

# **Finding Gravitational Waves from the Early Universe**

**Eiichiro Komatsu (Max Planck Institute for Astrophysics)**

***Monthly Research Colloquia, Agenzia Spaziale Italiana (MoRe-ASI), May 17, 2021***

# Let's find Gravitational Waves (GW)!

*But how? The detection method depends on the GW frequency.*

- **Laser interferometers on the ground: deca- to kilo Hz** (*LIGO, VIRGO, ..., ET*)

- The wavelength  $\sim$  the size of Earth

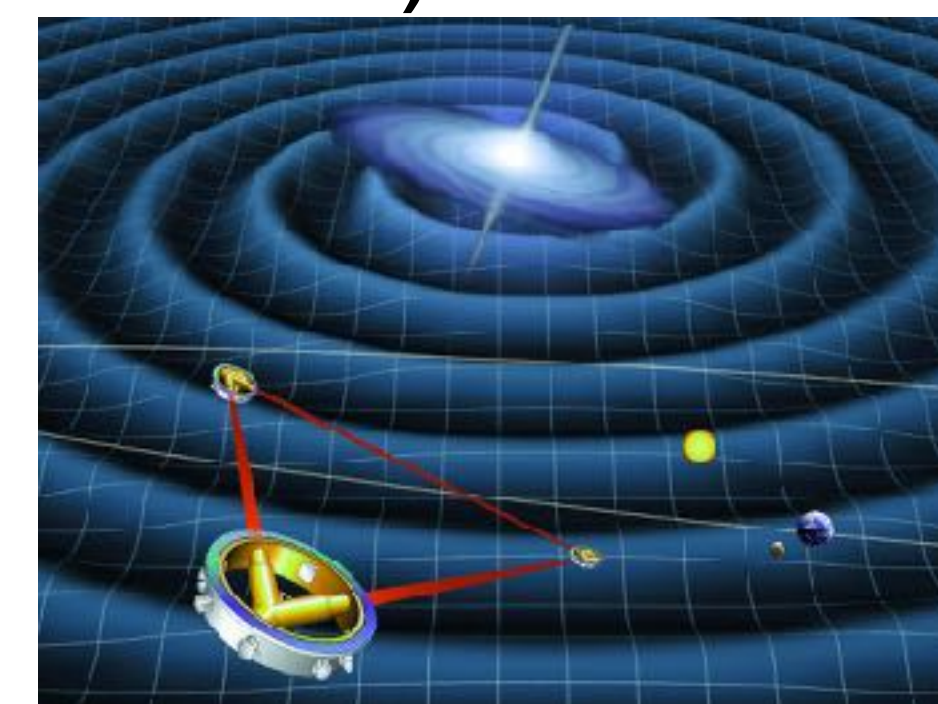


- **Laser interferometers in space: milli Hz** (*LISA*), deci Hz (future mission?)

- The wavelength  $\sim$  Astronomical Unit

- **Pulsar timing arrays: nano Hz** (*EPTA, SKA*)

- The wavelength  $\sim$  the size of the Milky Way



- **Cosmic microwave background: atto Hz** (*WMAP, Planck, LiteBIRD*)

- The wavelength  $\sim$  **billions of light years!**

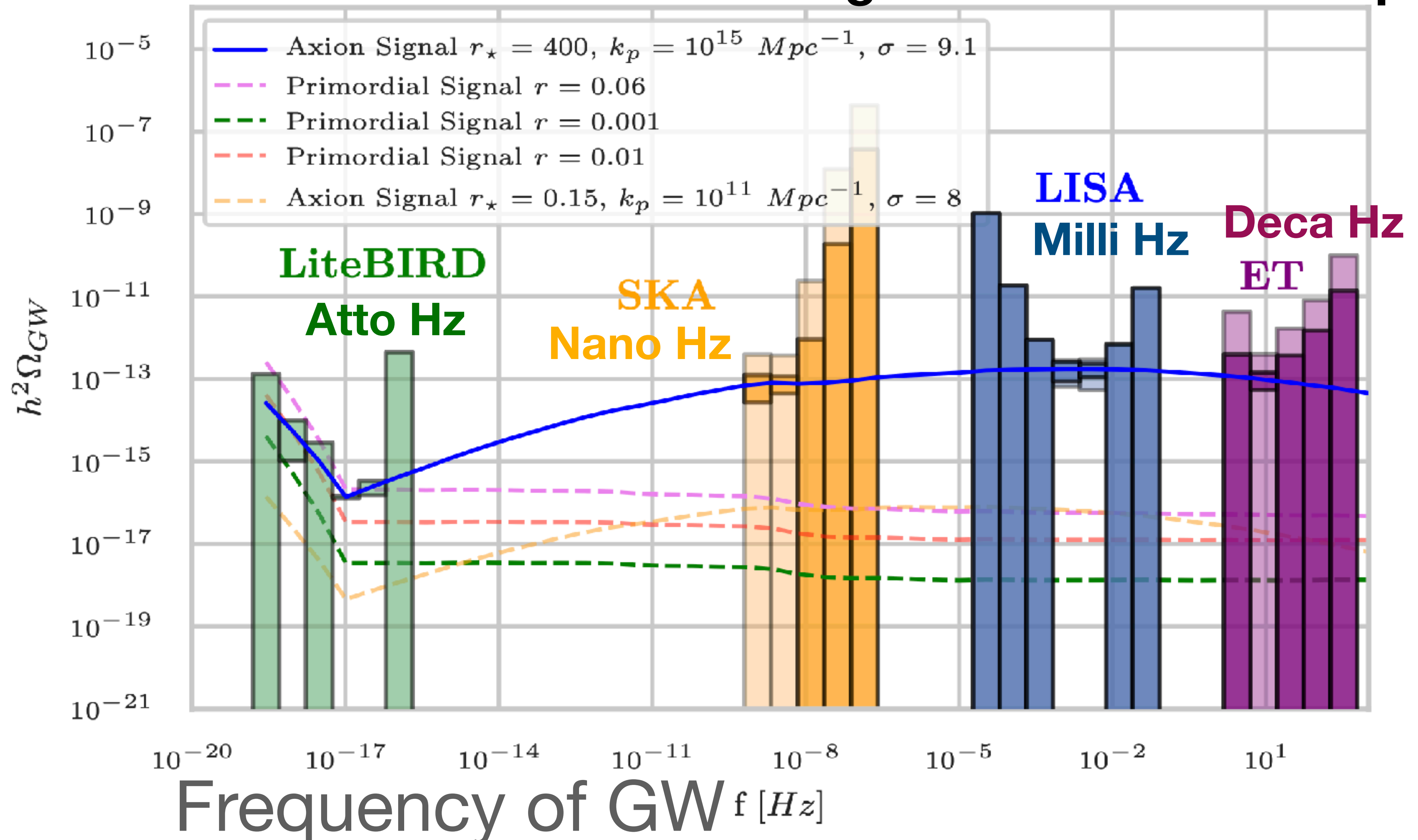




# GWs from the early Universe are everywhere!

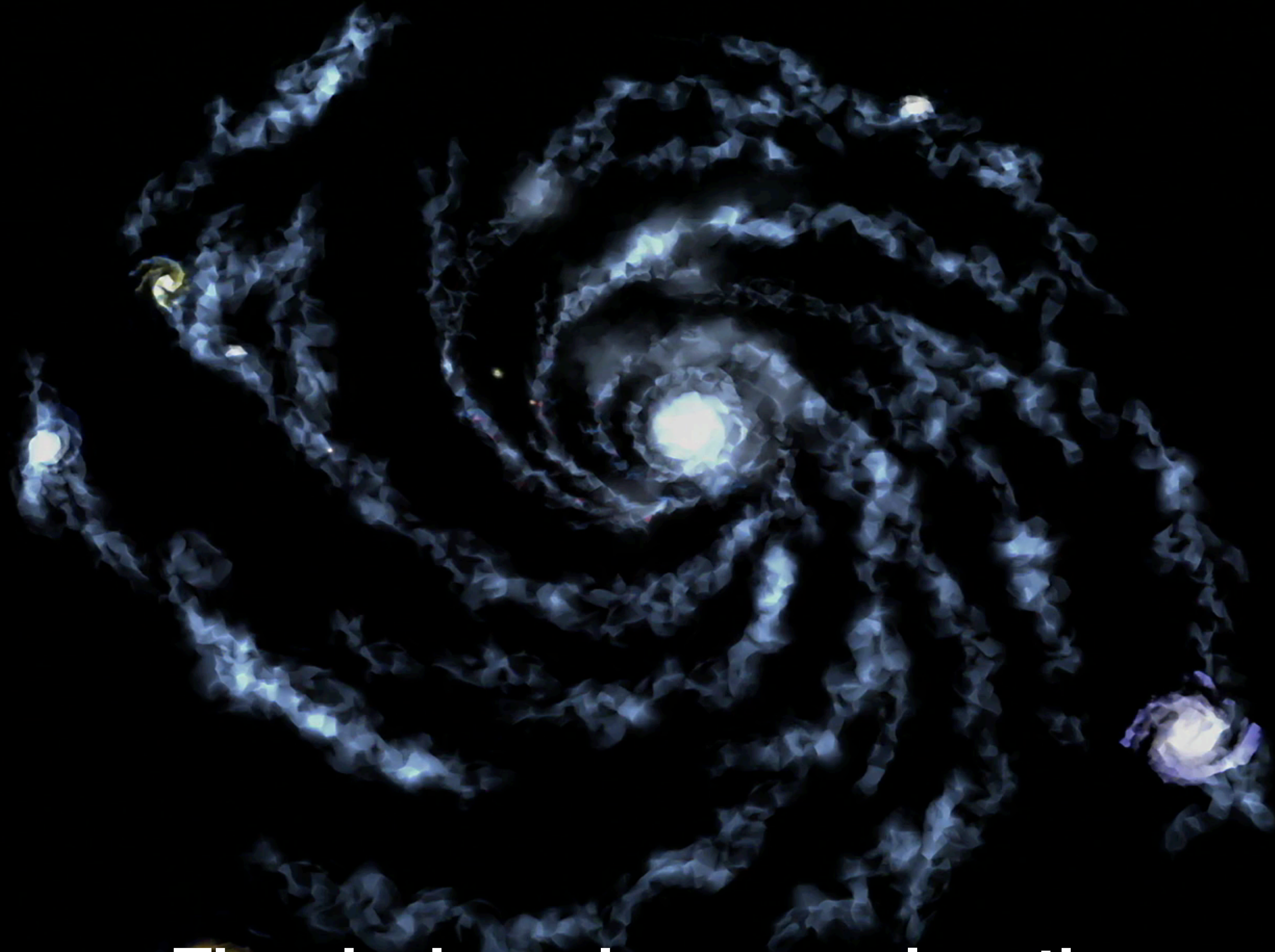
We can measure it across 21 orders of magnitude in the GW frequency

Energy Density of GW  
today





Credit: WMAP Science Team

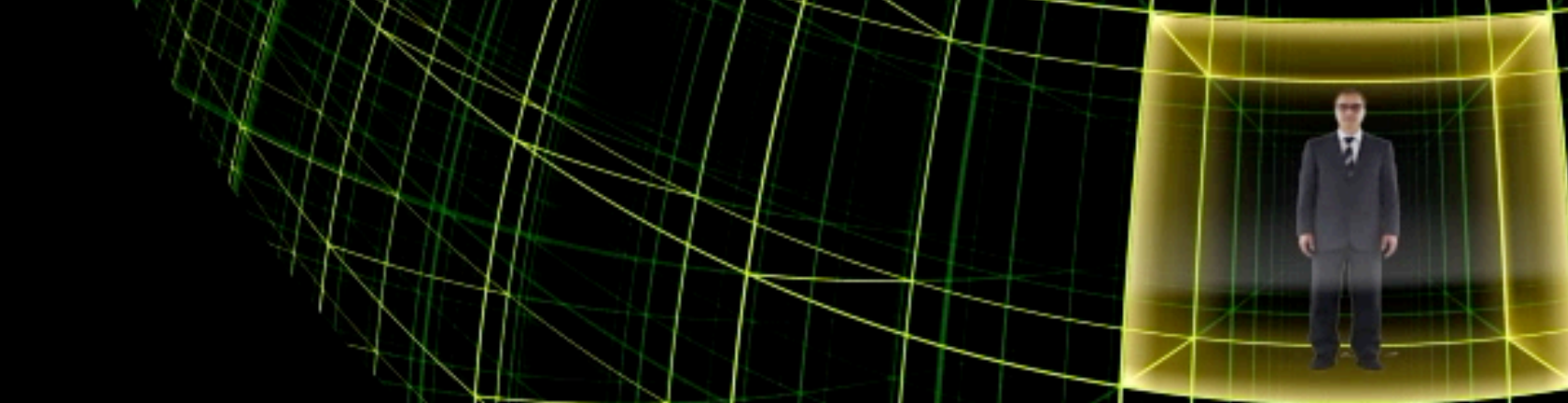


## The sky in various wavelengths

Visible -> Near Infrared -> Far Infrared -> Submillimeter -> Microwave



# Where did the CMB we see today come from?



155

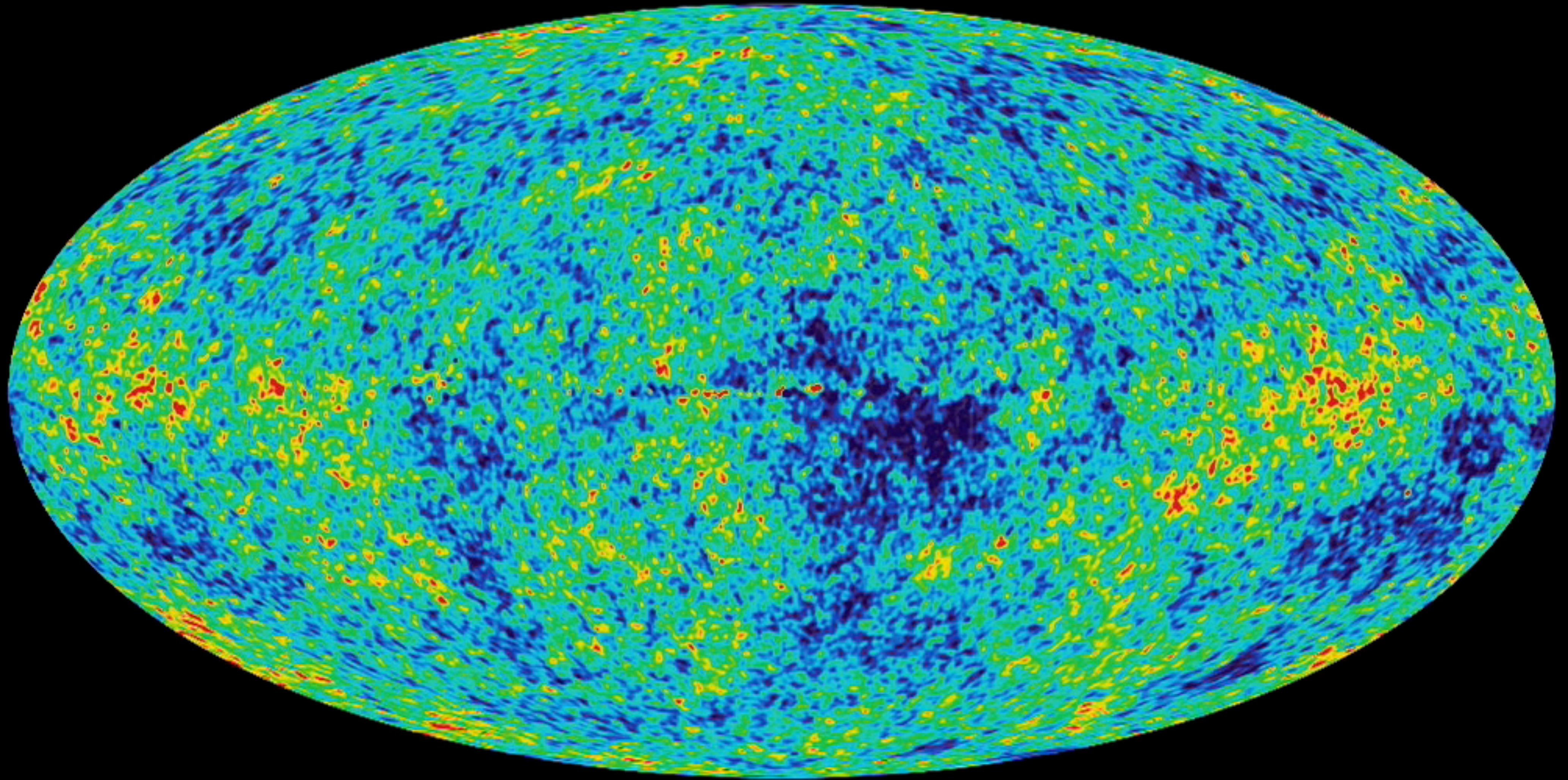
S07C04\_DM

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From “HORIZON”



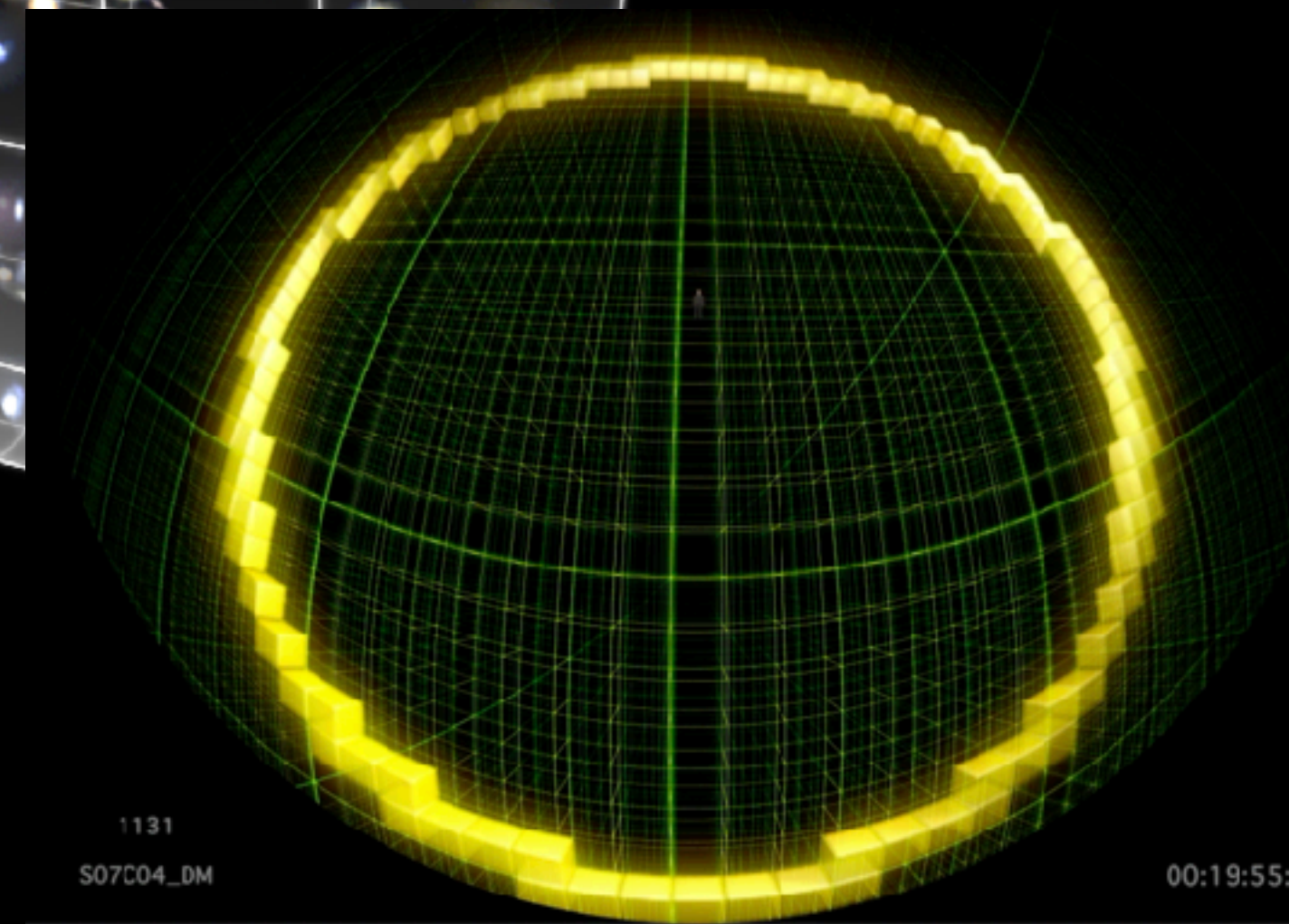
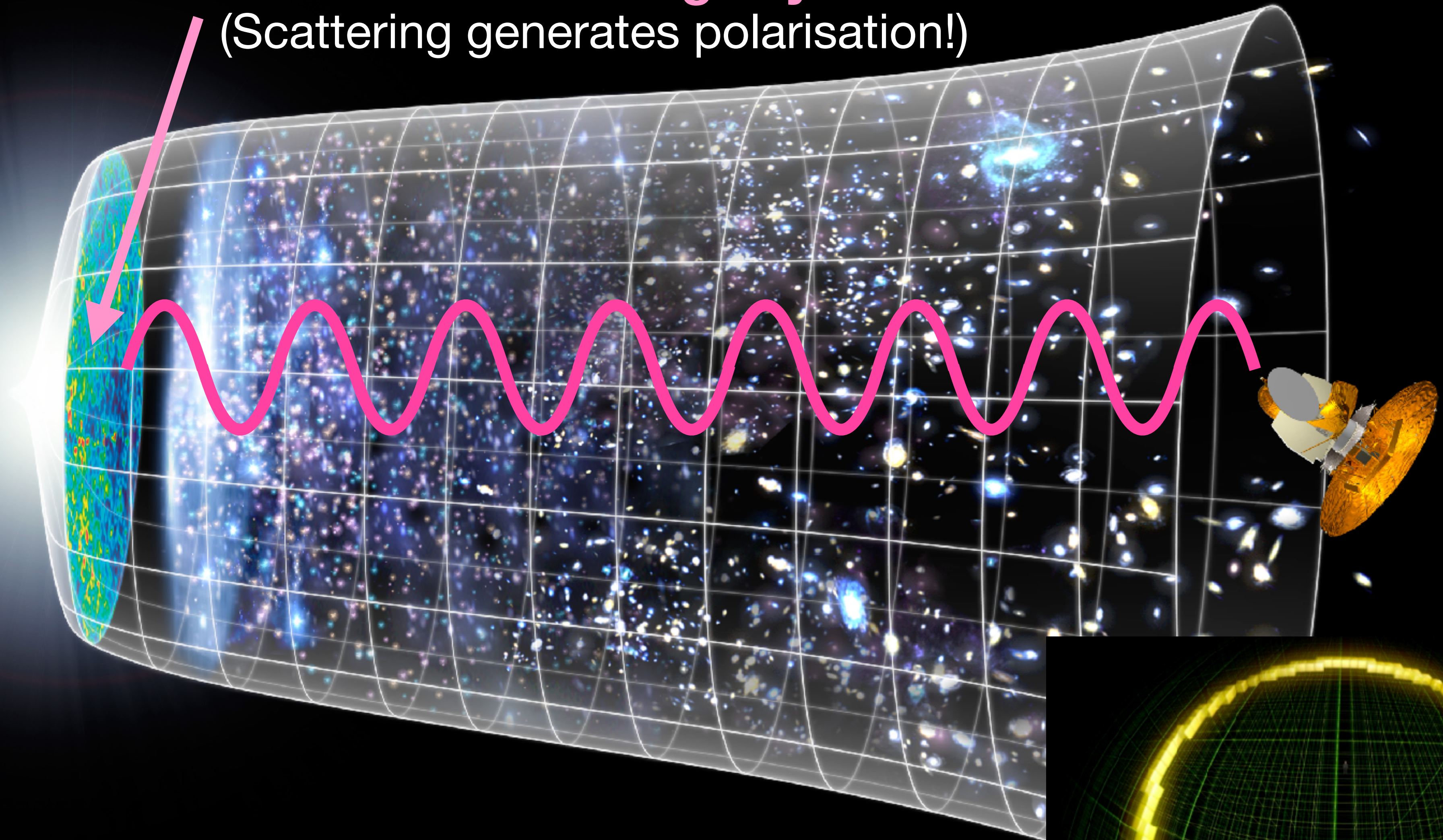
Credit: WMAP Science Team





# The surface of “last scattering” by electrons

(Scattering generates polarisation!)

