

*Hefei, September 2009*

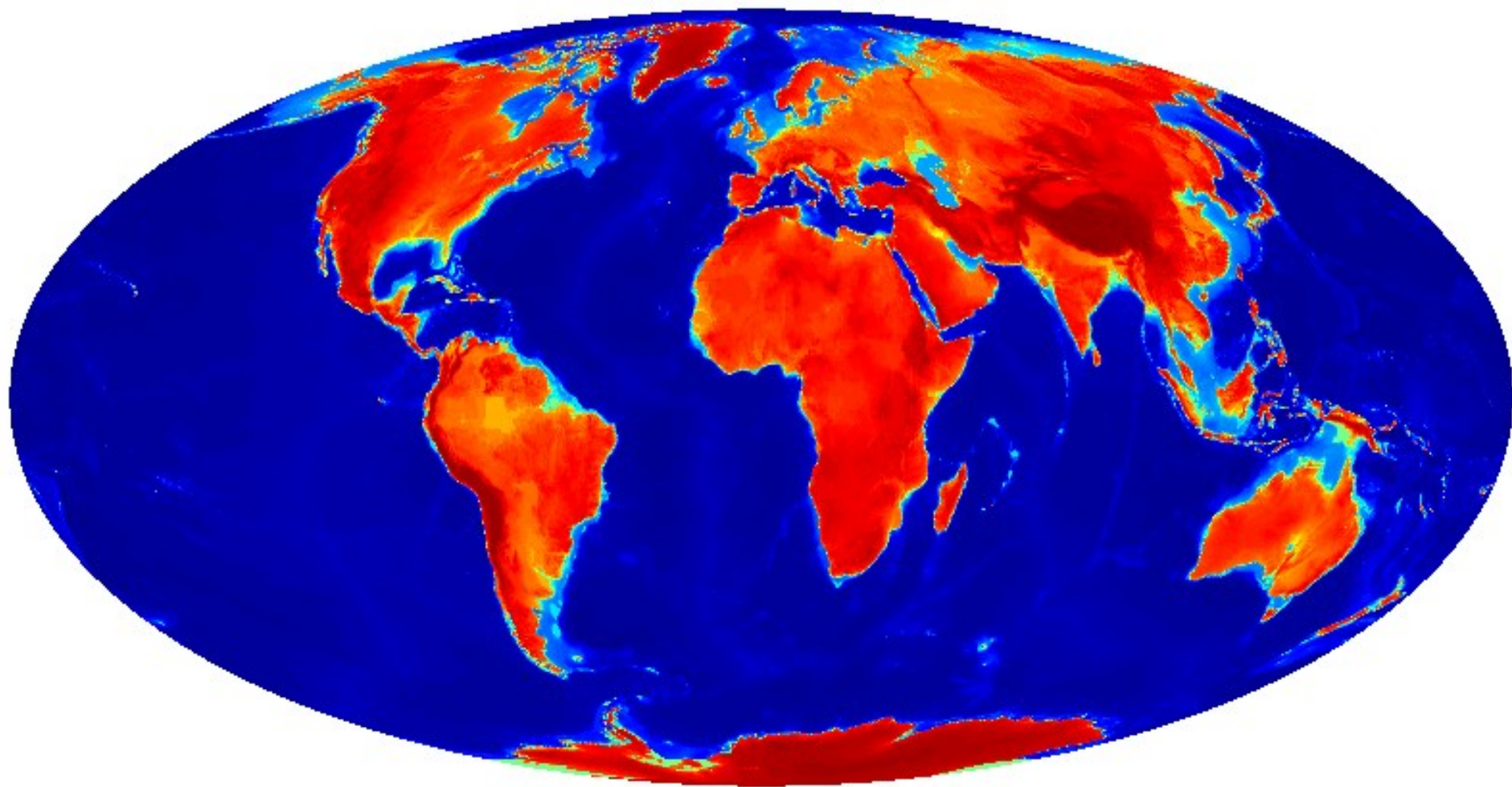
# **Understanding our Universe:** Ideas and opportunities in modern astrophysics

*Simon White*

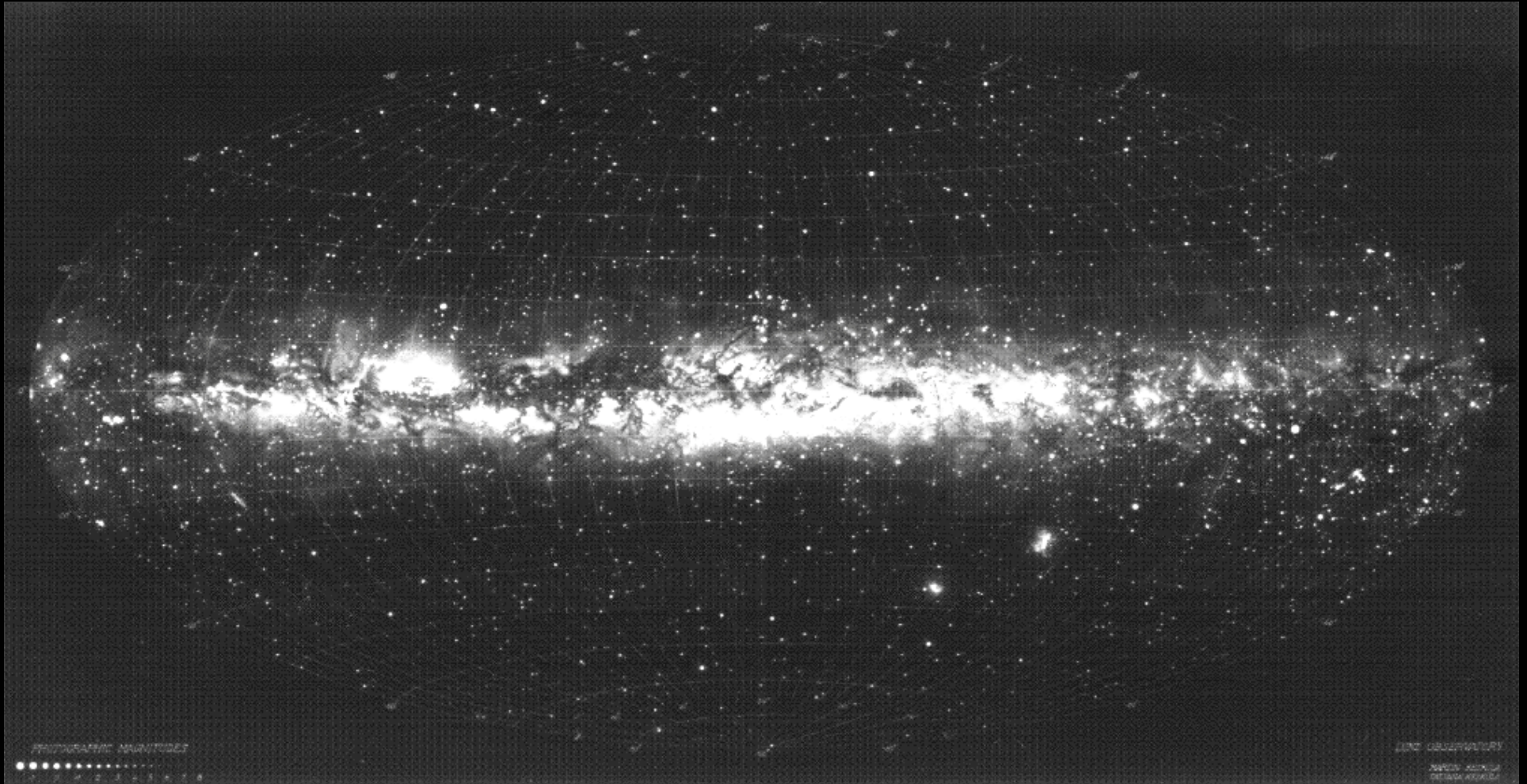
*Max Planck Institute for Astrophysics*



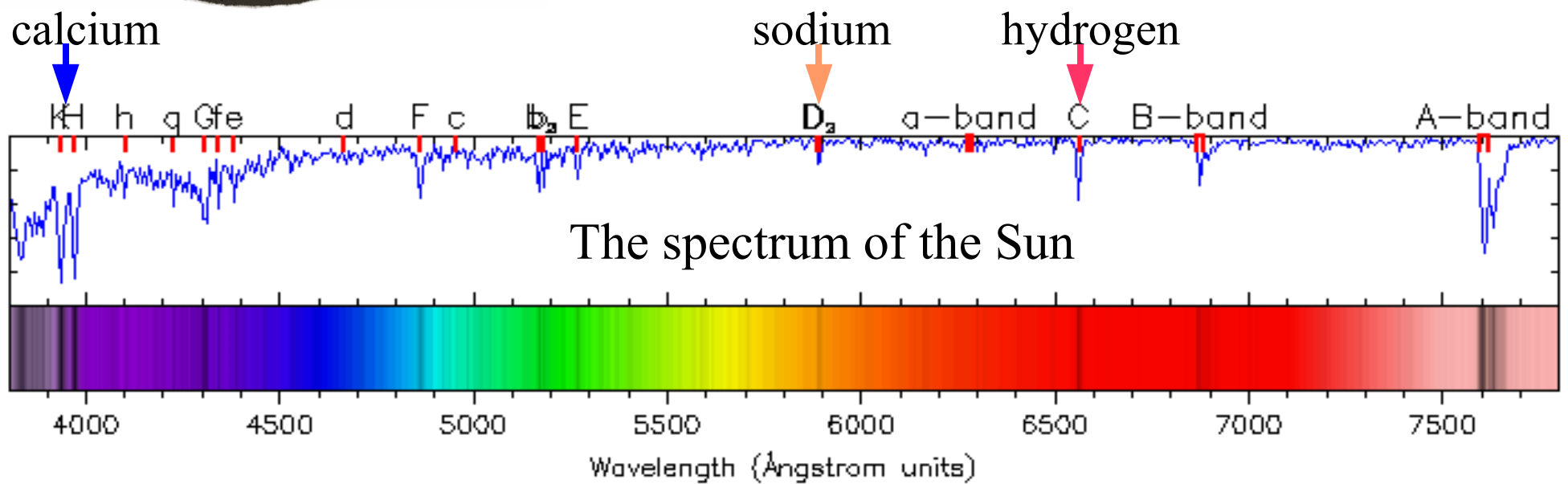
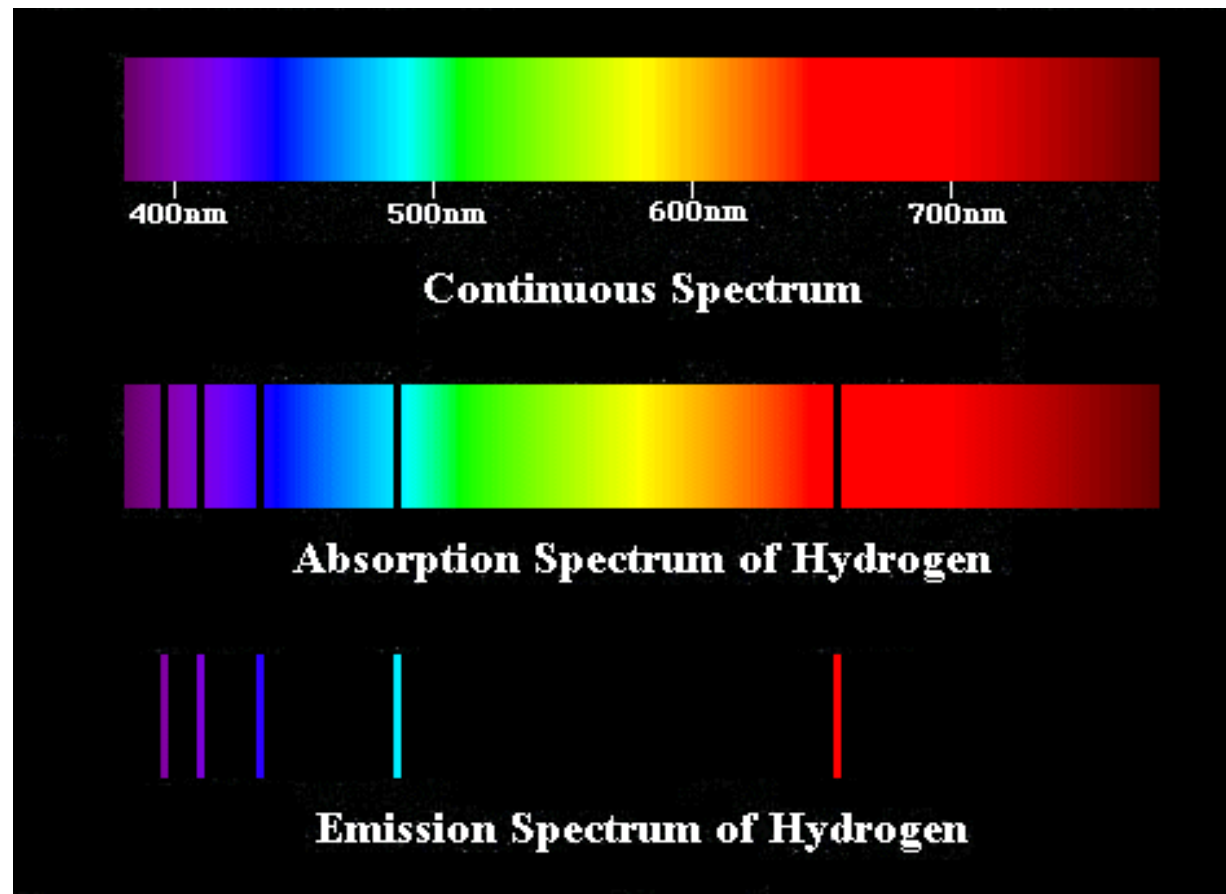
What can we know about things we cannot touch?



# Star map of the whole sky



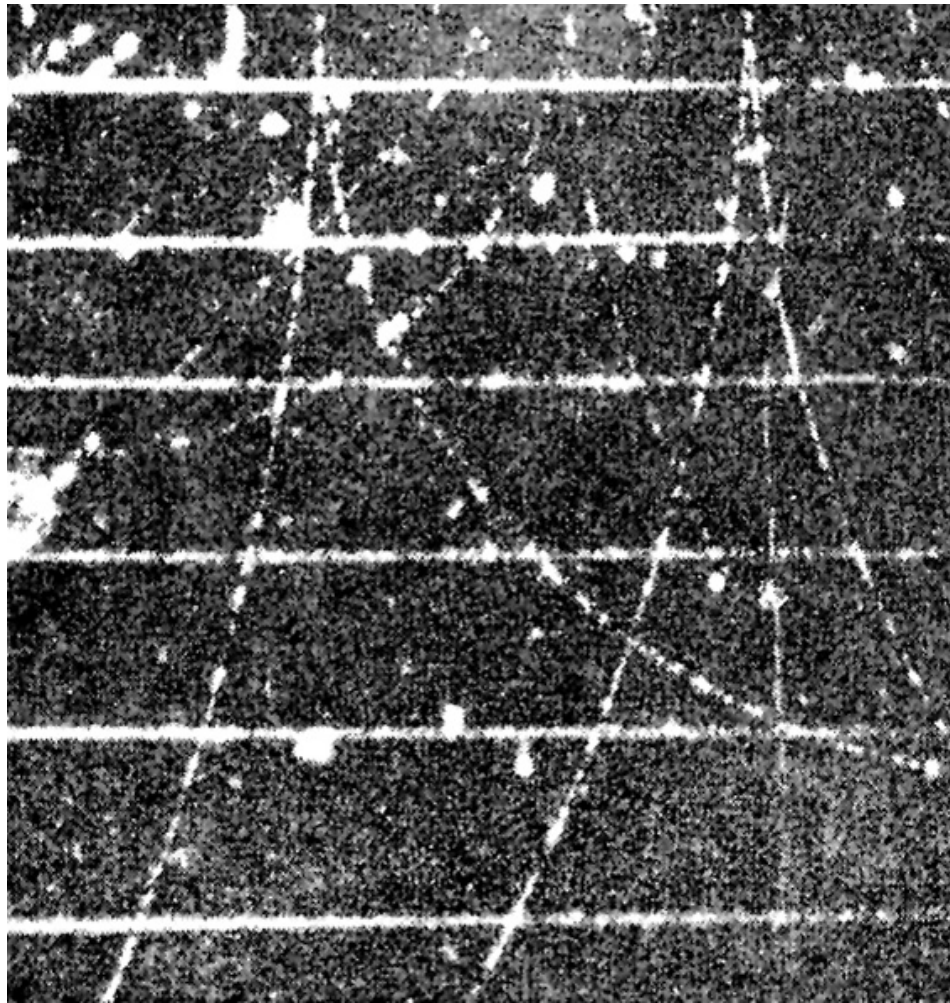






What can we know about things we cannot see?





$e^-$

$e^-$

$e^-$



$e^+$



What can we know about things that we cannot see or touch?

