# $$
I_{\mathrm{CS}}=\int \mathrm{d}^{4} x \sqrt{-g}\left(-\frac{\alpha}{4 f} \chi F \widetilde{F}\right)
$$ 

In search of new physics for the Universe
The lecture slides are available at
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## Topics

## From the syllabus

1. What is parity symmetry?
2. Chern-Simons interaction
3. Parity violation 1: Cosmic inflation

$$
I_{\mathrm{CS}}=\int \mathrm{d}^{4} x \sqrt{-g}\left(-\frac{\alpha}{4 f} \chi F \widetilde{F}\right)
$$

4. Parity violation 2: Dark matter
5. Parity violation 3: Dark energy
6. Light propagation: birefringence
7. Physics of polarization of the cosmic microwave background
8. Recent observational results, their implications, and future prospects


### 2.1 Parity Symmetry in Electromagnetism (EM)

## Maxwell's Equations

## In Heaviside units and c=1

$$
\begin{aligned}
\nabla \cdot \mathbf{E} & =\rho, & -\dot{\mathbf{E}}+\nabla \times \mathbf{B}=\mathbf{j} \\
\nabla \cdot \mathbf{B} & =0, & \dot{\mathbf{B}}+\nabla \times \mathbf{E}=0
\end{aligned}
$$

- These equations are invariant under both spatial translation and rotation.
- They are also invariant under parity transformation, if $\mathbf{E}$ and $\mathbf{j}$ are vectors, $\rho$ is a scalar, and $\mathbf{B}$ is a pseudovector.


## Parity-flipping Maxwell's Equations

## In Heaviside units and c=1

$$
\begin{aligned}
(-\nabla) \cdot(-\mathbf{E})=\rho, & & -(-\dot{\mathbf{E}})+(-\nabla) \times \mathbf{B}=(-\mathbf{j}) \\
(-\nabla) \cdot \mathbf{B}=0, & & \dot{\mathbf{B}}+(-\nabla) \times(-\mathbf{E})=0
\end{aligned}
$$

- These equations are invariant under both spatial translation and rotation.
- They are also invariant under parity transformation, if $\mathbf{E}$ and $\mathbf{j}$ are vectors, $\rho$ is a scalar, and $\mathbf{B}$ is a pseudovector.


## Parity-flipping Maxwell's Equations

## In Heaviside units and c=1

$$
\begin{array}{rlr}
(-\nabla) \cdot(-\mathbf{E})= & \rho, \quad-(-\dot{\mathbf{E}})+(-\nabla) \times \mathbf{B}=(-\mathbf{j}) \\
(-\nabla) \cdot \mathbf{B}= & 0, & \dot{\mathbf{B}}+(-\nabla) \times(-\mathbf{E})=0 \\
& \underset{\substack{\text { If there is a magnetic monopole } \\
\text { it must be a pseudoscalar! }}}{ }
\end{array}
$$

- They are also invariant under parity transformation, if $\mathbf{E}$ and $\mathbf{j}$ are vectors, $\rho$ is a scalar, and $\mathbf{B}$ is a pseudovector.


## Simplifying Maxwell's Equations

## Let's go 4D.

$$
\begin{aligned}
& \nabla \cdot \mathbf{E}=\rho, \quad-\dot{\mathbf{E}}+\nabla \times \mathbf{B}=\mathbf{j} \\
& \hline \nabla \cdot \mathbf{B}=0, \quad \dot{\mathbf{B}}+\nabla \times \mathbf{E}=0
\end{aligned}
$$

- These equations can be written in a compact form as

$$
\partial_{\nu} F^{\mu \nu}=j^{\mu} \quad \partial_{\nu} \tilde{F}^{\mu \nu}=0
$$

$$
\mu=0,1,2,3, \quad j^{\mu}=(\rho, \mathbf{j}), \quad \partial_{\mu}=\partial / \partial x^{\mu}, \quad x^{\mu}=(t, \mathbf{x})
$$

## Antisymmetric Field Strength Tensor, Fuv

$F^{\mu v}=-F^{v} \mu$

$$
F^{\mu \nu}=\left(\begin{array}{cccc}
0 & E_{x} & E_{y} & E_{z} \\
-E_{x} & 0 & B_{z} & -B_{y} \\
-E_{y} & -B_{z} & 0 & B_{x} \\
-E_{z} & B_{y} & -B_{x} & 0
\end{array}\right)
$$