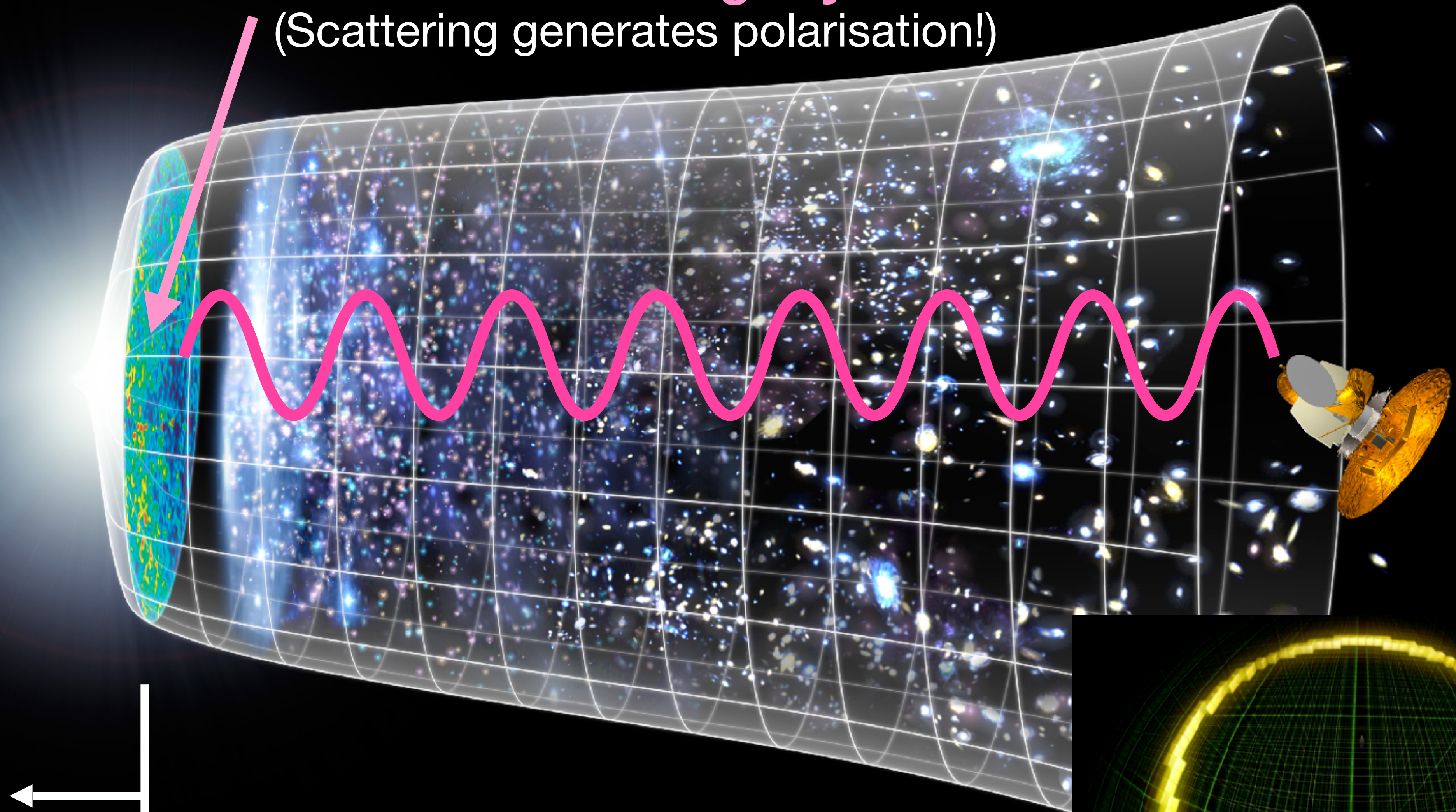


Not shown: The cosmological redshift due to the expansion of the Universe

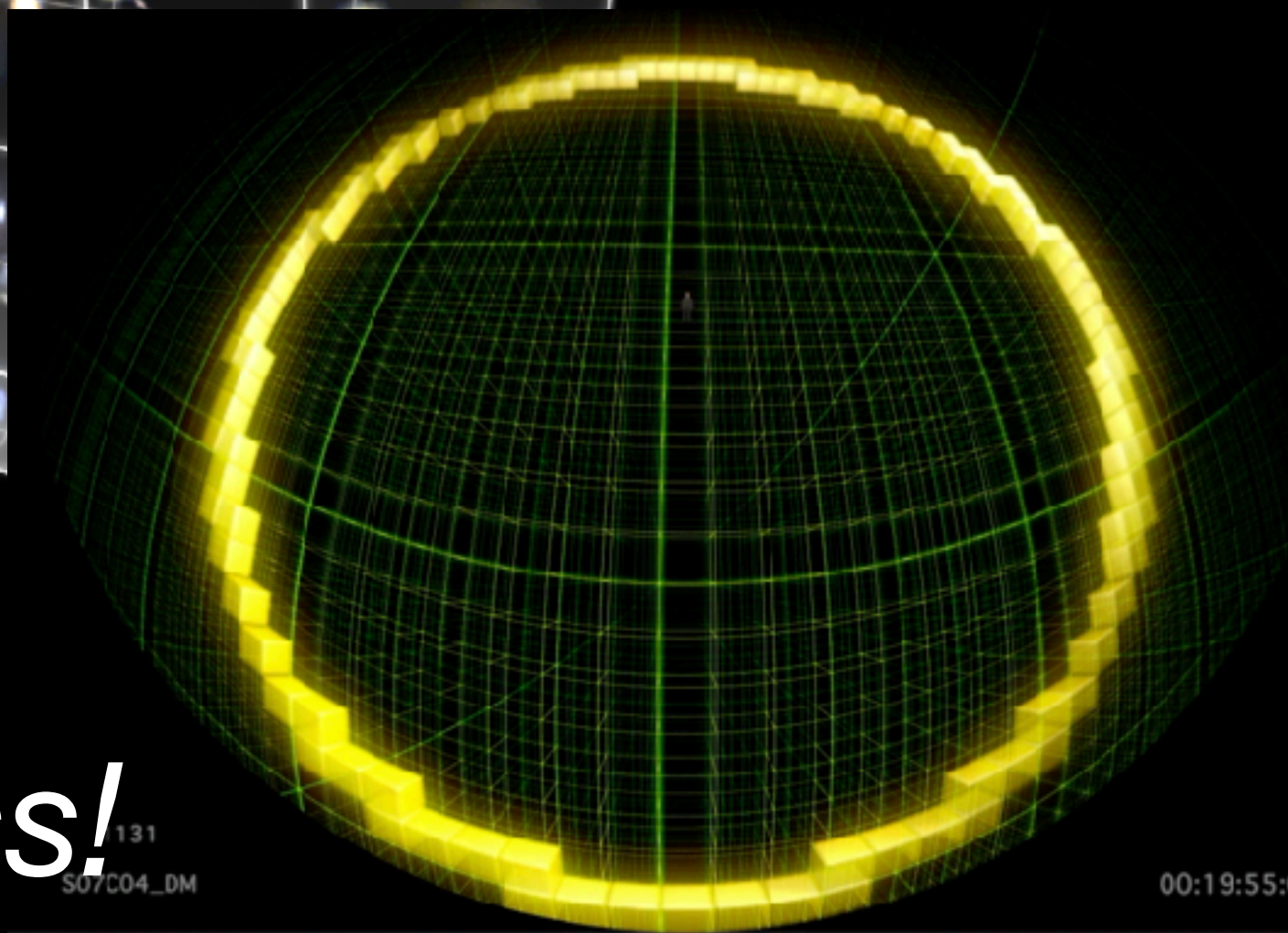


# The surface of “last scattering” by electrons

(Scattering generates polarisation!)



How do we “see” beyond this “wall”? *Laws of physics!*





**Before we talk about the GW,  
let's talk about the sound waves  
(scalar modes)**



# Gravitational Field Equations (Einstein's Eq.)

Credit: WMAP Science Team

$$\nabla^2 \Psi = 4\pi G a^2 \sum_{\alpha} \left[ \delta \rho_{\alpha} - \frac{3\dot{a}}{a} (\bar{\rho}_{\alpha} + \bar{P}_{\alpha}) \delta u_{\alpha} \right],$$

$$\partial_i \partial_j (\Phi - \Psi) = -8\pi G a^2 \partial_i \partial_j \sum_{\alpha} \pi_{\alpha},$$

## Energy Conservation

$$\frac{\partial}{\partial t} (\delta \rho_{\gamma} / \bar{\rho}_{\gamma}) - \frac{4q^2}{3a^2} \delta u_{\gamma} = 4\dot{\Psi},$$

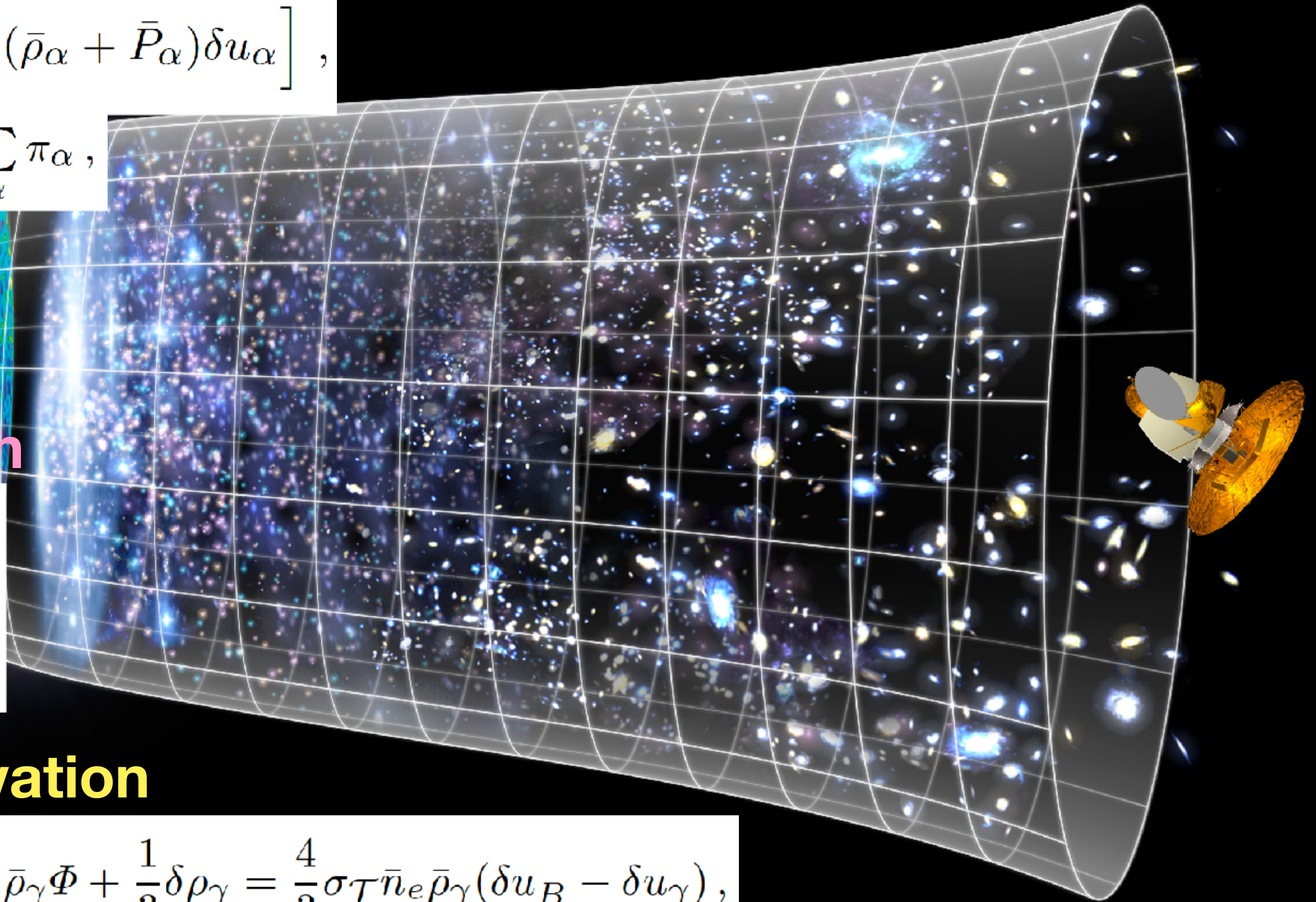
$$\frac{\partial}{\partial t} (\delta \rho_B / \bar{\rho}_B) - \frac{q^2}{a^2} \delta u_B = 3\dot{\Psi},$$

## Momentum Conservation

$$\frac{4}{3} \frac{\partial}{\partial t} (\bar{\rho}_{\gamma} \delta u_{\alpha}) + \frac{4\dot{a}}{a} \bar{\rho}_{\gamma} \delta u_{\gamma} + \frac{4}{3} \bar{\rho}_{\gamma} \Phi + \frac{1}{3} \delta \rho_{\gamma} = \frac{4}{3} \sigma_T \bar{n}_e \bar{\rho}_{\gamma} (\delta u_B - \delta u_{\gamma}),$$

$$\frac{\partial}{\partial t} (\bar{\rho}_B \delta u_B) + \frac{3\dot{a}}{a} \bar{\rho}_B \delta u_B + \bar{\rho}_B \Phi = -\frac{4}{3} \sigma_T \bar{n}_e \bar{\rho}_{\gamma} (\delta u_B - \delta u_{\gamma}),$$

*Laws of physics!*





Full-dome movie for planetarium

Director: Hiromitsu Kohsaka



HORIZON :Beyond the Edge of the Visible Universe [Trailer]



Gravitational Field Equations

+

Energy Conservation

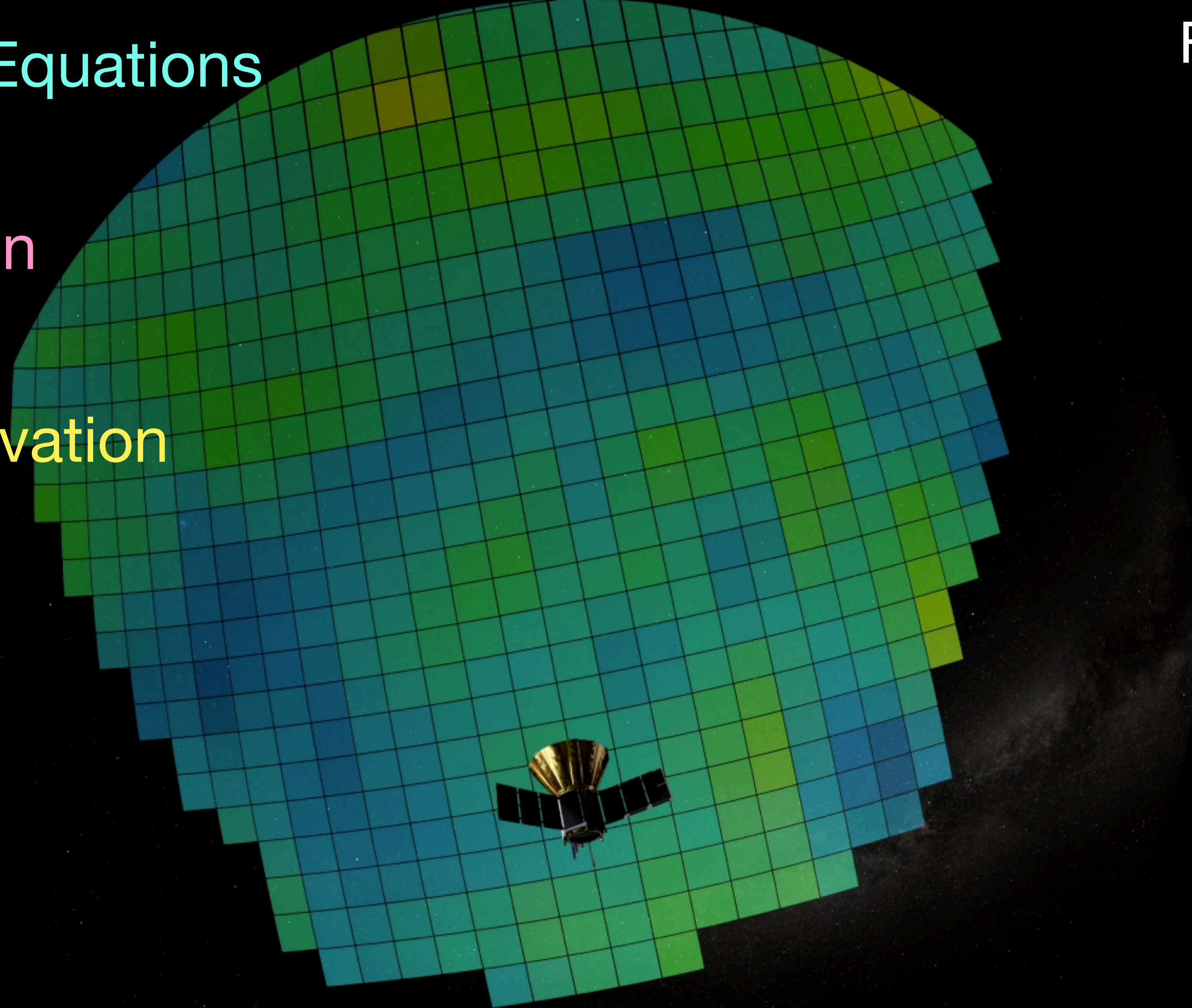
+

Momentum Conservation

||

**Sound Waves!**

From “HORIZON”