# The Coma Galaxy Cluster

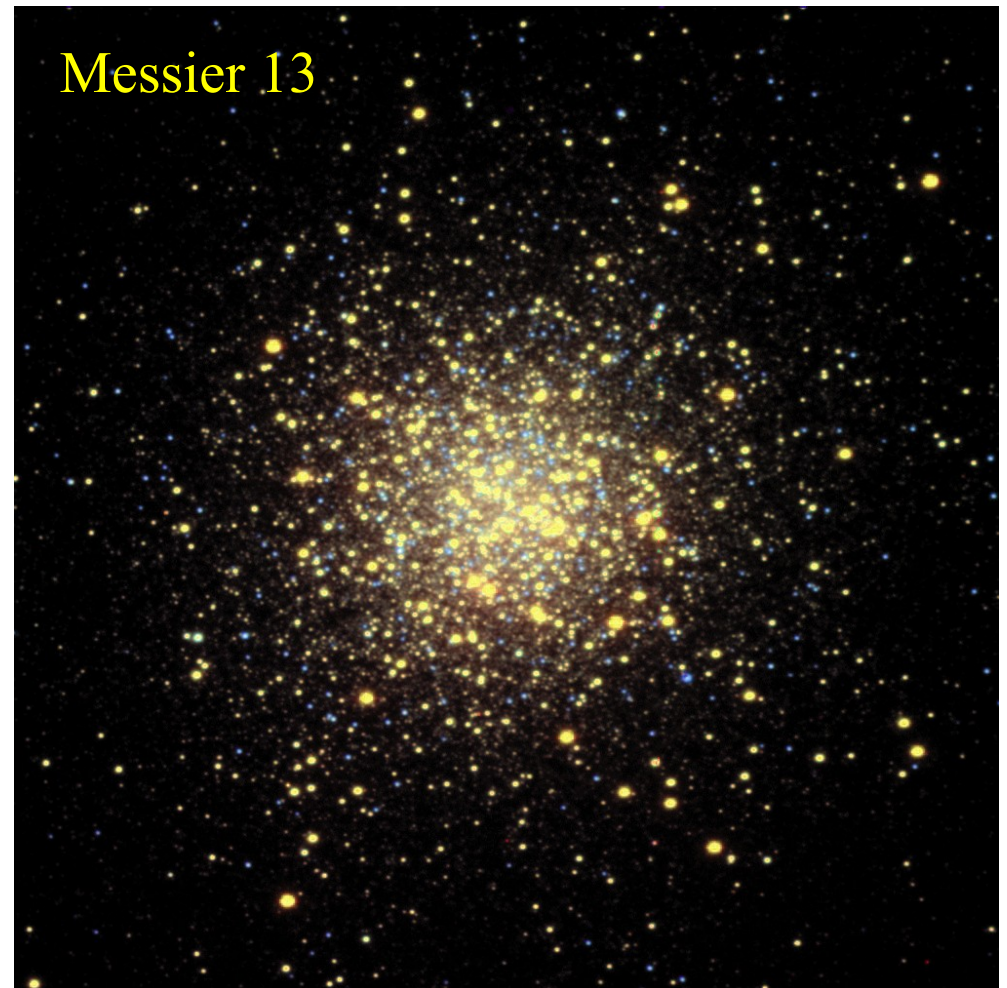


Fritz Zwicky



What can we know about processes which act over billions of years when we live for only seventy years?

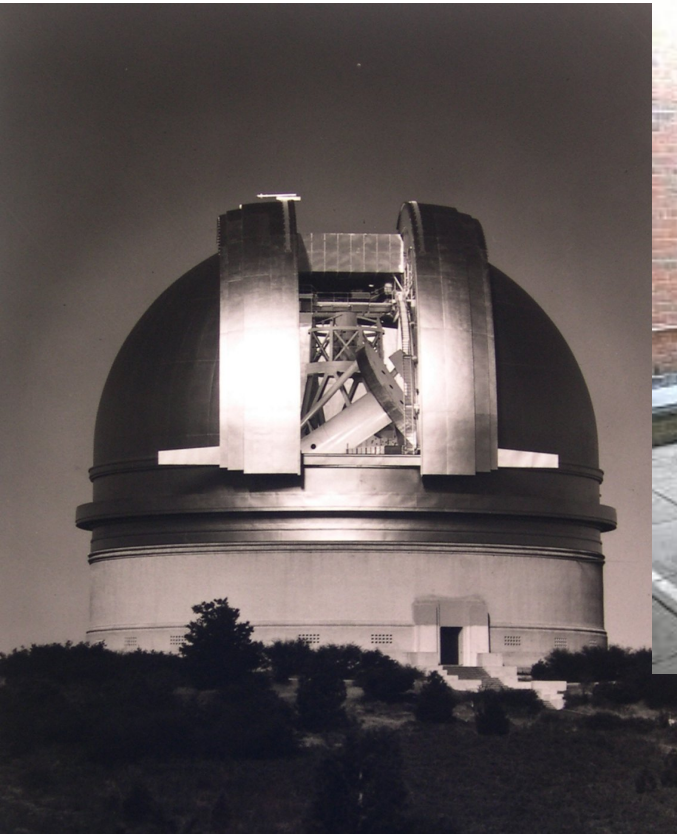
# Do archaeology!



Use old objects to find out what the Universe was like when they were young.



Use telescopes as time machines - look directly into the past

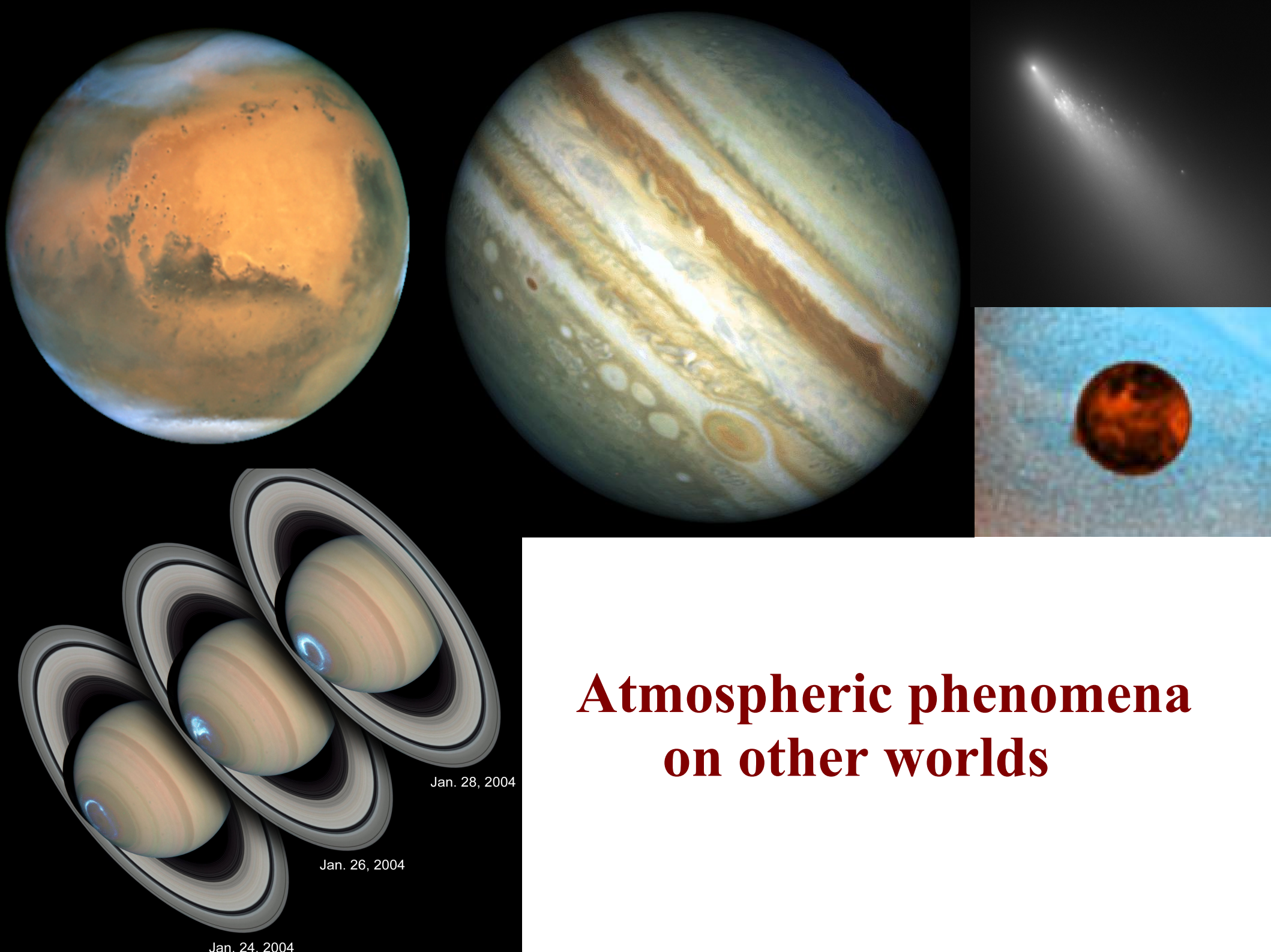


We see objects as they were when the light left them,  
not as they are today

# Astrophysics

- puts our Earth in its cosmic context
- allows us to study physics under *very* extreme conditions
- has taught us the origin of the elements
- has shown us what our Universe is made of
- allows us to see back in time, almost to the beginning
- makes cosmic evolution directly visible to us
- provides insight into our long-term future
- is unveiling new worlds, and (perhaps) other life

