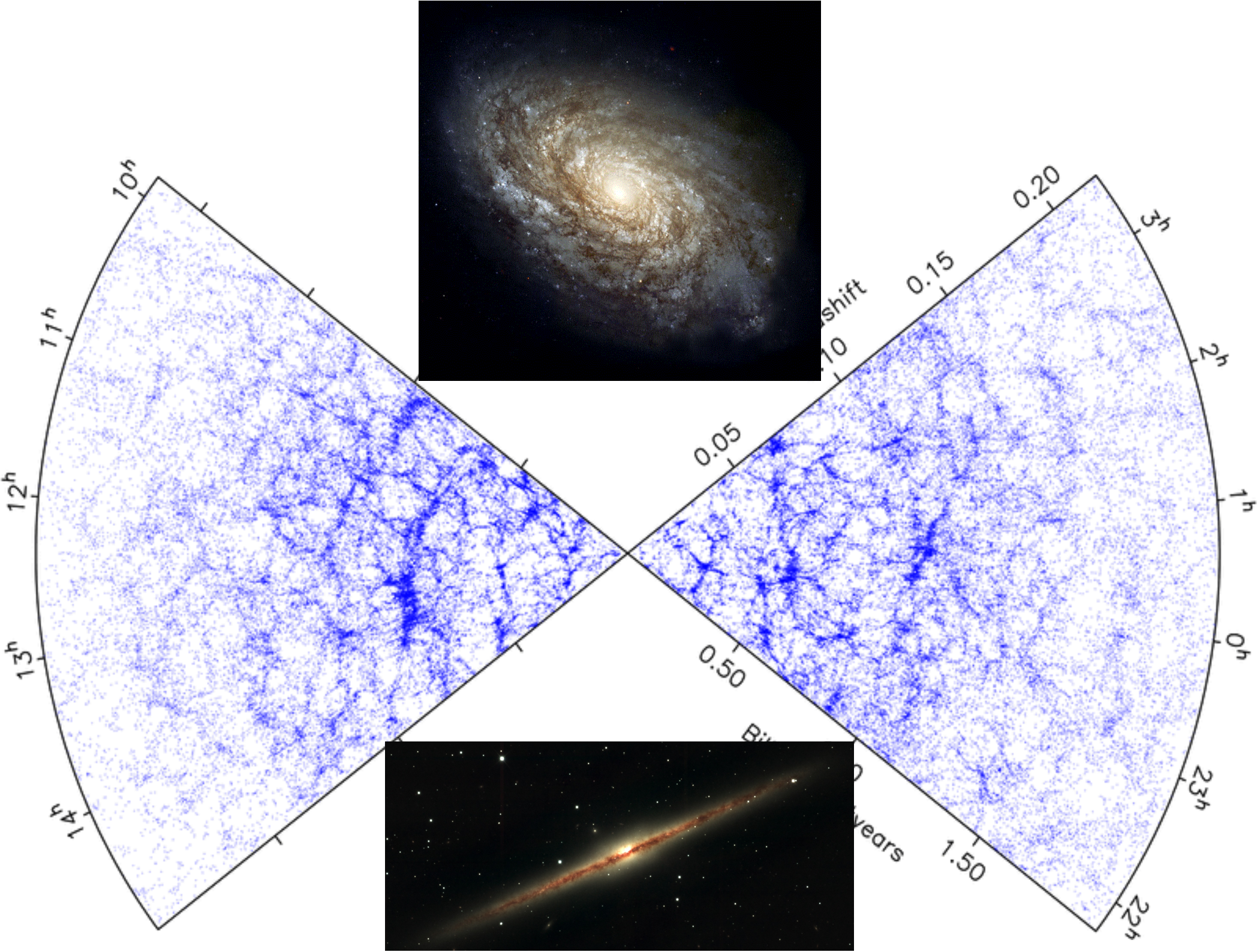
*Planck* results strongly support the idea that all structure originated from quantum zero-point fluctuations of the very early vacuum

