The lecture slides are available at https://wwwmpa.mpa-garching.mpg.de/~komatsu/ lectures--reviews.html

Parity Violation in Cosmology In search of new physics for the Universe

Lecture 1

Eiichiro Komatsu (Max Planck Institute for Astrophysics) Universe+ "Topics in Theoretical Cosmology", MPP July 22, 2025

Reference: EK, Nature Rev. Phys. 4, 452 (2022)





MAX-PLANCK-INSTITUT FÜR ASTROPHYSIK





Overarching Theme Let's find new physics!

- the standard model of elementary particles and fields.
 - What is dark matter (CDM)?
 - What is dark energy (Λ) ?
 - Why is the spatial geometry of the Universe Euclidean (flat)?
 - What powered the Big Bang?

The current cosmological model (*flat ACDM*) requires new physics beyond

Overarching Theme There are many ideas, but how can we make progress?

- The current cosmological model (*flat ACDM*) requires new physics beyond the standard model of elementary particles and fields.
 - What is dark matter (CDM)? => CDM, WDM, FDM, ...
 - What is dark energy (/)? => Dynamical field, modified gravity, quantum gravity, ...
 - Why is the spatial geometry of the Universe Euclidean (*flat*)? => Inflation, contracting universe, ...
 - What powered the Big Bang? => Scalar field, gauge field, ...

New in cosmology! Ove Violation of parity symmetry may hold the answer to these fundamental questions. There

- The current cosmological model (*flat ACDM*) requires new physics beyond the standard model of elementary particles and fields.
 - What is dark matter $(CDM)? => CDM, WDM, FDM, \dots$
 - What is dark energy (Λ) ? => Dynamical field, modified gravity, quantum gravity, ...
 - Why is the spatial geometry of the Universe Euclidean (*flat*)? = Inflation, contracting universe, ...
 - What powered the Big Bang? => Scalar field, gauge field, ...

Reference: nature reviews physics

About the journal \sim Explore content \checkmark

Available also at <u>nature > nature reviews physics > review articles > article</u> arXiv:2202.13919

Review Article | Published: 18 May 2022 New physics from the polarized light of the cosmic microwave background Key Words:

Eiichiro Komatsu 🗠

Nature Reviews Physics 4, 452–469 (2022) Cite this article

Nature Rev. Phys. 4, 452 (2022)

Publish with us \checkmark Subscribe

> **Cosmic Microwave Background** (CMB) Polarization **Parity Symmetry** 3.



The Plan [13:15–15:00] 2 x 45 minutes

- Lecture 1 "Known Physics" [45 min]
- Lecture 2 "New Physics" [45 min]

https://wwwmpa.mpa-garching.mpg.de/~komatsu/lectures--

<u>Break and problem solving (voluntary) [15 min]</u> The lecture slides are available at

reviews.html



1.1 Parity

Probing Parity Symmetry Definition

- **Parity transformation = Inversion of all spatial coordinates**
 - $(x, y, z) \rightarrow (-x, -y, -z)$

- Parity symmetry in physics states:

• The laws of physics are invariant under inversion of all spatial coordinates.

Violation of parity symmetry = The laws of physics are <u>not</u> invariant under...

But, who cares about coordinates? The key is the coordinate transformation

- should not depend on how we chart the world with coordinates."
 - Yes, that is absolutely correct.
- contains useful information.

• You may say, "Coordinates are just a convenient mathematical tool. Physics

 Coordinate transformations are different. The underlying physical principle does not depend on the choice of coordinates. However, "how a physical system appears to change from one coordinate system to another" often

Continuous Coordinate Transformation - 1 Spatial translation and homogeneity

- answer (to within the uncertainty).
- experiment does not change.
 - There is no special location in space => homogeneity.
 - This even implies that the total momentum is conserved!
 - Noether's theorem

• We do an experiment in Munich, and repeat it in Tokyo. We find the same

• This is evidence for invariance under spatial translation. We shift spatial coordinates by a constant vector $c, x \rightarrow x + c$, and the physics relevant to the



https://mathshistory.st-andrews.ac.uk



Continuous Coordinate Transformation - 2 Spatial rotation and isotropy

- apparatus at different angles. We find the same answer (to within the uncertainty).
- physics relevant to the experiment does not change.
 - There is no special direction in space => isotropy.
 - This even implies that the total angular momentum is conserved!
 - Noether's theorem