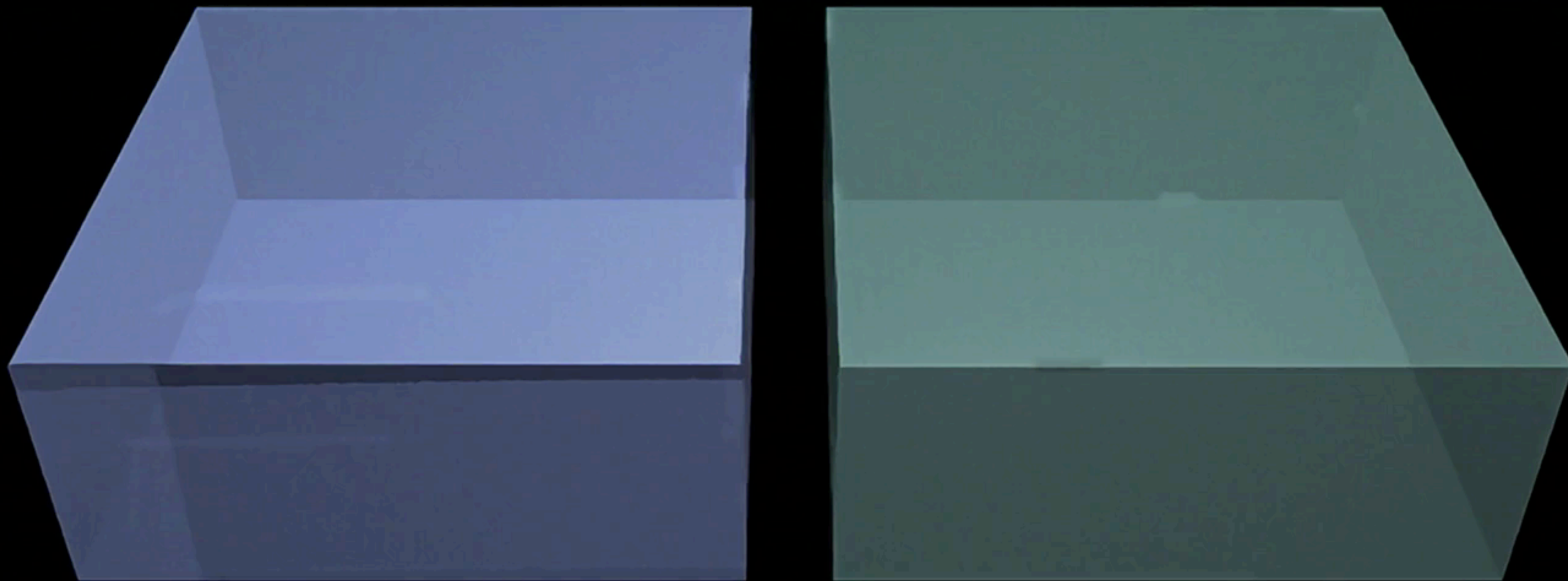


# Zuppa di Miso Cosmica

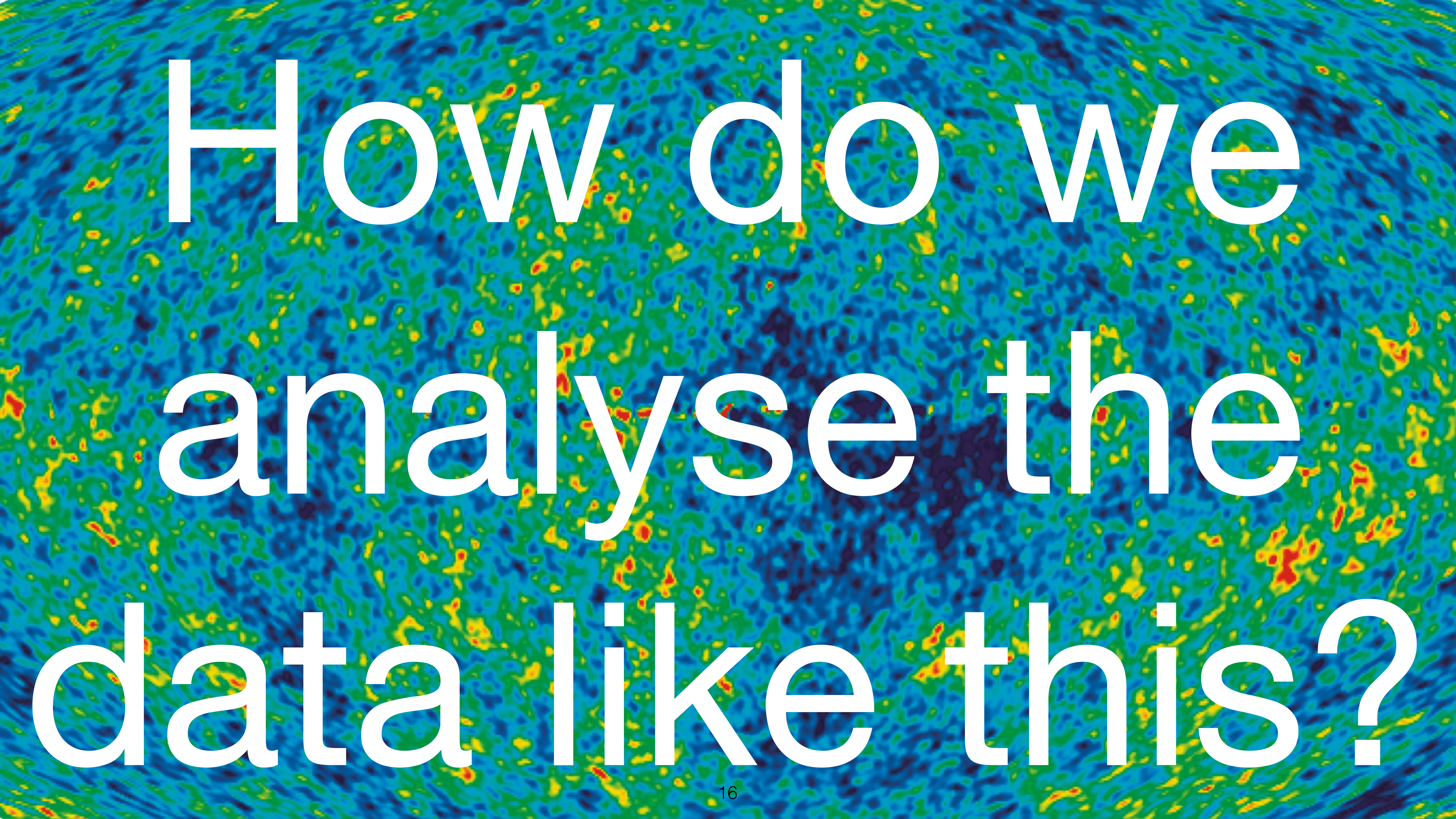
- When matter and radiation were hotter than 3000 K, matter was completely ionised. The Universe was filled with plasma, which behaves just like a soup
- Think about a Miso soup (if you know what it is). Imagine throwing Tofus into a Miso soup, while changing the density of Miso
- And imagine watching how ripples are created and propagate throughout the soup



Credit: WMAP Science Team





The background of the slide is a Cosmic Microwave Background (CMB) fluctuation map. It shows a complex pattern of temperature variations across the sky, with colors ranging from dark blue (cooler) to red and yellow (warmer). The pattern consists of numerous small, irregular patches and larger-scale structures, representing the primordial density fluctuations in the early universe.

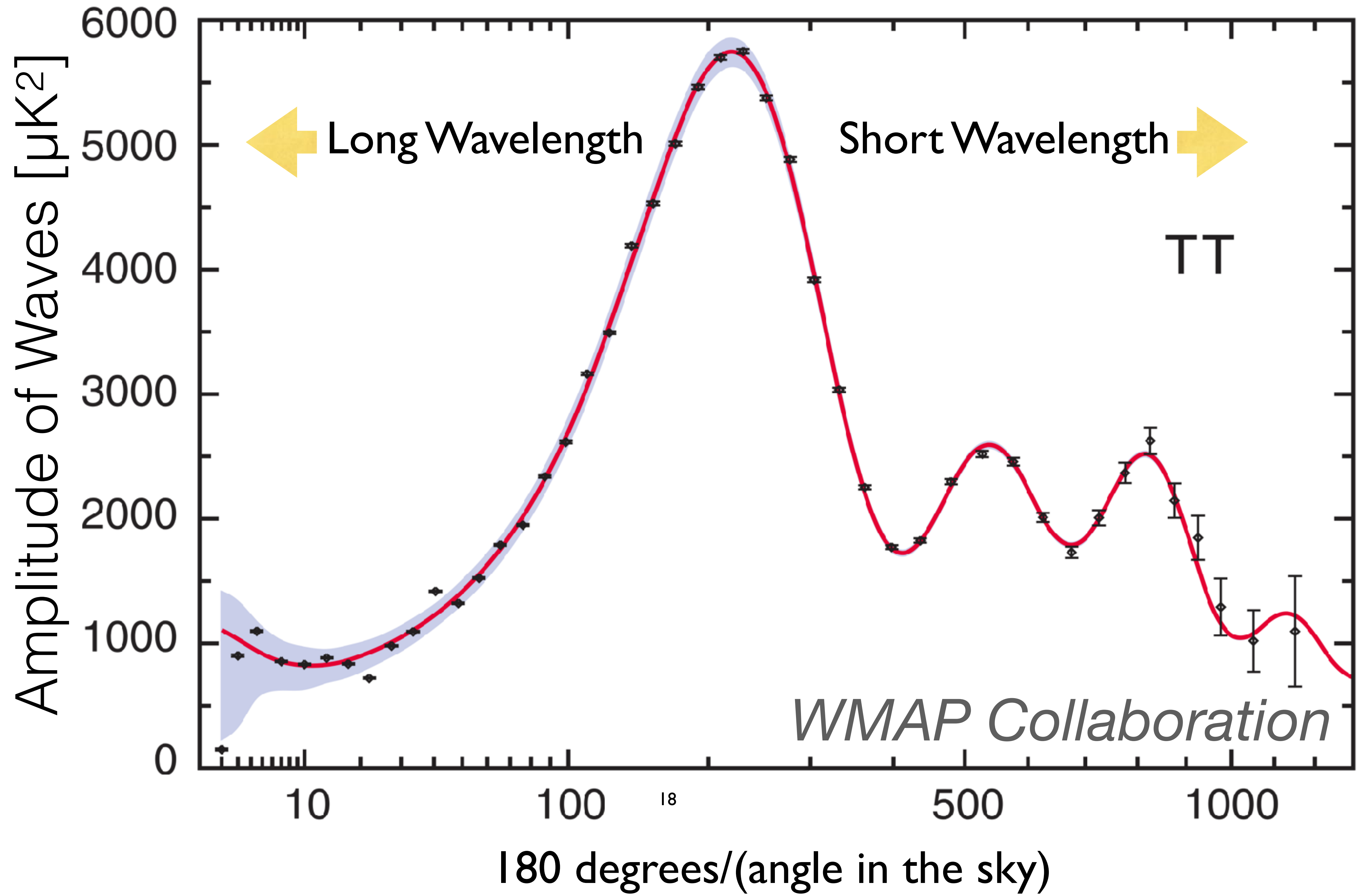
How do we  
analyse the  
data like this?



# Data Analysis

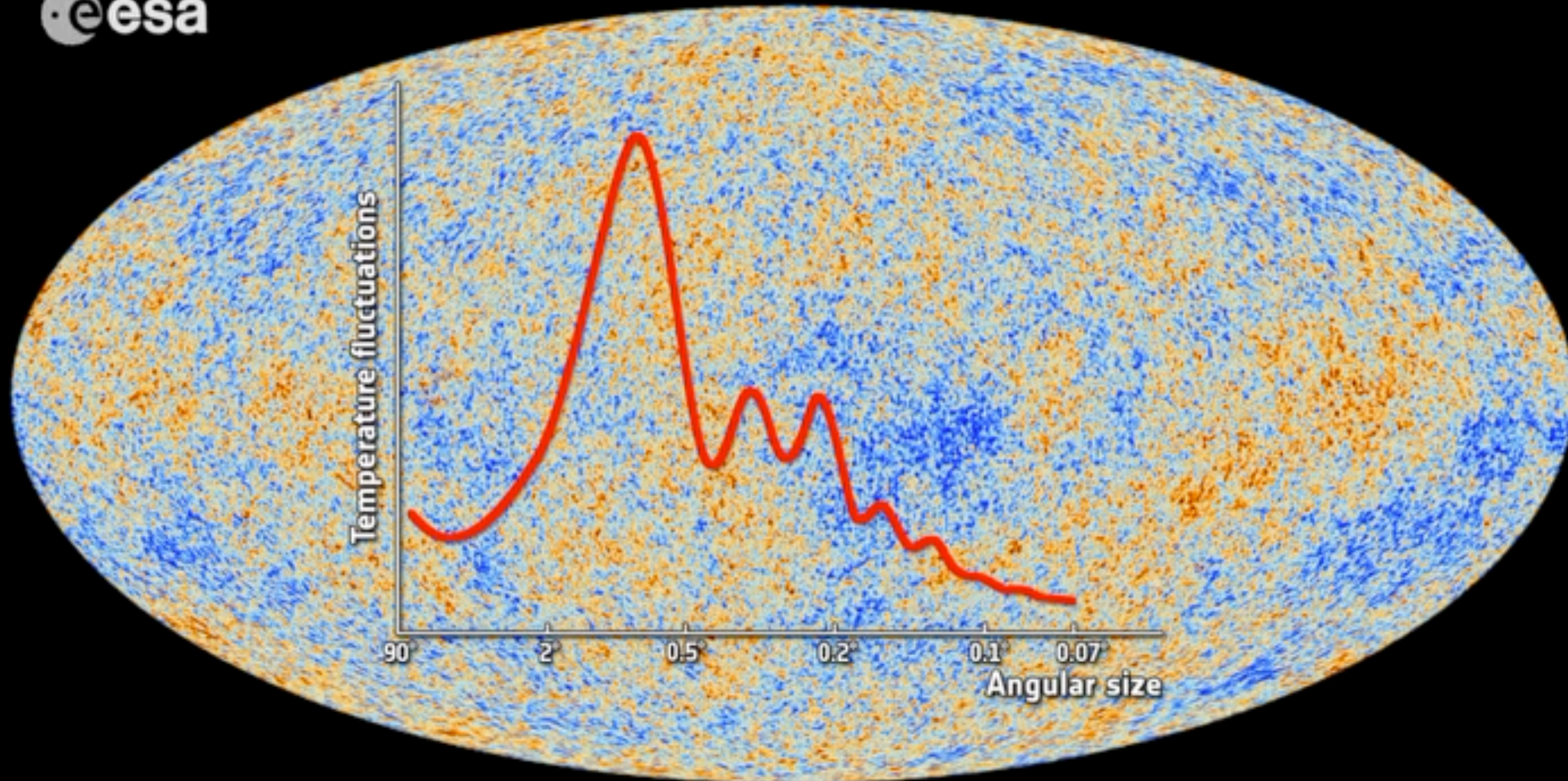
- Decompose temperature fluctuations in the sky into a set of waves with various wavelengths
- Make a diagram showing the strength of each wavelength: **Power Spectrum**







# Power Spectrum, Explained







The Royal Swedish Academy of Sciences has decided to award  
the 2019 Nobel Prize in Physics to

**JAMES PEEBLES**

"for theoretical discoveries in physical cosmology"

# James Peebles Facts

# Sound waves in the fireball Universe, predicted in 1970



James Peebles  
The Nobel Prize in Physics 2019

Born: 1935, Winnipeg, Canada

Affiliation at the time of the award: I  
Princeton, NJ, USA

Prize motivation: "for theoretical dis  
cosmology."

Prize share: 1/2

THE ASTROPHYSICAL JOURNAL, 162:815–836, December 1970

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## PRIMEVAL ADIABATIC PERTURBATION IN AN EXPANDING UNIVERSE\*

P. J. E. PEEBLES†

Joseph Henry Laboratories, Princeton University

AND

J. T. YU‡

Goddard Institute for Space Studies, NASA, New York

*Received 1970 January 5; revised 1970 April 1*







# Sound waves in the fireball Universe, predicted in 1970

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## SMALL-SCALE FLUCTUATIONS OF RELIC RADIATION\*

R. A. SUNYAEV and YA. B. ZELDOVICH

*Institute of Applied Mathematics, Academy of Sciences of the U.S.S.R., Moscow, U.S.S.R.*

(Received 11 September, 1969)

The Franklin Institute  
of Physics



and told me that I am lazy.





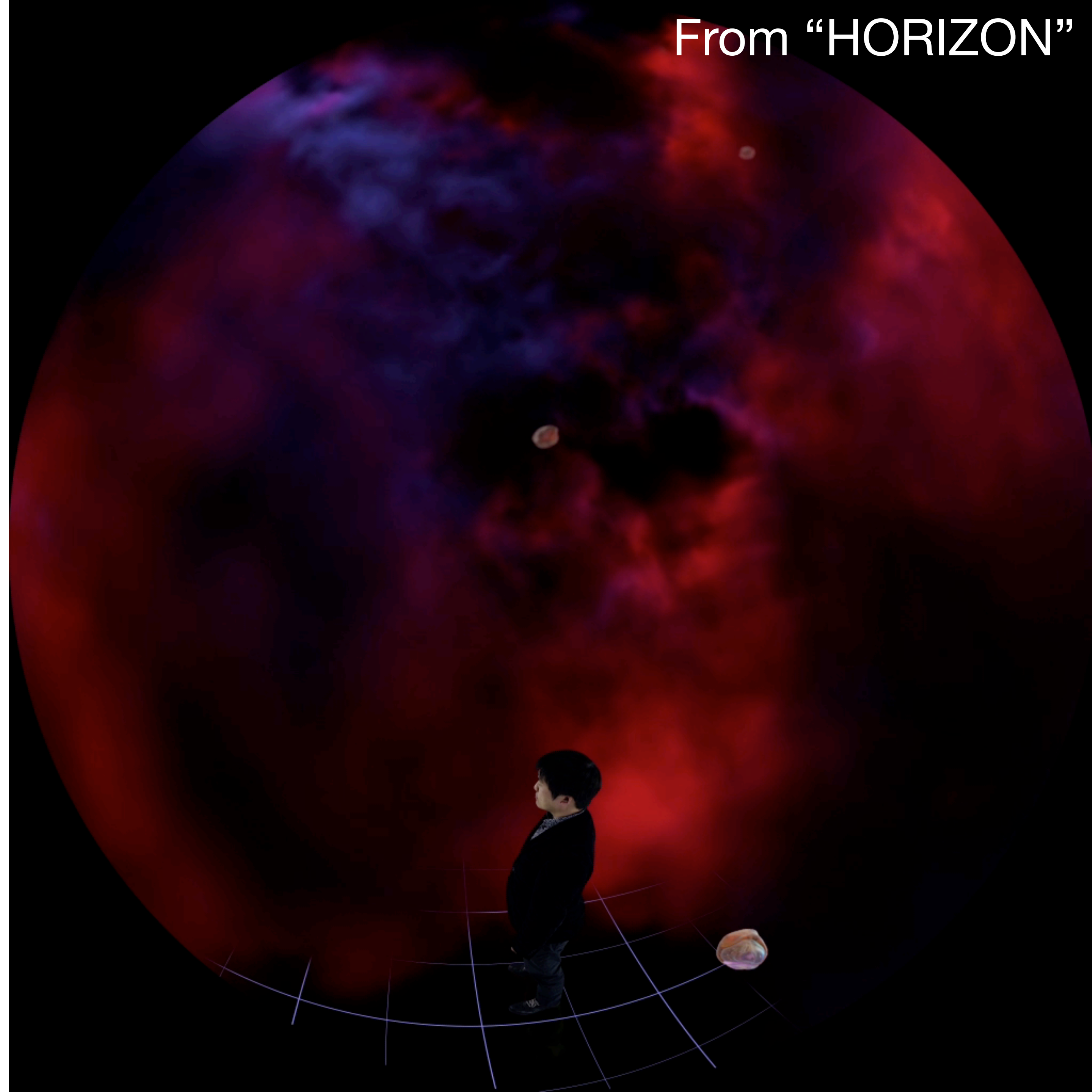
# Determine the composition of the Universe

## The Universe as a “hot soup”

- The power spectrum allows us to determine the composition of the Universe, such as the density of atoms, dark matter, and dark energy.



- **Definitive evidence for non-baryonic nature of dark matter!**





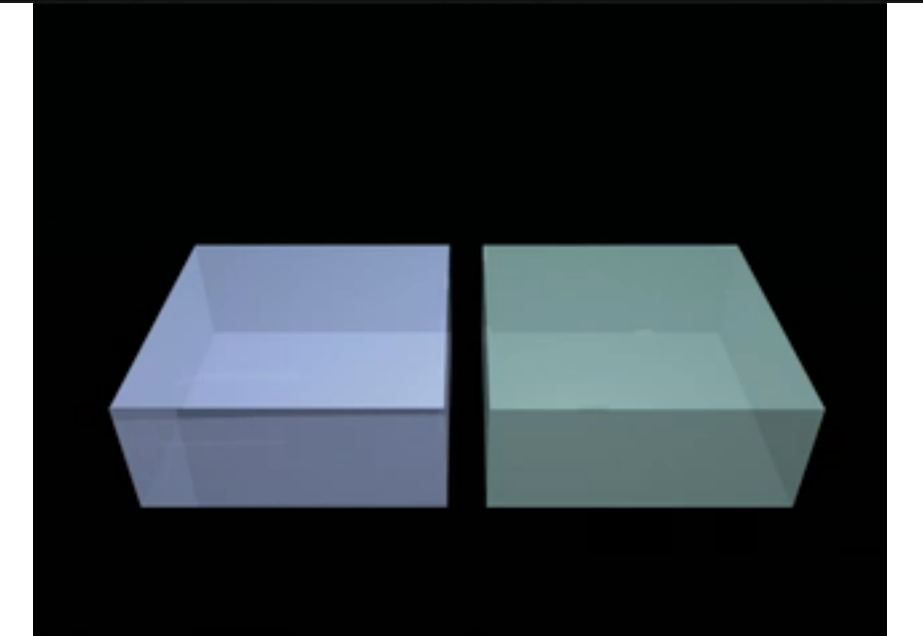
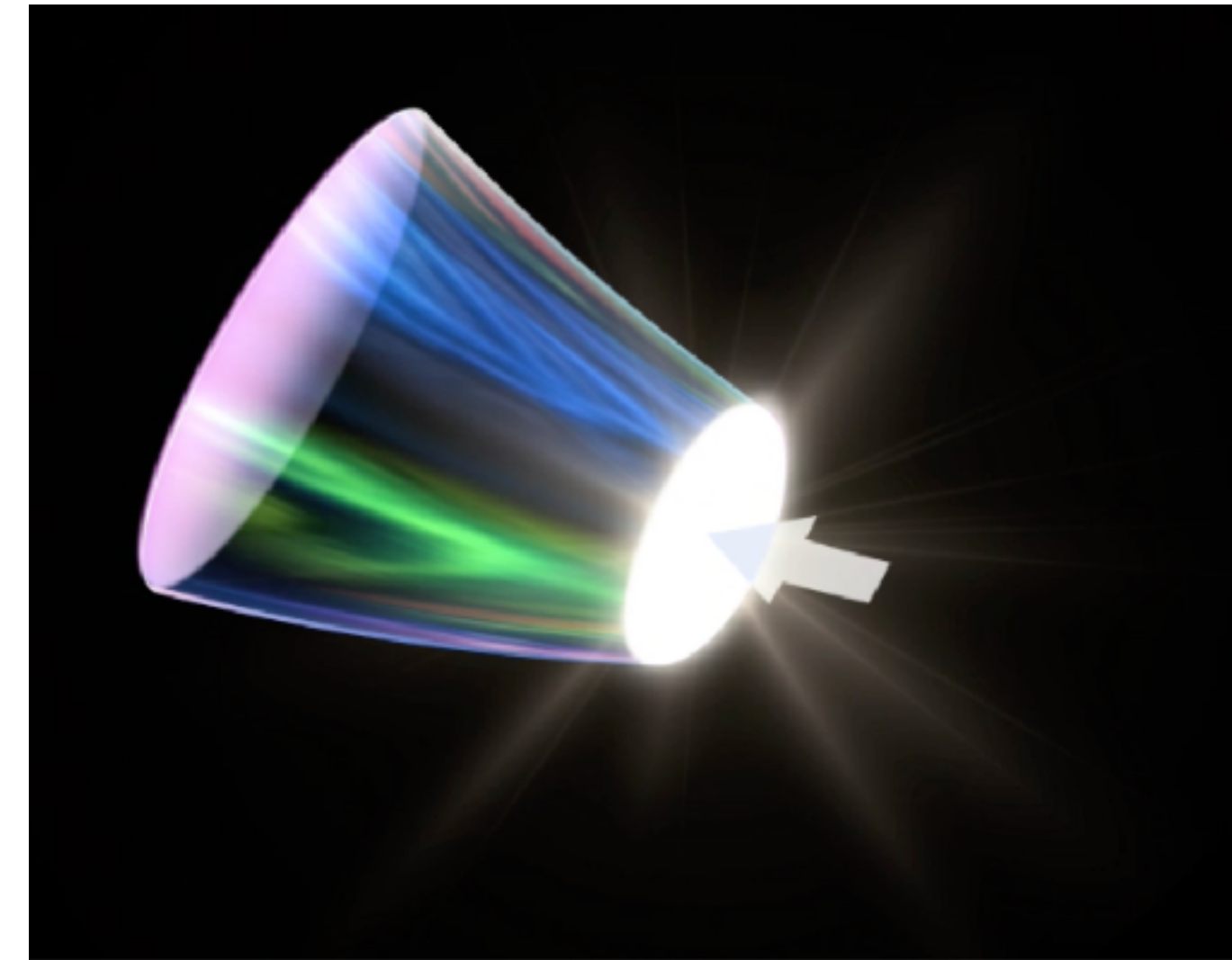
# “Let’s give some impact to the beginning of this model”

- What gave the initial fluctuation to the cosmic hot soup?