# Large-scale structure in the present-day Universe

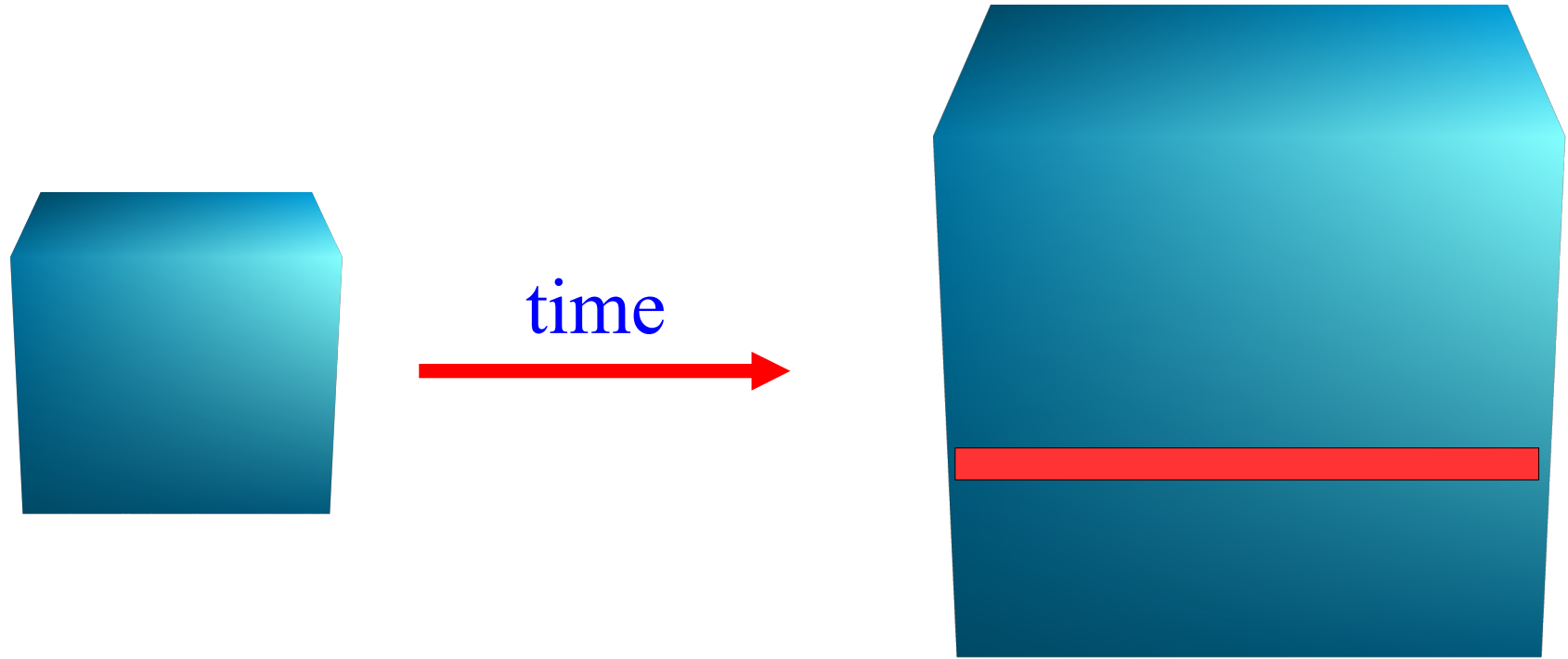
2dF Galaxy Redshift Survey







# How to follow cosmic evolution in a supercomputer

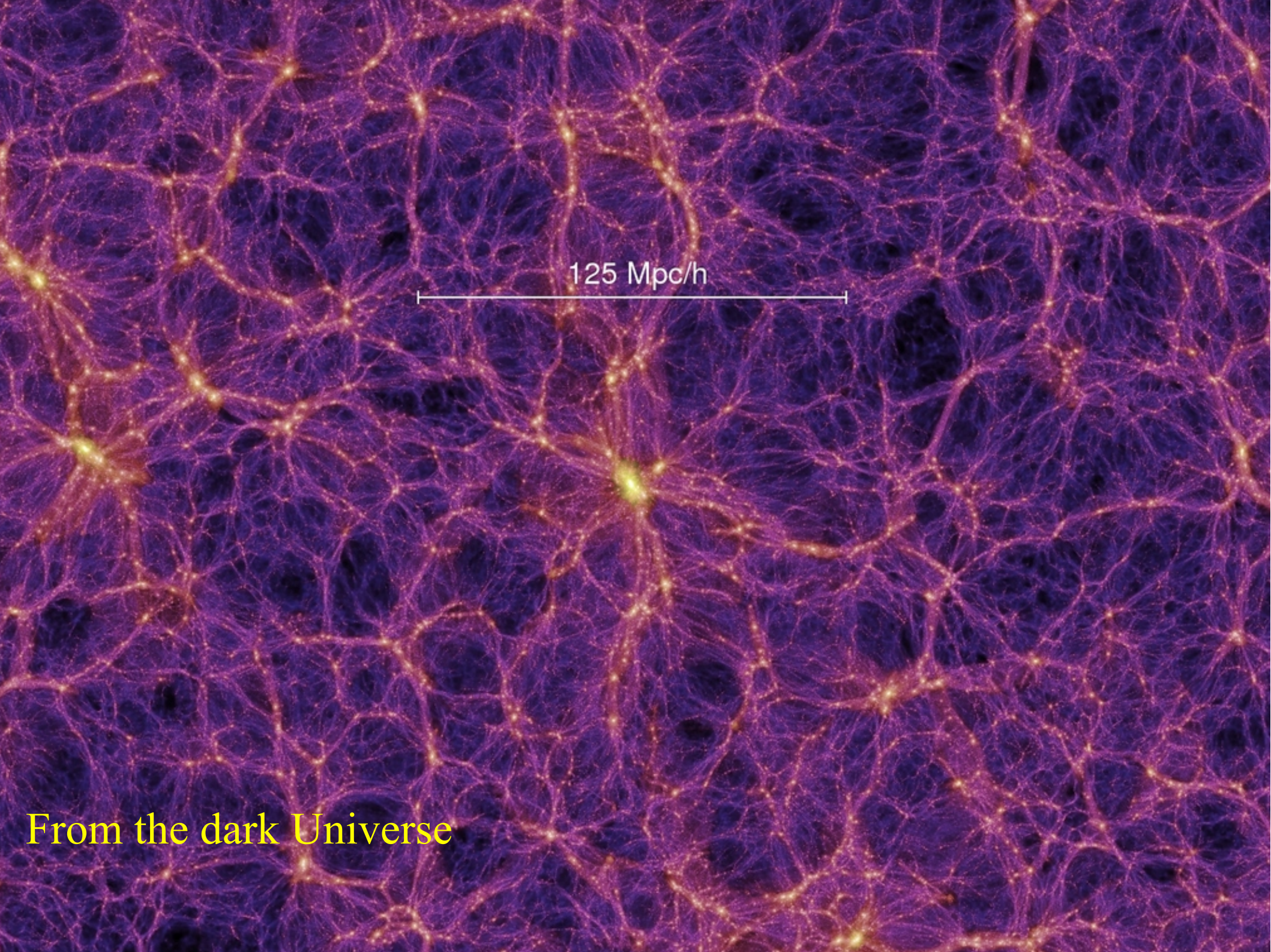


- Follow the material in a cube which expands with the Universe
- Start 400,000 years after the Big Bang
- Set initial conditions to match microwave background structure
- Calculate the evolution forwards to the present day

# Images of the Dark Matter distribution in a virtual Universe

- Evolution of structure in a thin, expanding slice
- A zoom from the entire visible Universe into a galaxy cluster
- A flight through the dark Universe