• We do an experiment. We repeat it a few times after rotating the experimental

• This is evidence for invariance under spatial rotation. We rotate spatial coordinates by $x \rightarrow Rx$, where R is a 3-dimensional rotation matrix, and the



https://mathshistory.st-andrews.ac.uk

Parity: Discrete Coordinate Transformation

observe its mirror image(*) with equal probability?"

where only one of (x,y,z) is flipped.

• We ask, "When we observe a certain phenomenon in nature, do we also

• (*) "Mirror image" is an ambiguous word. A parity transformation is $(x, y, z) \rightarrow x$ (-x, -y, -z), whereas a "mirror image" often refers to, e.g., $(x, y, z) \rightarrow (-x, y, z)$,



Do we also observe this with equal probability?



Note that this is not full parity transformation, as only one axis is flipped.



Parity and Rotation

- Parity transformation $(x \rightarrow x)$ and 3d rotation $(x \rightarrow Rx)$ are different.
 - R is a continuous transformation and the determinant of R is det(R) = +1.
 - Parity is a discrete transformation and the determinant is -1, as







 \boldsymbol{y} 7

z**Parity = Mirror + 2d Rotation** One may think of parity transformation as a mirror in one of the coordinates (e.g., $z \rightarrow -z$) and 2d rotation by π in the others. • Let's demonstrate it!









Rotation ln x-y





1.2 Pseudovector, Pseudoscalar

Parity Transformation: Vector E.g., momentum, electric field



- **p** is the same vector, written using two different basis vectors.

$$\hat{e}'_{y} + p'_{z}\hat{e}'_{z}$$
$$p'_{y}\hat{e}_{y} - p'_{z}\hat{e}_{z}$$

• Therefore, **p**'s components are transformed as $(p'_x, p'_y, p'_z) = (-p_x, -p_y, -p_z)$



Parity Transformation: Pseudovector E.g., angular momentum, magnetic field

- - change sign. Thus, their products do not change, e.g.,

 $L'_x = Y'p'_z - Z'p'_u$ $= (-Y)(-p_z) - (-Z)$ $=L_x$

• Orbital angular momentum, $\mathbf{L} = \mathbf{r} \times \mathbf{p}$, is a *pseudovector*. Its components do <u>not</u> change under parity transformation: $(L'_x, L'_y, L'_z) = (L_x, L_y, L_z)$

• Both $\mathbf{r} = (X, Y, Z)$ and $\mathbf{p} = (p_x, p_y, p_z)$ are vectors whose components



Parity Transformation: Pseudoscalar How to test parity symmetry?

- A dot product of a vector and a pseudovector is a pseudoscalar.
 - Like a scalar, a pseudoscalar is invariant under rotation.
 - But, a pseudoscalar changes sign under parity transformation.
- Experimental test of parity symmetry: Construct a pseudoscalar and see if the average value is zero. If not, the system violates parity symmetry!
 - <u>Example</u>: a dot product of particle A's momentum and particle B's angular momentum: $\mathbf{p}_A \cdot \mathbf{L}_B$. Measure this and average over many trials. Does the average vanish, $\langle \mathbf{p}_A \cdot \mathbf{L}_B \rangle = 0$?



1.3 Discovery of Parity Violation in β -decay (weak interaction)

\sim **SOO** L) \mathcal{O} S etters to, l Review, J Physic

Experimental Test of Parity Conservation in Beta Decay*

C. S. WU, Columbia University, New York, New York

AND

E. AMBLER, R. W. HAYWARD, D. D. HOPPES, AND R. P. HUDSON, National Bureau of Standards, Washington, D. C. (Received January 15, 1957)

 \mathbf{T} N a recent paper¹ on the question of parity in weak I interactions, Lee and Yang critically surveyed the experimental information concerning this question and reached the conclusion that there is no existing evidence either to support or to refute parity conservation in weak interactions. They proposed a number of experiments on beta decays and hyperon and meson decays which would provide the necessary evidence for parity conservation or nonconservation. In beta decay, one could measure the angular distribution of the electrons coming from beta decays of polarized nuclei. If an asymmetry in the





The Wu Experiment of β-decay ${}^{60}Co -> {}^{60}Ni + e^{-} + \overline{v}_e + 2\gamma$



- parity symmetry is respected in β -decay.