- Equivalently,

$$
\underset{\substack{F^{0 i}=E_{i} \\
F^{i j}=\epsilon^{i j k} B_{k} \\
F^{12}=B_{z}, F^{23}=B_{x}, F^{31}=B_{y} \\
F^{21}=-B_{z}, F^{32}=-B_{x}, F^{13}=-B_{y}}}{\substack{\text { Levi-Civita } \\
\text { symbol }}} \begin{gathered}
t i j k(\mathrm{i}, \mathrm{j}, \mathrm{k}) \text { is even permutation of }(1,2,3) \\
-1 \text { if }(\mathrm{i}, \mathrm{j}, \mathrm{k}) \text { is odd permutation of }(1,2,3) \\
0 \text { otherwise } \\
\epsilon^{123}=1, \epsilon^{132}=-1, \epsilon^{312}=1, \ldots
\end{gathered}
$$

## Antisymmetric Field Strength Tensor, $F_{\mu \mathrm{v}}$

## $F_{\mu \mathrm{v}}=-F_{\mathrm{v} \mu}$

$$
F_{\mu \nu}=\eta_{\mu \alpha} \eta_{\nu \beta} F^{\alpha \beta} \quad \text { where } \eta_{\mu \alpha}=\operatorname{diag}(-1,1,1,1)
$$

$$
F_{\mu \nu}=\left(\begin{array}{cccc}
0 & -E_{x} & -E_{y} & -E_{z} \\
E_{x} & 0 & B_{z} & -B_{y} \\
E_{y} & -B_{z} & 0 & B_{x} \\
E_{z} & B_{y} & -B_{x} & 0
\end{array}\right)
$$

- Therefore,

$$
F^{2} \equiv F_{\mu \nu} F^{\mu \nu}=2(\mathbf{B} \cdot \mathbf{B}-\mathbf{E} \cdot \mathbf{E})
$$

This is a scalar and is invariant under parity transformation.

Dual Field Strength Tensor, $\tilde{F}^{\mathrm{Av}}$
$\tilde{A}_{\tilde{w}}=-\tilde{F}$

$$
\tilde{F}^{\mu \nu}=\frac{1}{2} \epsilon^{\mu \nu \alpha \beta} F_{\alpha \beta} \text { where } \underset{\substack{\text { Levi-cicita } \\ \text { symbol }}}{\epsilon^{\mu \nu \alpha \beta}}=\{
$$

$$
+1 \begin{gathered}
\text { if }(\mu, v, v, \beta) \text { is even } \\
\text { of }(0,1,2,3)
\end{gathered}
$$

$$
-1^{\text {if }(\mu, v, a, \beta) \text { is odd perm. }} \begin{gathered}
\text { of }(0,1,2,3)
\end{gathered}
$$

0 otherwise

$$
\tilde{F}^{\mu \nu}=\left(\begin{array}{cccc}
0 & B_{x} & B_{y} & B_{z} \\
-B_{x} & 0 & -E_{z} & E_{y} \\
-B_{y} & E_{z} & 0 & -E_{x} \\
-B_{z} & -E_{y} & E_{x} & 0
\end{array}\right) \Rightarrow \begin{aligned}
& \text { Equivalently, } \\
& \tilde{F}^{0 i}=B_{i} \\
& \tilde{F}^{i j}=-\epsilon^{i j k} E_{k}
\end{aligned}
$$

- Therefore,

$$
F \tilde{F} \equiv F_{\mu \nu} \tilde{F}^{\mu \nu}=-4 \mathbf{B} \cdot \mathbf{E}
$$

### 2.2 Action Principle for EM

$$
\partial_{\nu} F^{\mu \nu}=j^{\mu}, \quad \partial_{\nu} \tilde{F}^{\mu \nu}=0
$$

## What is the action that gives Maxwell's equations?

 In Heaviside units and $c=1$- The answer is

$$
I=-\frac{1}{4} \int d^{4} x F^{2}+\int d^{4} x A_{\mu} j^{\mu} \quad d^{4} x=d t d^{3} \mathbf{x}
$$

with
Vector potential

$$
F_{\mu \nu}=\partial_{\mu} A_{\nu}-\partial_{\nu} A_{\mu} \quad \text { where } A_{\mu}=(-\phi, \mathbf{A})
$$

