#### CMB Polarisation: Toward an Observational Proof of Cosmic Inflation

Eiichiro Komatsu, Max-Planck-Institut für Astrophysik Colloquium, ICTP, October 22, 2014

#### March 17, 2014

BICEP2's announcement



#### First Direct Evidence of Cosmic Inflation

Release No.: 2014-05

For Release: Monday, March 17, 2014 - 10:45am



Cambridge, MA - Almost 14 billion years ago, the universe we inhabit burst into existence in an extraordinary event that initiated the Big Bang. In the first fleeting fraction of a second, the universe expanded exponentially, stretching far beyond the view of our best telescopes. All this, of course, was just theory.

Researchers from the BICEP2 collaboration today announced the first direct evidence for this cosmic inflation. Their data also represent the first images of gravitational waves, or ripples in space-time. These waves have been described as the "first tremors of the Big Bang." Finally, the data confirm a deep connection between quantum mechanics and general relativity.

"Detecting this signal is one of the most important goals in cosmology today. A lot of work by a lot of people has led up to this point," said John Kovac (Harvard-Smithsonian Center for Astrophysics), leader of the BICEP2 collaboration.



allaboration today announced the first direct evidence for this cosmic inflation. Their

#### 宇宙誕生直後の瞬間膨張、インフレーション初観測 米チ ーム

2014年3月18日05時00分

# Signature of Cosmic Inflation in the Sky [?]

**BICEP2: B signal** 

**BICEP2** Collaboration



One of the goals of this presentation is to help you understand what this figure is actually showing

Breakthroughs in Cosmological Research Over the Last Two Decades

 We can actually see the physical condition of the universe when it was very young

From "Cosmic Voyage"

## Sky in Optical (~0.5µm)

courtesy University of Arizona

## Sky in Microwave (~1mm)

courtesy University of Arizona

#### From Samtleben et al. (2007)



#### Arno Penzias & Robert Wilson, 1965

#### A MEASUREMENT OF EXCESS ANTENNA TEMPERATURE AT 4080 Mc/s

Measurements of the effective zenith noise temperature of the 20-foot horn-reflector antenna (Crawford, Hogg, and Hunt 1961) at the Crawford Hill Laboratory, Holmdel, New Jersey, at 4080 Mc/s have yielded a value about 3.5° K higher than expected. This excess temperature is, within the limits of our observations, isotropic, unpolarized, and free from seasonal variations (July, 1964–April, 1965). A possible explanation for the observed excess noise temperature is the one given by Dicke, Peebles, Roll, and Wilkinson (1965) in a companion letter in this issue.

May 13, 1965 Bell Telephone Laboratories, Inc Crawford Hill, Holmdel, New Jersey



#### •lsotropic

A. A. PENZIAS R. W. WILSON

#### 1:25 model at Deutsches Museum



## The REAL back-end system of the Penzias-Wilson experiment, exhibited at Deutsches Museum







#### May 20, 1964 CMB "Discovered"

Der Ra zei Ö B H V E s

#### 4" to 17 0" ach dian 5 10 1 EDT HUSS A with Cdl 1.4 protion took

Schreiberaufzeichnung der ersten Messung des Mikrowellenhintergrundes am 20.5.1964

Recording of the first measurement of cosmic microwave background radiation taken on 5/20/1964.



Smoot et al. (1992)

#### COBE/DMR, 1992



#### A spare unit of COBE/DMR ( $\lambda$ =1cm)





## COBE to WMAP



2001







#### **WMAP** Spacecraft

#### **Radiative Cooling: No Cryogenic System**



## WMAP Science Team





## Outstanding Questions

- Where does anisotropy in CMB temperature come from?
  - This is the origin of galaxies, stars, planets, and everything else we see around us, including ourselves
- The leading idea: quantum fluctuations in vacuum, stretched to cosmological length scales by a rapid exponential expansion of the universe called "cosmic inflation" in the very early universe

Starobinsky (1980); Sato (1981); Guth (1981); Linde (1982); Albrecht & Steinhardt (1982)

## Cosmic Inflation

- In a tiny fraction of a second, the size of an atomic nucleus became the size of the Solar System
  - In 10<sup>-36</sup> second, space was stretched by at least a factor of 10<sup>26</sup>

## Stretching Micro to Macro

Quantum fluctuations on microscopic scales



## Inflation!



• Become macroscopic, classical fluctuations

#### Key Predictions of Inflation

 Fluctuations we observe today in CMB and the matter distribution originate from quantum fluctuations generated during inflation





tensor mode

scalar

mode



# We measure distortions in space

• A distance between two points in space

$$d\ell^2 = a^2(t)[1 + 2\zeta(\mathbf{x}, t)][\delta_{ij} + h_{ij}(\mathbf{x}, t)]dx^i dx^j$$