Wu et al. (1957)

• Electrons must be emitted with equal probability in all directions relative to J, if

• This was not observed: $\langle \mathbf{p}_e \cdot \mathbf{J} \rangle \neq 0$. Parity symmetry is violated in β -decay!





"This Month in Physics History", APS News, October 2022 **Initial reaction** Many physicists did not believe it initially.

- To Lee and Yang's theoretical paper on parity violation in β -decay:
- To Wu's discovery paper:
 - exciting. How sure is this news?)
- left and right!



• Wolfgang Pauli said, "Ich glaube aber nicht, daß der Herrgott ein schwacher Linkshänder ist" (I do not believe that the Lord is a weak left-hander).

• Wolfgang Pauli said, "Sehr aufregend. Wie sicher ist die Nachricht?" (Verv

This was shocking news. The weak interaction distinguishes between

• In this lecture we ask, "Does the Universe distinguish between left and right?"



1.4 Helicity

Momentum and spin of a massless particle



handed) states as



• Imagine a photon (γ) traveling at the speed of light with a momentum vector p.

• The photon is a spin-1 particle. The spin angular momentum vector **S** is a pseudovector. We then define "parallel" (right-handed) and "antiparallel" (left-

- p "Parallel" (right-handed): $\mathbf{S} \cdot \mathbf{p} > 0$
 - \rightarrow p "Antiparallel" (left-handed): $\mathbf{S} \cdot \mathbf{p} < 0$



Momentum and spin of a massless particle

Imagine a photon (y) traveling at the speed of light with a momentum vector \mathbf{p} .



• The photon is a spin-1 particle. The spin angular momentum vector S is a pseudovector. We then define "parallel" (right-handed) and "antiparallel" (lefthanded) states as



- $\mathbf{S} \rightarrow \mathbf{p}$ "Parallel" (right-handed): $\mathbf{S} \cdot \mathbf{p} > 0$
 - • "Antiparallel" (left-handed): $\mathbf{S} \cdot \mathbf{p} < 0$





Momentum and spin of a massless particle



handed) states as



Imagine a photon (γ) traveling at the speed of light with a momentum vector p.

• The photon is a spin-1 particle. The spin angular momentum vector S is a pseudovector. We then define "parallel" (right-handed) and "antiparallel" (left-

 $\mathbf{S} \rightarrow \mathbf{p}$ "Parallel" (right-handed): $\mathbf{S} \cdot \mathbf{p} > 0$

P "Antiparallel" (left-handed): $\mathbf{S} \cdot \mathbf{p} < 0$ 30



Helicity is a pseudoscalar Party transformation changes "right-handed" to "left-handed" and vice versa

• For massless particles, we define the "helicity", λ , as

S'=S

• For a photon, $\lambda = \pm 1$.





- λ is a pseudoscalar because it is a product of a momentum vector (p) and a spin pseudovector (S).
 - On the other hand, "scalar", such as \mathbf{p}^2 and \mathbf{S}^2 , does not change sign.
- For a graviton, $\lambda = \pm 2$.
- Asymmetry between $\lambda = \pm 1$ and ± 2 is the sign of parity violation!

Helicity and circular polarization

- $\lambda = \pm 1$ states of a photon describe circular polarization.
 - $\lambda = +1$: Right-handed circular polarization
 - $\lambda = -1$: Left-handed circular polarization

- Linear polarization is given by a super position of $\lambda = \pm 1$ states with the equal number of photons. Linear polarization carries no angular momentum!
 - Circularly polarized light carries angular momentum, which is equal to $(N_{+} - N_{-})\hbar$ where N_{+} is the number of $\lambda = \pm 1$ photons.