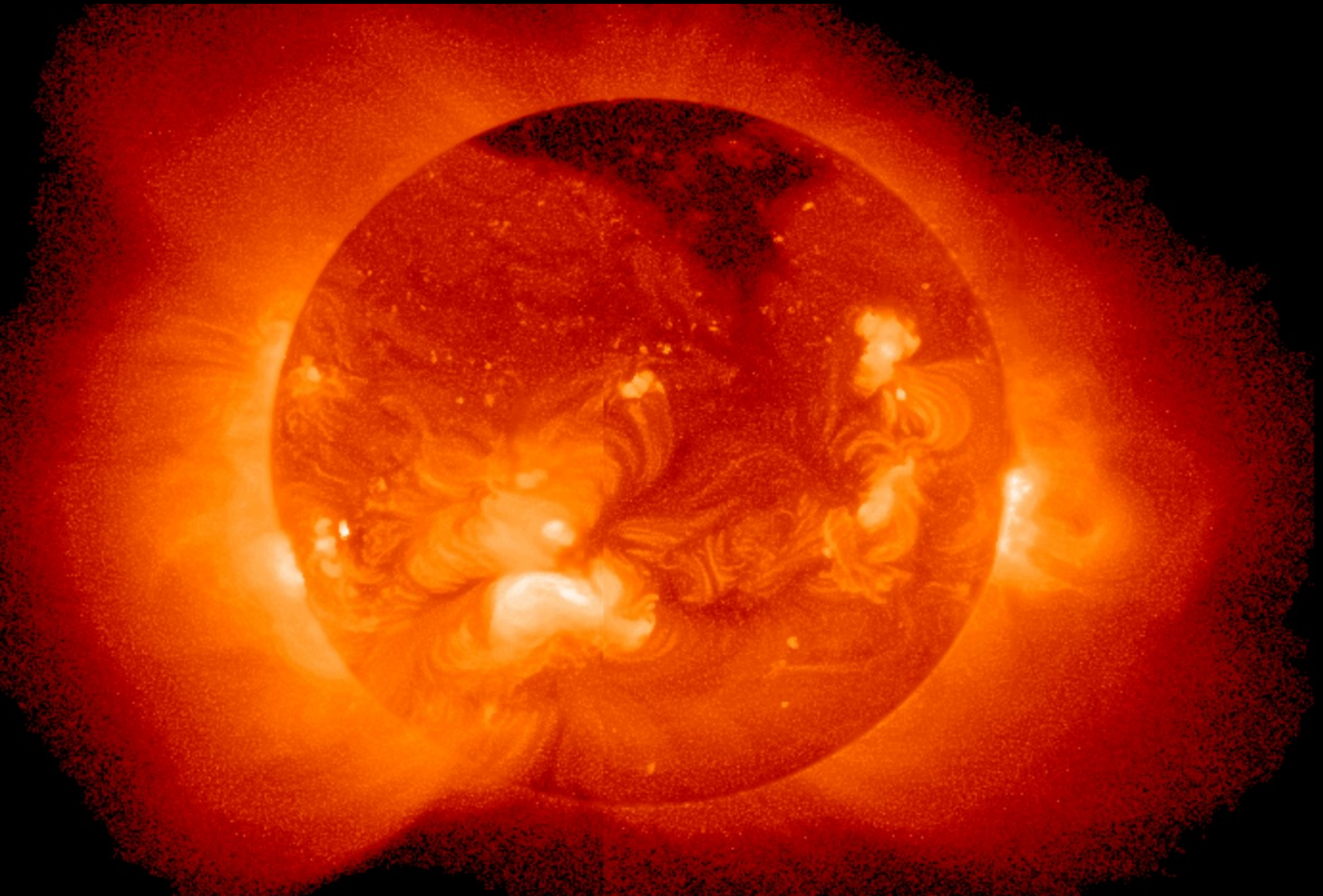




# Atmospheric phenomena on other worlds

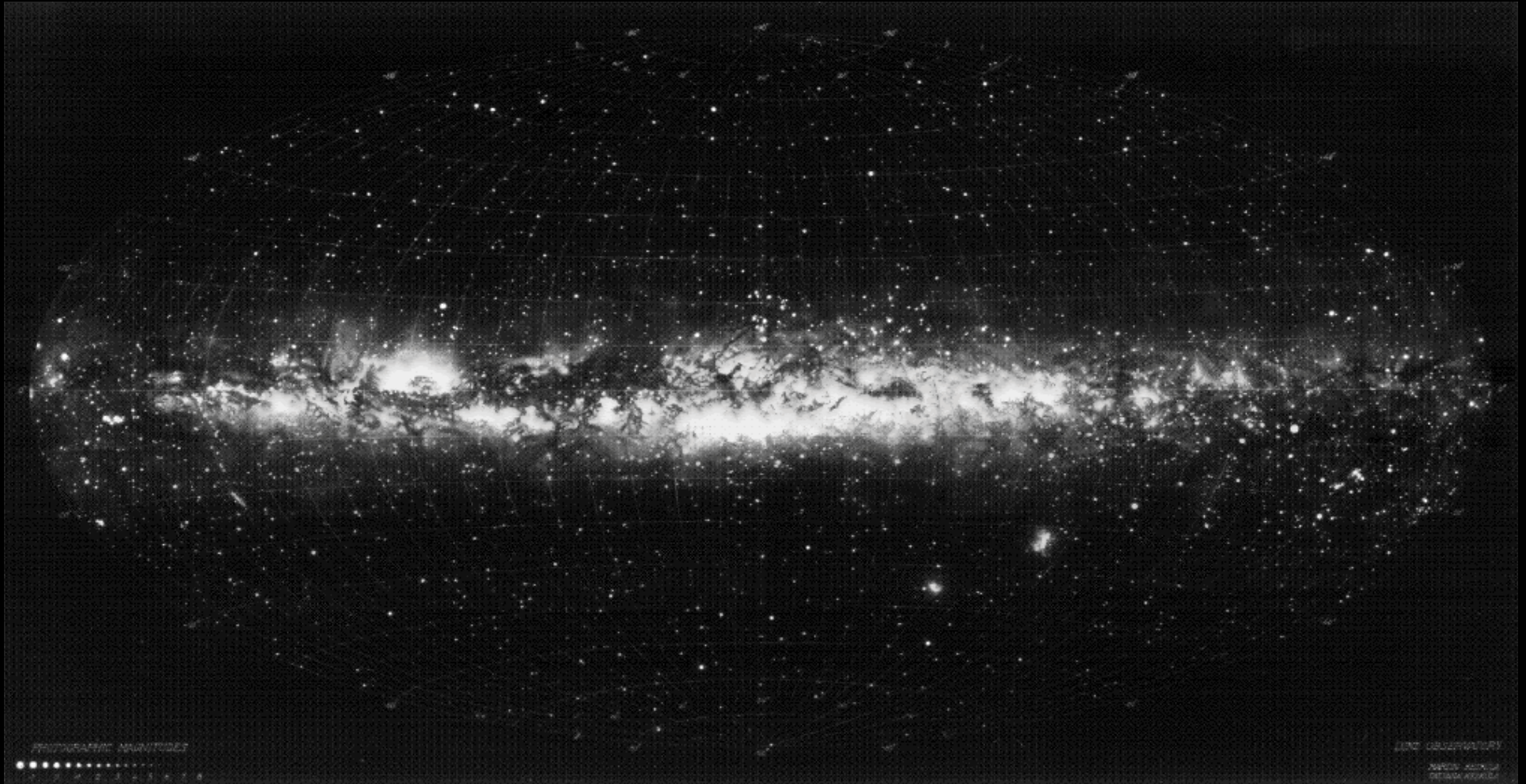


# X-ray image of the Sun



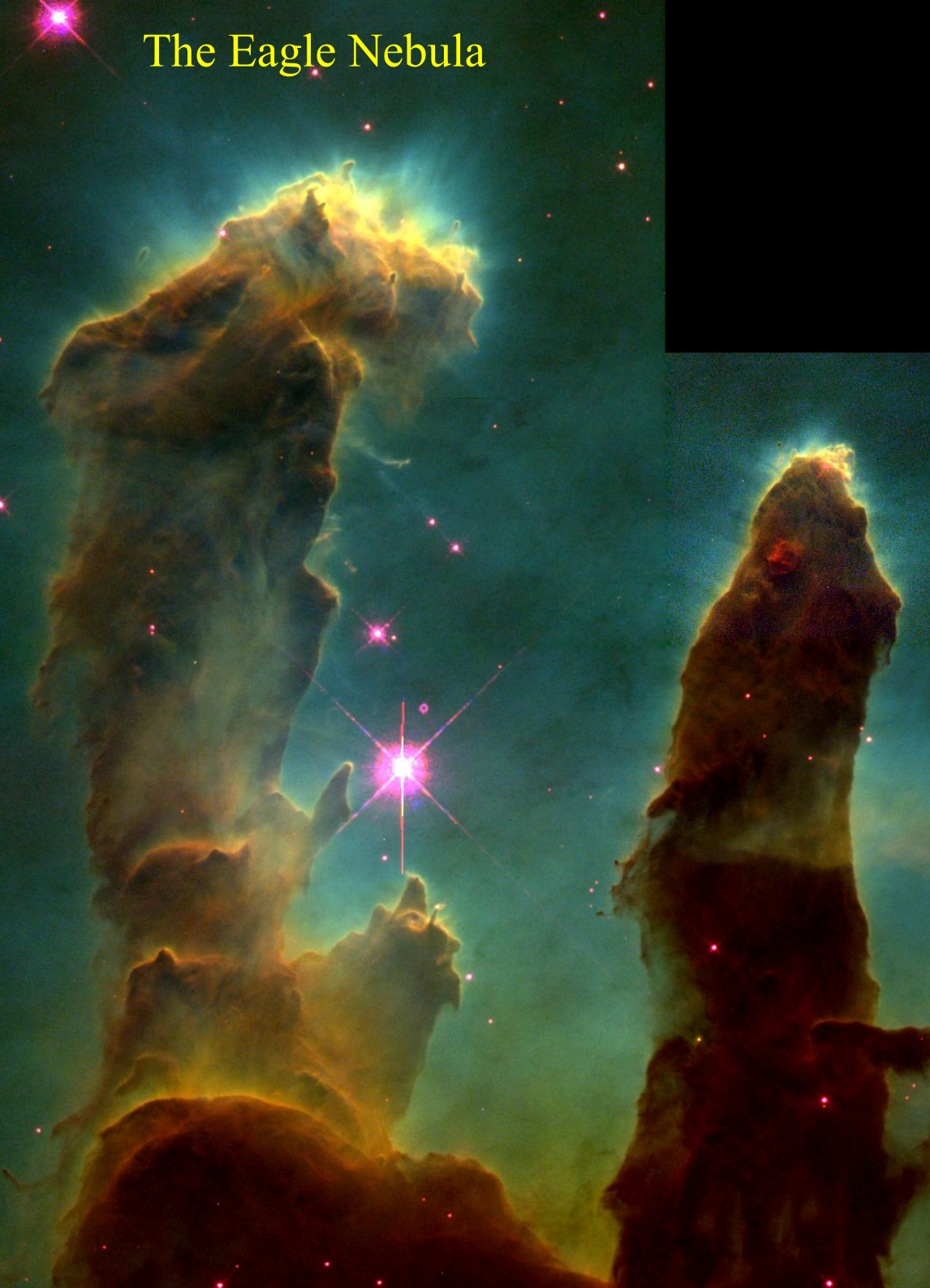


# Star map of the whole sky



to 10,000 light years

The Eagle Nebula

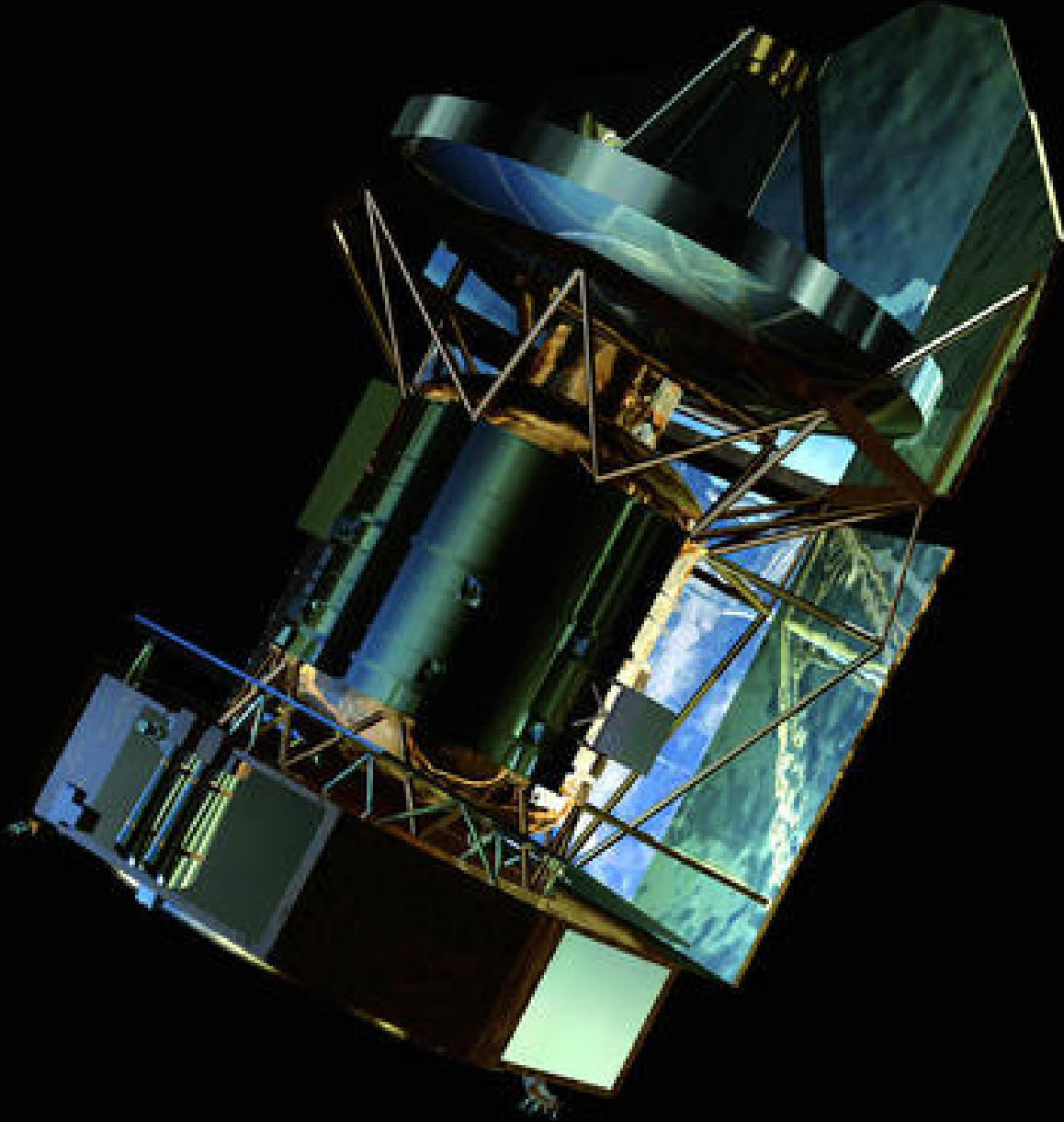


Proplyds in Orion



**The birth  
of stars**





# Herschel

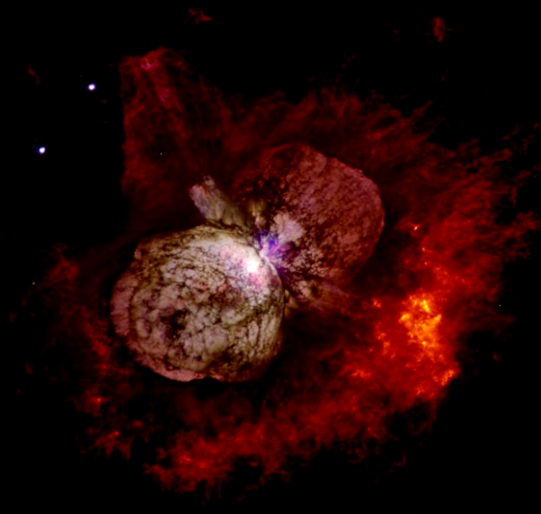
The newest space  
telescope, launch  
May 2009