- ζ: "curvature perturbation" (scalar mode)
  - Perturbation to the determinant of the spatial metric
- h<sub>ij</sub>: "gravitational waves" (tensor mode)
  - Perturbation that does not change the determinant (area)



# Tensor-to-scalar Ratio $\langle h_{ij}h^{ij}\rangle$

 We really want to find this quantity! The current upper bound: r<0.1 [WMAP & Planck]</li>

#### Heisenberg's Uncertainty Principle

- You can borrow energy from vacuum, if you promise to return it immediately
- [Energy you can borrow] x [Time you borrow] = constant

#### Heisenberg's Uncertainty Principle

- [Energy you can borrow] x [Time you borrow] = constant
- Suppose that the distance between two points increases in proportion to a(t) [which is called the scale factor] by the expansion of the universe
- Define the "expansion rate of the universe" as

$$H \equiv \frac{\dot{a}}{a}$$
 [This has units of 1/time]

# Fluctuations are proportional to H

 [Energy you can borrow] x [Time you borrow] = constant

• 
$$H \equiv \frac{\dot{a}}{a}$$
 [This has units of 1/time]

- Then, both ζ and h<sub>ij</sub> are proportional to H
- Inflation occurs in 10<sup>-36</sup> second this is such a short period of time that you can borrow a lot of energy!
  H during inflation in energy units is 10<sup>14</sup> GeV

\*WMAP 9-year Results (2012) and Planck 2013 Results

## Key Predictions of Inflation

- Inflation must end; thus, H slowly decreases with time
  - This means that the amplitude of fluctuations on larger scales is bigger than those on smaller scales. This has now been observed\*
- The origin of fluctuations is quantum. The wave function of vacuum fluctuations of a free field is a Gaussian. CMB anisotropy is Gaussian to better than 0.1% precision\*
  - There exist ultra long-wavelength primordial gravitational waves. This is yet to be found. How can we find this?

#### CMB Polarisation



• CMB is [weakly] polarised!





#### WMAP Collaboration

#### 23 GHz



#### Stokes U



#### WMAP Collaboration

## 23 GHz [13 mm]


## 33 GHz [9.1 mm]



#### Stokes Q

## 41 GHz [7.3 mm]



#### Stokes Q

## 61 GHz [4.9 mm]





## 94 GHz [3.2 mm]



#### Stokes Q

## How many components?

- CMB:  $T_v \sim v^0$
- Synchrotron:  $T_v \sim v^{-3}$
- Dust:  $T_v \sim v^2$
- Therefore, we need at least 3 frequencies to separate them

## Seeing polarisation in the WMAP data

- Average polarisation data around cold and hot temperature spots
- Outside of the Galaxy mask [not shown], there are 11536 hot spots and 11752 cold spots
- Averaging them beats the noise down





# Radial and tangential polarisation around temperature spots

- This shows polarisation generated by the plasma flowing into gravitational potentials
- Signatures of the "scalar mode" fluctuations in polarisation
- These patterns are called "Emodes"

#### Planck Collaboration





Seljak & Zaldarriaga (1997); Kamionkowski et al. (1997)

## E and B modes



- Density fluctuations
   [scalar modes] can
   only generate E modes
- Gravitational waves can generate both E and B modes

E mode

B mode

#### Physics of CMB Polarisation



- Necessary and sufficient conditions for generating polarisation in CMB:
  - Thomson scattering
  - Quadrupolar temperature anisotropy around an electron

#### Origin of Quadrupole

- Scalar perturbations: motion of electrons with respect to photons
- Tensor perturbations: gravitational waves



• What do they do to the distance between particles?



• Anisotropic stretching of space generates quadrupole temperature anisotropy. How?





Contraction of space -> temperature rises



#### propagation direction of GW



Polarisation directions perpendicular/parallel to the wavenumber vector -> E mode polarisation

#### propagation direction of GW

h<sub>x</sub>=cos(kx)





× ///// × ///// × ///// //

Polarisation directions 45 degrees tilted from to the wavenumber vector -> **B mode polarisation** 

### Important note:

- Definition of h<sub>+</sub> and h<sub>x</sub> depends on coordinates, but definition of E- and B-mode polarisation does not depend on coordinates
- Therefore, h<sub>+</sub> does not always give E; h<sub>x</sub> does not always give B
  - The important point is that h<sub>+</sub> and h<sub>x</sub> always coexist. When a linear combination of h<sub>+</sub> and h<sub>x</sub> produces E, another combination produces B