This means that circular polarization is a more fundamental state for photons.

1.5 Parity Symmetry in Electromagnetism (EM)

Throughout this talk, I will assume homogeneity and isotropy of space (invariance under 3d translation and rotation).

Maxwell's Equations In Minkowski space, Heaviside units and c=1

$$abla \cdot \mathbf{E} =
ho \,, \qquad -\dot{\mathbf{E}}$$
 $abla \cdot \mathbf{B} = 0 \,, \qquad \dot{\mathbf{B}}$

 These equations are invariant under Poincaré transformation (spatial) translation and rotation and Lorentz boost).

$\mathbf{D} + \nabla \times \mathbf{B} = \mathbf{j}$ $\mathbf{S} + \nabla \times \mathbf{E} = 0$



Throughout this talk, I will assume homogeneity and isotropy of space (invariance under 3d translation and rotation).

Parity-flipping Maxwell's Equations In Minkowski space, Heaviside units and c=1

$$(-\nabla) \cdot (-\mathbf{E}) = \rho, \quad -(-\dot{\mathbf{E}}) + (-\nabla) \times \mathbf{B} = (-\mathbf{j})$$
$$(-\nabla) \cdot \mathbf{B} = 0, \qquad \dot{\mathbf{B}} + (-\nabla) \times (-\mathbf{E}) = 0$$

- These equations are invariant under Poincaré transformation (spatial) translation and rotation and Lorentz boost).
- a scalar, and **B** is a pseudovector.

They are also invariant under parity transformation, if E and j are vectors, ρ is



Throughout this talk, I will assume homogeneity and isotropy of space (invariance under 3d translation and rotation).

Maxwell's Equations in a covariant form









Antisymmetric Field Strength Tensor, F^{µv} $F^{\mu\nu} = -F^{\nu\mu}$



• Equivalently,

$$F^{0i} = E_i$$
$$F^{ij} = \epsilon^{ijk} B_k$$

 $F^{12} = B_z, F^{23} = B_x, F^{31} = B_y$ $F^{21} = -B_z, F^{32} = -B_x, F^{13} = -B_y$ symbol

 $\epsilon^{ijk} = \begin{cases} +1 \text{ if } (i,j,k) \text{ is even permutation of } (1,2,3) \\ -1 \text{ if } (i,j,k) \text{ is odd permutation of } (1,2,3) \\ \text{Levi-Civita} \end{cases}$ otherwise $\epsilon^{123} = 1, \epsilon^{132} = -1, \epsilon^{312} = 1, \dots$



Antisymmetric Field Strength Tensor, F_{µv} $F_{\mu\nu} = -F_{\nu\mu}$

$$F_{\mu
u} = \eta_{\mulpha}\eta_{
ueta}F^{lphaeta}$$
 whe



Therefore,

 $\equiv F_{\mu\nu}F^{\mu\nu} = 2(\mathbf{B}\cdot\mathbf{B} - \mathbf{E}\cdot\mathbf{E})$ This is a *scalar* and is invariant under parity transformation.

ere $\eta_{\mu\alpha} = diag(-1, 1, 1, 1)$



Dual Field Strength T $\widetilde{F}^{\mu\nu} = -\widetilde{F}^{\nu\mu}$ $\widetilde{F}^{\mu\nu} = \frac{1}{2} \epsilon^{\mu\nu\alpha\beta} F_{\alpha\beta} \quad \text{where} \quad \epsilon_{L}$



• Therefore,

 $F\bar{F} \equiv F_{\mu\nu}\tilde{F}^{\mu\nu} = -4\mathbf{B}\cdot\mathbf{E}$

$$e^{\mu\nu\alpha\beta} = \begin{cases} +1 & \text{if } (\mu,\nu,\alpha,\beta) \text{ is even } p \\ +1 & \text{of } (0,1,2,3) \\ -1 & \text{if } (\mu,\nu,\alpha,\beta) \text{ is odd } p \\ -1 & \text{of } (0,1,2,3) \\ 0 & \text{otherwise} \end{cases}$$

$$B_{y} \quad B_{z} \\ B_{z} \quad E_{z} \quad E_{y} \\ 0 & -E_{x} \\ D_{x} \quad 0 \end{cases} \bullet \text{Equivalently,}$$

$$\tilde{F}^{0i} = B_{i} \\ \tilde{F}^{ij} = -\epsilon^{ijk}$$

This is a *pseudoscalar* and changes sign under parity transformation!