*Mukhanov & Chibisov (1981); Hawking (1982); Starobinsky (1982); Guth & Pi (1982);  
Bardeen, Turner & Steinhardt (1983)*

## Leading Idea:

- Quantum mechanics at work in the early Universe
  - “*We all came from quantum fluctuations*”
- But, how did the quantum fluctuation on the *microscopic* scale become *macroscopic* over large distances?
- **What is the missing link between the small and large scales?**





**Gravity + Quantum**

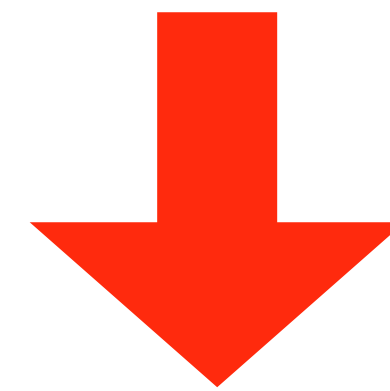
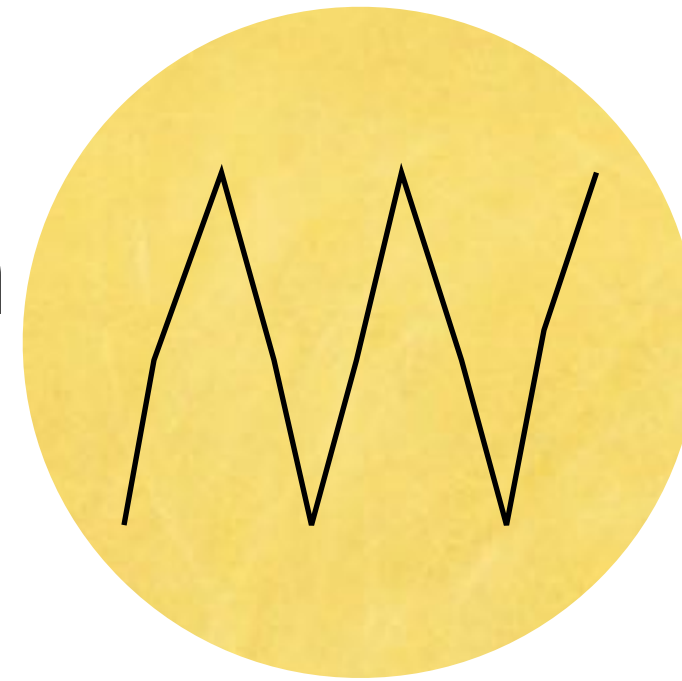
**= The origin of all the structures  
we see in the Universe**



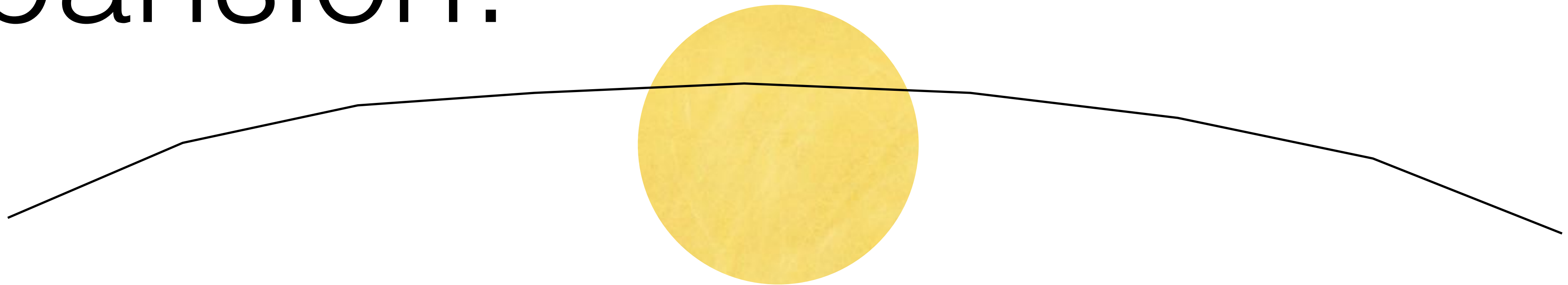
*Starobinsky (1980); Sato (1981); Guth (1981); Linde (1982); Albrecht & Steinhardt (1982)*

# Cosmic Inflation

Quantum mechanical fluctuation  
on microscopic scales



Exponential  
Expansion!



- Exponential expansion (inflation) stretches the wavelength of quantum fluctuations to cosmological scales



**What? How can we believe such  
a statement?**



# We have accumulated very good evidence so far

## The next step: Primordial Gravitational Waves

- Since the first discovery of the CMB temperature fluctuation by COBE in 1992, we have made a tremendous progress in making much more detailed measurements of the CMB over the last three decades.
- Three space missions, COBE (NASA) -> WMAP (NASA) -> Planck (ESA), as well as a host of ground-based and balloon-borne experiments. **Truly the global community effort!**
- What more do we want? **Primordial gravitational waves.** (Starobinsky 1979)
- Why more evidence? Because “*the extraordinary claim requires extraordinary evidence*” (Carl Sagan)

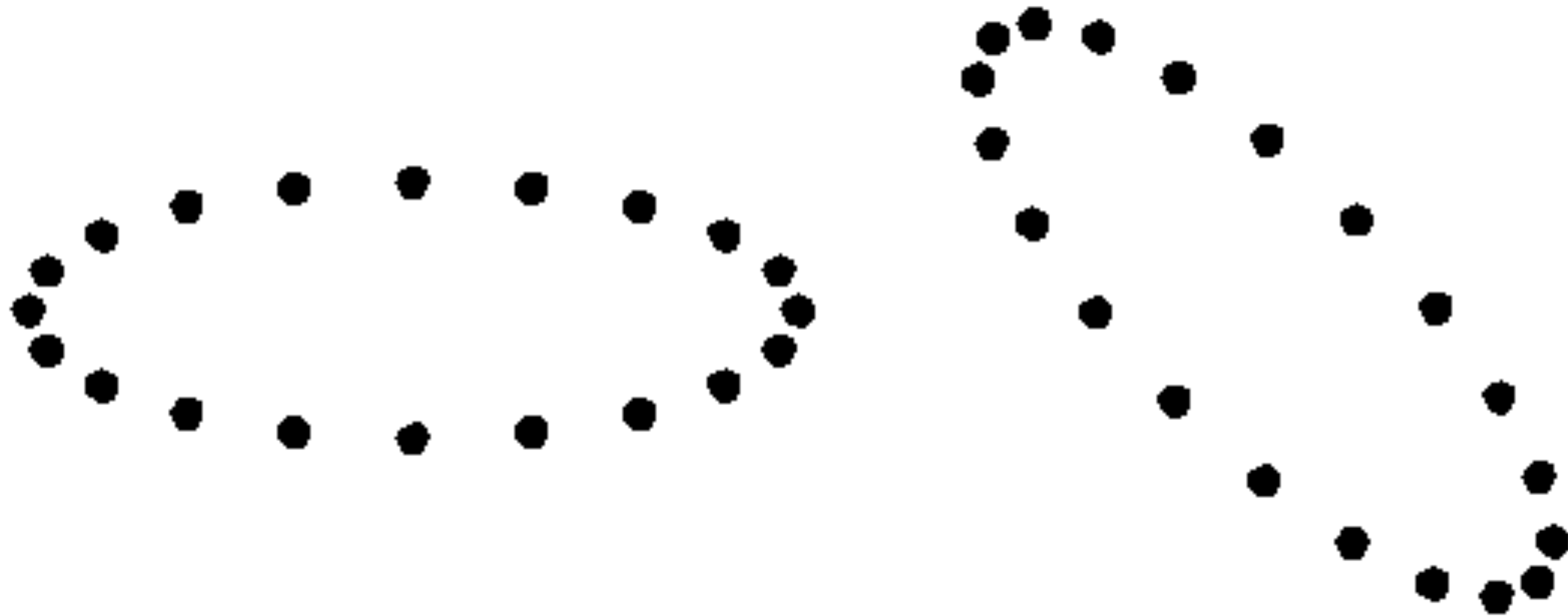


**Let's talk about the GW  
(tensor modes)**



# Gravitational waves are coming towards you!

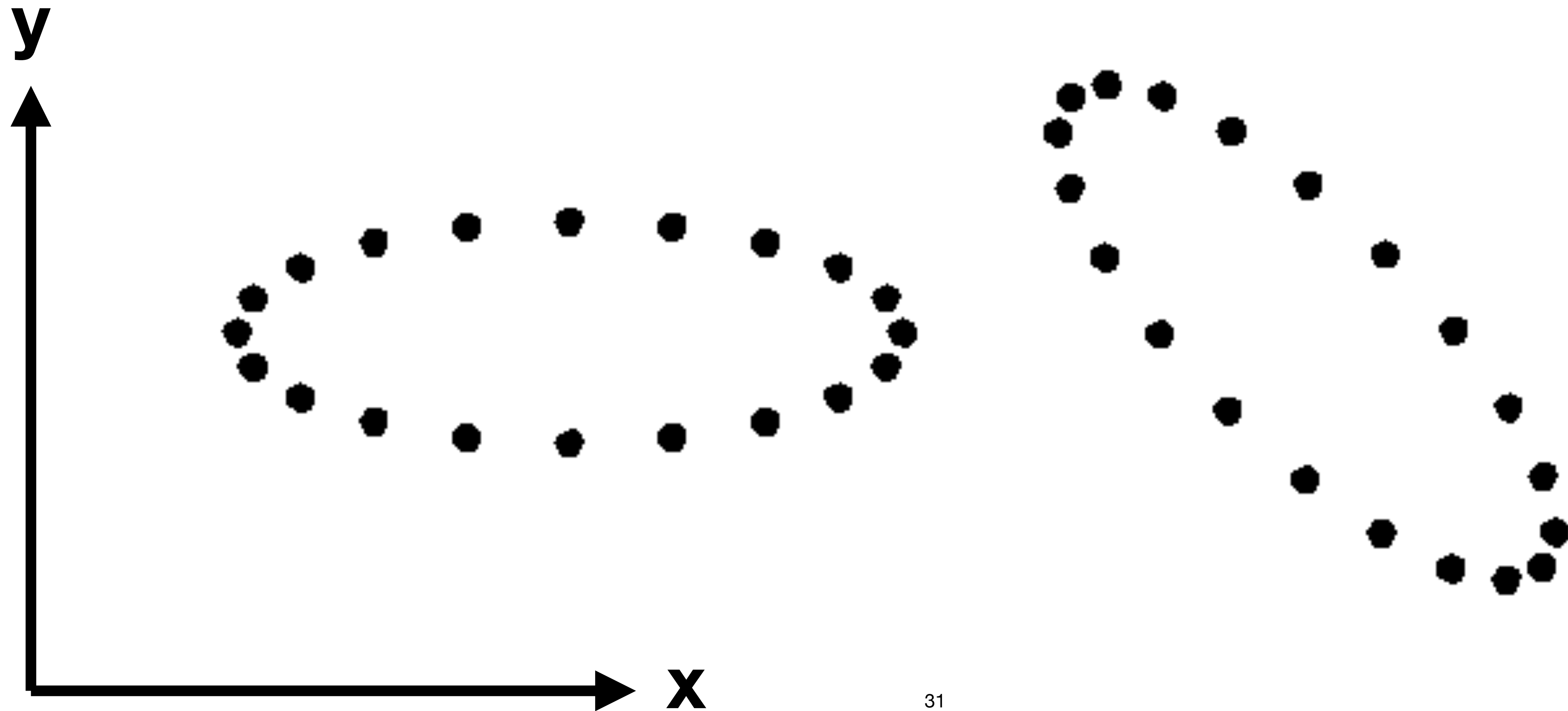
To visualise the waves, watch motion of test particles.





# Gravitational waves are coming towards you!

To visualise the waves, watch motion of test particles.





# Distance between two points

- In Cartesian coordinates, the distance between two points in Euclidean space is

$$ds^2 = dx^2 + dy^2 + dz^2$$

- To include the isotropic expansion of space,

$$ds^2 = \boxed{a^2(t)}(dx^2 + dy^2 + dz^2)$$

Scale Factor



# Distortion in space

- Compact notation using Kronecker's delta symbol:

$$ds^2 = a^2(t) \sum_{i=1}^3 \sum_{j=1}^3 \delta_{ij} dx^i dx^j$$

$x = (x, y, z)$

$\delta_{ij} = 1$  for  $i=j$ ;  
 $\delta_{ij} = 0$  otherwise

- To include distortion in space,

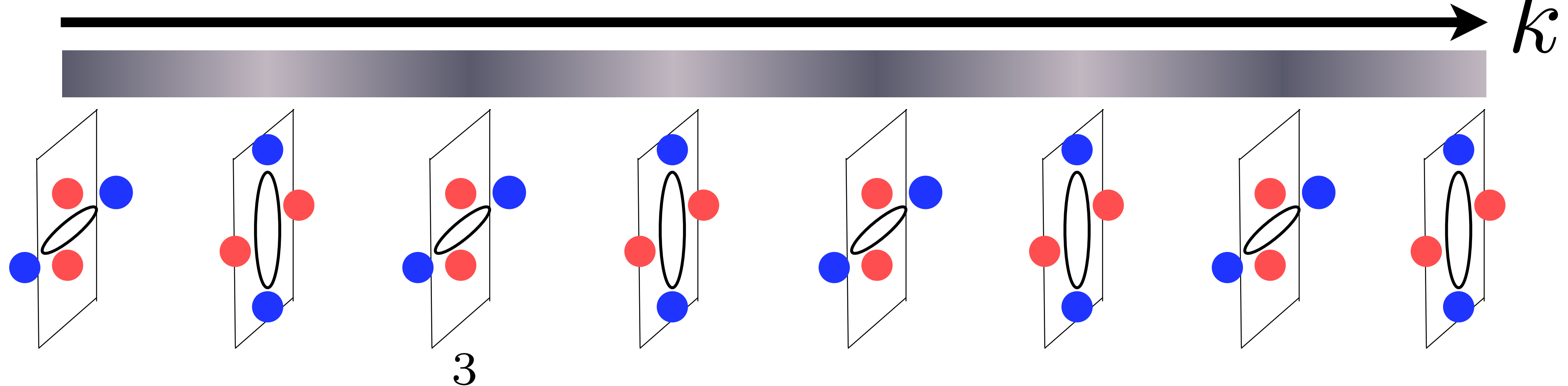
$$ds^2 = a^2 \sum_{i=1}^3 \sum_{j=1}^3 (\delta_{ij} + \boxed{h_{ij}}) dx^i dx^j$$

**Distortion in space!**



# Four conditions for gravitational waves

- The gravitational wave shall be transverse.
  - The direction of distortion is perpendicular to the propagation direction  $\vec{k}$



Thus,

$$\sum_{i=1}^3 k^i h_{ij} = 0$$

3 conditions for  $h_{ij}$

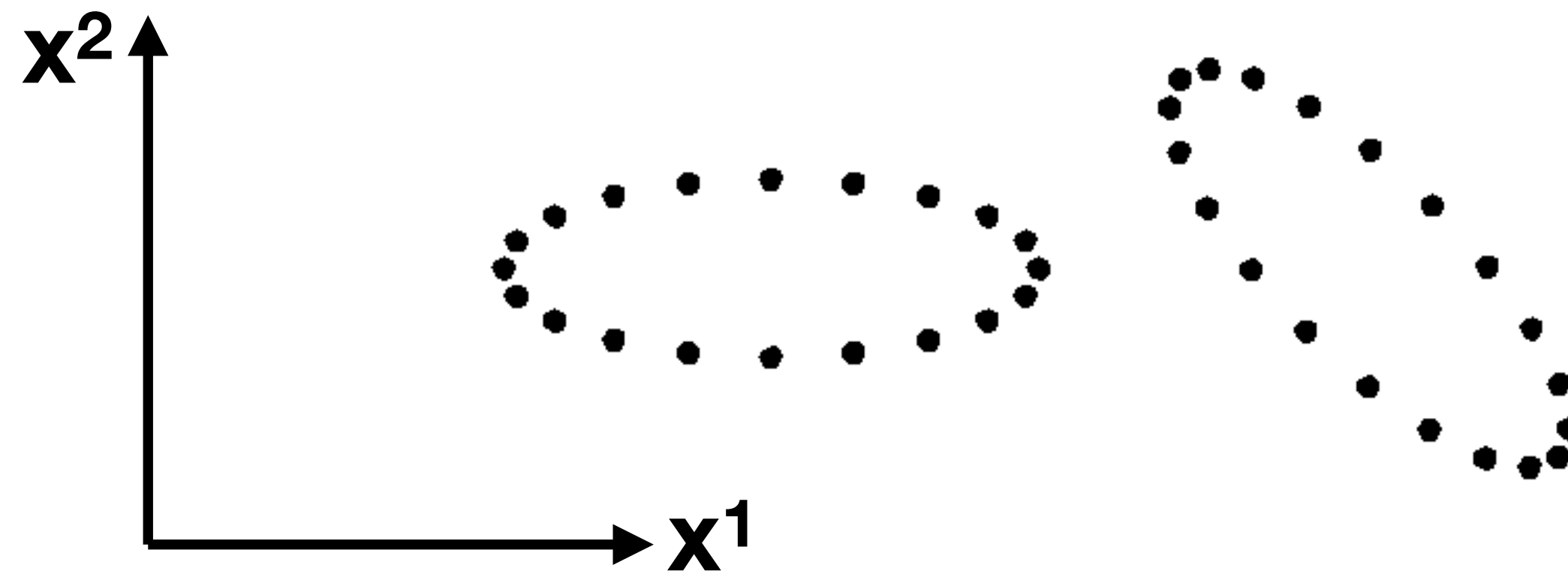


# Four conditions for gravitational waves

- The gravitational wave shall not change the area

- The determinant of  $\delta_{ij} + h_{ij}$  is 1

$$ds^2 = a^2 \sum_{i=1}^3 \sum_{j=1}^3 (\delta_{ij} + h_{ij}) dx^i dx^j$$



Thus, 
$$\sum_{i=1}^3 h_{ii} = 0$$

1 condition for  $h_{ij}$

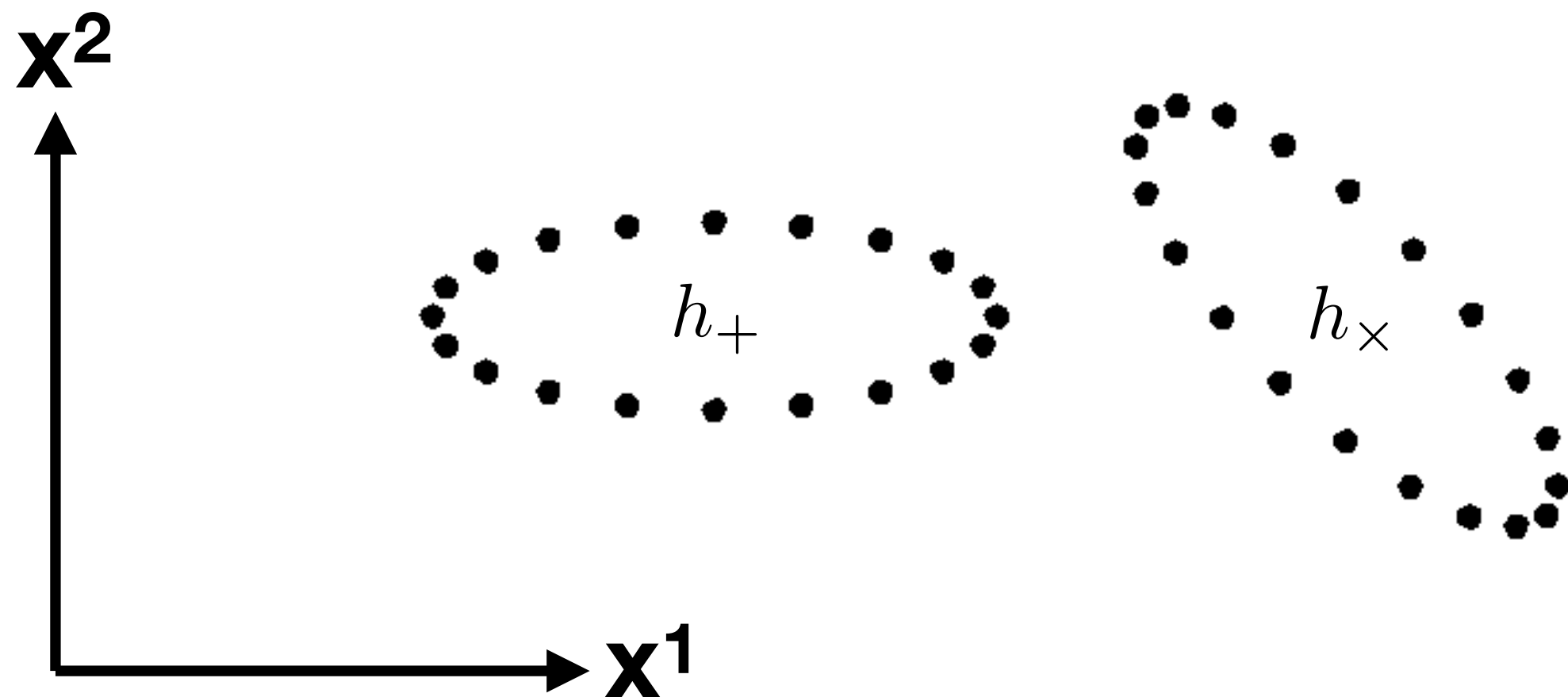


# 6 – 4 = 2 degrees of freedom for GW

We call them “plus” and “cross” modes

- The symmetric matrix  $h_{ij}$  has 6 components, but there are 4 conditions. Thus, we have two degrees of freedom.
- If the GW propagates in the  $x^3=z$  axis, non-vanishing components of  $h_{ij}$  are

$$h_{ij} = \begin{pmatrix} h_+ & h_\times & 0 \\ h_\times & -h_+ & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