Crab Nebula



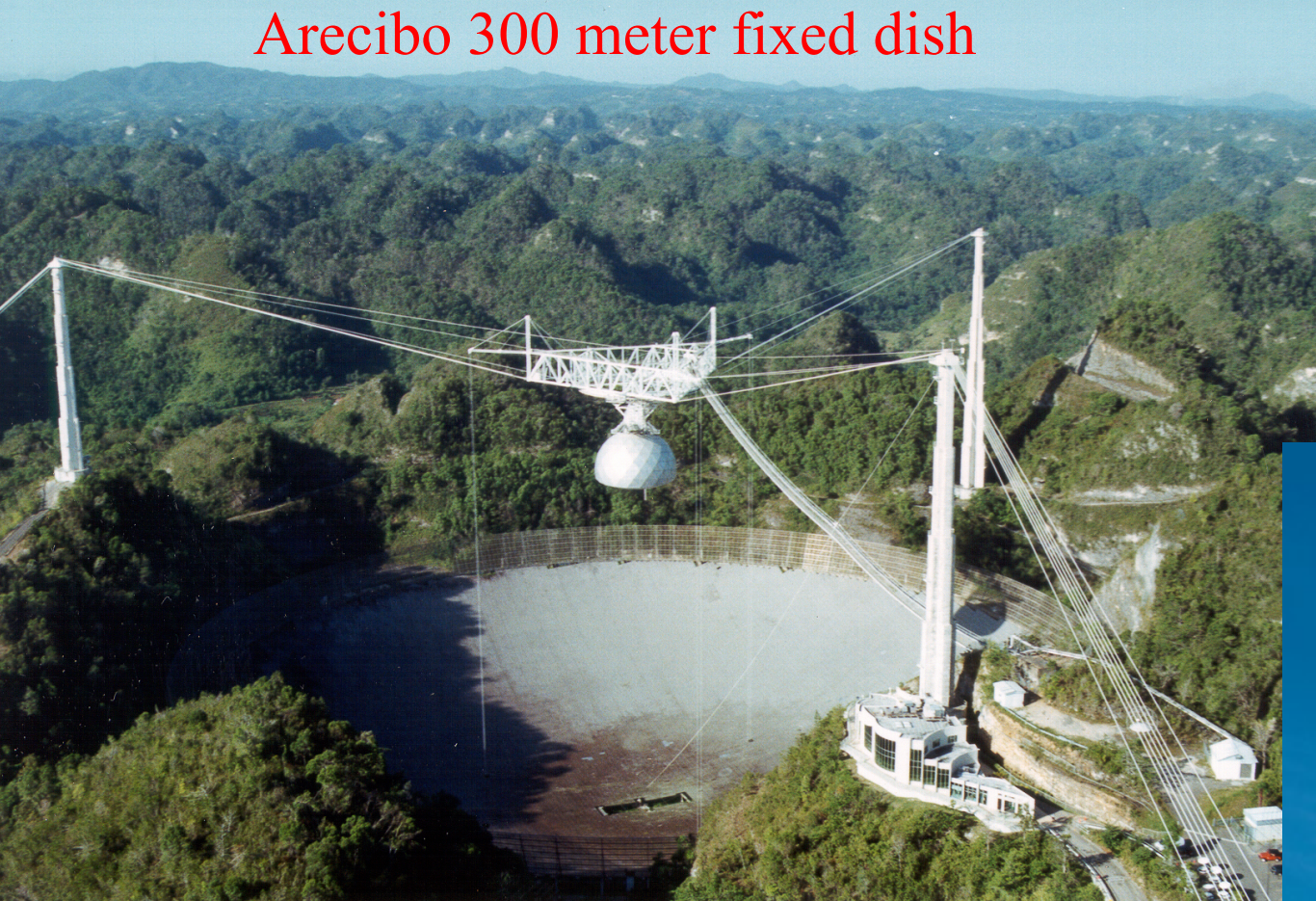
## The death of stars

$\eta$  Carina





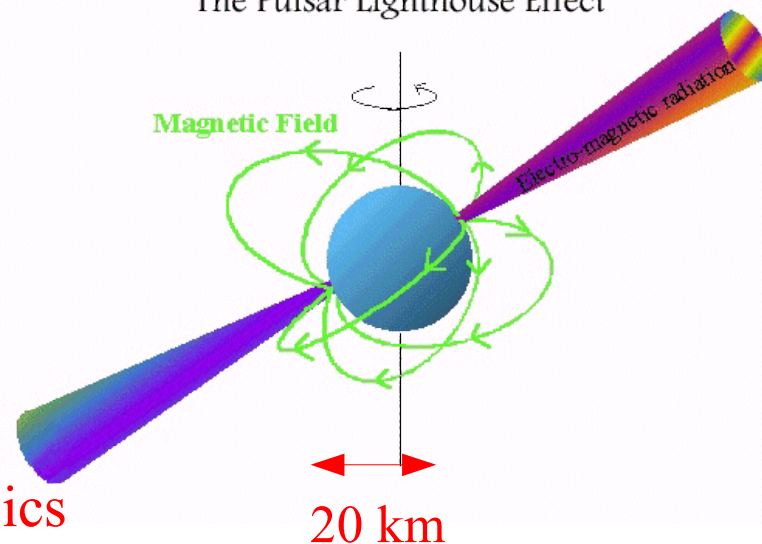
Arecibo 300 meter fixed dish



# Pulsar Hunters



The Pulsar Lighthouse Effect



## NEUTRON STAR

Mass of the Sun

Size of Hefei

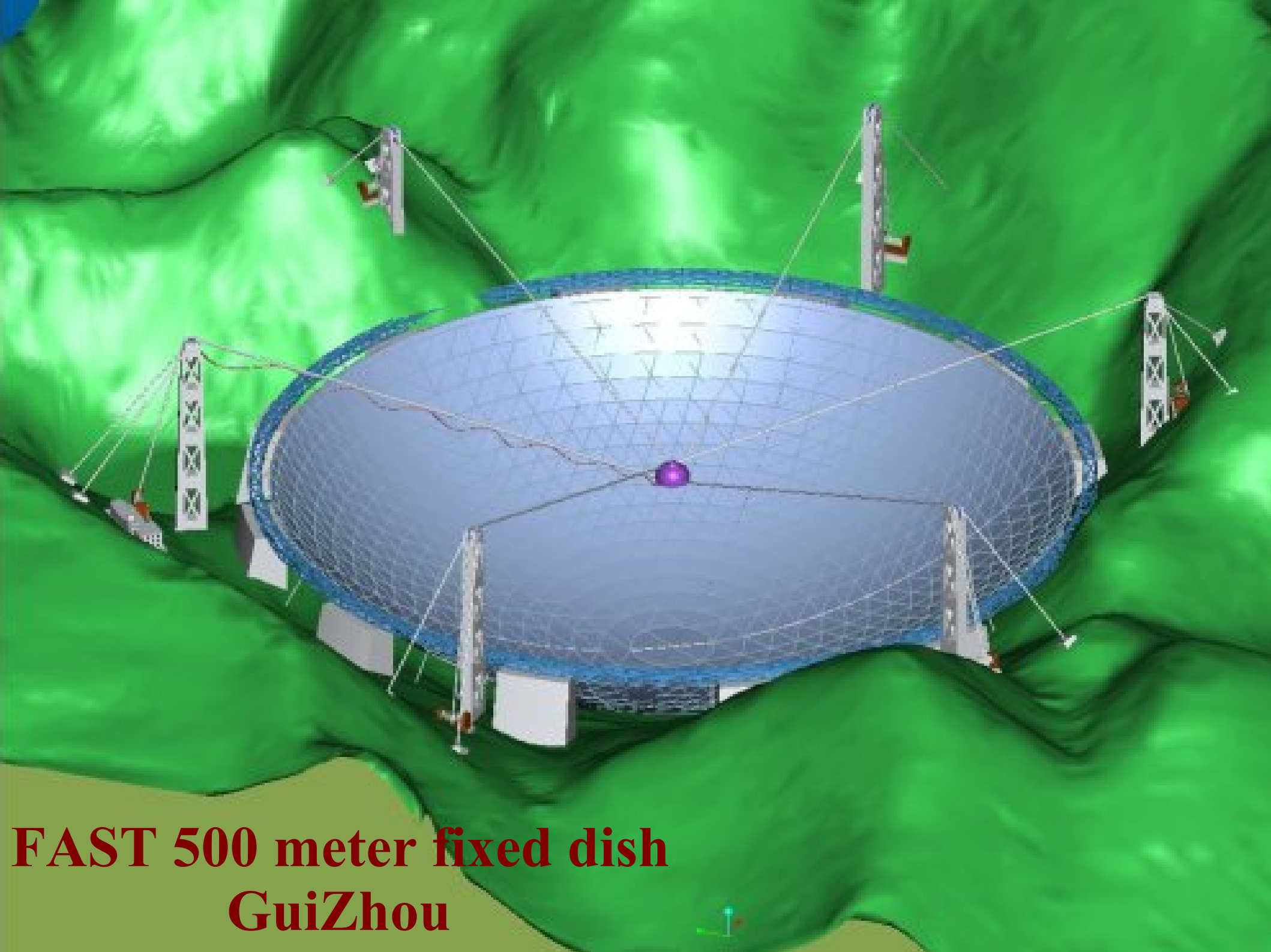
Made of neutrons

Almost a black hole

Nobel Prizes in Physics  
1974, 1996

Effelsberg 100 meter

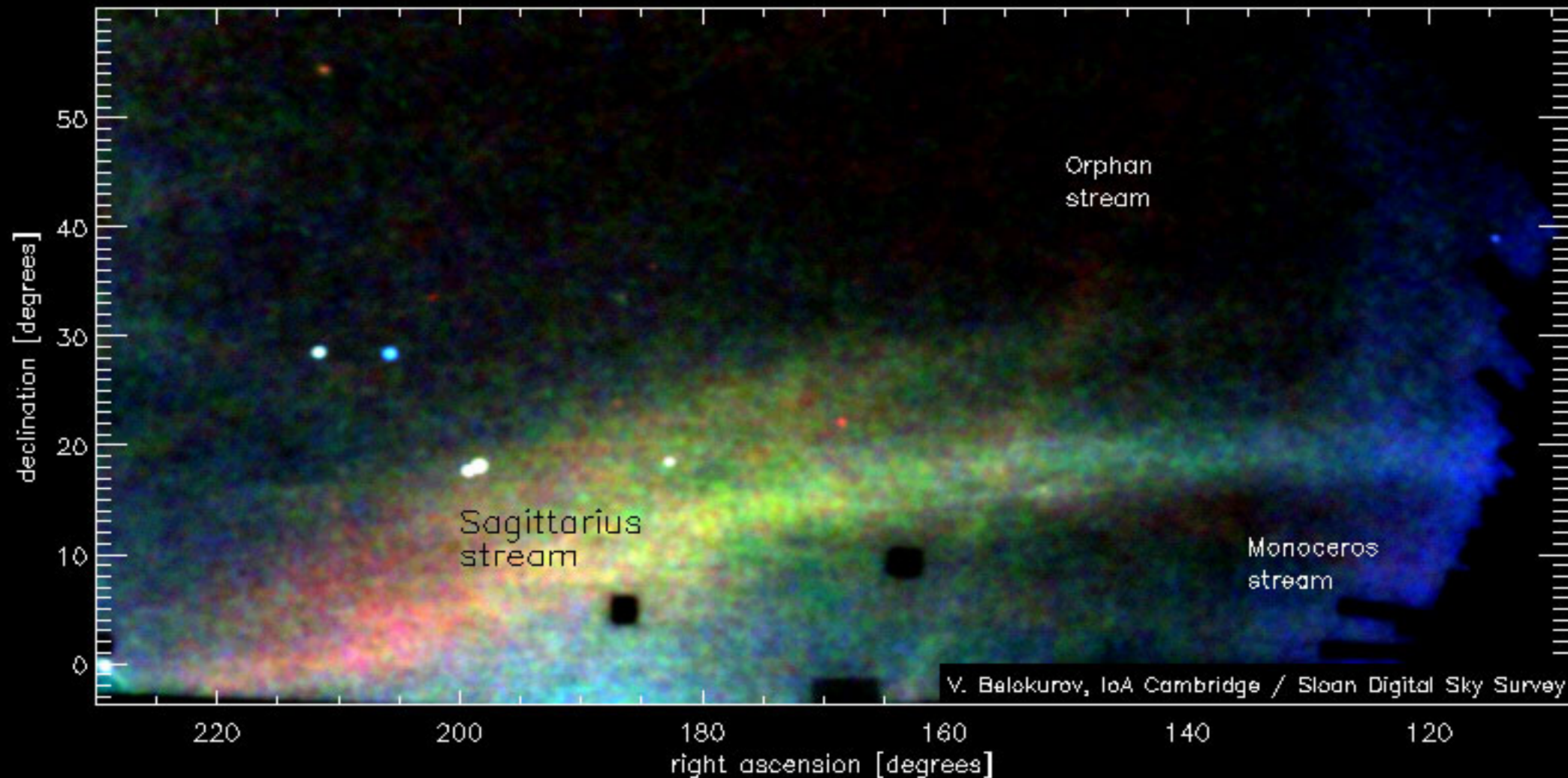




**FAST 500 meter fixed dish**  
**GuiZhou**

# The Field of Streams

Stars in the outer parts of our own Galaxy  
The remains of old, tidally shredded satellite galaxies





# Sloan Digital Sky Survey Telescope, New Mexico



# LAMOST telescope

## Hebei





# The Andromeda Nebula: our nearest big neighbor



to 2,000,000 light years



# Spiral galaxies

M101



NGC 5907





# A Galactic Traffic Accident....





....and its result



Elliptical galaxy NGC 4458



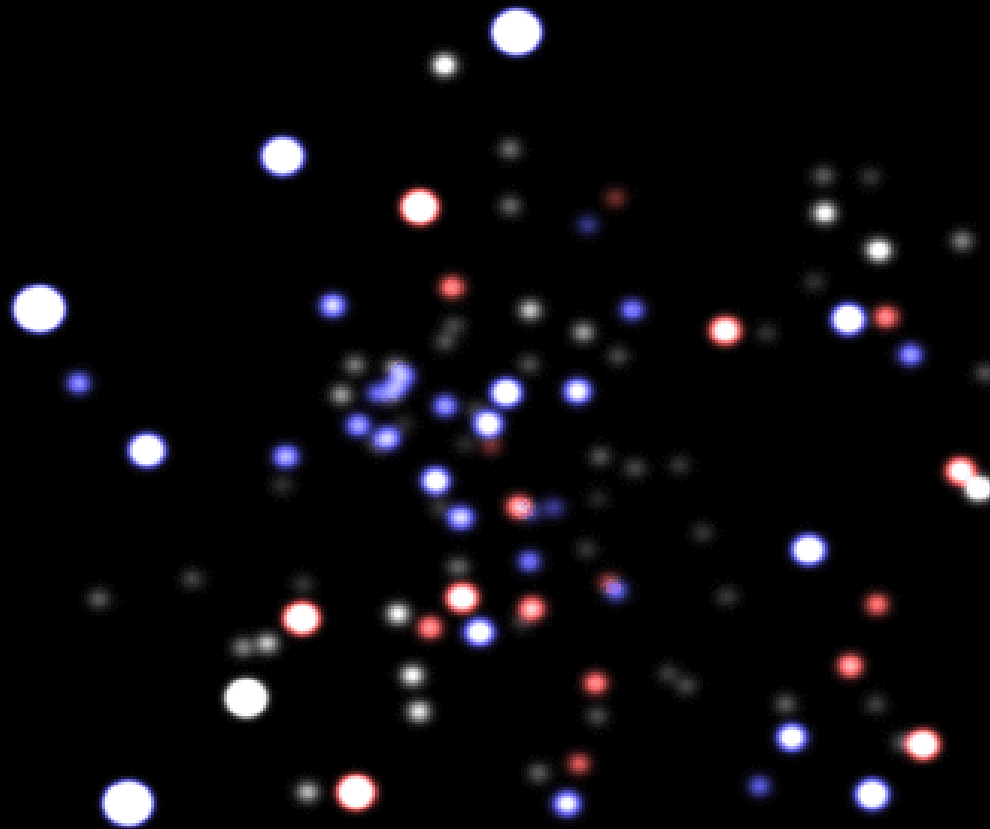
Jet from M87's black hole



# Infrared images of the centre of the Milky Way

1992 to 2008

The stars orbit a central black hole 4 million times the Sun's mass



← 0.5 light-years →

# The galaxy cluster, Abell 2218 a gravitational telescope

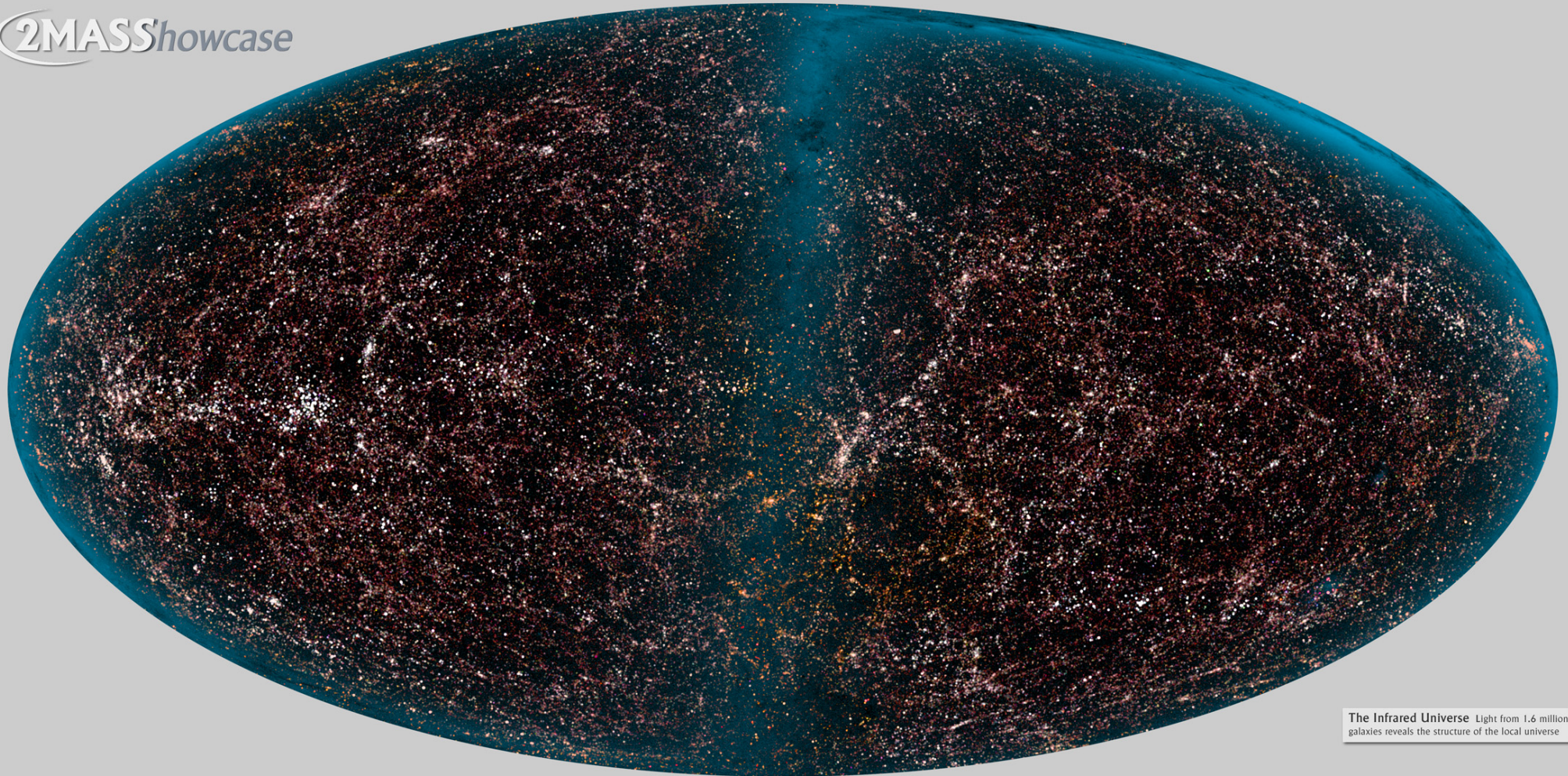


to 4 billion light years (cluster), 15 billion light year (background)



# Map of galaxies across the whole sky

2MASS*showcase*



**The Infrared Universe** Light from 1.6 million galaxies reveals the structure of the local universe

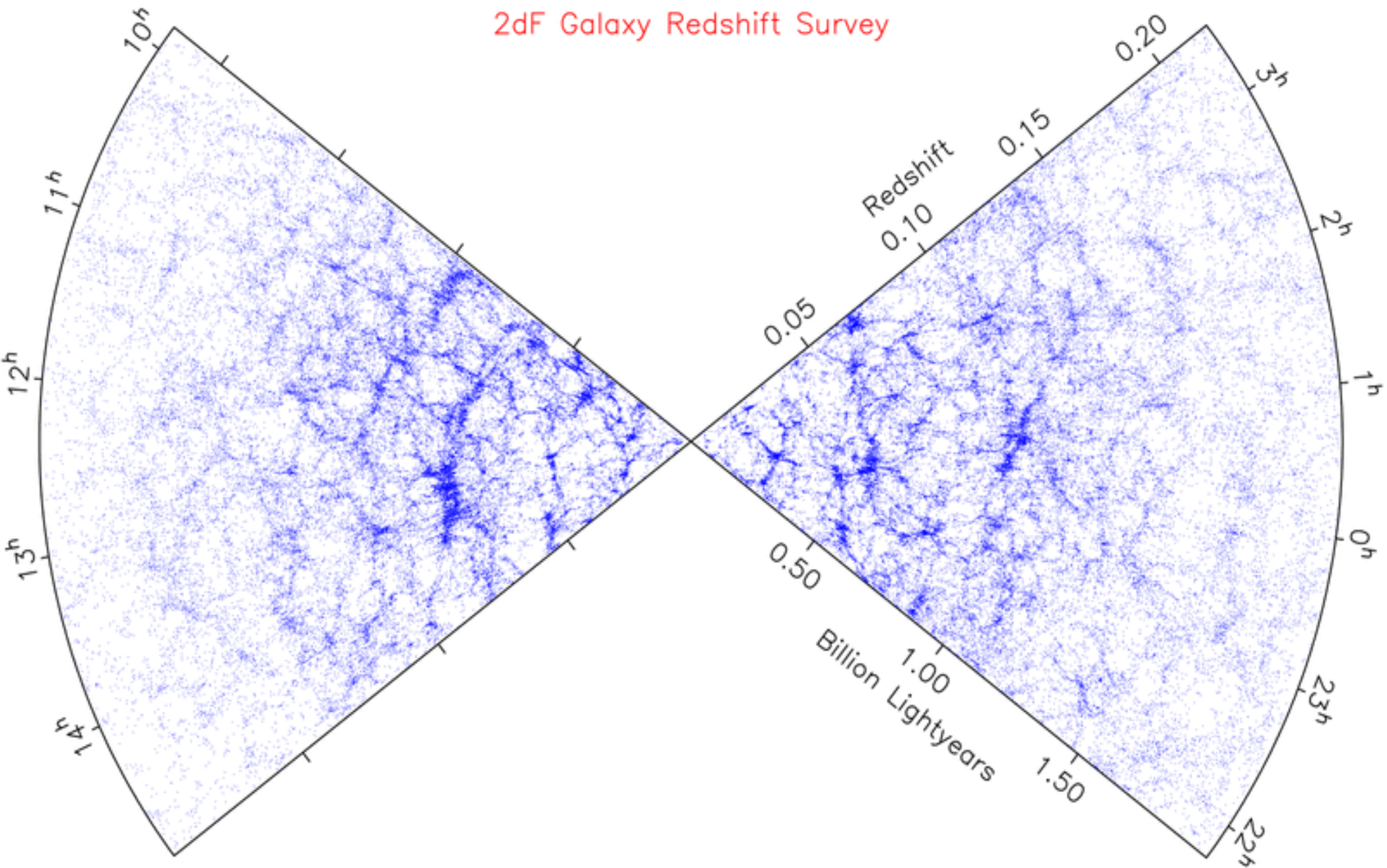
Two Micron All Sky Survey Image Mosaic: Infrared Processing and Analysis Center/Caltech & University of Massachusetts