1.6 Action Principle for EM

In Heaviside units and c=1

• The answer is

 $I=-rac{1}{\Lambda}\int d^4x\;F^2+\int d^4x\;A_\mu j^\mu \quad d^4x=dtd^3{f x}$

with

Vector potential $F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$ where $A_{\mu} = (-\phi, \mathbf{A})$ Therefore, $\begin{cases} F_{i0} = \partial_i A_0 - \dot{A}_i = E_i \\ F_{ij} = \partial_i A_j - \partial_j A_i = \epsilon_{ijk} B_k \end{cases} \Rightarrow \mathbf{E} = -\nabla \phi - \dot{\mathbf{A}} \\ \mathbf{B} = \nabla \times \mathbf{A} \end{cases}$

$\partial_{\nu}F^{\mu\nu} = j^{\mu}, \quad \partial_{\nu}\tilde{F}^{\mu\nu} = 0$ What is the action that gives Maxwell's equations?





In Heaviside units and c=1

• The answer is

$$I = -\frac{1}{4} \int d^4x \ F^2 + \int$$

with

 $F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$ where $A_{\mu} = (-\phi, \mathbf{A})$

One set of Maxwell's equations is simply given by the definition of $F_{\mu\nu}$: $\partial_{\nu}\tilde{F}^{\mu\nu} = \partial_{\nu}\frac{1}{2}\epsilon^{\mu\nu\alpha\beta}(\partial_{\alpha}A_{\beta} - \partial_{\beta}A_{\alpha}) = \epsilon^{\mu\nu\alpha\beta}\partial_{\nu}\partial_{\alpha}A_{\beta} = 0$

$\partial_{\nu}F^{\mu\nu} = j^{\mu}, \quad \checkmark \partial_{\nu}\tilde{F}^{\mu\nu} = 0$ What is the action that gives Maxwell's equations?

 $d^4x A_{\mu}j^{\mu} d^4x = dtd^3\mathbf{x}$







In Heaviside units and c=1

$$I = -\frac{1}{4} \int d^4x \ F^2 + \int$$

stationary point. For a small change in $A_{\mu} \rightarrow A_{\mu} + \delta A_{\mu}$, the corresponding change in $I \rightarrow I + \delta I$ is also small.

$$\delta I = \int d^4 x \; F^{\mu
u} \partial_
u (\delta A_\mu) + \int \mathbf{Integration} \ \int d^4 x \; (-\partial_
u F^{\mu
u} + j^\mu) \delta .$$



 $\int d^4x A_{\mu} j^{\mu}$

• The idea: The equation of motion for A_{μ} is the path that gives a

 $\int d^4x \ (\delta A_{\mu}) j^{\mu} \qquad \stackrel{\text{Hint:}}{\stackrel{\delta}{\delta}(F^2)} = 2F^{\mu\nu} \delta F_{\mu\nu} \\ = -4F^{\mu\nu} \partial_{\nu} (\delta A_{\mu})$ $A_{\mu} = 0 \implies \partial_{\nu} F^{\mu\nu} = j^{\mu}$



1.7 Propagation of EM Waves

Finding symmetries in the action It is like a treasure hunt!

$$I = -\frac{1}{4} \int d^4x \ F^2 + \int$$

- This action is invariant under spatial translation, rotation, and parity transformation.
- It is also invariant under the following "gauge transformation",

$$A_{\mu} \to A_{\mu} + \partial_{\mu} f$$

• Here, *f* is an arbitrary scalar function.