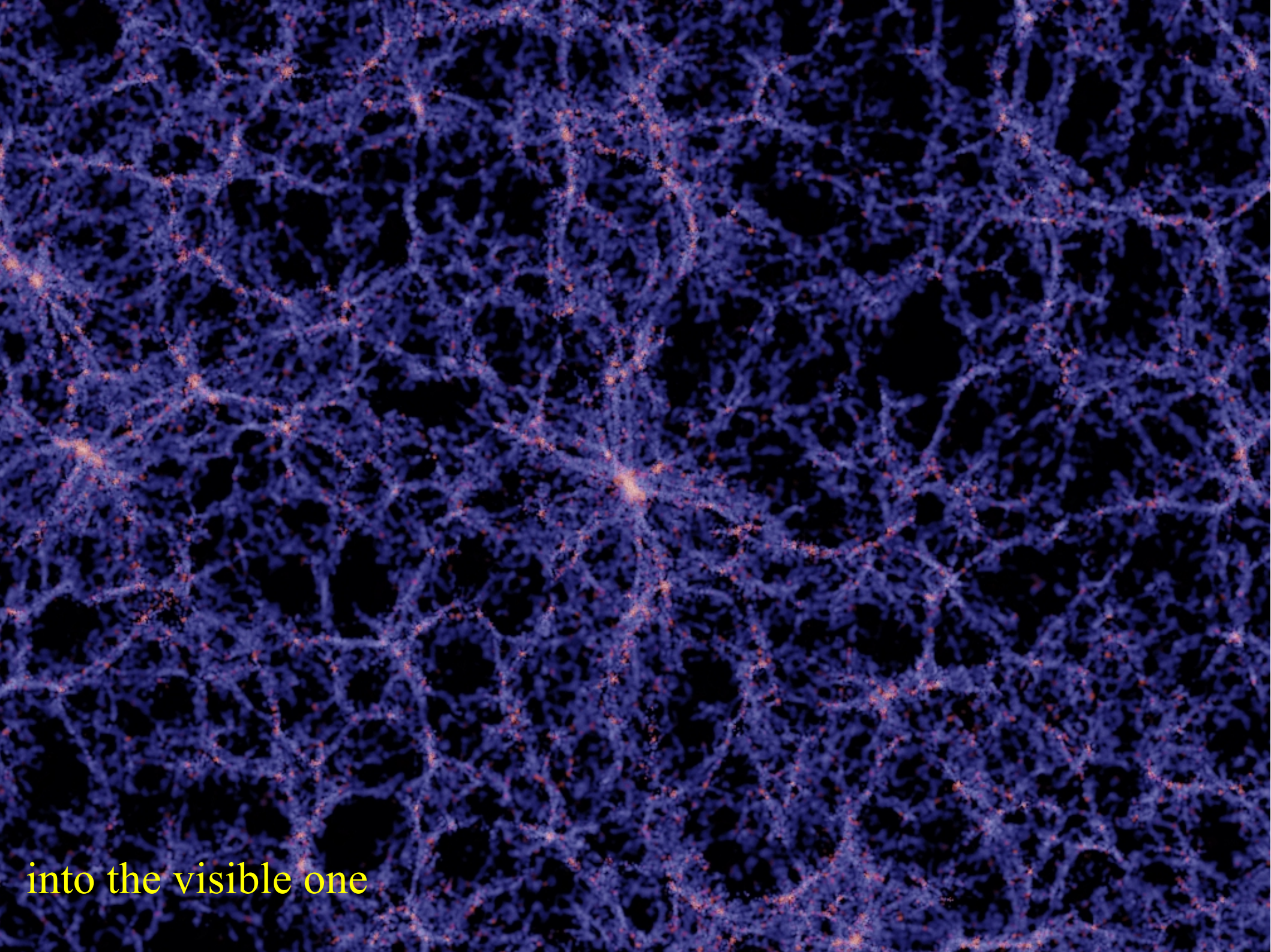




125 Mpc/h

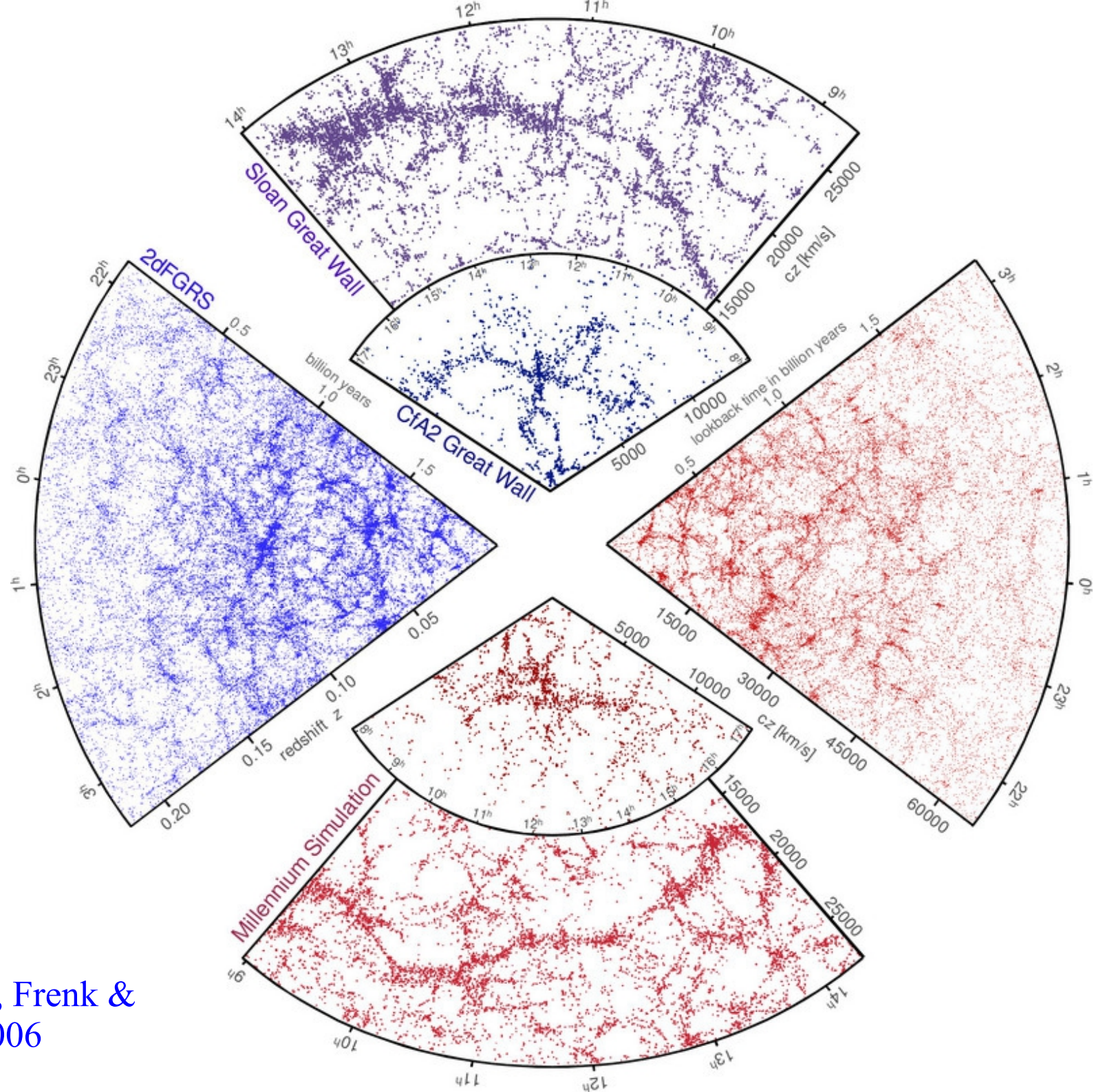
From the dark Universe