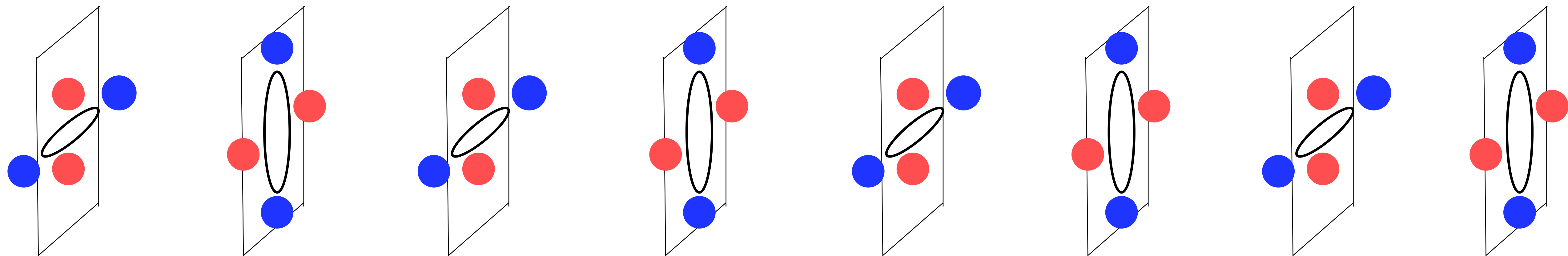




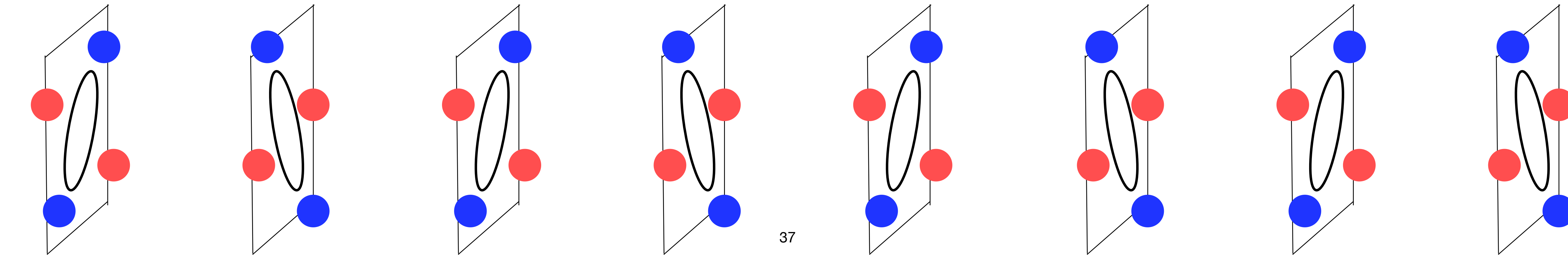
Propagation direction of GW  $\vec{k}$



$h_+ = \cos(kz)$



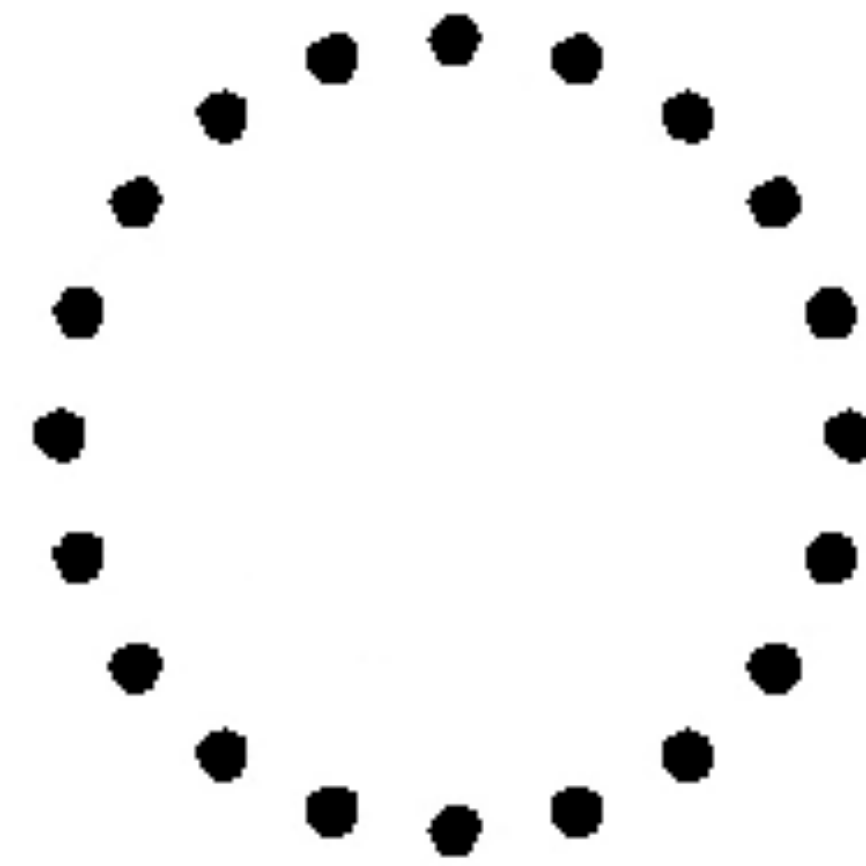
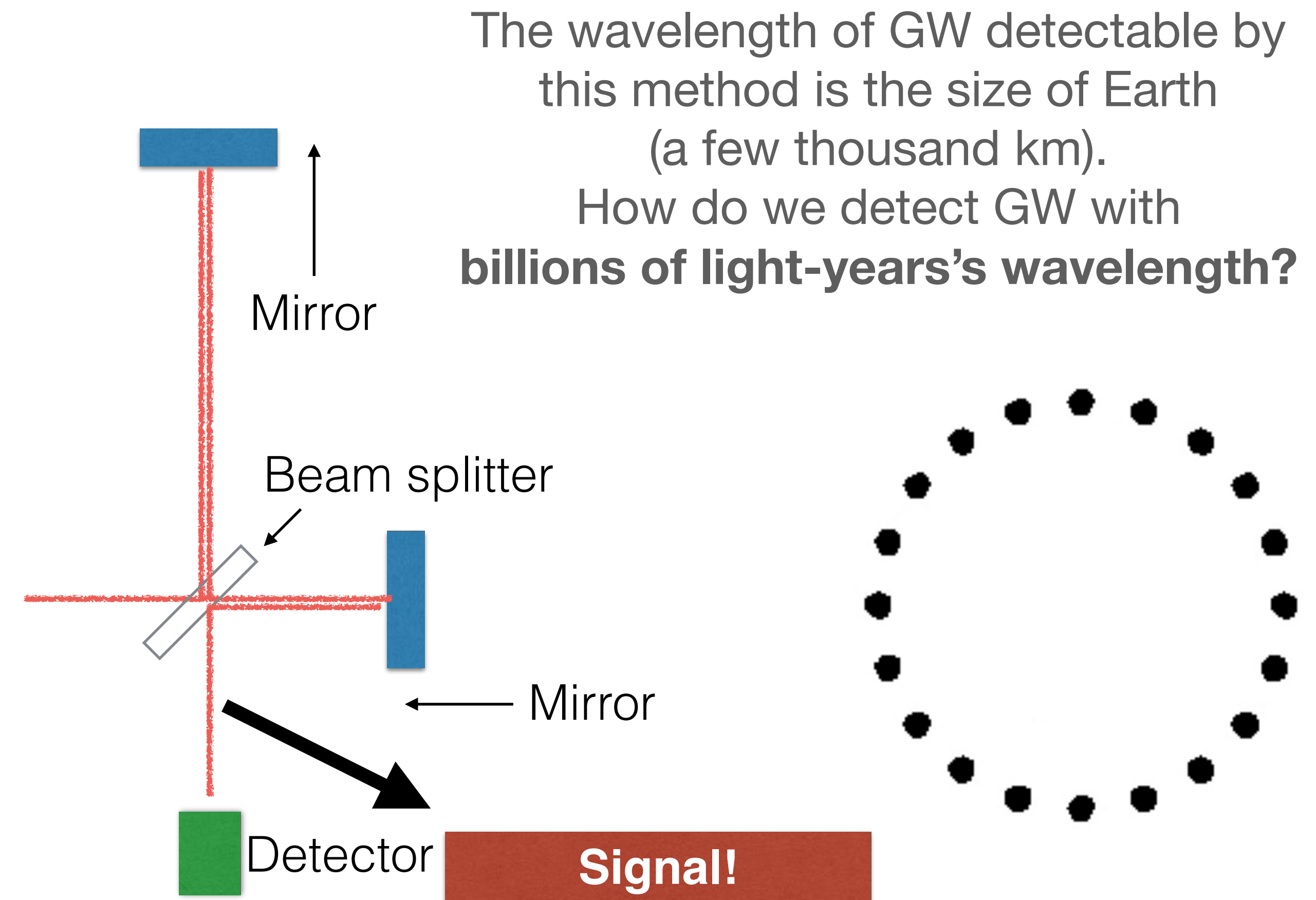
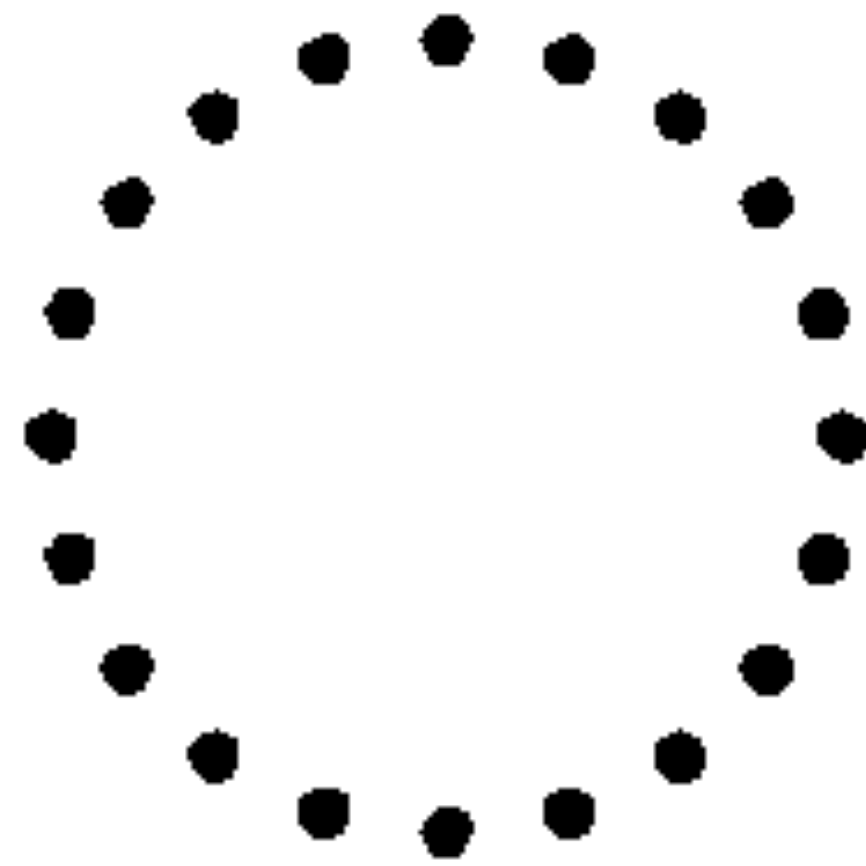
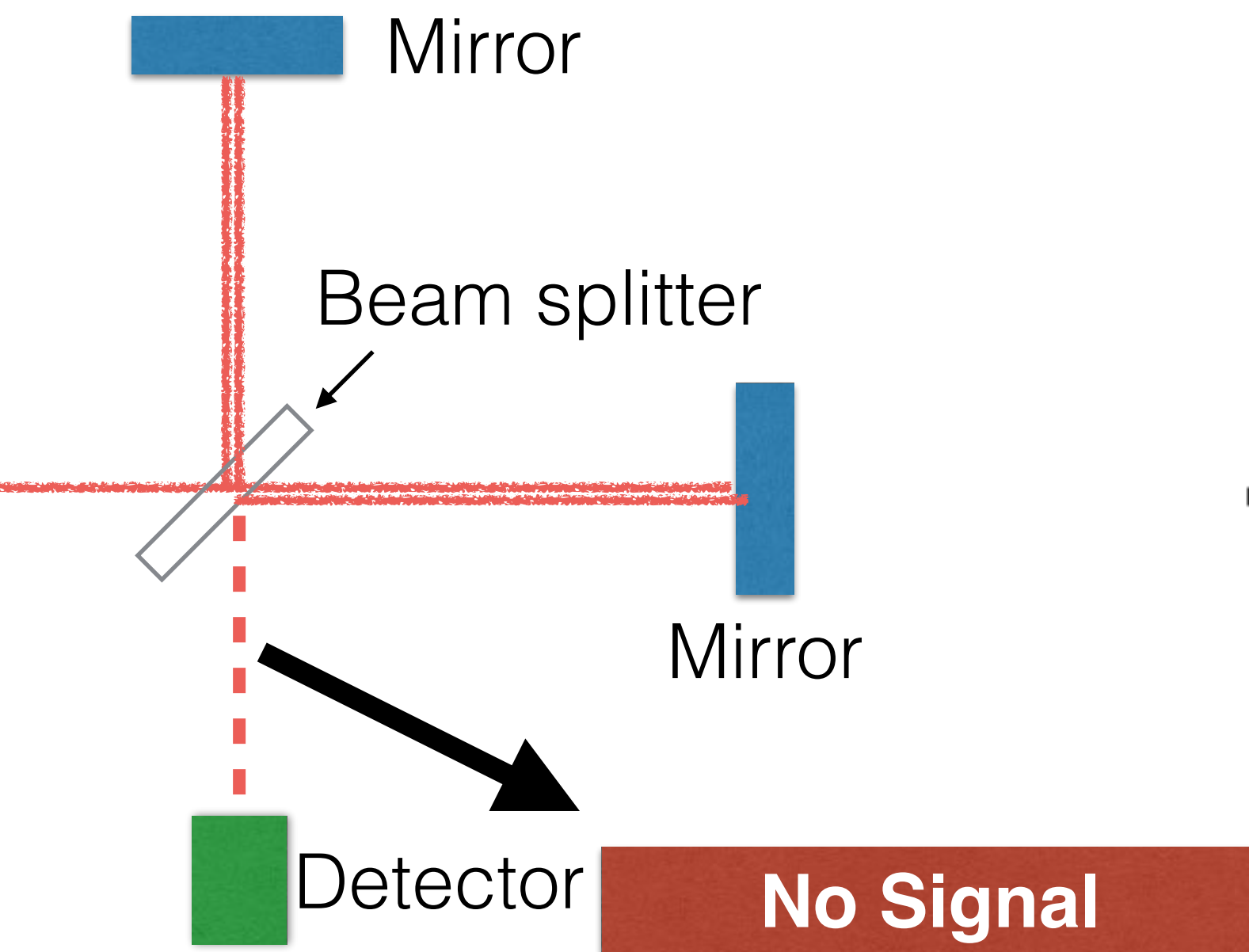
$h_x = \cos(kz)$





# How to detect GW?

## Laser interferometer technique, used by LIGO and VIRGO

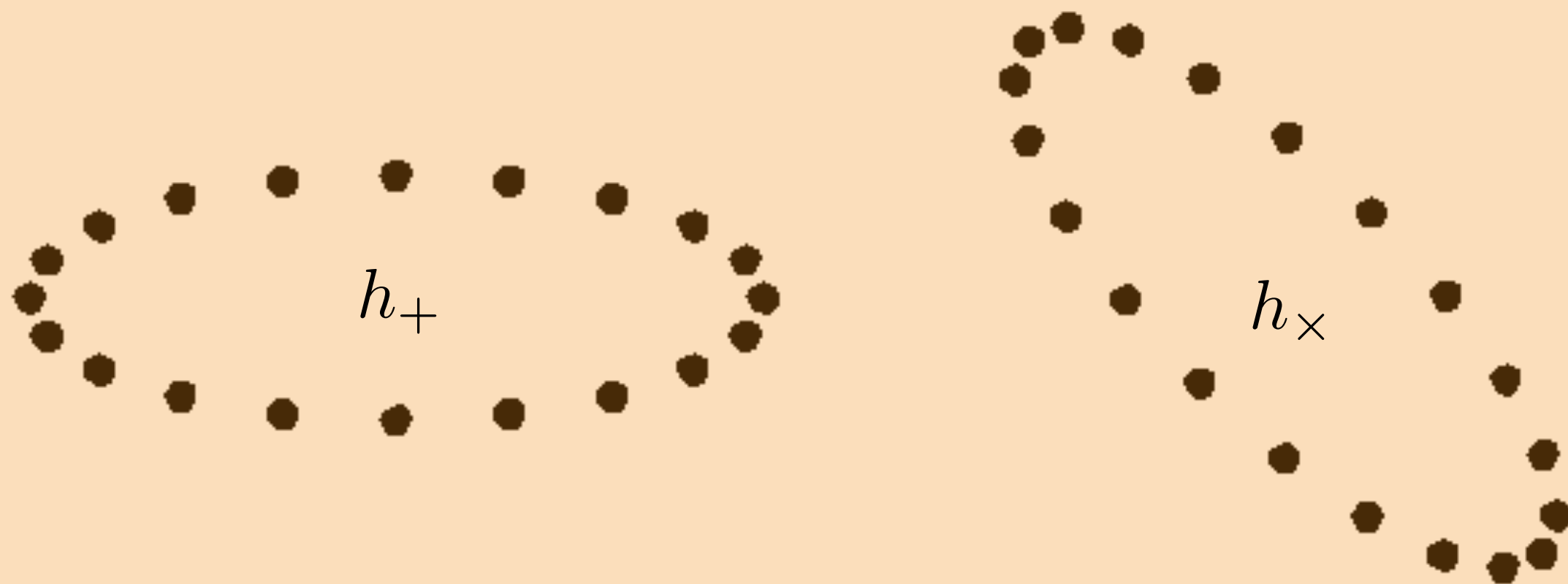




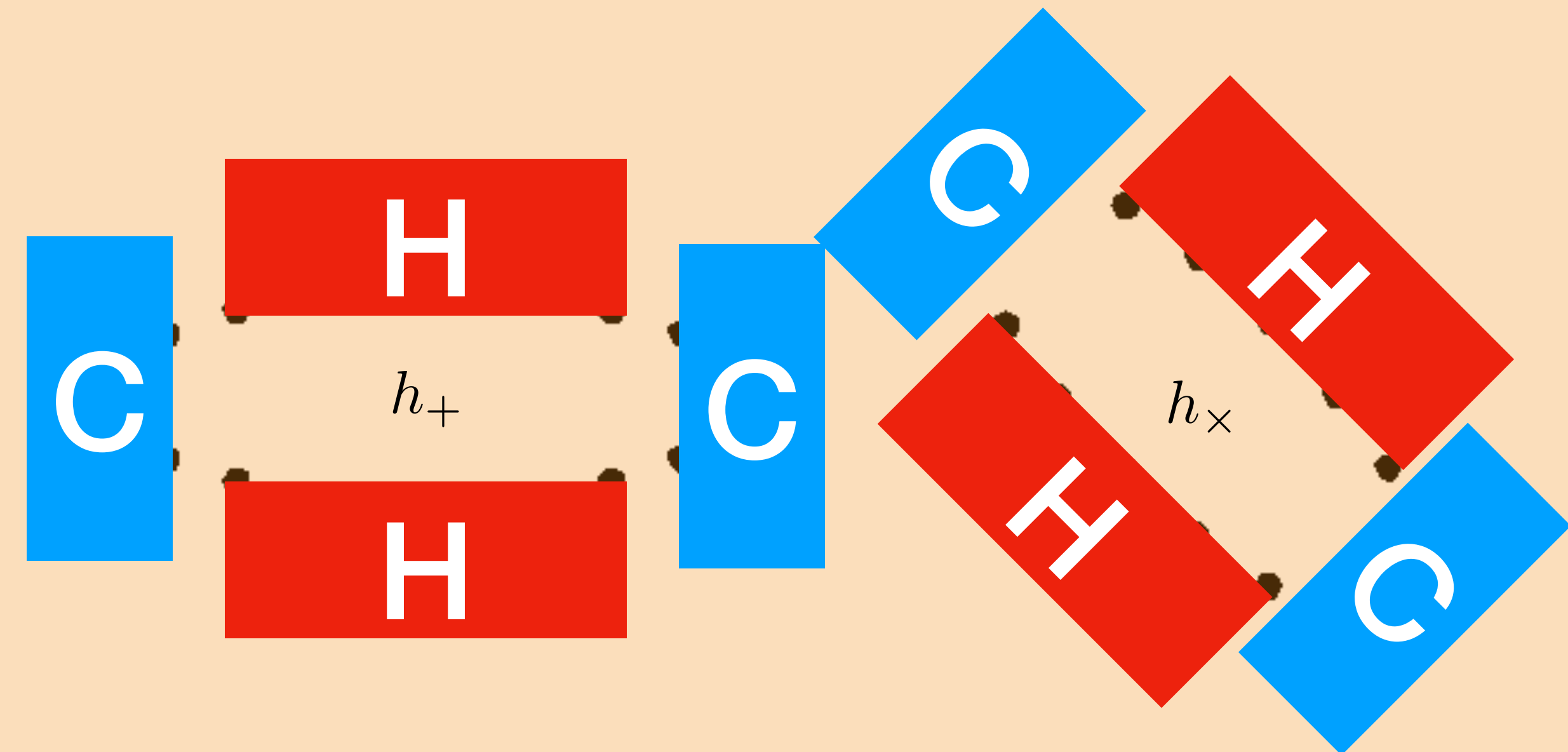
# Detecting GW by CMB

Quadrupole temperature anisotropy generated by red- and blue-shifting of photons

Isotropic radiation field (CMB)