 $\int d^4x A_{\mu} j^{\mu}$

Integration by parts Hint: $\int d^4x \ (\partial_\mu f) j^\mu \stackrel{\bullet}{=} - \int d^4x \ f \partial_\mu j^\mu = 0$ due to the charge conservation: $\partial_{\mu} j^{\mu} = 0 \implies \dot{\rho} + \nabla \cdot \mathbf{j} = 0$ 45



Finding symmetries in the action It is like a treasure hunt!

$$I = -\frac{1}{4} \int d^4x \ F^2 + \int$$

- This action is invariant under spatial translation, rotation, and parity transformation.

$$\phi \to \phi - f$$

 It is also invariant under the following "gauge transformation", Integration by parts Hint: $\int d^4x \ (\partial_\mu f) j^\mu \stackrel{\bullet}{=} - \int d^4x \ f \partial_\mu j^\mu = 0$ $\mathbf{A} \to \mathbf{A} + \nabla f$ • Here, f is an arbitrary scalar function. due to the charge conservation: $\partial_{\mu}j^{\mu} = 0 \implies \dot{\rho} + \nabla \cdot \mathbf{j} = 0$ 46

 $\int d^4x A_{\mu} j^{\mu}$



Warm-up: The wave equation for A^µ Maxwell's equations in vacuum

Maxwel

I's equations in vacuum
$$\partial_{\nu} F^{\mu\nu} = 0$$
 gives

$$-\Box A^{\mu} + \eta^{\mu\alpha} \partial_{\alpha} (\partial_{\nu} A^{\nu}) = 0$$

$$\Box = \eta^{\alpha\beta} \partial_{\alpha} \partial_{\beta} = -\frac{\partial^2}{\partial t^2} + \nabla^2 \qquad \eta^{\alpha\beta} = \text{diag}(-1, 1)$$

$$A^{\mu} = \eta^{\mu\alpha} A_{\alpha} = (\phi, \mathbf{A})$$
"Lorenz gauge condition"

where

equations in vacuum
$$\partial_{\nu} F^{\mu\nu} = 0$$
 gives

$$\Box A^{\mu} + \eta^{\mu\alpha} \partial_{\alpha} (\partial_{\nu} A^{\nu}) = 0$$

$$\Box = \eta^{\alpha\beta} \partial_{\alpha} \partial_{\beta} = -\frac{\partial^2}{\partial t^2} + \nabla^2 \qquad \eta^{\alpha\beta} = \text{diag}(-1)$$

$$u^{\mu} = \eta^{\mu\alpha} A_{\alpha} = (\phi, \mathbf{A})$$

s equations in vacuum
$$\partial_{\nu} F^{\mu\nu} = 0$$
 gives
 $-\Box A^{\mu} + \eta^{\mu\alpha} \partial_{\alpha} (\partial_{\nu} A^{\nu}) = 0$
 $\Box = \eta^{\alpha\beta} \partial_{\alpha} \partial_{\beta} = -\frac{\partial^2}{\partial t^2} + \nabla^2 \qquad \eta^{\alpha\beta} = \text{diag}(-1)$
 $A^{\mu} = \eta^{\mu\alpha} A_{\alpha} = (\phi, \mathbf{A})$

by choosing $\Box f = -\partial_{\nu}A^{\nu}$ in A_{μ}

• Now, using invariance under the gauge transformation, we can set $\partial_{\nu}A^{\nu}=0$

$$ightarrow A_{\mu} + \partial_{\mu} f$$
 . Then . . .





Warm-up: The wave equation for A^µ The use case of the gauge invariance

Maxwell

's equations in vacuum
$$\partial_{\nu} F^{\mu\nu} = 0$$
 and $\partial_{\nu} A^{\nu} = 0$ gives

$$\Box A^{\mu} = 0 \quad \Longrightarrow \quad \text{The equation for a wave} \text{traveling at the speed of light!}$$

$$\Box = \eta^{\alpha\beta} \partial_{\alpha} \partial_{\beta} = -\frac{\partial^2}{\partial t^2} + \nabla^2$$

$$A^{\mu} = \eta^{\mu\alpha} A_{\alpha} = (\phi, \mathbf{A}) \text{ with } \dot{\phi} + \nabla \cdot \mathbf{A} = 0$$