to 1,000,000,000 light years



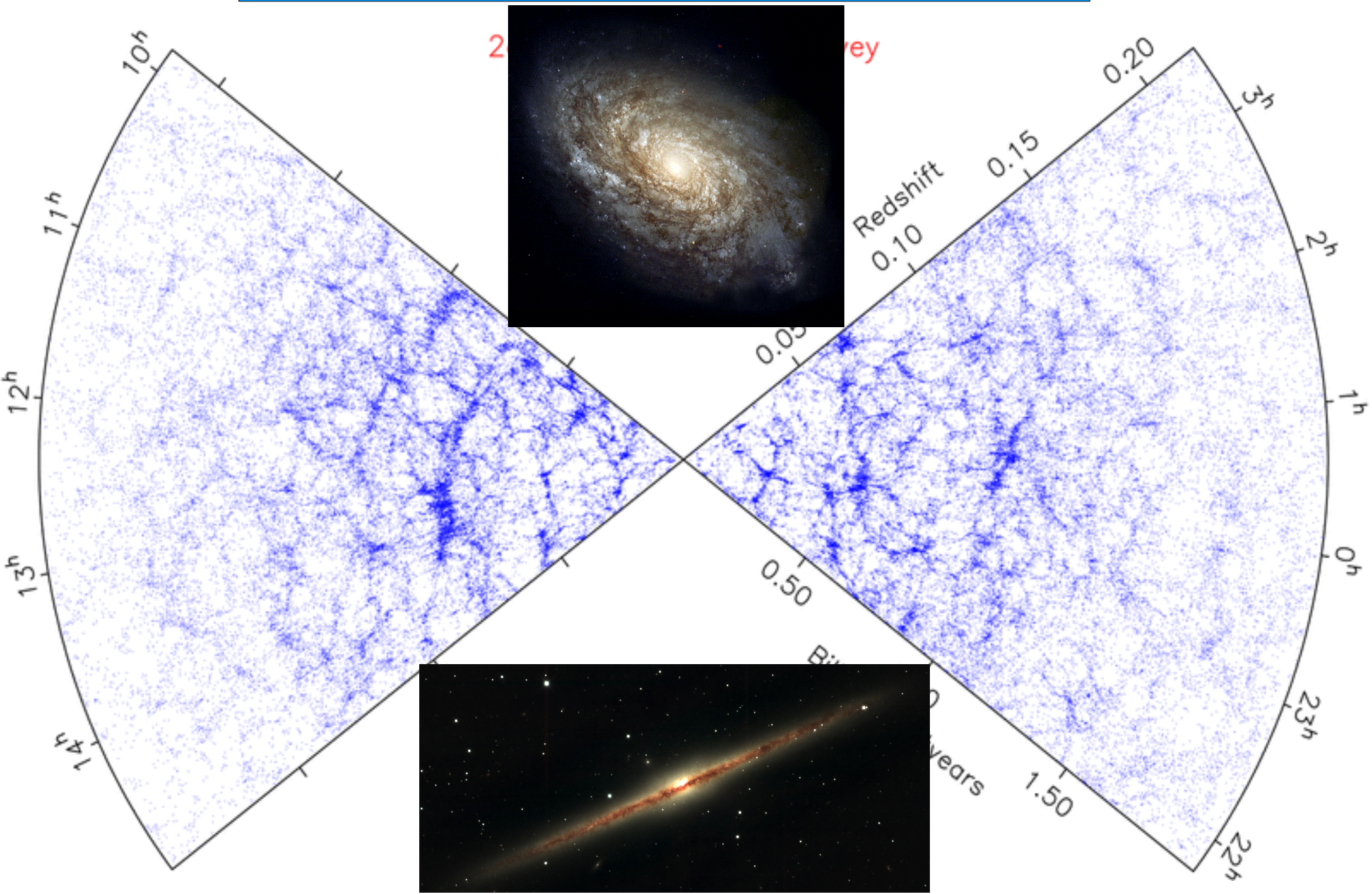
# Nearby large-scale structure

2dF Galaxy Redshift Survey





# Nearby large-scale structure





The deepest  
photo ever made

A 300 hour  
exposure with the  
Hubble Space  
Telescope

Galaxies seen  
back 90% in time  
to the Big Bang

to more than 30,000,000,000 light years



# Hubble Space Telescope



# European Southern Observatory: Very Large Telescope





# The Very Large Array: New Mexico, United States



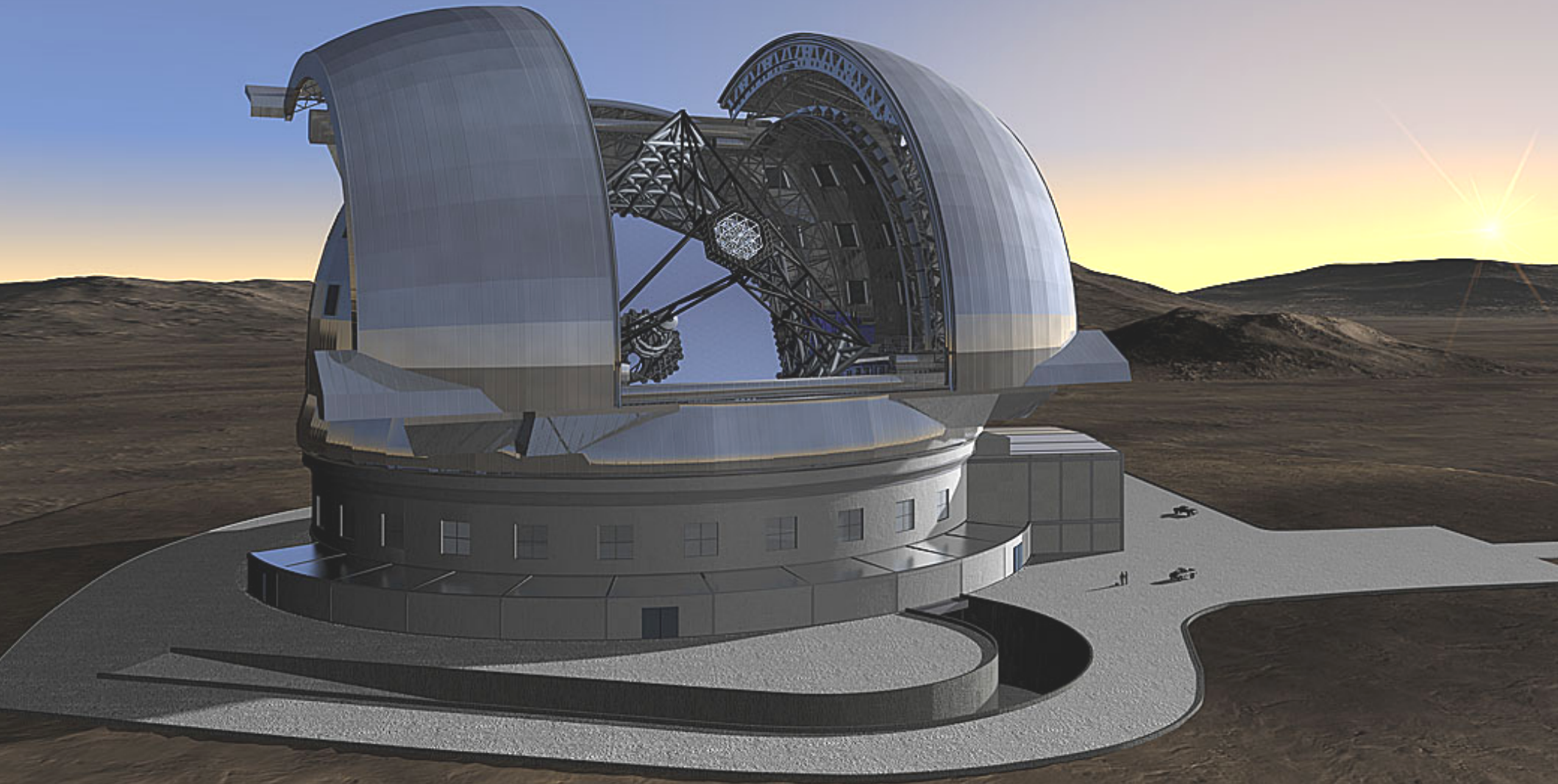
# **The Atacama Large Millimeter Array**

**5000 meter altitude, Chajnantor Plateau, Chile**  
**Europe/US/Japan collaboration**





# Planned 42m European Extremely Large Telescope



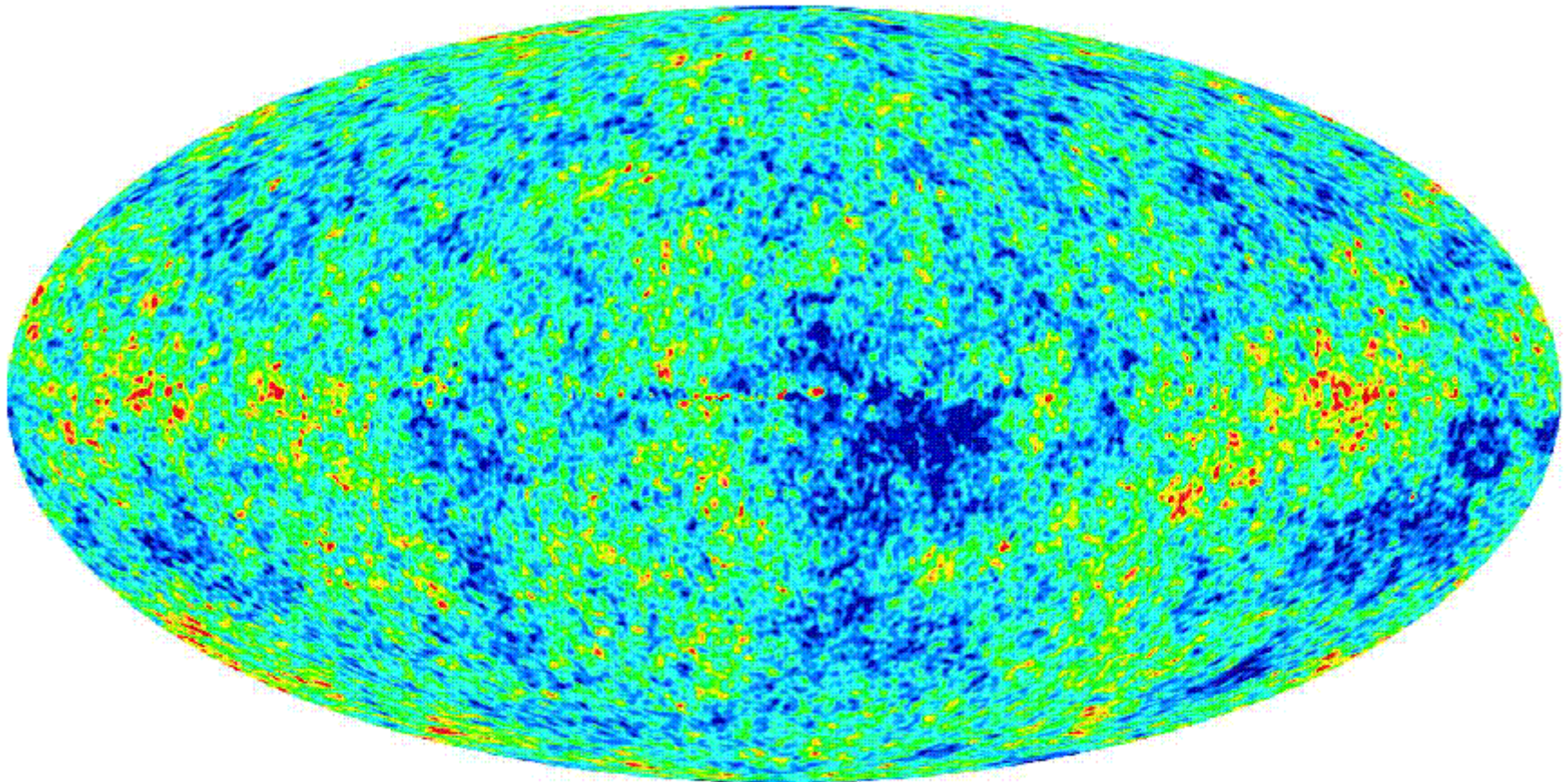




**The 21cm Array Telescope, Xinjiang**



# Map of the Cosmic Microwave Background Our Universe at the age of 400,000 years



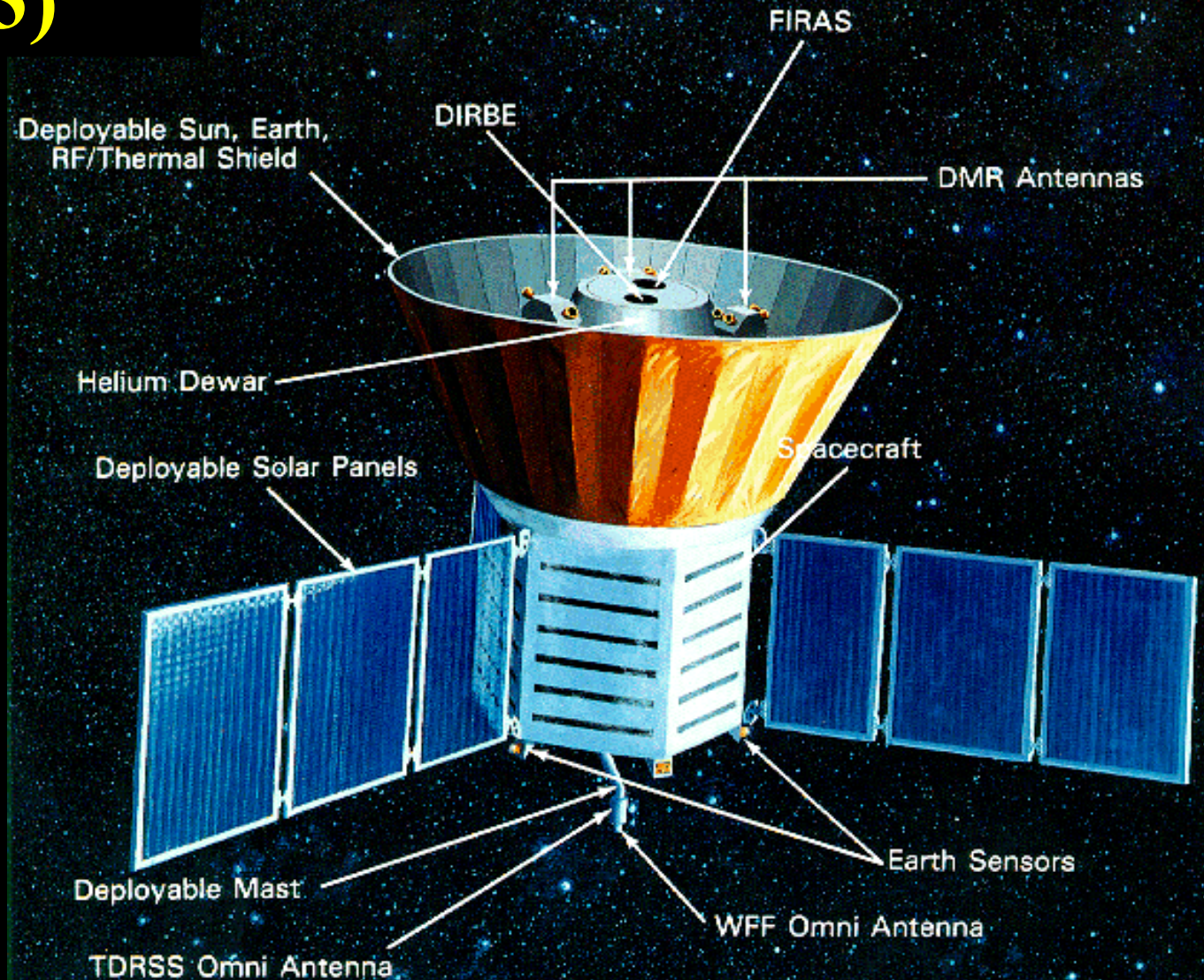
Structure which is 40 billion light-years away today



# The COBE satellite (1989 - 1993)

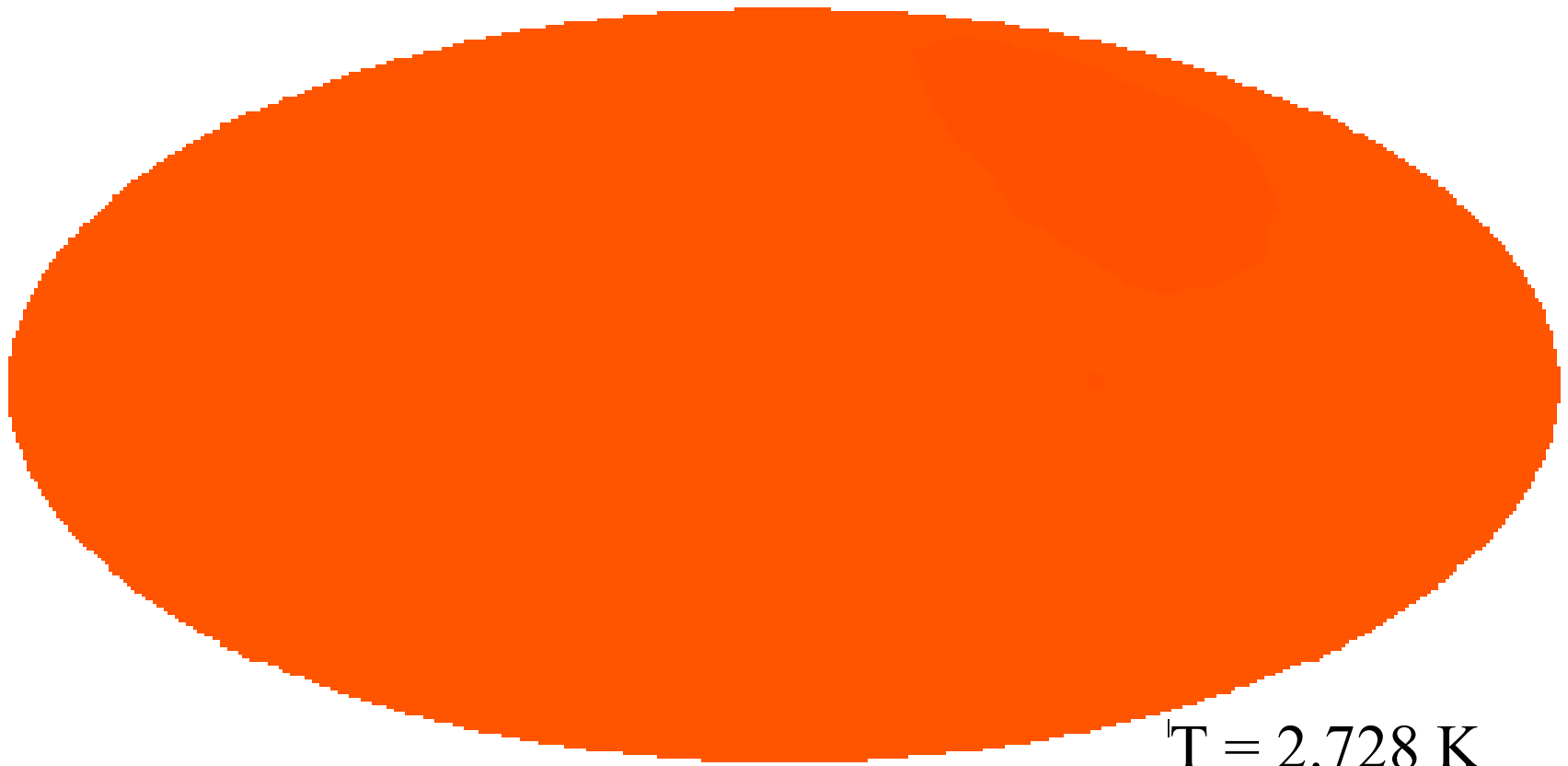
- Two instruments made maps of the whole sky in microwaves and in infrared radiation
- One instrument took a precise spectrum of the sky in microwaves