Isotropic radiation field (CMB)

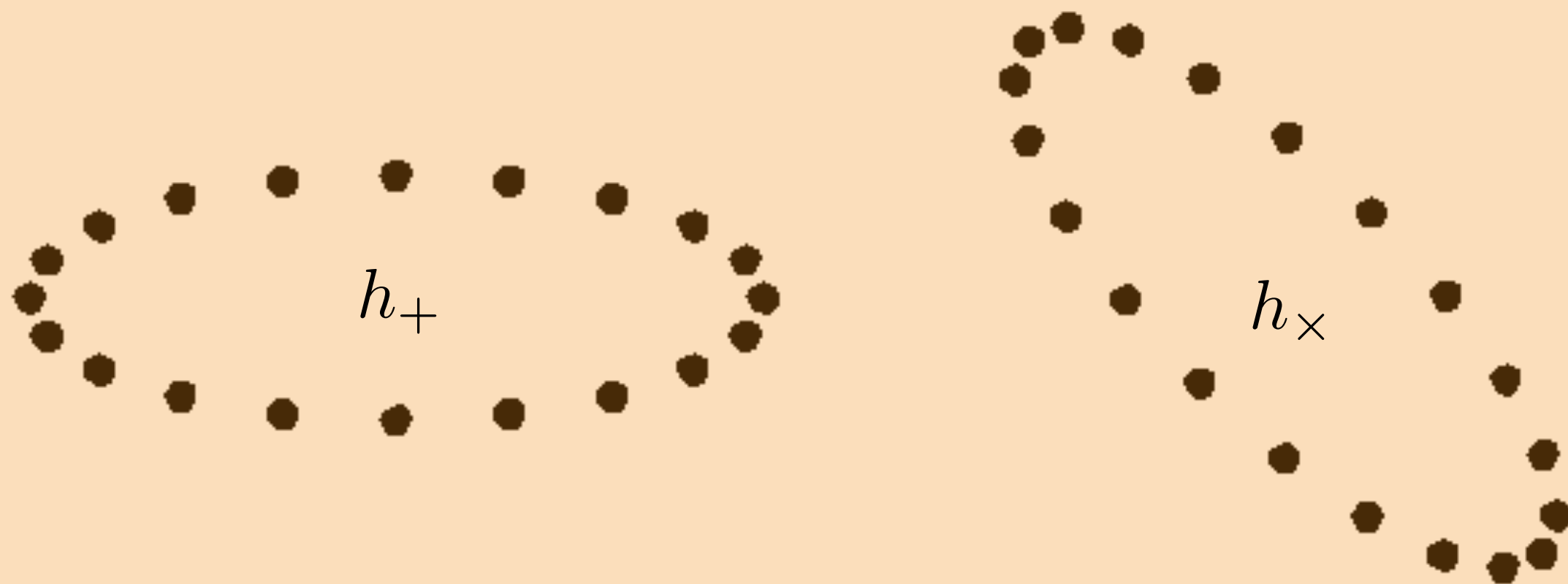




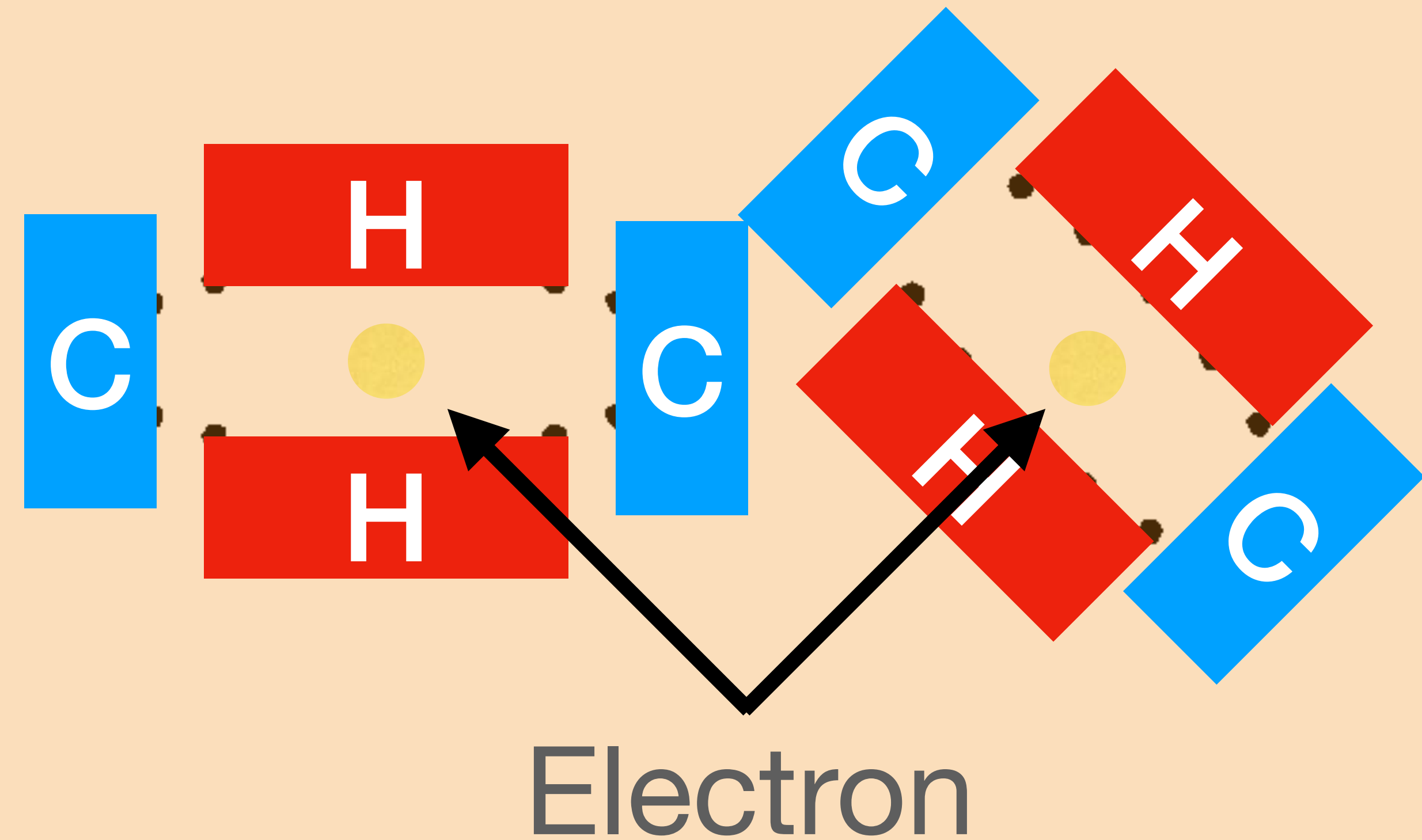
# Detecting GW by CMB

Quadrupole temperature anisotropy generated by red- and blue-shifting of photons

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Isotropic radiation field (CMB)

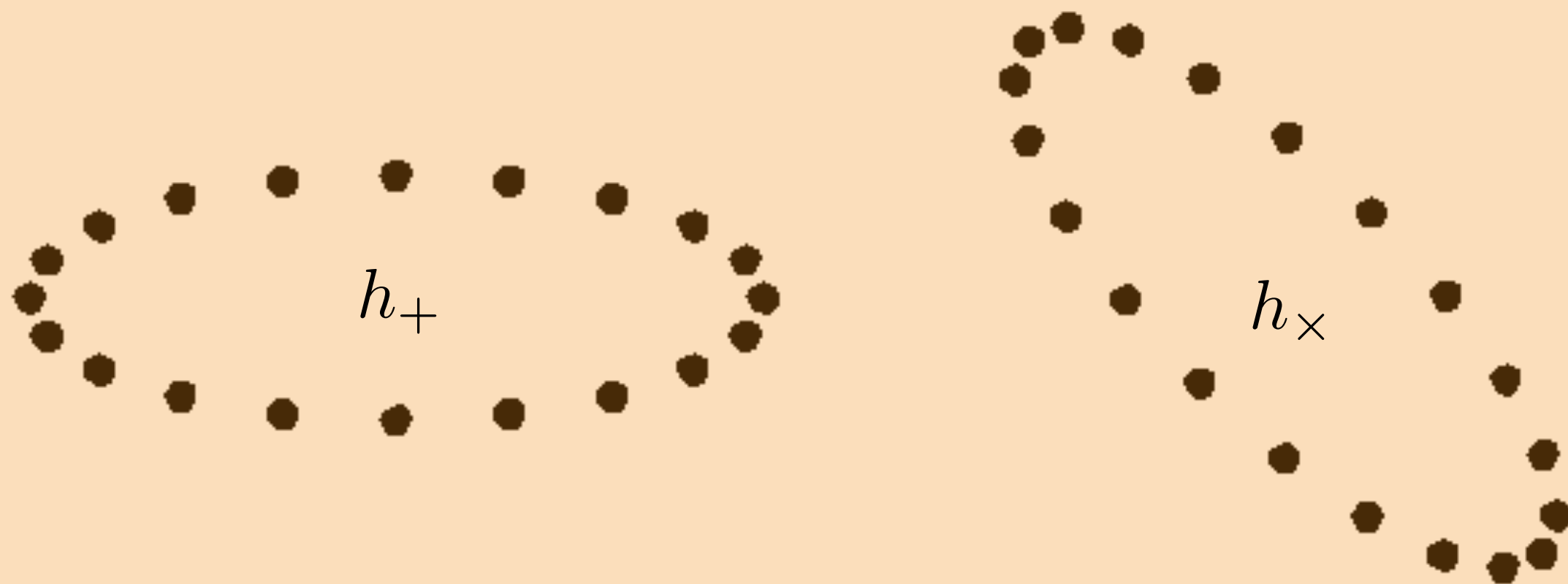




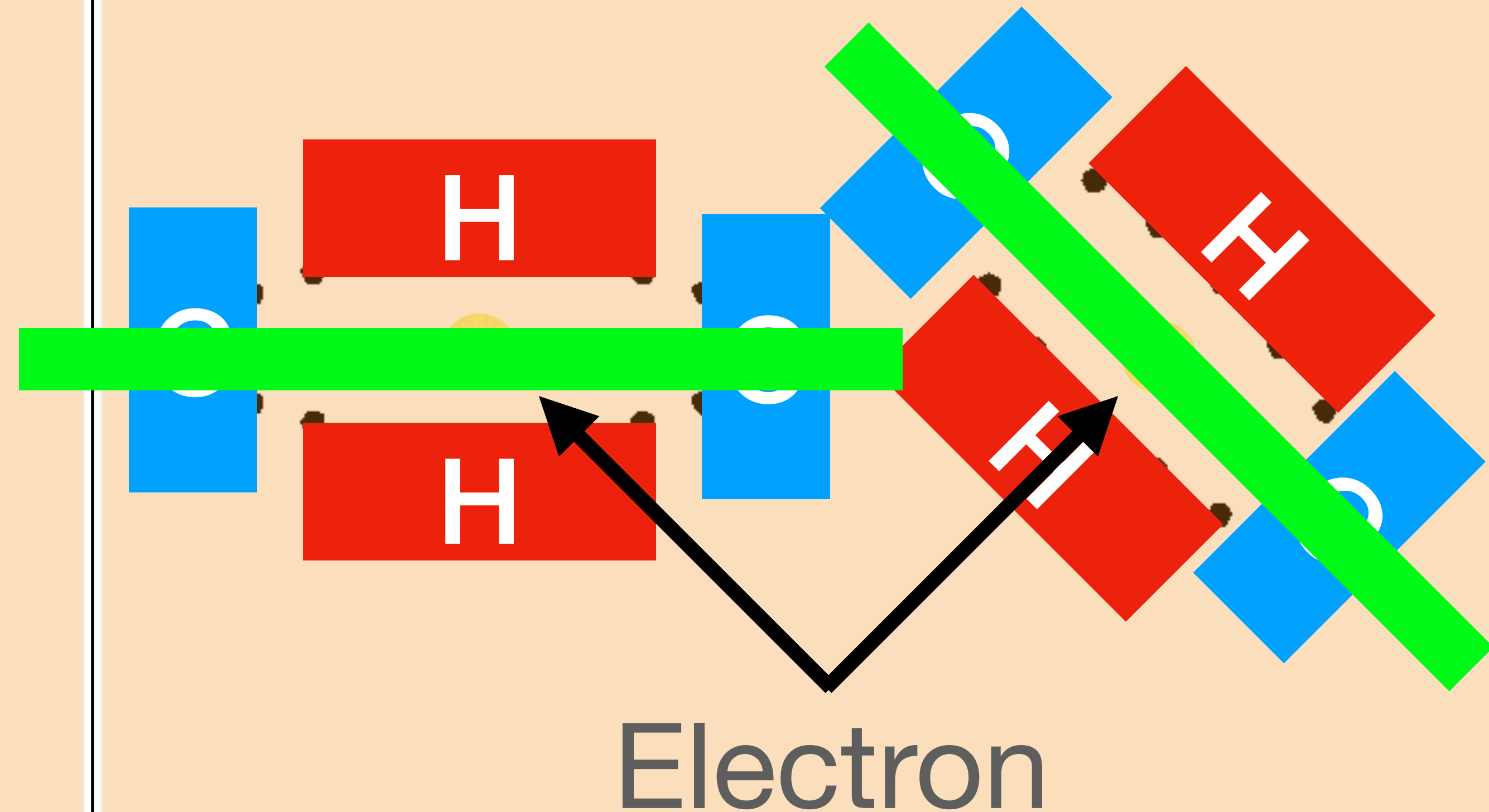
# Detecting GW by CMB *Polarisation*

Quadrupole temperature anisotropy scattered by an electron

Isotropic radiation field (CMB)



Isotropic radiation field (CMB)





Credit: TALEX





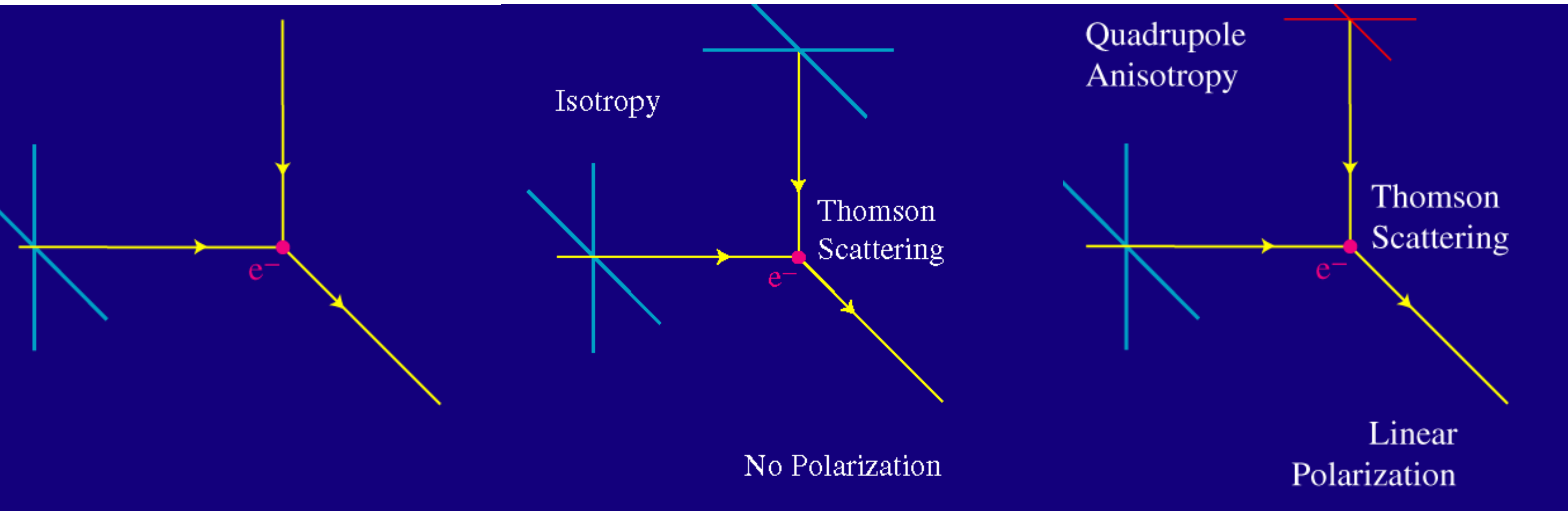
Credit: TALEX





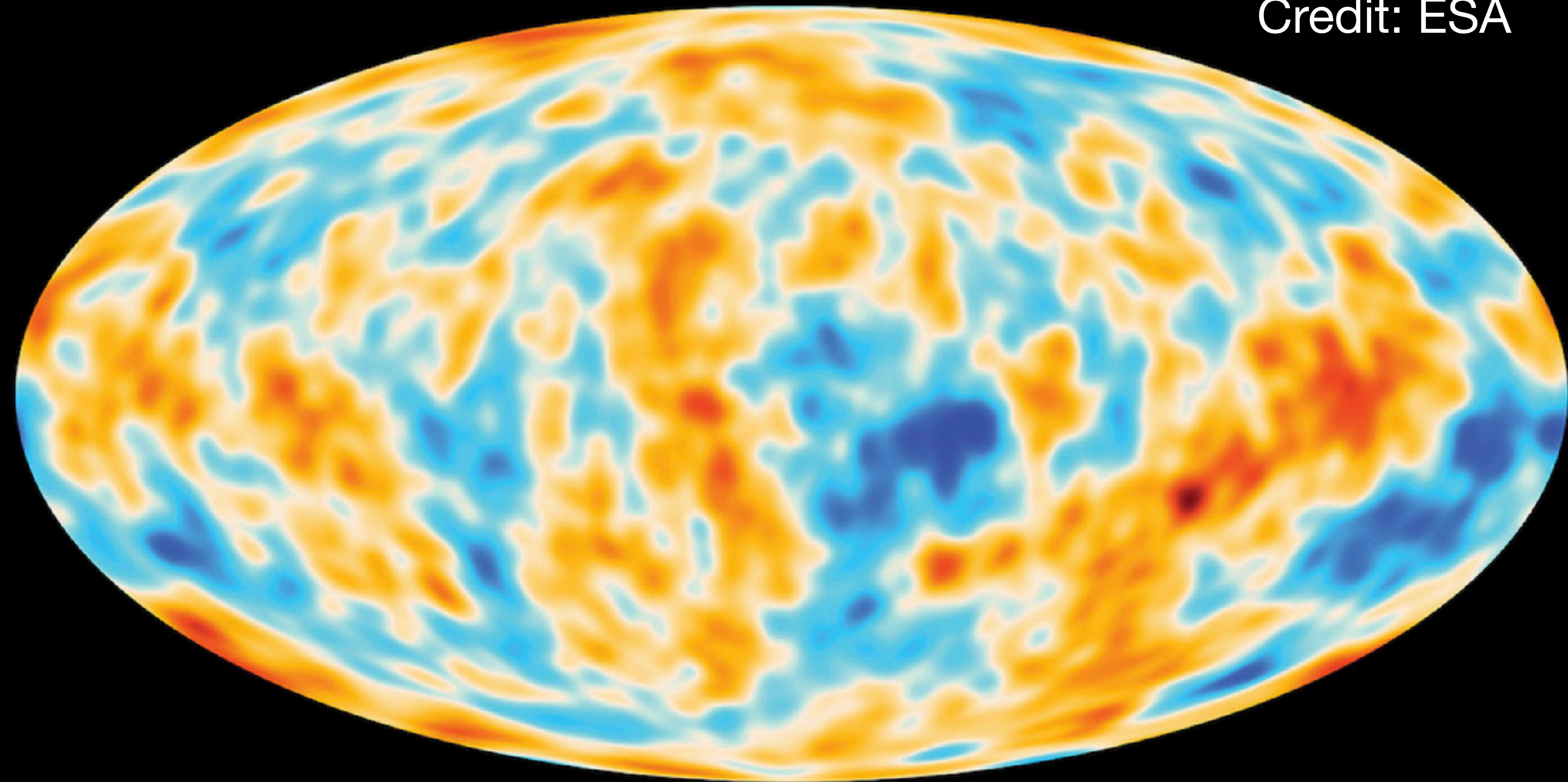
# Physics of CMB Polarisation

Necessary and sufficient condition: Scattering and Quadrupole Anisotropy





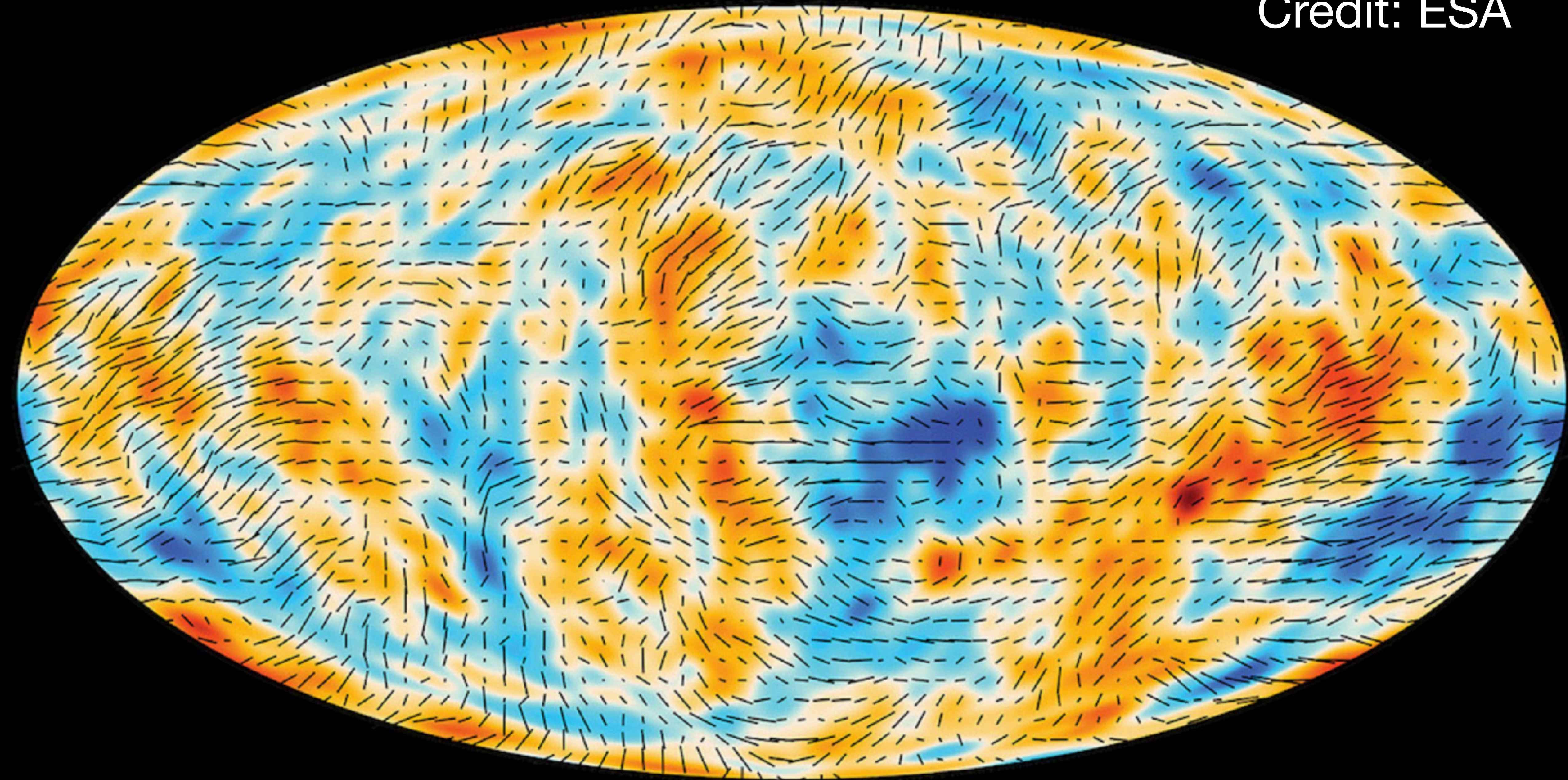
Credit: ESA



Temperature (smoothed)