Therefore, $\left\{\begin{array}{l}F_{i 0}=\partial_{i} A_{0}-\dot{A}_{i}=E_{i} \\ F_{i j}=\partial_{i} A_{j}-\partial_{j} A_{i}=\epsilon_{i j k} B_{k}\end{array} \quad \mathbf{E}=-\nabla \phi-\dot{\mathbf{A}}=\begin{array}{l}\mathbf{B}=\nabla \times \mathbf{A}\end{array}\right.$

$$
\partial_{\nu} F^{\mu \nu}=j^{\mu}, \quad \checkmark \partial_{\nu} \tilde{F}^{\mu \nu}=0
$$

## What is the action that gives Maxwell's equations?

 In Heaviside units and $c=1$- The answer is

$$
I=-\frac{1}{4} \int d^{4} x F^{2}+\int d^{4} x A_{\mu} j^{\mu} \quad d^{4} x=d t d^{3} \mathbf{x}
$$

with

$$
F_{\mu \nu}=\partial_{\mu} A_{\nu}-\partial_{\nu} A_{\mu} \quad \text { where } A_{\mu}=(-\phi, \mathbf{A})
$$

One set of Maxwell's equations is simply given by the definition of $F_{\mu v}$ :

$$
\partial_{\nu} \tilde{F}^{\mu \nu}=\partial_{\nu} \frac{1}{2} \epsilon^{\mu \nu \alpha \beta}\left(\partial_{\alpha} A_{\beta}-\partial_{\beta} A_{\alpha}\right)=\epsilon^{\mu \nu \alpha \beta} \partial_{\nu} \partial_{\alpha} A_{\beta}=0
$$

## $\checkmark \partial_{\nu} F^{\mu \nu}=j^{\mu}, \quad \partial_{\nu} \tilde{F}^{\mu \nu}=0$

## What is the action that gives Maxwell's equations?

In Heaviside units and $c=1$

$$
I=-\frac{1}{4} \int d^{4} x F^{2}+\int d^{4} x A_{\mu} j^{\mu}
$$

- The idea: The equation of motion for $A_{\mu}$ is the path that gives a stationary point. For a small change in $A_{\mu} \rightarrow>A_{\mu}+\delta A_{\mu}$, the corresponding change in $I->I+\delta /$ is also small.

$$
\begin{aligned}
& \delta I=\int d^{4} x F^{\mu \nu} \partial_{\nu}\left(\delta A_{\mu}\right)+\int d^{4} x\left(\delta A_{\mu}\right) j^{\mu \quad} \begin{aligned}
\frac{\text { Hint: }}{\delta\left(F^{2}\right)} & =2 F^{\mu \nu} \delta F_{\mu \nu} \\
& =-4 F^{\mu \nu} \partial_{\nu}\left(\delta A_{\mu}\right)
\end{aligned} \\
& =\int d^{4} x\left(-\partial_{\nu} F^{\mu \nu}+j^{\mu}\right) \delta A_{\mu}=0 \rightarrow \partial_{\nu} F^{\mu \nu}=j^{\mu}
\end{aligned}
$$

## Finding symmetries in the action

## It is like a treasure hunt!

$$
I=-\frac{1}{4} \int d^{4} x F^{2}+\int d^{4} x A_{\mu} j^{\mu}
$$

- This action is invariant under spatial translation, rotation, and parity transformation.
- It is also invariant under the following "gauge transformation",

$$
A_{\mu} \rightarrow A_{\mu}+\partial_{\mu} f
$$

Integration

$$
\int d^{4} x\left(\partial_{\mu} f\right) j^{\mu} \stackrel{\text { Hint: }}{=}-\int d^{4} x f \partial_{\mu} j^{\mu}=0
$$

- Here, $f$ is an arbitrary scalar function.