where

equations in vacuum
$$\partial_{\nu} F^{\mu\nu} = 0$$
 and $\partial_{\nu} A^{\nu} = 0$ gives
 $A^{\mu} = 0$ \longrightarrow The equation for a wave
traveling at the speed of light!
 $\Box = \eta^{\alpha\beta}\partial_{\alpha}\partial_{\beta} = -\frac{\partial^2}{\partial t^2} + \nabla^2$
 $\mu^{\mu} = \eta^{\mu\alpha}A_{\alpha} = (\phi, \mathbf{A})$ with $\dot{\phi} + \nabla \cdot \mathbf{A} = 0$

s equations in vacuum
$$\partial_{\nu} F^{\mu\nu} = 0$$
 and $\partial_{\nu} A^{\nu} = 0$ gives

$$\Box A^{\mu} = 0 \quad \clubsuit \quad \text{The equation for a wave} \\ \text{traveling at the speed of light!} \\ \Box = \eta^{\alpha\beta} \partial_{\alpha} \partial_{\beta} = -\frac{\partial^2}{\partial t^2} + \nabla^2 \\ A^{\mu} = \eta^{\mu\alpha} A_{\alpha} = (\phi, \mathbf{A}) \text{ with } \dot{\phi} + \nabla \cdot \mathbf{A} = 0$$

• The number of degrees of freedom for A^{μ} is 3 due to the Lorenz gauge condition.



Physical degrees of freedom of EM waves 3? 2?

- We know that photons must have only 2 helicity states, λ=±1 (two circular polarization states).
 - Shouldn't the number of physical degrees of freedom be 2, instead of 3? The answer is yes.
- The Lorenz gauge does not fully specify A^µ. We can still add

$$A_{\mu}
ightarrow A_{\mu} + \partial_{\mu} f_2$$
 which satisfies $\Box f_2 = 0$

Choosing f₂ will fully specify A^µ. This leaves 2 degrees of freedom.

EM waves must be transverse

• It is common to choose f and f₂ such that



- We have 2 conditions. That leaves 2 degrees of freedom for EM waves.
- This choice is consistent with the Lorenz gauge condition $\phi +
 abla \cdot \mathbf{A} = 0$
- $\nabla \cdot \mathbf{A} = 0$ requires that the EM wave be *transverse*, i.e., the change in **A** is perpendicular to the direction of propagation of the EM wave.
- We will use this condition throughout the lecture.

The arrows show directions of the electric field vector **E**.





Maxwell's equations in vacuum Expanding space, in Heaviside units and c=1

In expanding space, one obtains

$$\mathbf{A}'' - \nabla^2 \mathbf{A} = 0 \quad \text{where} \quad \prime = \frac{\partial}{\partial \tau} = a \frac{\partial}{\partial t}$$

- The distance between two points in 4d spacetime: $ds^2 = \eta_{\mu\nu} dx^{\mu} dx^{\nu} = -dt^2 + d\mathbf{x}^2$ $\rightarrow -dt^2 + a^2(t)d\mathbf{x}^2 = a^2(\tau)$ $= g_{\mu\nu}dx^{\mu}dx^{\nu}$ with $x^{\mu} =$

• Remarkably, it takes the same form as in non-expanding space, except for the change of variables from the physical time, t, to the conformal time, τ .

$$[Non-expanding space] [-d au^2+d\mathbf{x}^2] \quad [Expanding space] \ (au,\mathbf{x}) \quad g_{\mu
u}=a^2{
m diag}(-1,\mathbf{1})=a^2$$



Recap: Lecture 1 Known Physics

- Transformation properties:

 - Pseudovector (such as angular momentum and spin) does not.

Parity transformation changes "right-handed" to "left-handed" and vice versa.

 Violation of parity symmetry: Nature distinguishes right- and left-handed states, which has been observed in β -decay. *How about the Universe?*

Vector (such as momentum) changes sign under parity transformation.

Pseudoscalar (such as helicity) does => Key to testing for parity violation!

Problem Set Prelude to the next lecture

Show that FF is a total derivative and can be written as

 $F_{\mu\nu}\tilde{F}^{\mu\nu} = 2\partial_{\mu}(A_{\nu}\tilde{F}^{\mu\nu})$