into the visible one

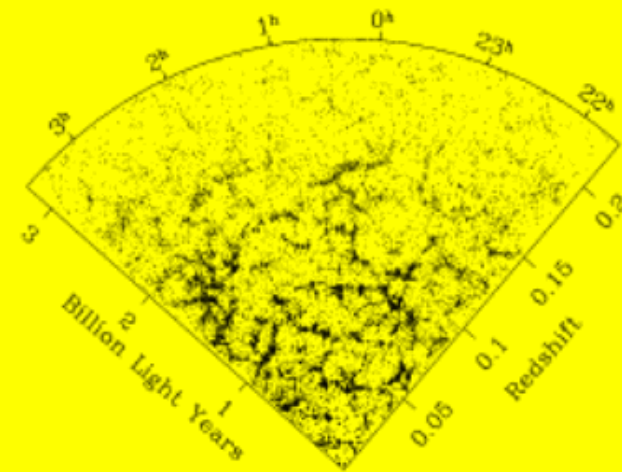
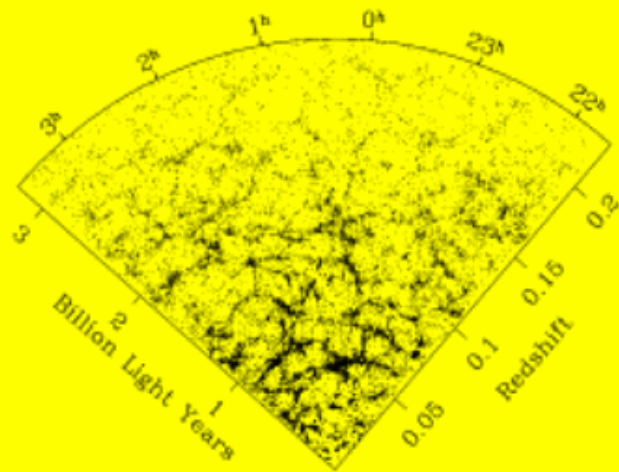
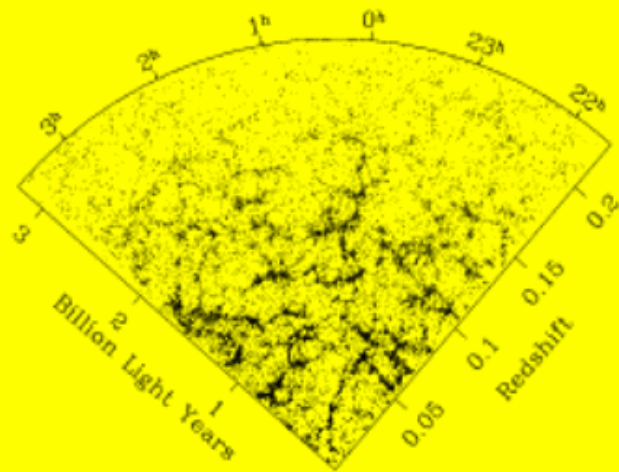