$$
\partial_{\mu} j^{\mu}=0 \rightarrow \dot{\rho}+\nabla \cdot \mathbf{j}=0
$$

## Finding symmetries in the action

## It is like a treasure hunt!

$$
I=-\frac{1}{4} \int d^{4} x F^{2}+\int d^{4} x A_{\mu} j^{\mu}
$$

- This action is invariant under spatial translation, rotation, and parity transformation.
- It is also invariant under the following "gauge transformation",

$$
\begin{aligned}
\phi & \rightarrow \phi-\dot{f} \\
\mathbf{A} & \rightarrow \mathbf{A}+\nabla f
\end{aligned}
$$

- Here, $f$ is an arbitrary scalar function.

$$
\int d^{4} x\left(\partial_{\mu} f\right) j^{\mu} \stackrel{\downarrow}{=}-\int d^{4} x f \partial_{\mu} j^{\mu}=0
$$

$$
\partial_{\mu} j^{\mu}=0 \rightarrow \dot{\rho}+\nabla \cdot \mathbf{j}=0
$$

## Problem Set 2 <br> Playing with Maxwell

1. Derive Maxwell's equations from $\partial_{\nu} F^{\mu \nu}=j^{\mu}, \partial_{\nu} \tilde{F}^{\mu \nu}=0$.
2. Show that $F \tilde{F}$ is a total derivative and can be written as

$$
F_{\mu \nu} \tilde{F}^{\mu \nu}=2 \partial_{\mu}\left(A_{\nu} \tilde{F}^{\mu \nu}\right)
$$

### 2.3 F $\tilde{F}$ in the action

## $F \widetilde{F}$ in the action?

$$
I=-\frac{1}{4} \int d^{4} x F^{2}+\int d^{4} x A_{\mu} j^{\mu}
$$

- This action is sufficient to produce all of Maxwell's equations.
- Can we add $\int d^{4} x F \tilde{F}$ to the action?
- The answer is yes. However, this is only a surface term, since $F \tilde{F}$ is a total derivative (as shown in Problem Set 2).


## $F \widetilde{F}$ in the action

 Carroll, Field, Jackiw (1990)
## Chern-Simons term

- Consider $I_{\mathrm{CS}}=-\frac{1}{4} \alpha \int d^{4} x \theta F \tilde{F} \quad$ with $F \tilde{F}=2 \partial_{\mu}\left(A_{\nu} \tilde{F}^{\mu \nu}\right)$
- $\alpha$ : a dimensionless constant
- $\theta$ : a dimensionless pseudoscalar field
- This is not a surface term! Integration by parts gives

$$
I_{\mathrm{CS}}=\frac{1}{2} \alpha \int d^{4} x\left(\partial_{\mu} \theta\right) A_{\nu} \tilde{F}^{\mu \nu}
$$



- This is a special case of the so-called Chern-Simons term, $p_{\mu} A_{\nu} \tilde{F}^{\mu \nu}$

$$
\text { with } p_{\mu}=\partial_{\mu} \theta
$$

## Consistency with gauge invariance

## $p_{\mu}$ cannot be arbitrary

$$
I_{\mathrm{CS}}=\frac{1}{2} \alpha \int d^{4} x p_{\mu} A_{\nu} \tilde{F}^{\mu \nu}
$$

- This action is invariant under the gauge transformation, $A_{\nu} \rightarrow A_{\nu}+\partial_{\nu} f$ if $\partial_{\nu} p_{\mu}-\partial_{\mu} p_{\nu}=0 \quad$ Hint: Use integration by parts and the identity
- For example:

$$
\partial_{\nu} \tilde{F}^{\mu \nu}=0
$$

- $p_{\mu}$ is a constant vector and not dynamical.
- $p_{\mu}$ is a gradient of a dynamical (pseudo)scalar field, such as $p_{\mu}=\partial_{\mu} \theta$.


## Consistency with gauge invariance

## $p_{\mu}$ cannot be arbitrary

$$
I_{\mathrm{CS}}=\frac{1}{2} \alpha \int d^{4} x p_{\mu} A_{\nu} \tilde{F}^{\mu \nu}
$$

- This action is invariant under the gauge transformation, $A_{\nu} \rightarrow A_{\nu}+\partial_{\nu} f$ if $\partial_{\nu} p_{\mu}-\partial_{\mu} p_{\nu}=0 \quad$ Hint: Use integration by parts and the identity
- For example: This implies the presence of a preferred direction in spacetime

$$
\partial_{\nu} \tilde{F}^{\mu \nu}=0
$$

and violation of Lorentz invariance!