Nobel Prize in Physics  
2006



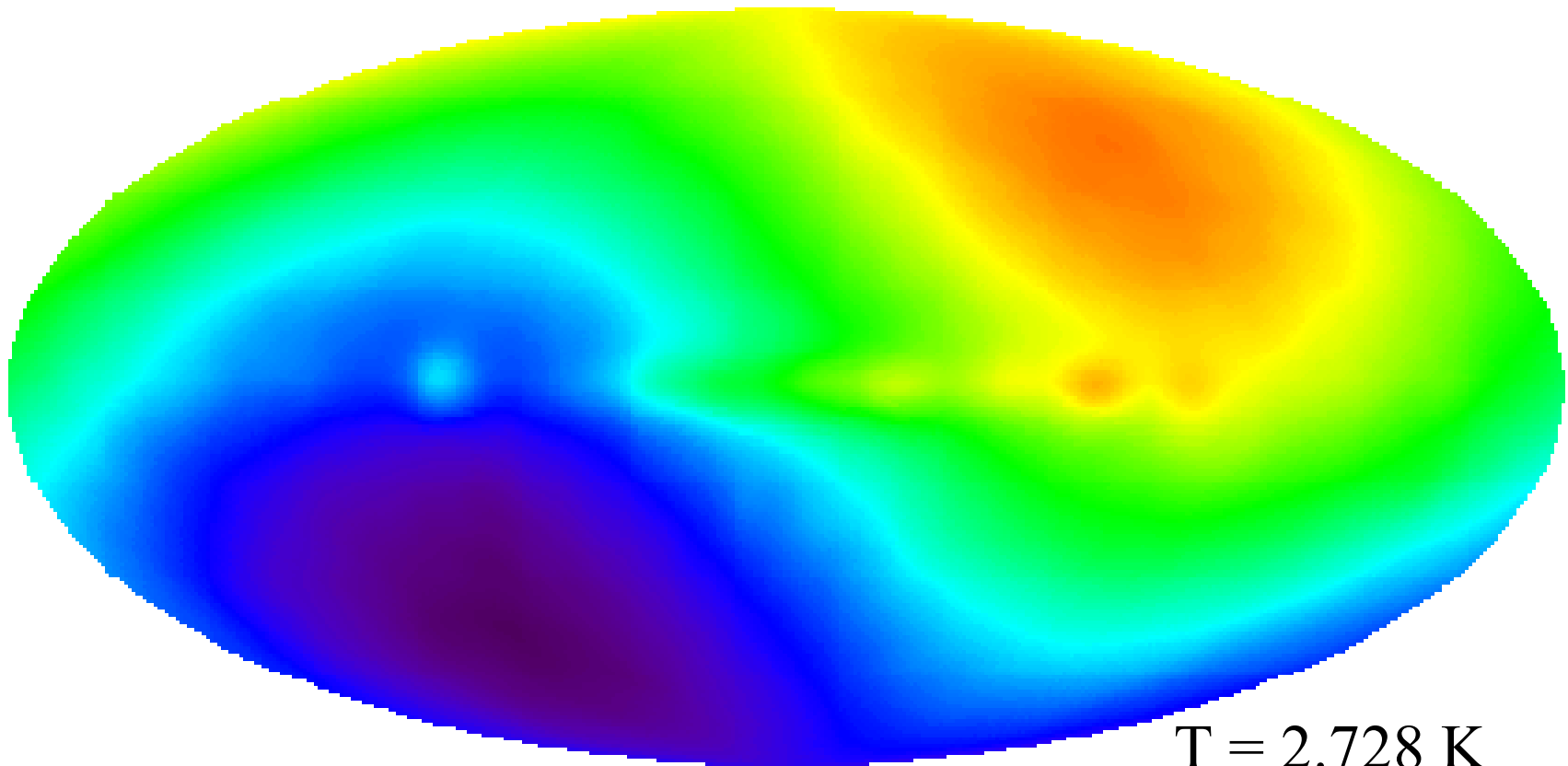


# COBE's temperature map of the entire sky



$T = 2.728 \text{ K}$   
 $\Delta T = 0.1 \text{ K}$

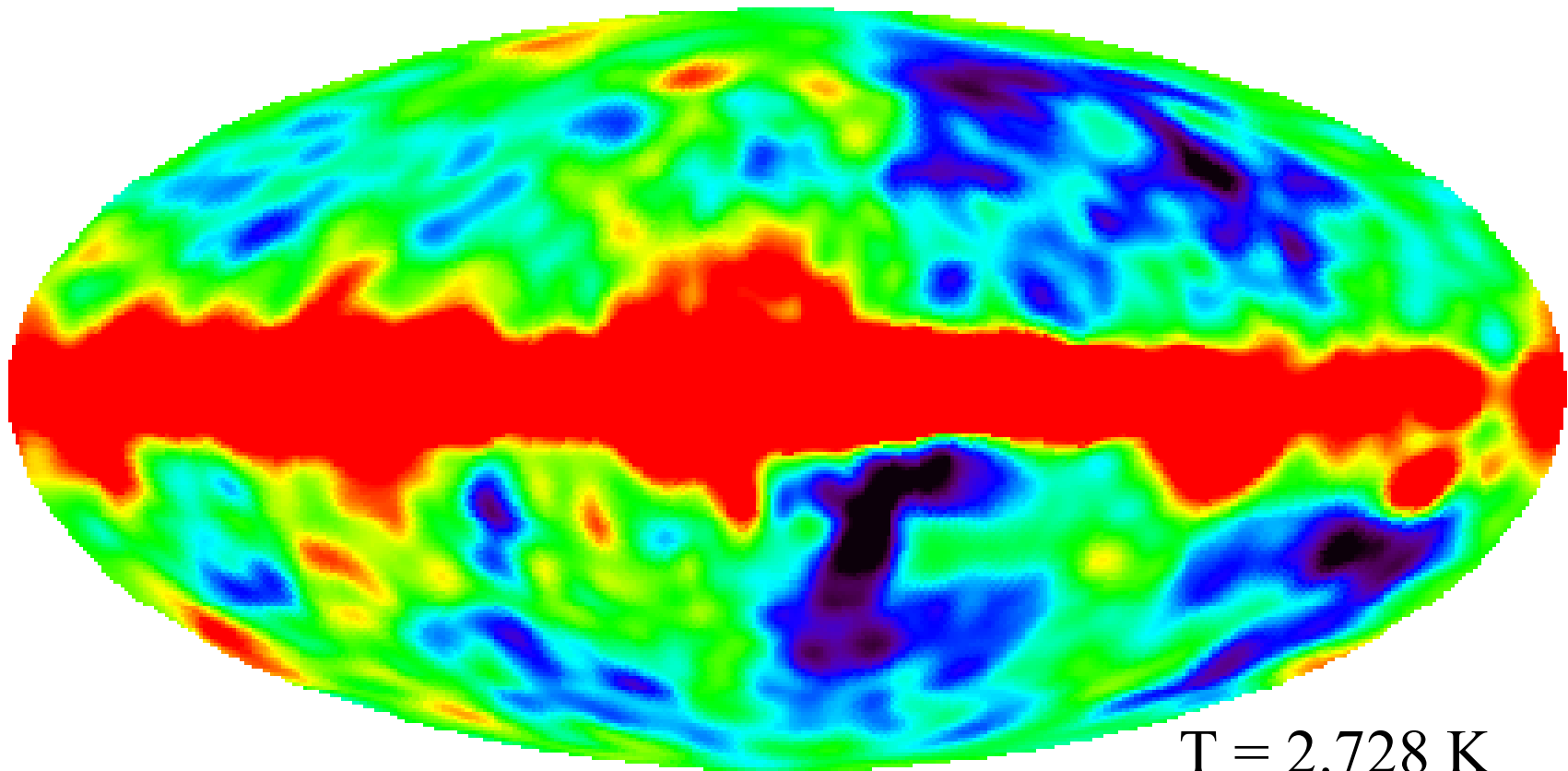
# COBE's temperature map of the entire sky



$T = 2.728 \text{ K}$   
 $\Delta T = 0.0034 \text{ K}$



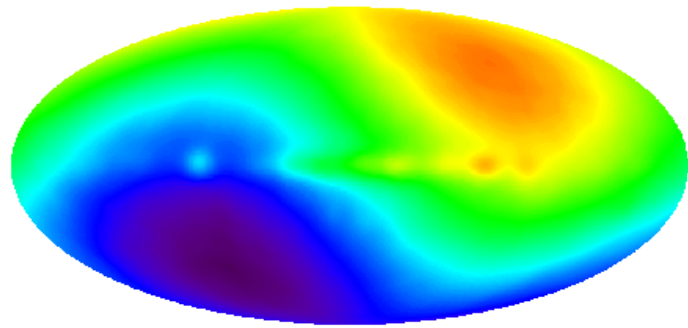
# COBE's temperature map of the entire sky



$$T = 2.728 \text{ K}$$

$$\Delta T = 0.00002 \text{ K}$$

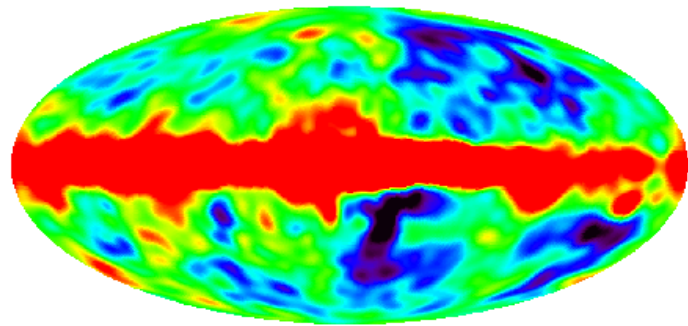
# Structure in the COBE map



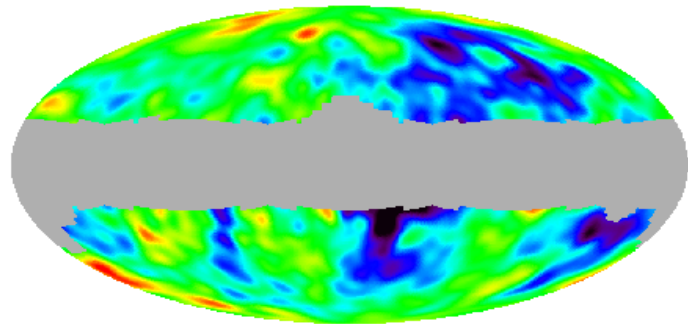
- One side of the sky is 'hot', the other is 'cold'

→ the Earth's motion through the Cosmos

$$V_{\text{Milky Way}} = 600 \text{ km/s}$$



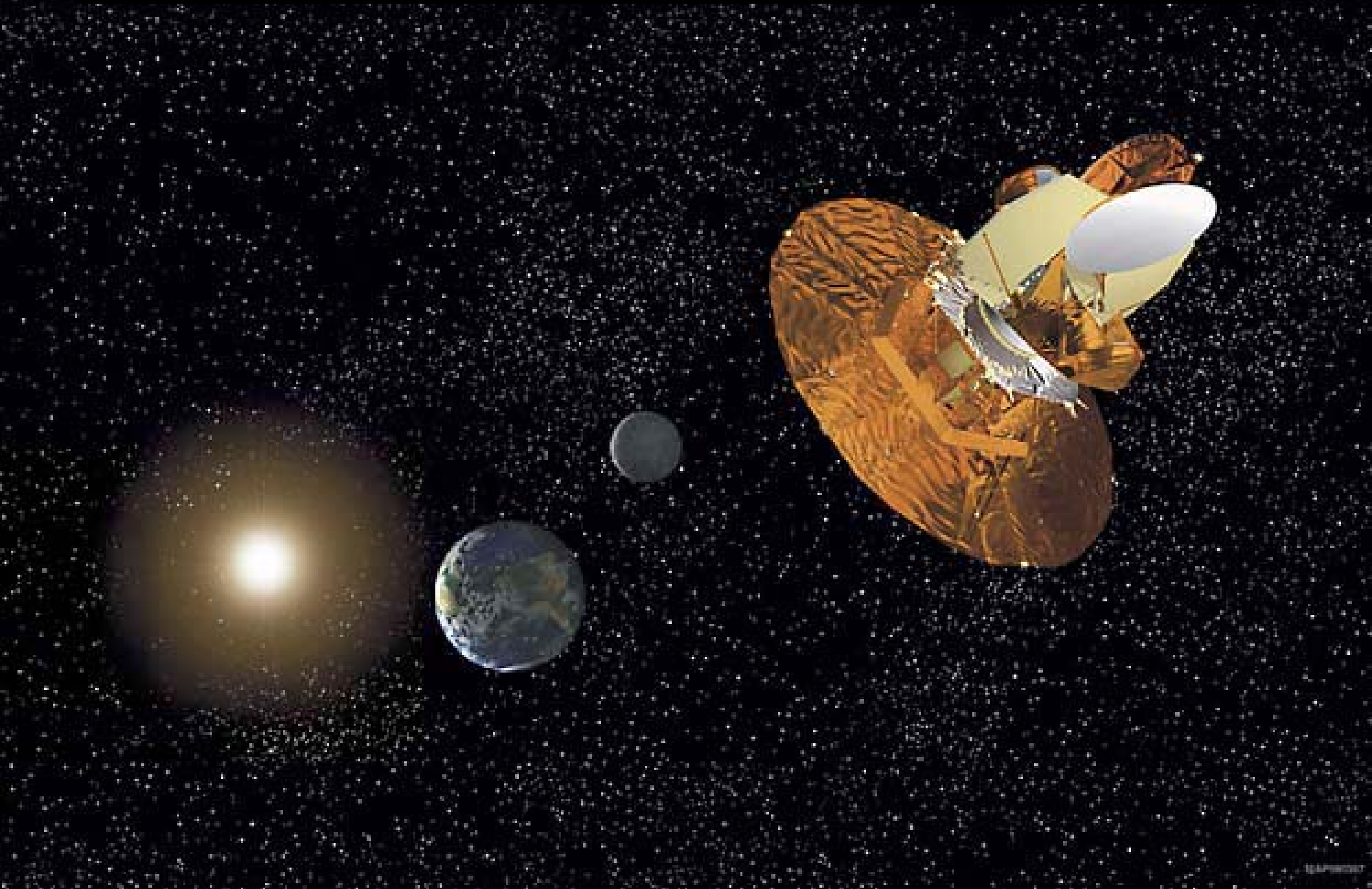
- Radiation from hot gas and dust in our own Milky Way