Credit: ESA

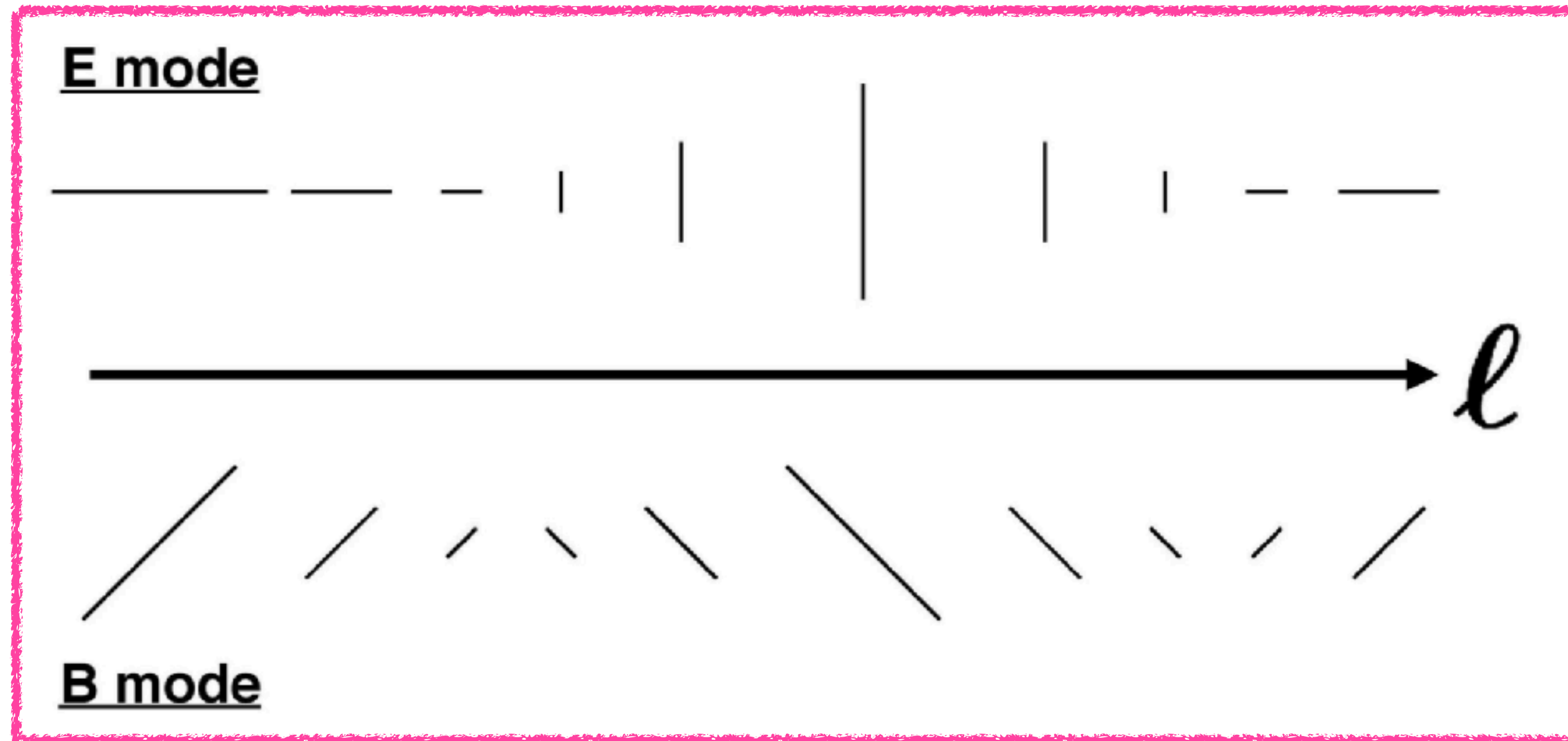


Temperature (smoothed) + Polarisation



# E- and B-mode decomposition

Concept defined in Fourier space



Direction of the Fourier  
wavenumber vector

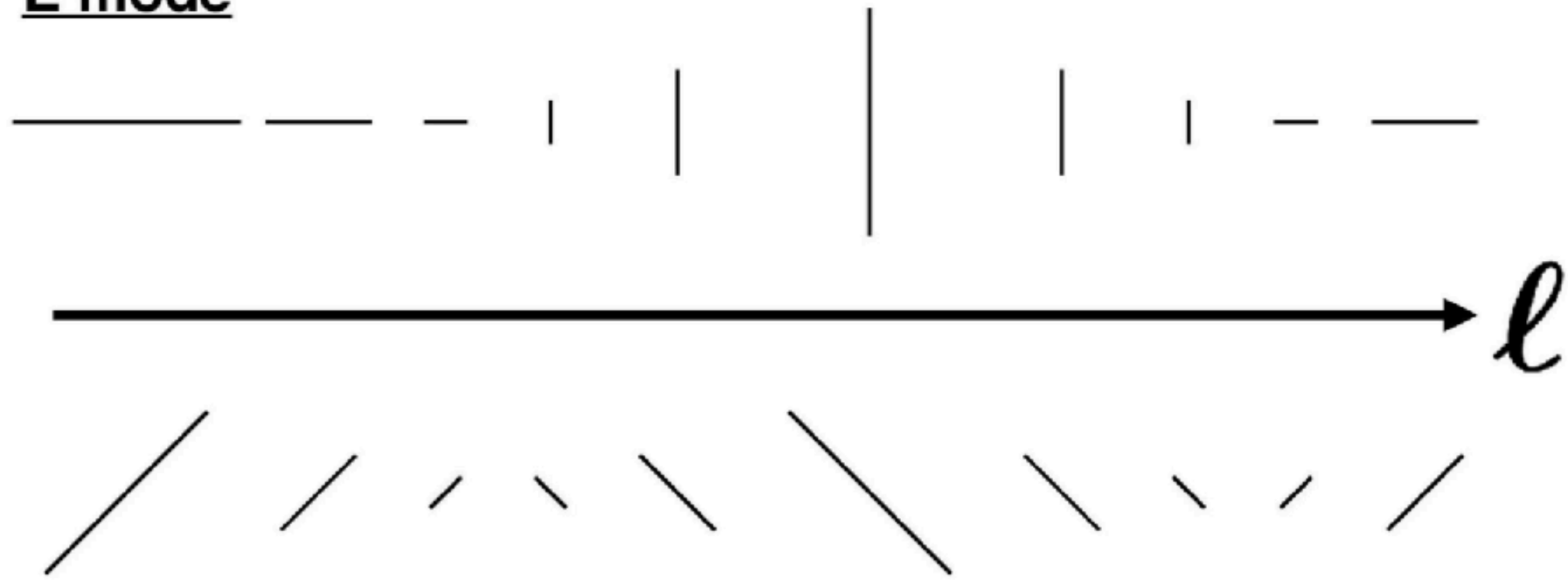
- **E-mode** : Polarisation directions are **parallel or perpendicular** to the wavenumber direction
- **B-mode** : Polarisation directions are **45 degrees tilted** w.r.t the wavenumber direction



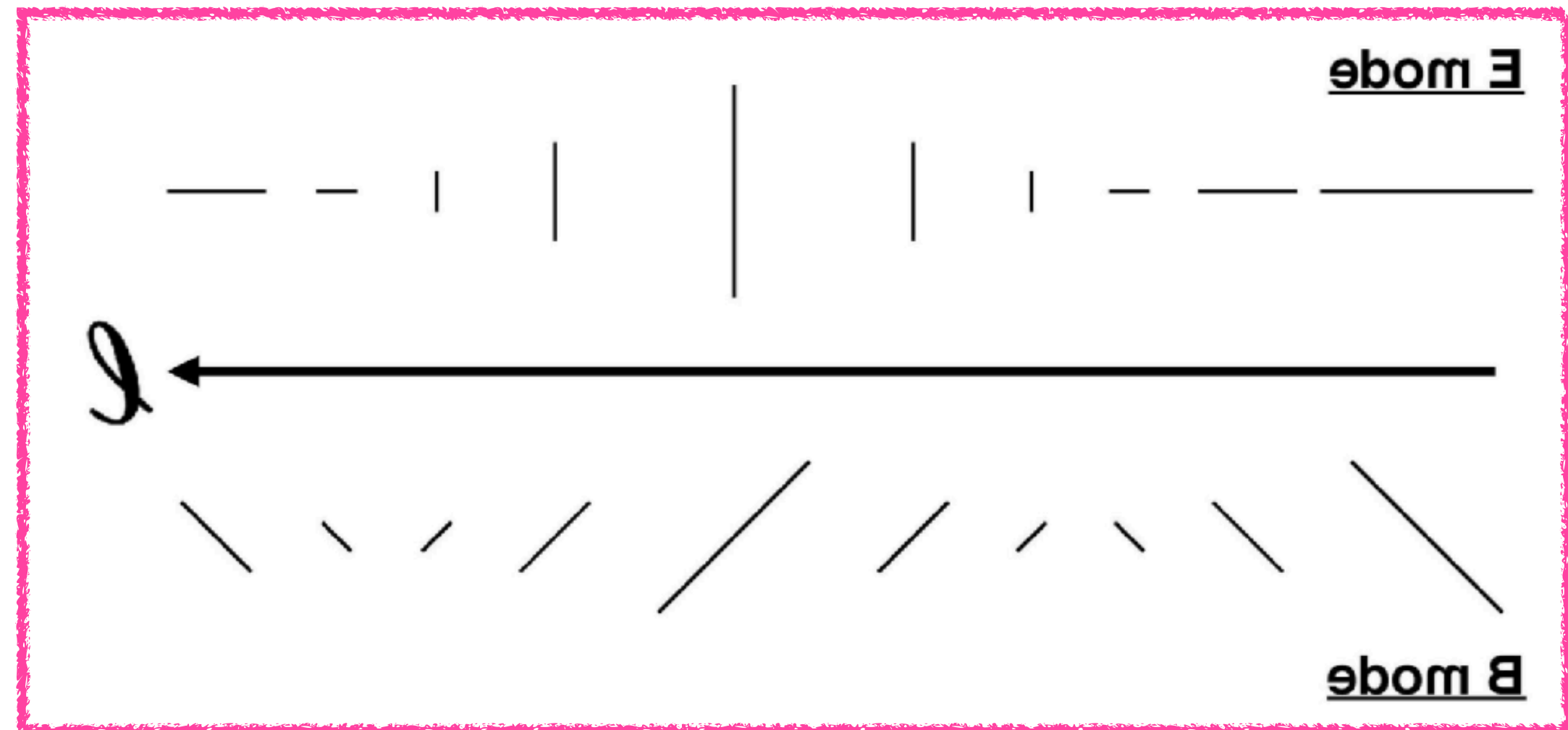
# Parity Flip

**E-mode remains the same, whereas B-mode changes the sign**

E mode



B mode



- Two-point correlation functions invariant under the parity flip are

$$\langle E_{\ell} E_{\ell'}^* \rangle = (2\pi)^2 \delta_D^{(2)}(\ell - \ell') C_{\ell}^{EE}$$

$$\langle B_{\ell} B_{\ell'}^* \rangle = (2\pi)^2 \delta_D^{(2)}(\ell - \ell') C_{\ell}^{BB}$$

$$\langle T_{\ell} E_{\ell'}^* \rangle = \langle T_{\ell}^* E_{\ell'} \rangle = (2\pi)^2 \delta_D^{(2)}(\ell - \ell') C_{\ell}^{TE}$$

- The other combinations  $\langle TB \rangle$  and  $\langle EB \rangle$  are not invariant under the parity flip.

- [Side Note]** We can use these combinations to probe parity-violating physics (e.g., axions)

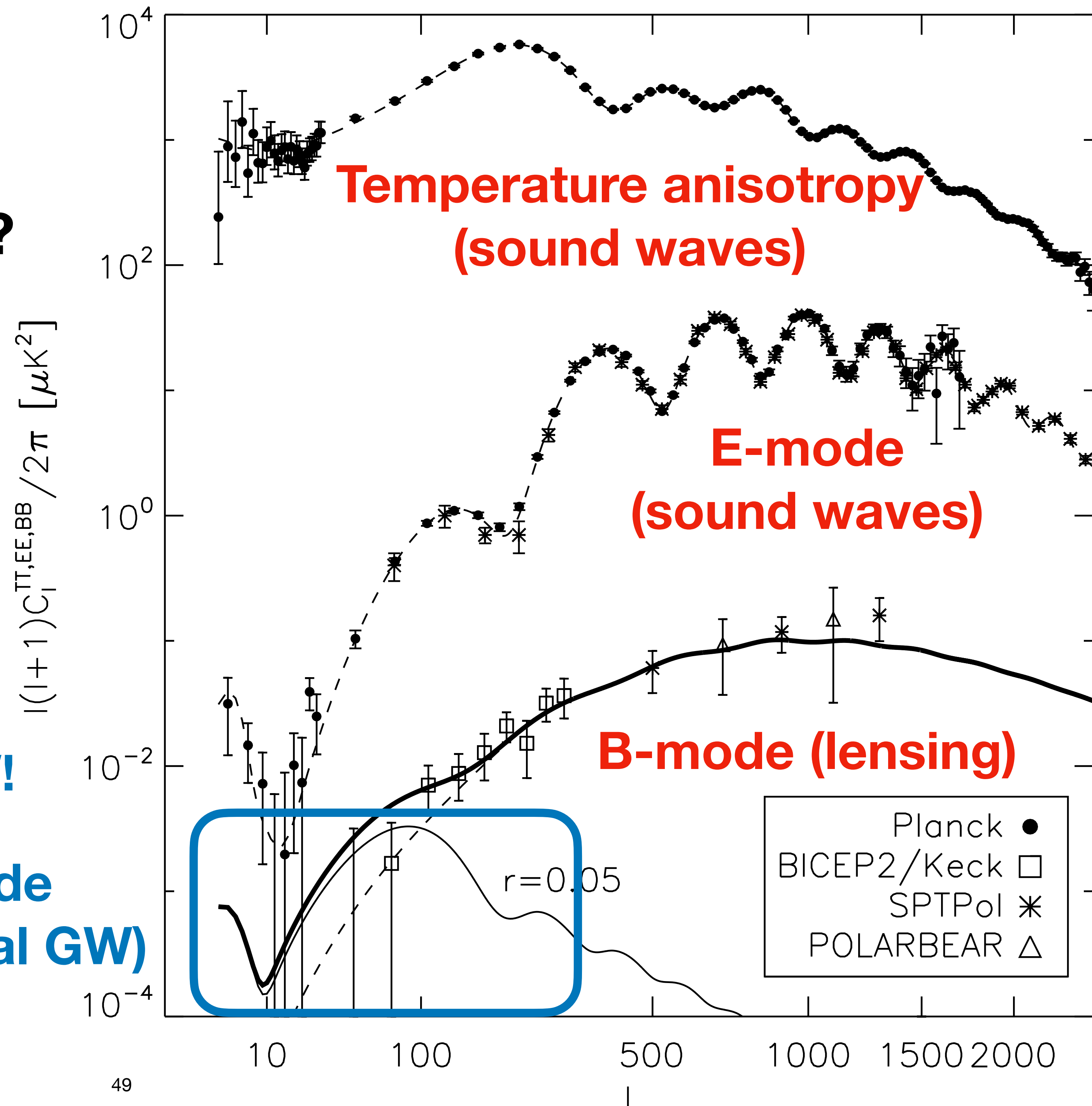


# Power Spectra

Where are we? What is next?

- The temperature and polarisation power spectra originating from **the scalar (density) fluctuation** have been measured.
- The next quest: **B-mode power spectrum from the primordial GW!**

**B-mode  
(Primordial GW)**





# Tensor-to-scalar Ratio

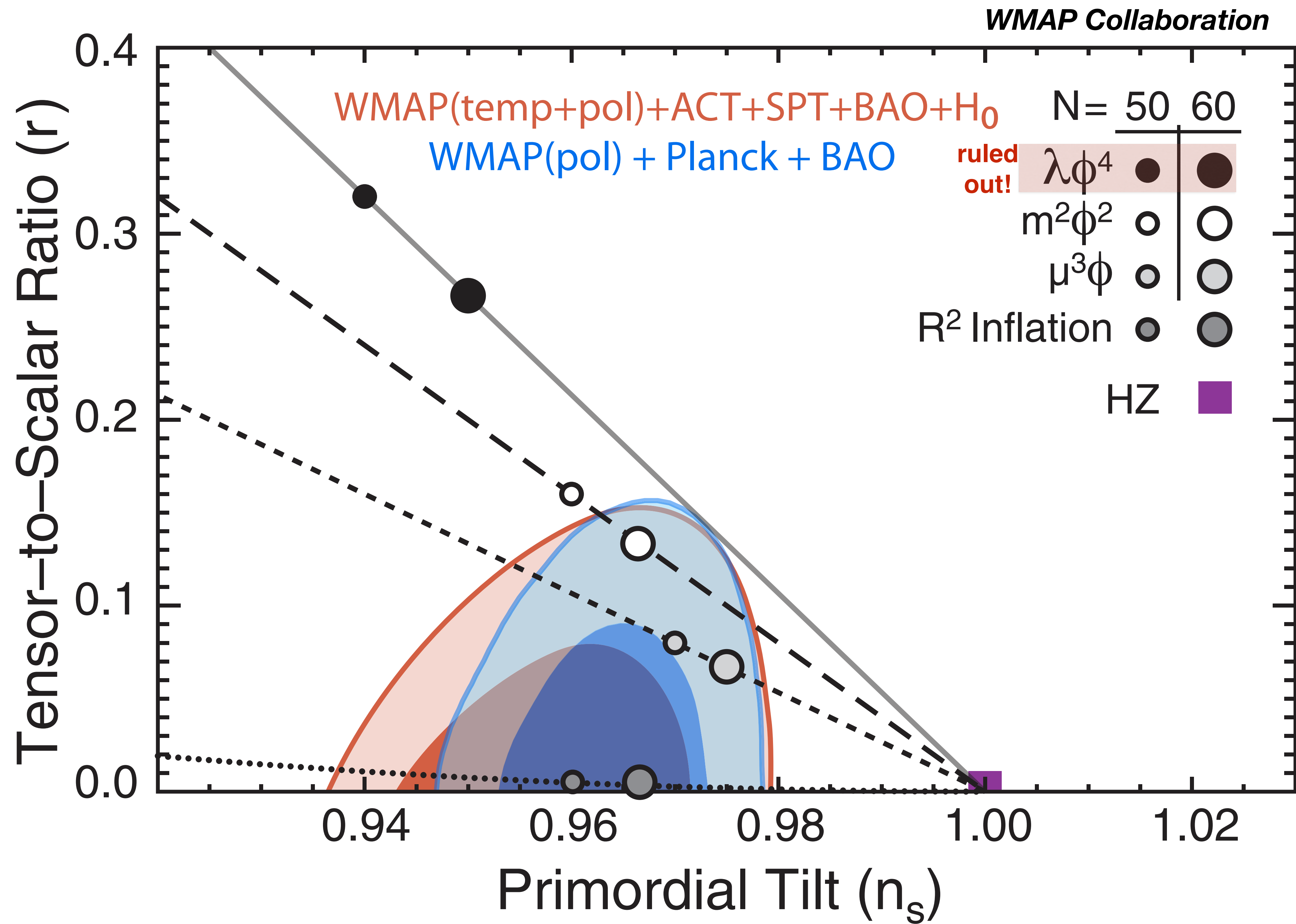
$$r \equiv \frac{\langle h_{ij} h^{ij} \rangle}{\langle \zeta^2 \rangle}$$