- $p_{\mu}$ is a constant vector and not dynamical, or
- $p_{\mu}$ is a gradient of a dynamical (pseudo)scalar field, such as $p_{\mu}=\partial_{\mu} \theta$.


## The main goals of this lecture series

## Let's find new physics!

- We will study the cosmological consequence of

$$
I_{\mathrm{CS}}=-\frac{1}{4} \alpha \int d^{4} x \theta F \tilde{F}
$$

- Specifically, we ask if $\theta$ is -
- active during cosmic inflation,
- responsible for dark matter, or
- responsible for dark energy.


## The main goals of this lecture series

## Let's find new physics!

- We will study the cosmological consequence of

$$
I_{\mathrm{CS}}=-\frac{1}{4} \alpha \int d^{4} x \theta F \tilde{F}
$$

- Specifically, we ask if $\theta$ is - . We will also study observational signatures in -
- active during cosmic inflation,
- responsible for dark matter, or
- responsible for dark energy.
- Cosmic microwave background,
- Gravitational waves, and
- Large-scale structure of the Universe.


## Is there a known example of this term in particle physics? Yes, a pion. <br> 

- A pion is a composite meson composed of a quark and an antiquark.
- A neutral pion, $\pi^{0}$, is composed of either $u \bar{u}$ or d $\bar{d}$, and is a pseudoscalar.
(Chinowsky \& Steinberger, 1954)
- $\pi^{0}$ is coupled to photons via Ics where
- $\theta=\pi^{0} / f_{\pi}$ with $f_{\pi} \sim 184 \mathrm{MeV}$ (pion decay constant)
- $a=2 a_{\mathrm{Em}} N_{\mathrm{c}} /(3 \pi)$ with $N_{\mathrm{c}}=3$ (the number of quark colors) and $\mathrm{a}_{\mathrm{Em}} \sim 1 / 137$ (EM fine structure constant)
- $\pi^{0}$ decays into 2 photons via this term, which has been observed. So, what we are going to study in this lecture is not completely crazy!


## Correction to Maxwell's equations

## In Heaviside units and c=1

$$
I=-\frac{1}{4} \int d^{4} x\left(F^{2}+\alpha \theta F \tilde{F}\right)+\int \begin{gathered}
d^{4} x=d t d^{3} \mathbf{x} \\
d^{4} x A_{\mu} j^{\mu} \\
\text { Hint: } \delta(F \tilde{F})=\epsilon^{\mu \nu \alpha \beta}\left(\delta F_{\mu \nu}\right) F_{\alpha \beta}
\end{gathered}
$$

- We now derive the correction to Maxwell's equations from
- Finding the path that gives a stationary point,

$$
=-4\left(\partial_{\nu} \delta A_{\mu}\right) \tilde{F}^{\mu \nu}
$$

$$
\begin{aligned}
& \delta I=\int d^{4} x\left(F^{\mu \nu}+\alpha \theta \tilde{F}^{\mu \nu}\right) \partial_{\substack{\nu \\
\boldsymbol{V}^{\text {Integration }} \text { by parts }}}\left(\delta A_{\mu}\right)+\int d^{4} x\left(\delta A_{\mu}\right) j^{\mu} \\
& =\int d^{4} x\left[\underline{-\partial_{\nu}\left(F^{\mu \nu}+\alpha \theta \tilde{F}^{\mu \nu}\right)+j^{\mu}}\right] \delta A_{\mu}^{\text {by pars }}=0
\end{aligned}
$$

## Correction to Maxwell's equations

## In Heaviside units and c=1

- Therefore, the correction to Maxwell's equations is Hint: $\partial_{\nu} \tilde{F}^{\mu \nu}=0$

$$
\partial_{\nu} F^{\mu \nu}+\alpha\left(\partial_{\nu} \theta\right) \tilde{F}^{\mu \nu}=j^{\mu}
$$

- The result is

$$
\begin{aligned}
& =\int d^{4} x\left[\underline{\left.-\partial_{\nu}\left(F^{\mu \nu}+\alpha \theta \tilde{F}^{\mu \nu}\right)+j^{\mu}\right]}\right]_{0}^{\text {by pats }} \delta A_{\mu}=0
\end{aligned}
$$