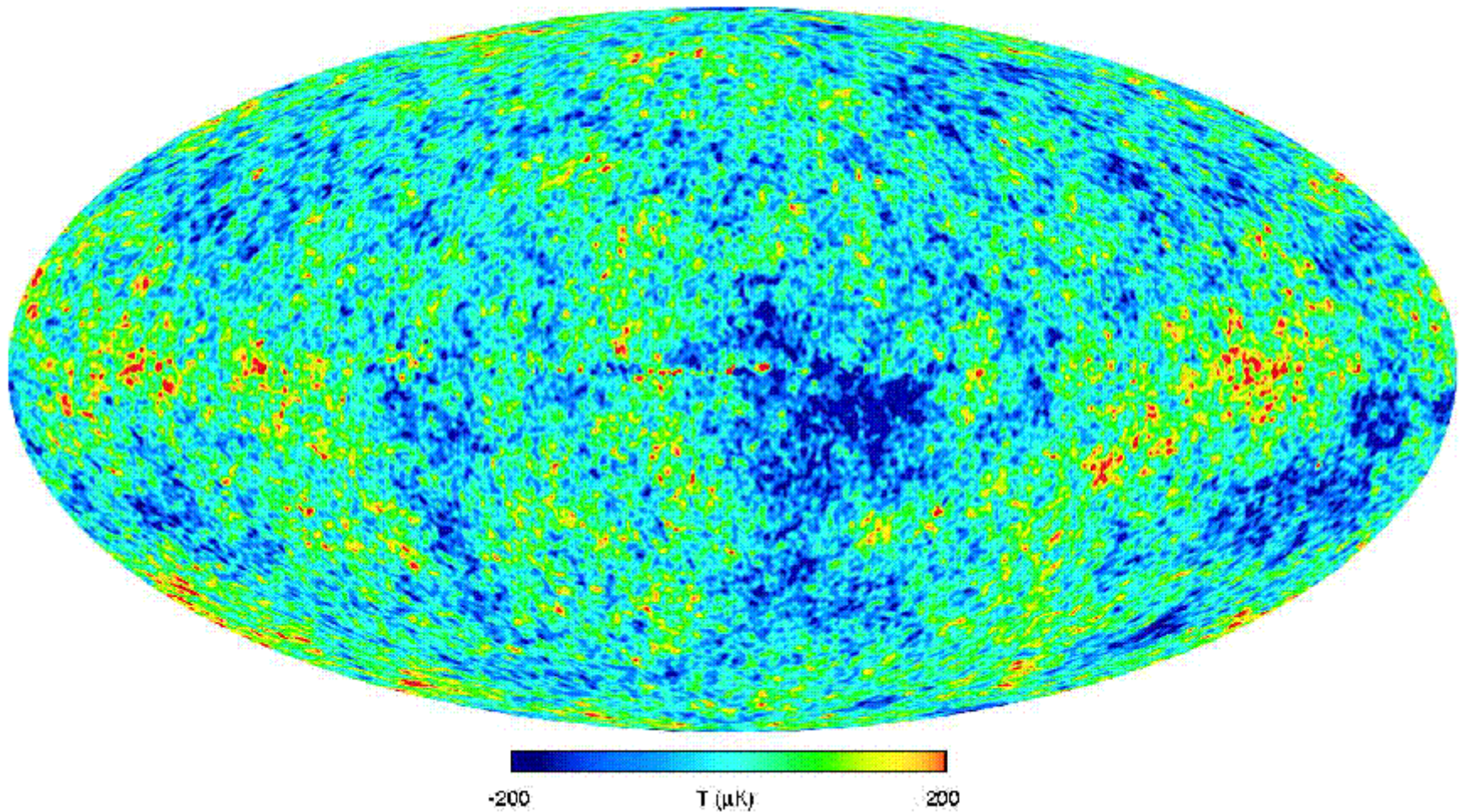
- Structure in the Microwave Background itself



# The *WMAP* Satellite at Lagrange-Point L2



# The *WMAP* of the whole CMB sky



Bennett et al 2003



# Structure in the Microwave Background

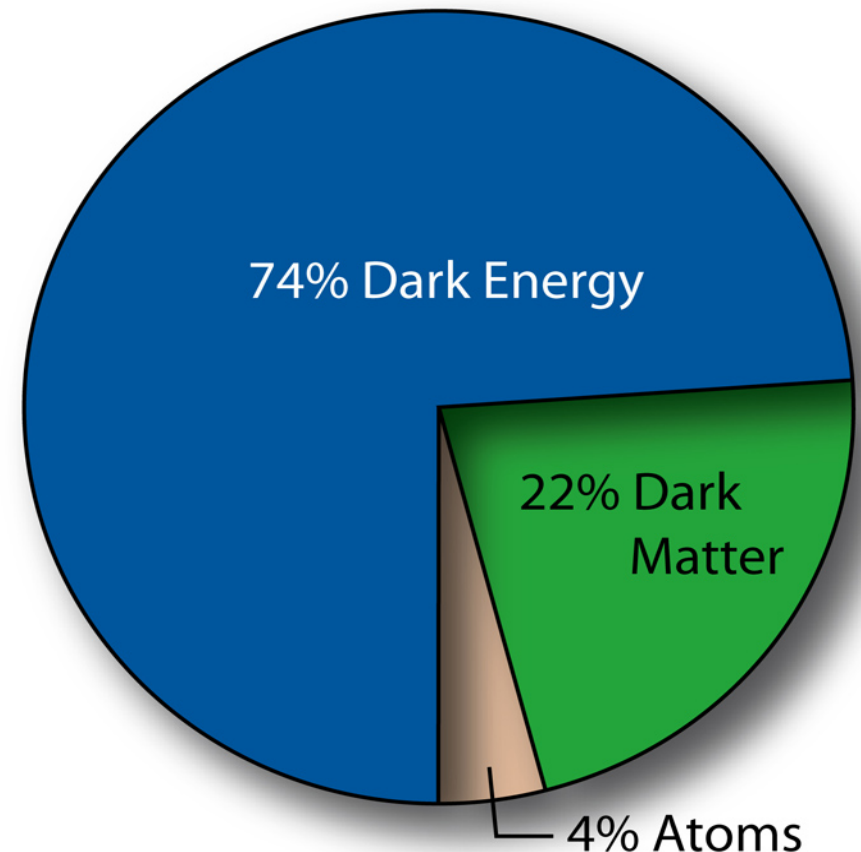
- The structure lies in cosmic “clouds”,  $\sim 4 \times 10^{10}$  l-yrs away
- It reflects weak “sound” waves,  $\Delta \sim 10^{-4}$ , in the clouds
- At the time the Universe was only 400,000 years old, and was 1,000 times smaller and 1,000 times hotter than today

The *pattern* of structure reflects

- A: The global geometry of the Universe
- B: The material content and thermal evolution of the Universe
- C: The process which generated the structure

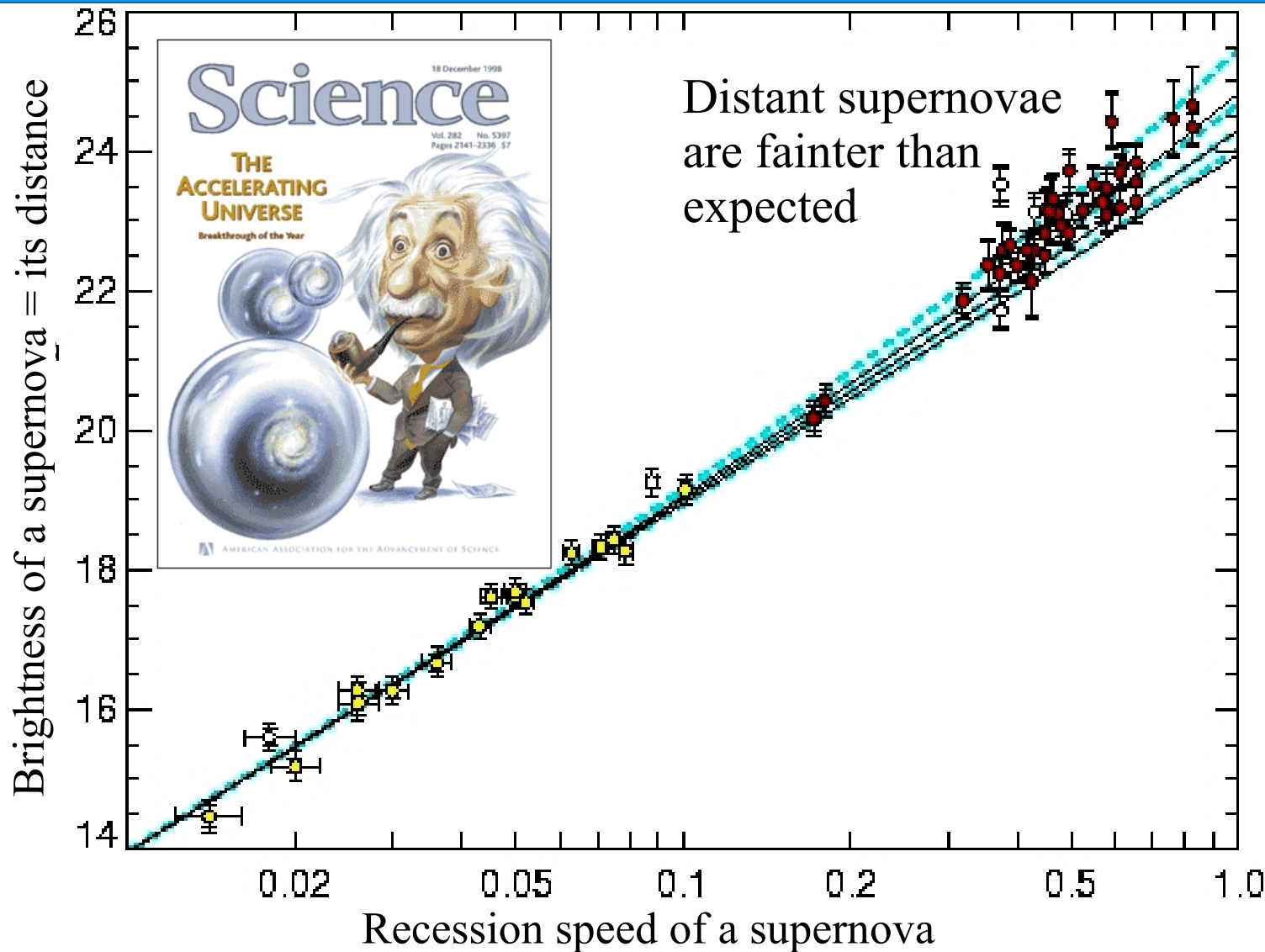
# What has WMAP taught us?

- Our Universe is flat -- its geometry is that described by Euclid
- Only a small fraction is made of ordinary matter -- about 4% today
- About 21% of today's Universe is non-baryonic Dark Matter  
neutralinos? axions? ...?
- About 75% is Dark Energy
- All structure was apparently produced by quantum fluctuations in the vacuum at a very early time



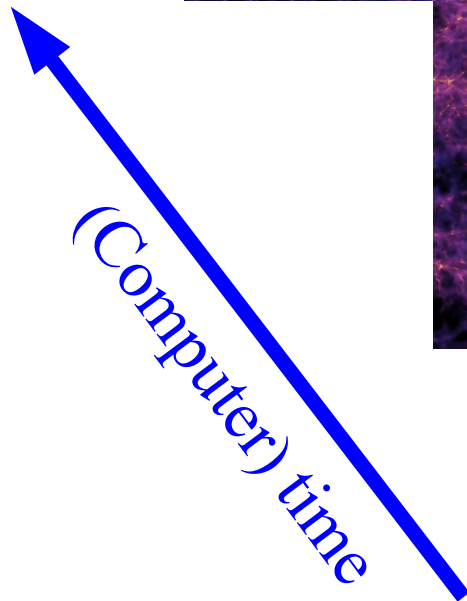
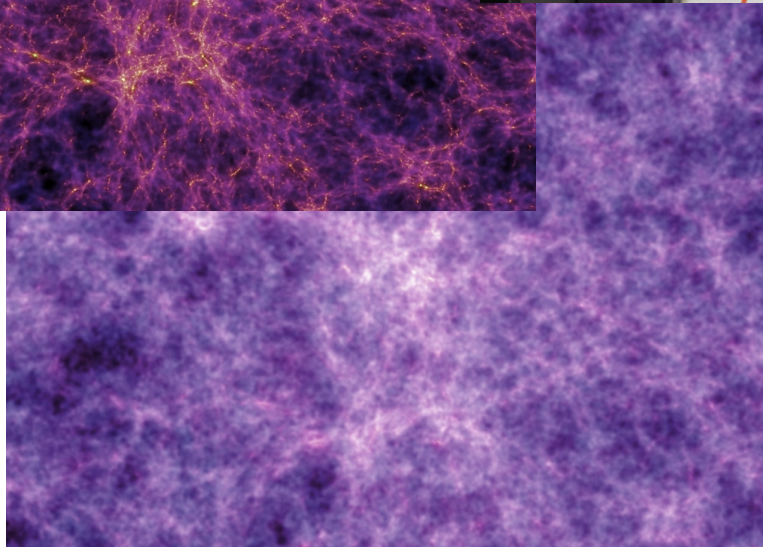
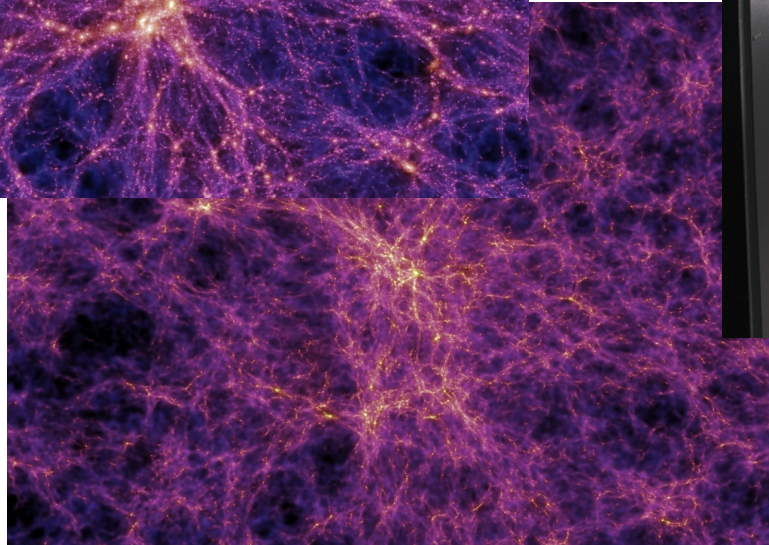
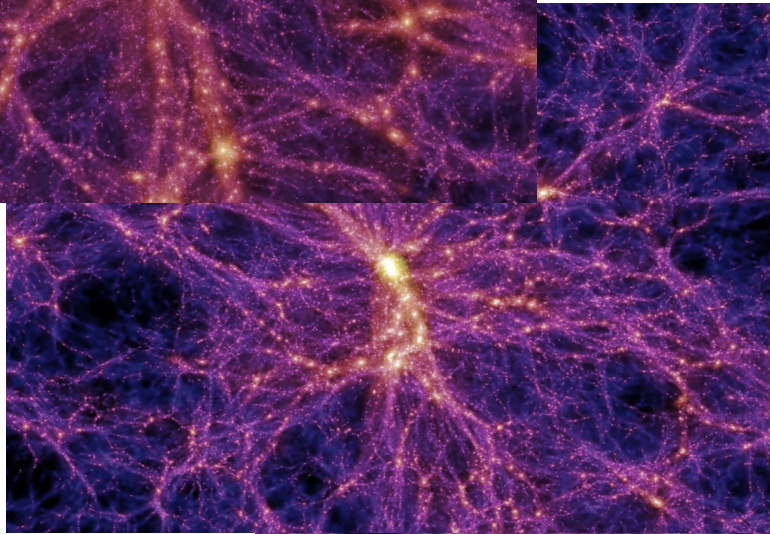
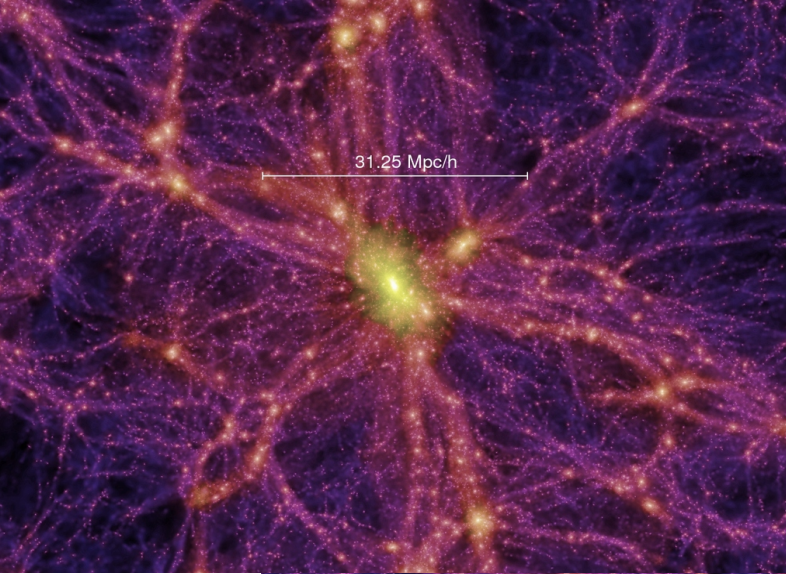


# The Universe expands faster today than in the past!



The cosmic expansion is accelerating! According to Einstein's theory of gravity, this requires negative pressure,  $p = -\rho$  --- Dark Energy