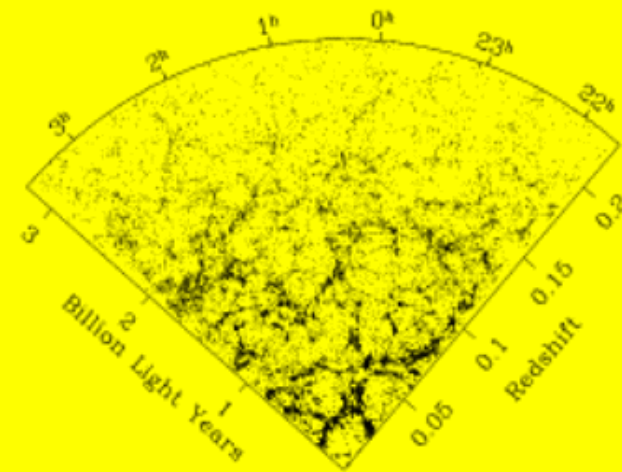
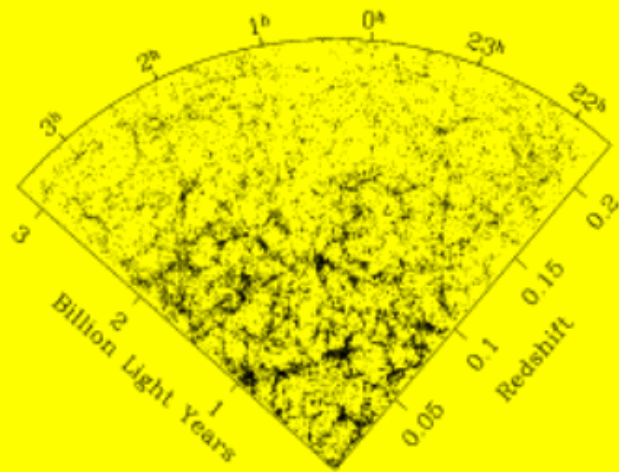
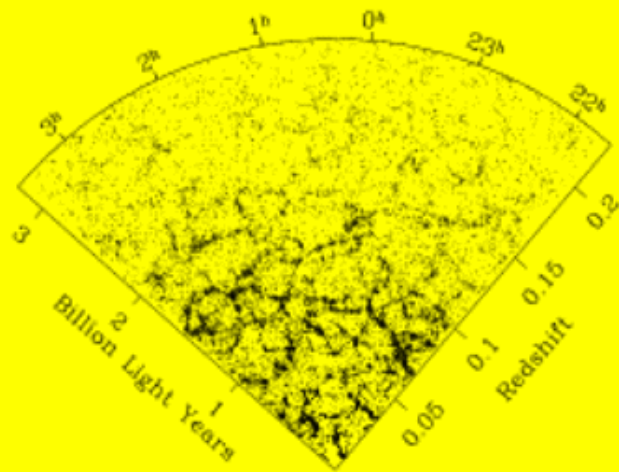