### 2.4 Parity Violation in EM Waves

## Warm-up: The wave equation for $A^{\mu}$

## Maxwell's equations in vacuum

- Maxwell's equations in vacuum $\partial_{\nu} F^{\mu \nu}=0$ gives

$$
-\square A^{\mu}+\eta^{\mu \alpha} \partial_{\alpha}\left(\partial_{\nu} A^{\nu}\right)=0
$$

where

$$
\begin{aligned}
\square & =\eta^{\alpha \beta} \partial_{\alpha} \partial_{\beta}=-\frac{\partial^{2}}{\partial t^{2}}+\nabla^{2} \\
A^{\mu} & =\eta^{\mu \alpha} A_{\alpha}=(\phi, \mathbf{A})
\end{aligned}
$$

$$
\eta^{\alpha \beta}=\operatorname{diag}(-1, \mathbf{1})
$$

"Lorenz gauge condition"

- Now, using invariance under the gauge transformation, we can set $\partial_{\nu} A^{\nu}=0$ by choosing $\square f=-\partial_{\nu} A^{\nu}$ in $A_{\mu} \rightarrow A_{\mu}+\partial_{\mu} f$. Then...


## Warm-up: The wave equation for $A^{\mu}$

## The use case of the gauge invariance

"Lorenz gauge condition"

- Maxwell's equations in vacuum $\partial_{\nu} F^{\mu \nu}=0$ and $\partial_{\nu} A^{\nu}=0$ gives

$$
\square A^{\mu}=0 \quad \square \begin{gathered}
\text { The equation for a wave } \\
\text { traveling at the speed of liah }
\end{gathered}
$$

where

$$
\begin{aligned}
\square & =\eta^{\alpha \beta} \partial_{\alpha} \partial_{\beta}=-\frac{\partial^{2}}{\partial t^{2}}+\nabla^{2} \\
A^{\mu} & =\eta^{\mu \alpha} A_{\alpha}=(\phi, \mathbf{A}) \text { with } \dot{\phi}+\nabla \cdot \mathbf{A}=0
\end{aligned}
$$

- The number of degrees of freedom for $A \mu$ is 3 due to the Lorenz gauge condition.


## Physical degrees of freedom of EM waves

## 3? 2?

- We know that photons must have only 2 helicity states, $\lambda= \pm 1$ (two circular polarization states).
- Shouldn't the number of physical degrees of freedom be 2 , instead of 3 ? The answer is yes.
- The Lorenz gauge does not fully specify $A^{\mu}$. We can still add

$$
A_{\mu} \rightarrow A_{\mu}+\partial_{\mu} f_{2}
$$

which satisfies $\square f_{2}=0$

- Choosing $f_{2}$ will fully specify $A^{\mu}$. This leaves 2 degrees of freedom.


## EM waves must be transverse

- It is common to choose $f$ and $f_{2}$ such that

$$
\phi=0, \quad \mathbf{\nabla}_{\text {in vacuum }} \cdot \boldsymbol{A}=\mathbf{0}_{\text {"Coulomb gauge condition" }}
$$

- This choice is consistent with the Lorenz gauge condition $\dot{\phi}+\nabla \cdot \mathbf{A}=0$
- $\nabla \cdot \mathbf{A}=0$ requires that the EM wave be transverse, i.e., the change in $\mathbf{A}$ is perpendicular to the direction of propagation of the EM wave.
- We will use this condition throughout the lecture.


## Correction to the EM wave equation

## With the Chern-Simons term

$$
\partial_{\nu} F^{\mu \nu}+\alpha\left(\partial_{\nu} \theta\right) \tilde{F}^{\mu \nu}=0
$$

- With $A^{0}=\phi=0$ in the Lorenz gauge, we find

$$
-\square A^{i}+\alpha\left(\partial_{\nu} \theta\right) \tilde{F}^{i \nu}=0
$$

$$
\begin{aligned}
\square & =\eta^{\alpha \beta} \partial_{\alpha} \partial_{\beta}=-\frac{\partial^{2}}{\partial t^{2}}+\nabla^{2} \\
A^{\mu} & =\eta^{\mu \alpha} A_{\alpha}=(\phi, \mathbf{A})
\end{aligned}
$$

$$
\ddot{\mathbf{A}}-\nabla^{2} \mathbf{A}+\underbrace{\alpha[-\dot{\theta}(\nabla \times \mathbf{A})+(\nabla \theta) \times \dot{\mathbf{A}}]}_{\text {Correction to the EM wave equation! }}=0
$$