# Virtual universes can run faster than the real Universe







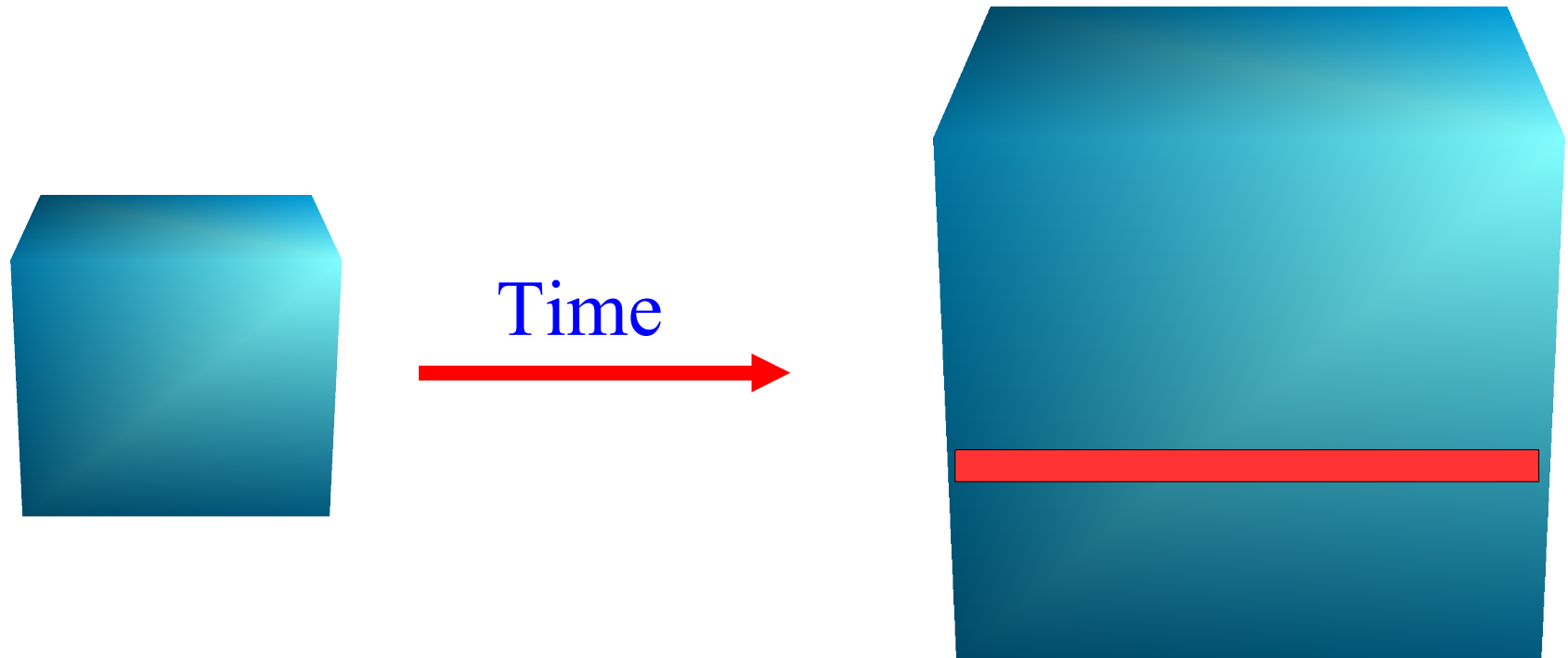
# Leibniz Computer Center, Munich



Since April 2007

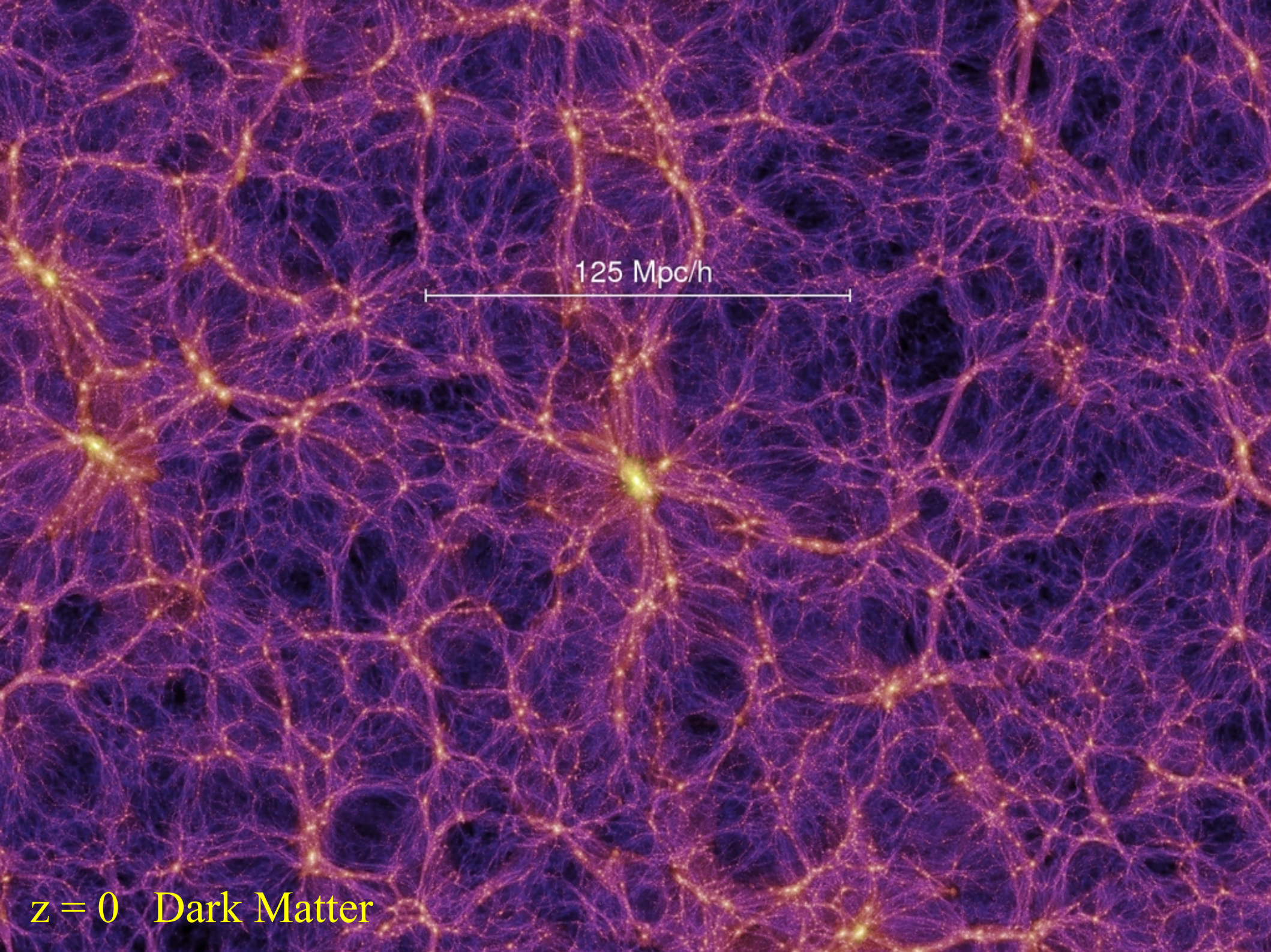
9728 core Altix-4700  
62.3 Tflop/s peak  
39 Tbyte hard memory  
600 Tbyte direct disk

# Evolving the Universe in a computer



- Follow the matter in an expanding cubic region
- Start 300,000 years after the Big Bang
- Match initial conditions to the observed Microwave Background
- Calculate evolution forward to the present day

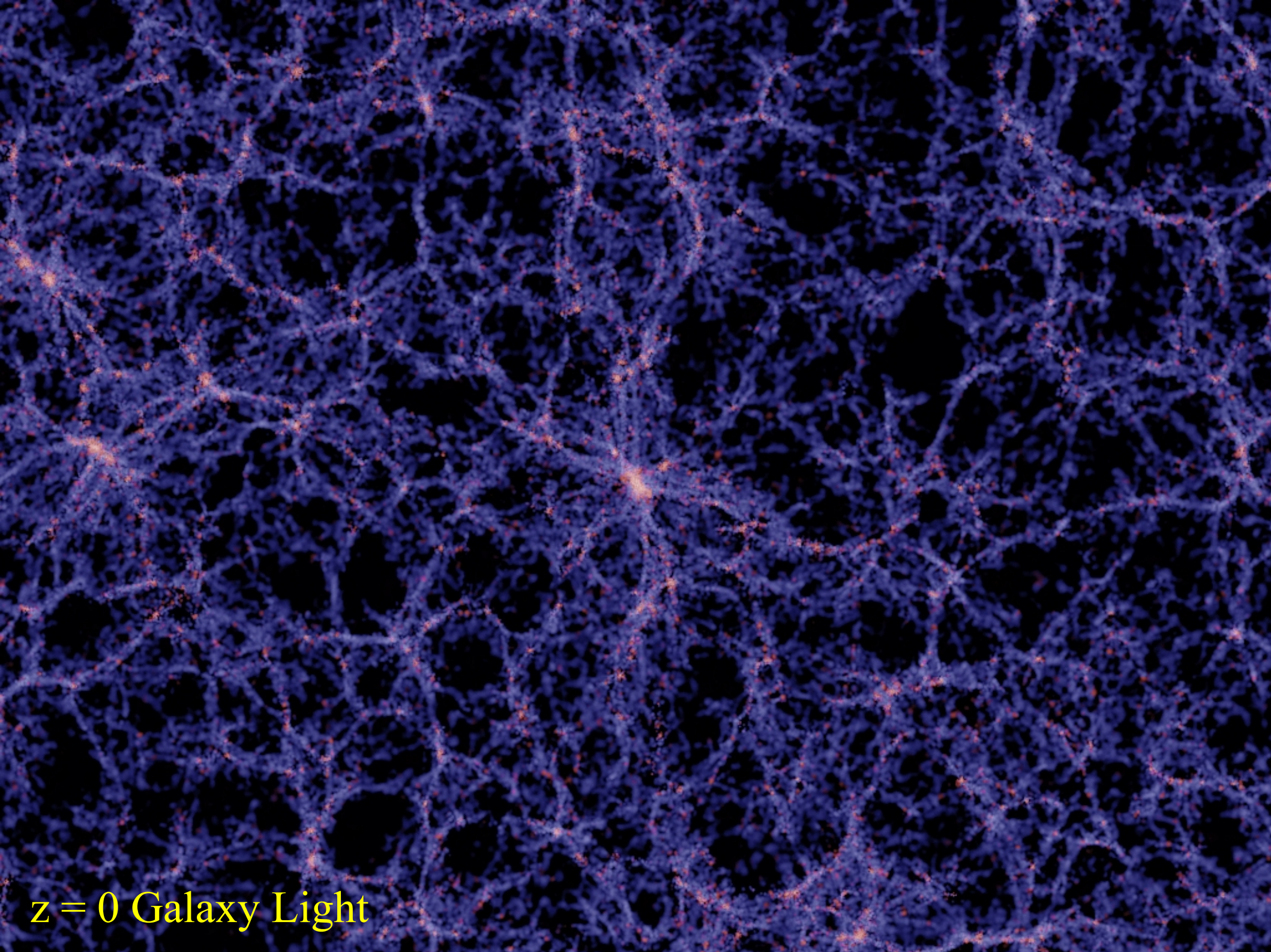




125 Mpc/h

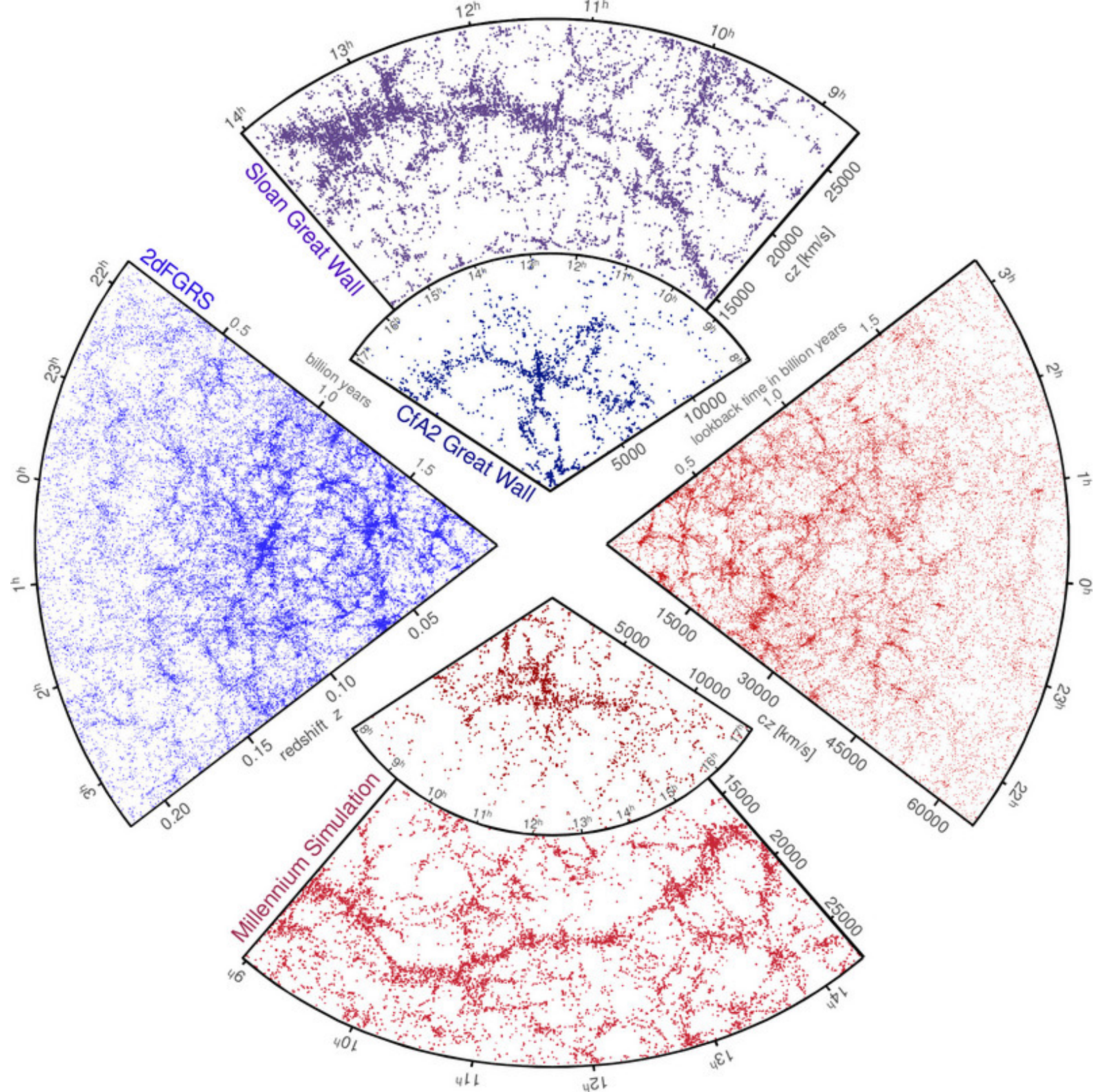
$z = 0$  Dark Matter





$z = 0$  Galaxy Light





# Astrophysics

- puts our Earth in its cosmic context
- allows us to study physics under *very* extreme conditions
- has taught us the origin of the elements
- has shown us what our Universe is made of
- allows us to see back in time, almost to the beginning
- makes cosmic evolution directly visible to us
- provides insight into our long-term future
- is unveiling new worlds, and (perhaps) other life



# The new USTC/MPS Astrophysics Centre

- A joint program of USTC, CAS and the Max Planck Society
- Involves all astronomy/astrophysics Max Planck Institutes
  - Max Planck Institute for Astrophysics
  - Max Planck Institute for extraterrestrial Physics
  - Max Planck Institute for Astronomy
  - Max Planck Institute for Radioastronomy
- Extended Graduate Research School at USTC teaching modern astronomy, astrophysics and cosmology with help of MPS staff
- Research projects joint with MPI's and involving extended stays at the partner institute in Germany (1 to 2 years)



MPI for  
Astrophysics

MPI for  
extraterrestrial Physics







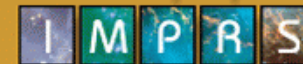
MPI for Radioastronomy



MPI for Astronomy



# The International Max-Planck Research School on Astrophysics at the Ludwig-Maximilians-University Munich



A Marie-Curie Early Stage Training Site of the European Commission

## - Basics

[About us](#)

[Where we are](#)

[IMPRS Lecturers](#)

[IMPRS](#)

[Committees](#)

## - Academic Matters

[Application &](#)

[Admission](#)

[PhD Projects](#)

[Current PhD](#)

[Supervisors](#)

[Degree](#)

[Requirements](#)



Summer in the city -Leopoldstraße on a sunny day in  
Munich



# Should you become a USTC/MPS astronomer?

- Contribute to our understanding of fundamental questions about our Universe and the objects which fill it
- Become part of a global research community
- Participate in China's rapidly growing programme to reach the forefront of modern astrophysics
- Keep ties to USTC, the CAS's university, and one of the best places in China for astrophysics
- Experience european research and life with the MPS, the world's highest ranked research institution in Space Science\*

\* Ahead of NASA, Harvard, Caltech/JPL and University of California, according to the US-based Institute of Scientific Information (2008).