Springel, Frenk &  
White 2006



# Virtual versus real Universes



## To conclude....

- Our Universe was born 13.7 billion years ago in a hot and almost uniform explosion -- the Big Bang
- All structure grew from quantum fluctuations of the early vacuum -- Everything has formed from “nothing”!
- Only 5% of today's Universe is made of ordinary matter
- About 27% is made of as yet unidentified elementary particles -- the Dark Matter
- About 68% consists of a new form of energy which accelerates the expansion of today's Universe -- the Dark Energy
- Galaxies and galaxy clusters, stars and planets formed out of sound waves in the primordial gas through the effects of Gravity



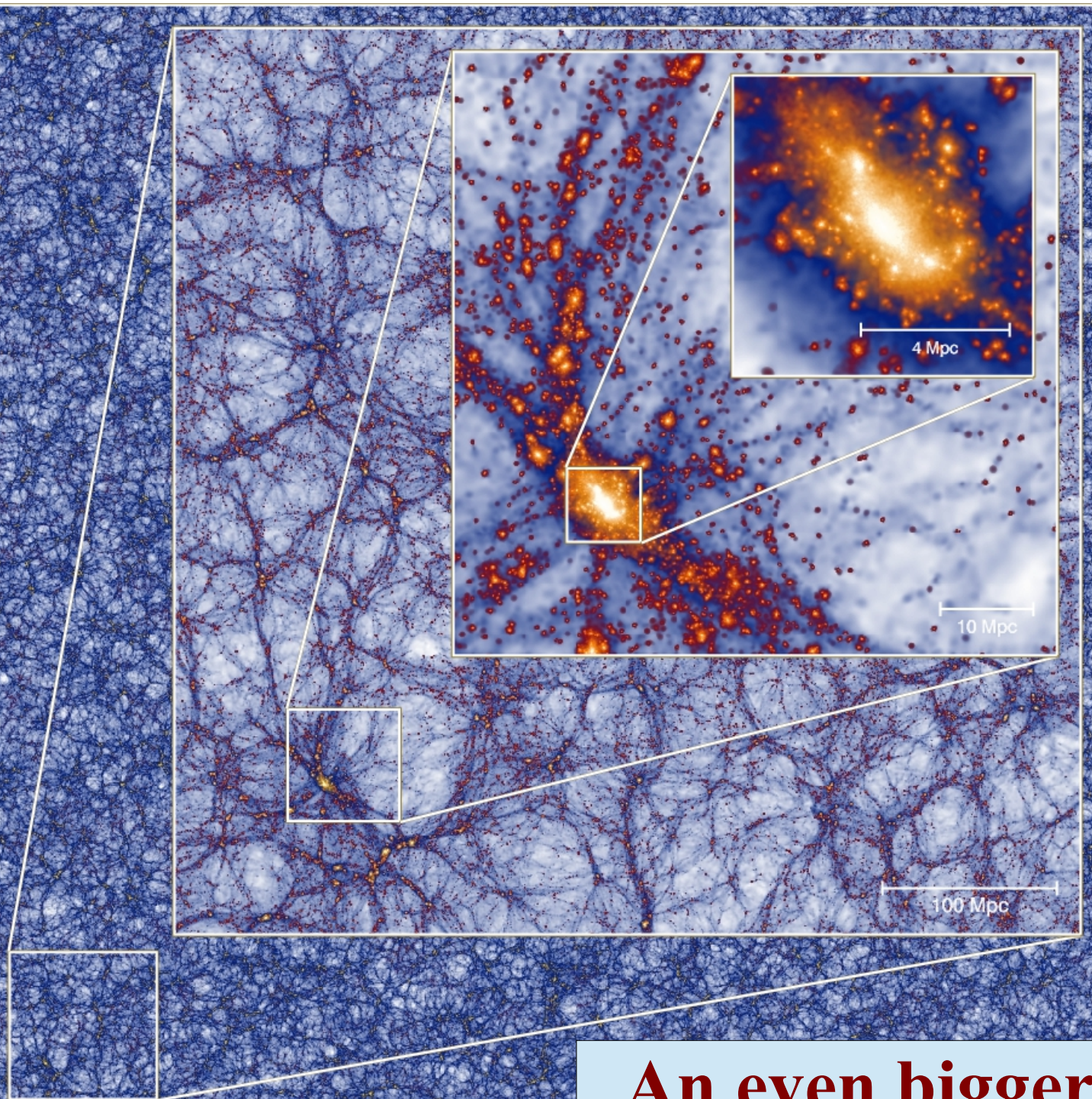


# Can we calculate into the future?

- What will happen to our Milky Way when it runs into the Andromeda nebula in about 3.5 billion years from now?
- What will become of the Earth?
- What will happen to humankind?