## Helicity Basis

## Going to Fourier space

- Fourier transform of $\mathbf{A}(t, \mathbf{x})$ is $\mathbf{A}(t, \mathbf{x})=(2 \pi)^{-3 / 2} \int d^{3} \mathbf{k} \quad \mathbf{A}_{\mathbf{k}}(t) e^{i \mathbf{k} \cdot \mathbf{x}}$
- The EM wave propagates in the direction of $\mathbf{k}$. The change in $\mathbf{A}_{\boldsymbol{k}}$ is perpendicular to $\mathbf{k}$.
"Coulomb gauge" $\nabla \cdot \mathbf{A}(t, \mathbf{x})=0 \rightarrow \mathbf{k} \cdot \mathbf{A}_{\mathbf{k}}(t)=0$
- Choose $\mathbf{k}$ to be on the $\mathrm{z}\left(=\mathrm{x}^{3}\right)$ axis. The helicity states, $\lambda= \pm 1$, are given for each Fourier mode by

$$
A_{ \pm}=\frac{A_{\mathbf{k}}^{1} \mp i A_{\mathbf{k}}^{2}}{\sqrt{2}}
$$



## Helicity Basis

## Transformation property under rotation

- To show that $A_{ \pm}$represents the helicity states, rotate the spatial coordinates around the $z$ axis in the right-handed system by an angle $\varphi$.
- The helicity states, $\lambda= \pm 1$, transform as


$$
\begin{aligned}
& A_{\lambda} \rightarrow A_{\lambda}^{\prime}=e^{i \lambda \varphi} A_{\lambda} \\
& \begin{array}{c}
\text { A+licity } \\
\text { A-: }^{\prime}: \text { Right-handed state }
\end{array}
\end{aligned}
$$



## Correction to the EM wave equation

## In the helicity basis

Note: $\mathbf{A}$ is a vector and $\theta$ is a pseudoscalar.

$$
\ddot{\mathbf{A}}-\nabla^{2} \mathbf{A}+\frac{\alpha[-\dot{\theta}(\nabla \times \mathbf{A})+(\nabla \theta) \times \dot{\mathbf{A}}]}{\text { Correction to the EM wave equation! }}=0
$$

- To simplify the analysis, assume that $\theta$ is homogeneous and depends only on time, $\theta(\mathrm{t}, \mathbf{x}) \rightarrow \bar{\theta}(\mathrm{t}$. Then in Fourier space

$$
\ddot{\mathbf{A}}_{\mathbf{k}}+k^{2} \mathbf{A}_{\mathbf{k}}-i \alpha \dot{\bar{\theta}}(\mathbf{k} \times \mathbf{A})=0
$$

$$
\ddot{A}_{ \pm}+\left(k^{2} \mp k \alpha \dot{\bar{\theta}}\right) A_{ \pm}=0
$$

## Recap: Day 2

- There are 2 scalars composed of $F_{\mu v}$ and $\tilde{F}_{\mu v}$.
- $F^{2}$ is a scalar, whereas $F \tilde{F}$ is a pseudoscalar.
- The action $\int d^{2} d^{3} \mathbf{x} F^{2}$ gives Maxwell's equations, whereas $\int{\mathrm{d} t d^{3} \mathbf{x}}_{\mathbf{x}} \boldsymbol{F} \tilde{F}$ is a surface term.
- A Chern-Simons interaction between a pseudoscalar field $\theta$ and photons, $\int d^{1}{ }^{3} \mathbf{x} \theta F \widetilde{F}$, modifies Maxwell's equations.
- The equation of motion for EM waves depends explicitly on helicity states. This is the signature of violation of parity symmetry! $\ddot{A}_{ \pm}+\left(k^{2} \mp k \alpha \dot{\bar{\theta}}\right) A_{ \pm}=0$
- What is the cosmological implication of this term? Let's find out!