## Signature of gravitational waves in the sky [?]



**<u>CAUTION</u>: we are NOT seeing a single plane wave propagating perpendicular to our line of sight</u>** 

## Signature of gravitational waves in the sky [?]

**BICEP2: B signal** 





The E-mode polarisation is totally dominated by the scalar-mode fluctuations [density waves]

## What is BICEP2?

- A small [26 cm] refractive telescope at South Pole
- 512 bolometers working at 150 GHz
- Previous: BICEP1 at 100 and 150 GHz [2006-2008]
- On-going: Keck Array = 5 x BICEP2 at 150 GHz [2011-2013] and additional detectors at 100 and 220 GHz [2014-]



BICEP



## How does BICEP2 measure polarisation?

• By taking the difference between two detectors (A&B), measuring two orthogonal polarisation states



Horizontal slots -> A detector Vertical slots -> B detector

These slots are co-located, so they look at approximately same positions in the sky

### Is the signal cosmological?

- Worries:
  - Is it from Galactic foreground emission, e.g., dust?
  - Is it from imperfections in the experiment, e.g., detector mismatches?



Eiichiro Komatsu March 14 near Munich @

If detection of the primordial B-modes were to be reported on Monday, I would like see:

[1] Detection (>3 sigma each) in more than one frequency, like 100 GHz and 150 GHz giving the same answers to within the error bars.

[2] Detection (could be a couple of sigmas each) in a few multipole bins, i.e., not in just one big multipole bin.

Then I will believe it!





#### Eiichiro Komatsu March 14 near Munich @

If detection of the primordial B-modes were to be reported on Monday, I would like see:



etection (>3 sigma each) in more than one frequency, like 100 GHz and Hz giving the same answers to within the error bars.

etection (could be a couple of sigmas each) in a few multipole bins, i.e., just one big multipole bin.

Then I will believe it!

## facebook

#### Analysis: Two-point Correlation Function

**BICEP2: B signal** 



#### BICEP2 Collaboration



No 100 GHz x 100 GHz [yet]

#### Can we rule out synchrotron or dust?



### Situation until a month ago

- No strong evidence that the detected signal is not cosmological
- No strong evidence that the detected signal is cosmological, either

#### September 22, 2014

Planck's Intermediate Paper on Dust



 Values of the "tensor-to-scalar ratio" equivalent to the B-mode power spectrum seen at various locations in the sky



- Planck measured the B-mode power spectrum at 353 GHz well
- •Extrapolating it down to 150 GHz appears to explain all of the signal seen by BICEP2...

## Current Situation

- Planck shows the evidence that the detected signal is not cosmological, but is due to dust
- No strong evidence that the detected signal is cosmological

#### We Can Do It! The search continues!!



1989–1993

2001-2010





### LiteBIRD

- Next-generation polarisation-sensitive microwave experiment. Target launch date: early 2020
- Led by Prof. Masashi Hazumi (KEK); a collaboration of ~70 scientists in Japan, USA, Canada, and Germany
- Singular goal: measurement of the primordial Bmode power spectrum with Err[r]=0.001
- 6 frequency bands between 50 and 320 GHz
## **LiteBIRD**

Lite (Light) Satellite for the Studies of B-mode Polarization and Inflation from Cosmic Background Radiation Detection

- Candidate for JAXA's future missions on "fundamental physics"
- Goal: Search for primordial gravitational waves to the lower bound of well-motivated inflationary models
- Full success: δr < 0.001 (δr is the total uncertainties on tensor-to-scalar ratio, which is a fundamental cosmology parameter related to the power of primordial gravitational waves)



## ESA's M4 Call is Out [Target Launch in 2025]

- We are working on the COrE+ mission proposal
  - COrE = Cosmic Origins Explorer
  - Original version not selected by M3
- The letter of intent has been sent, and the proposal is due mid January 2015
- The effort led by Paolo de Bernardis, Jacques Delabrouille, and Francois Bouchet

## COrE+: a sketch

- The previous definition of COrE+ is still being worked out. Heavily affected by BICEP2/Planck results, and a rather tight budget (450M Euro by ESA and perhaps 100M Euro by the European consortium) and weight limit (payload 800 kg)
- Still want **10x more sensitivity than Planck** with more frequency coverage, while maintaining comparable angular resolution
  - which means 5 times better angular resolution and many more frequencies than LiteBIRD
  - A near ultimate mission

## Conclusion

- The WMAP and Planck's temperature data provide strong evidence for the quantum origin of structures in the universe
- The next goal: unambiguous measurement of the primordial B-mode polarisation power spectrum
- LiteBIRD proposal: a B-mode CMB polarisation satellite in early 2020
- COrE+ proposal: more comprehensive (and last?)
  CMB satellite in late 2020