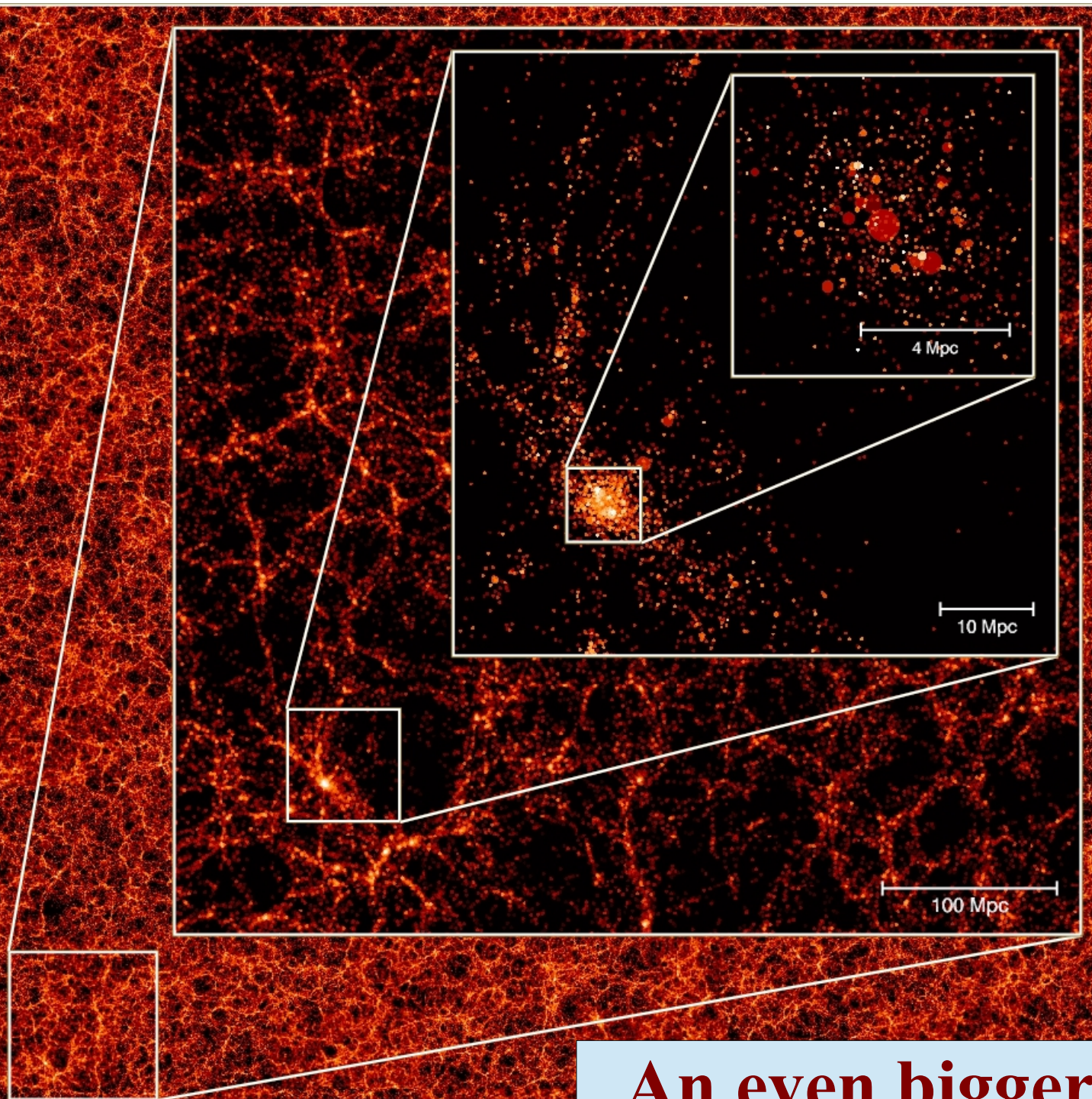






**An even bigger simulation:  
The Millennium XXL**





**An even bigger simulation:  
The Millennium XXL**