Scalar mode

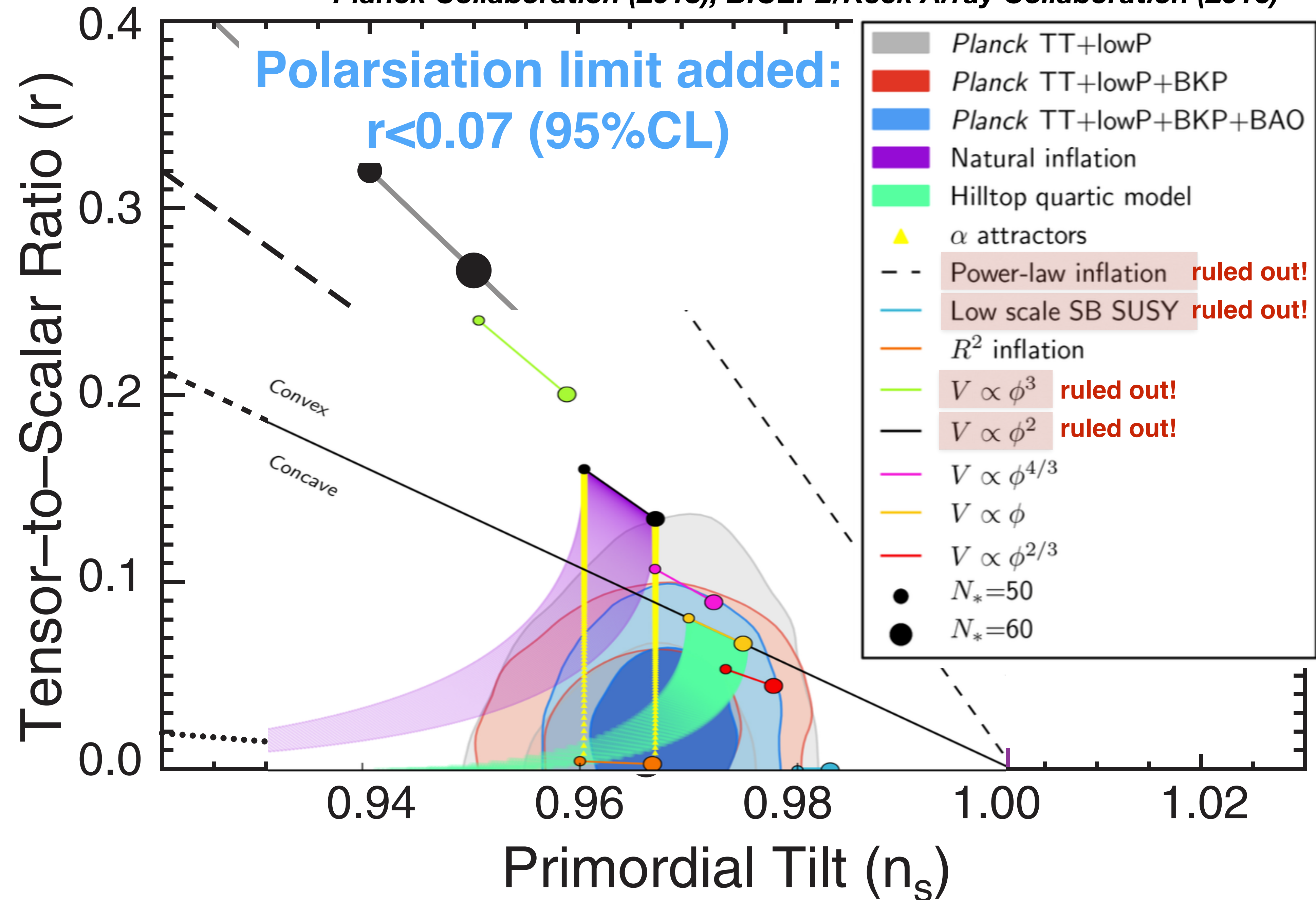
- We really want to find this! The current upper bound is  **$r < 0.06$**  (95%CL)

BICEP2/Keck Array Collaboration (2018)



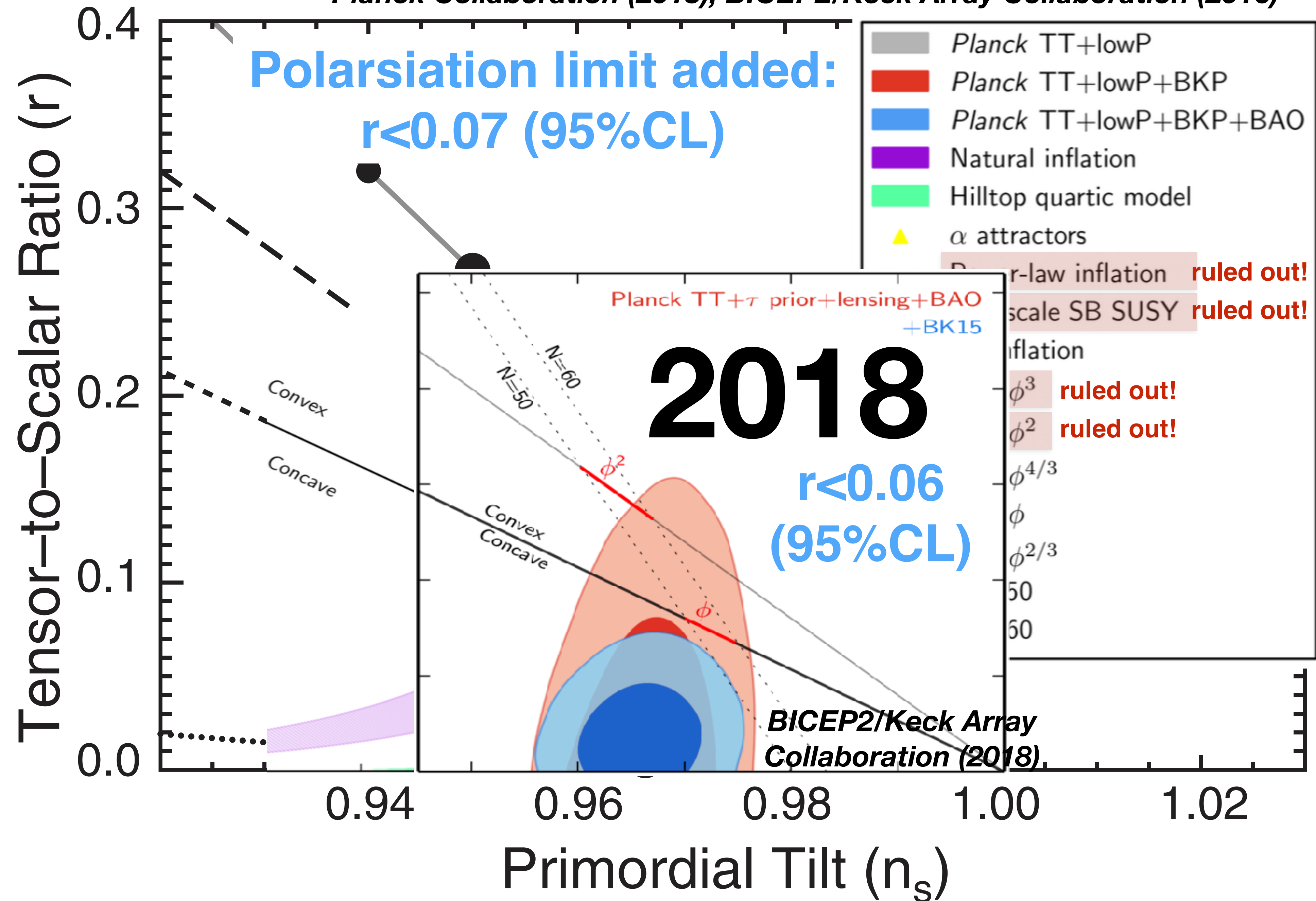








Planck Collaboration (2015); BICEP2/Keck Array Collaboration (2016)





# Experimental Landscape



# CMB Stages

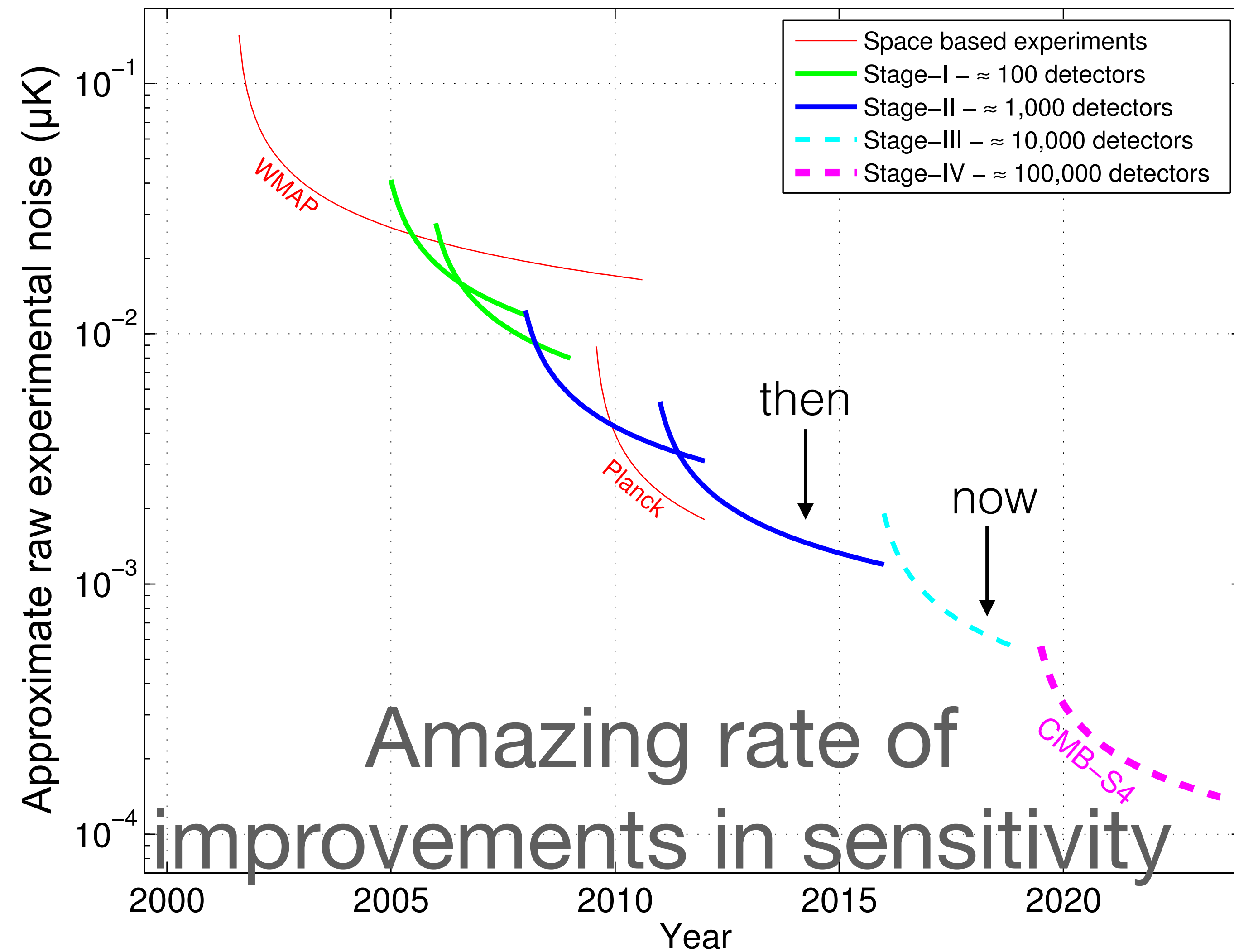


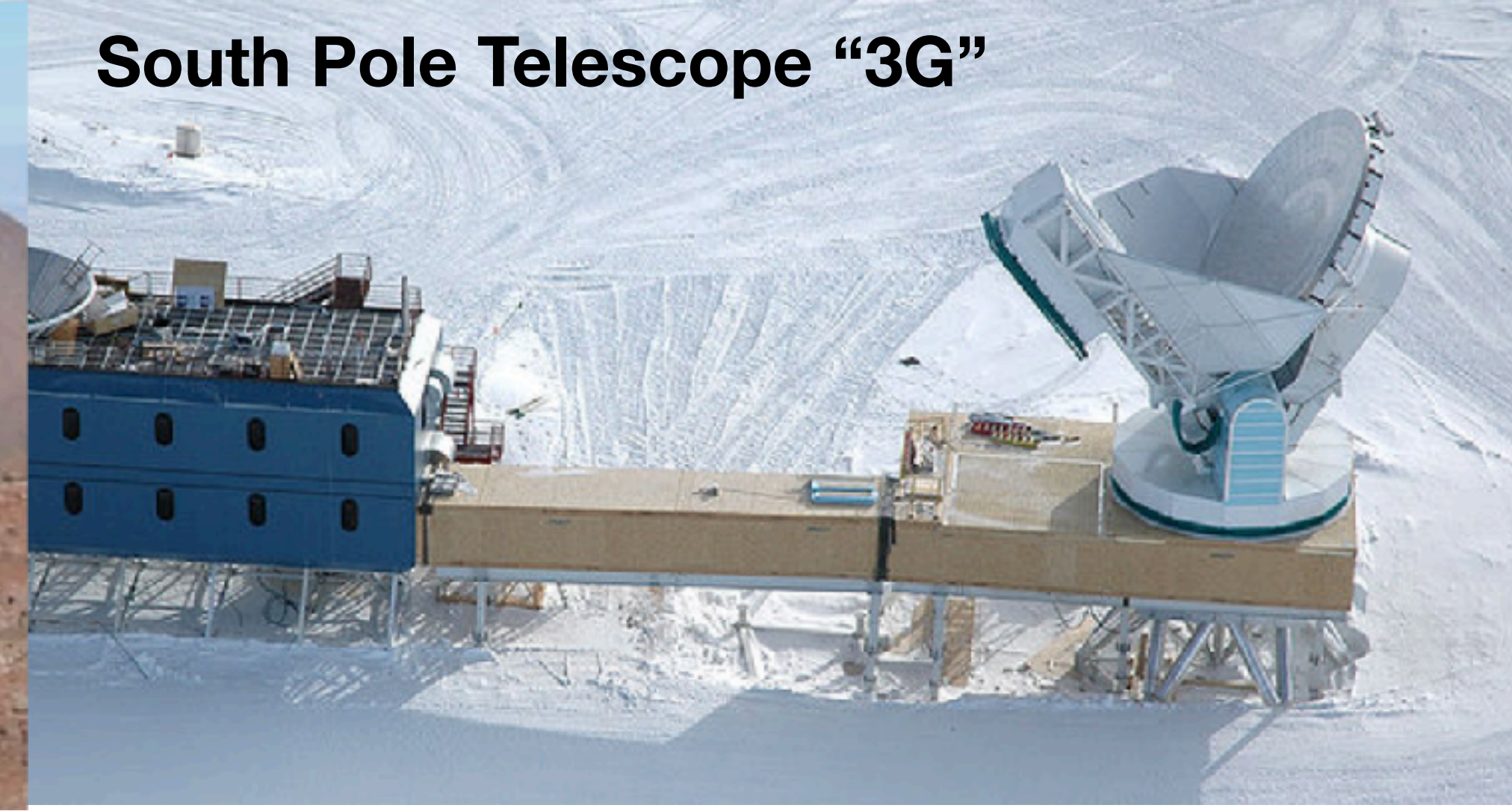
Figure by Clem Pryke for 2013 Snowmass documents



**Advanced Atacama  
Cosmology Telescope**

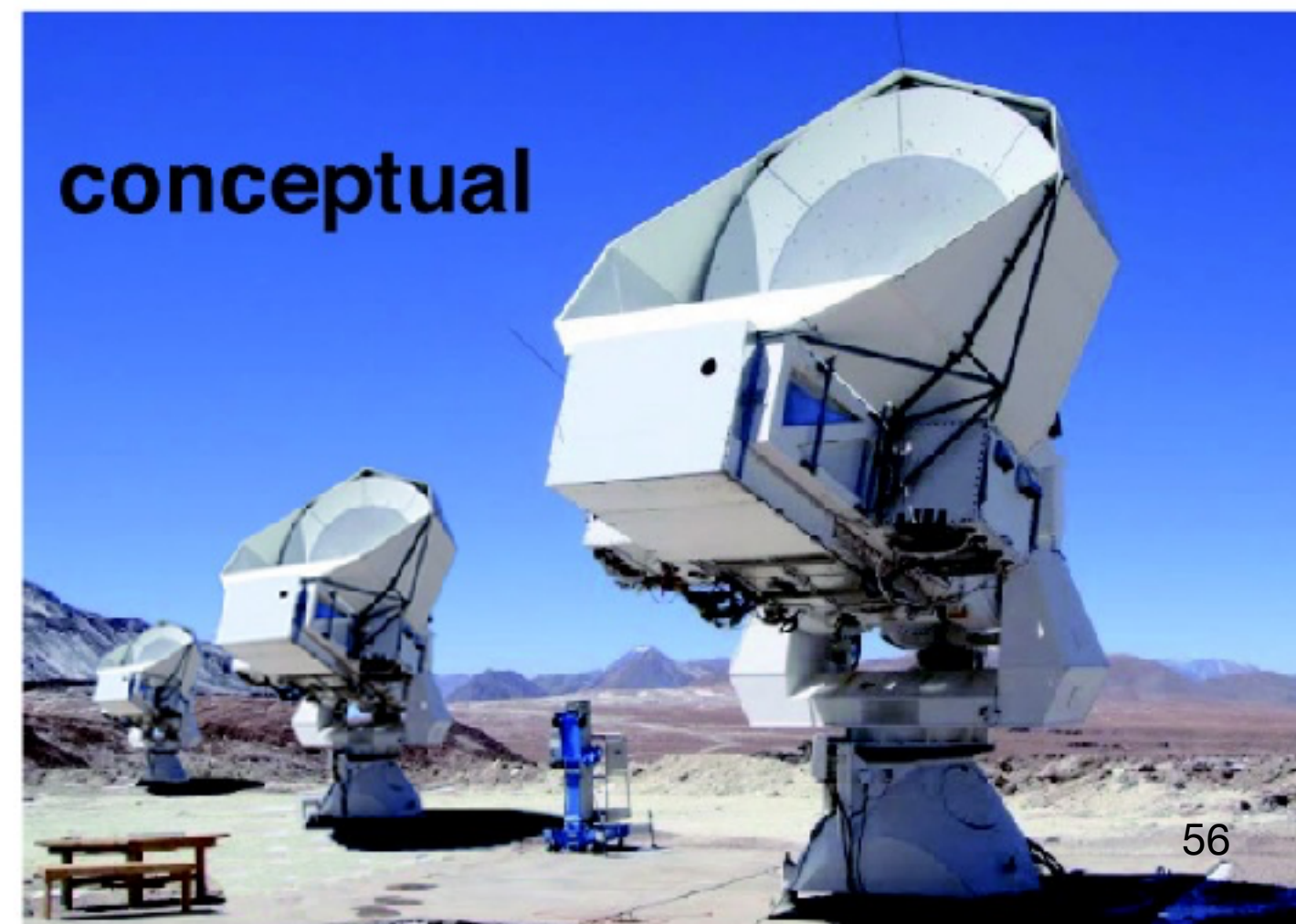


**South Pole Telescope “3G”**



# On-going Ground-based Experiments

**The Simons Array**



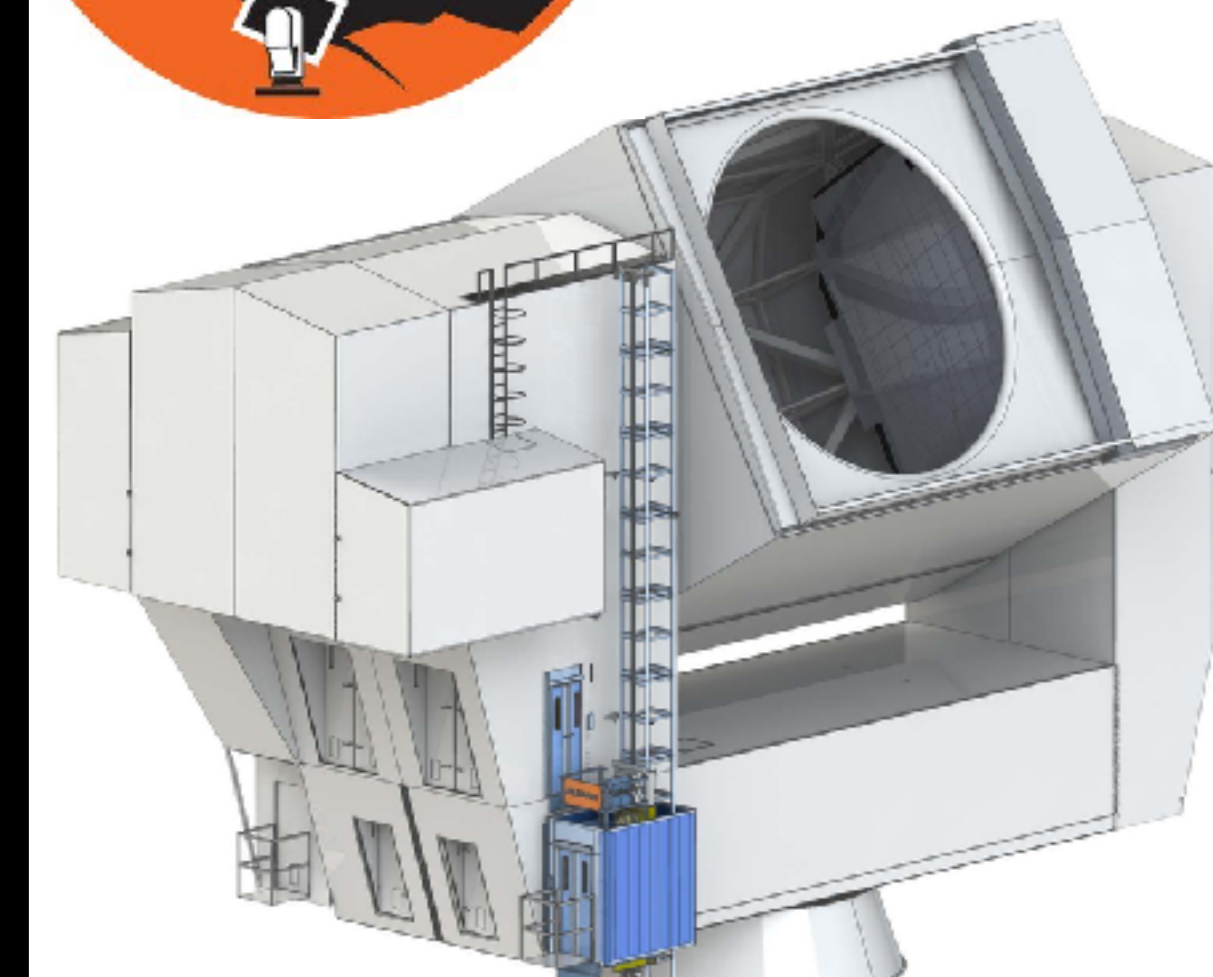
**BICEP/Keck Array**



**CLASS**





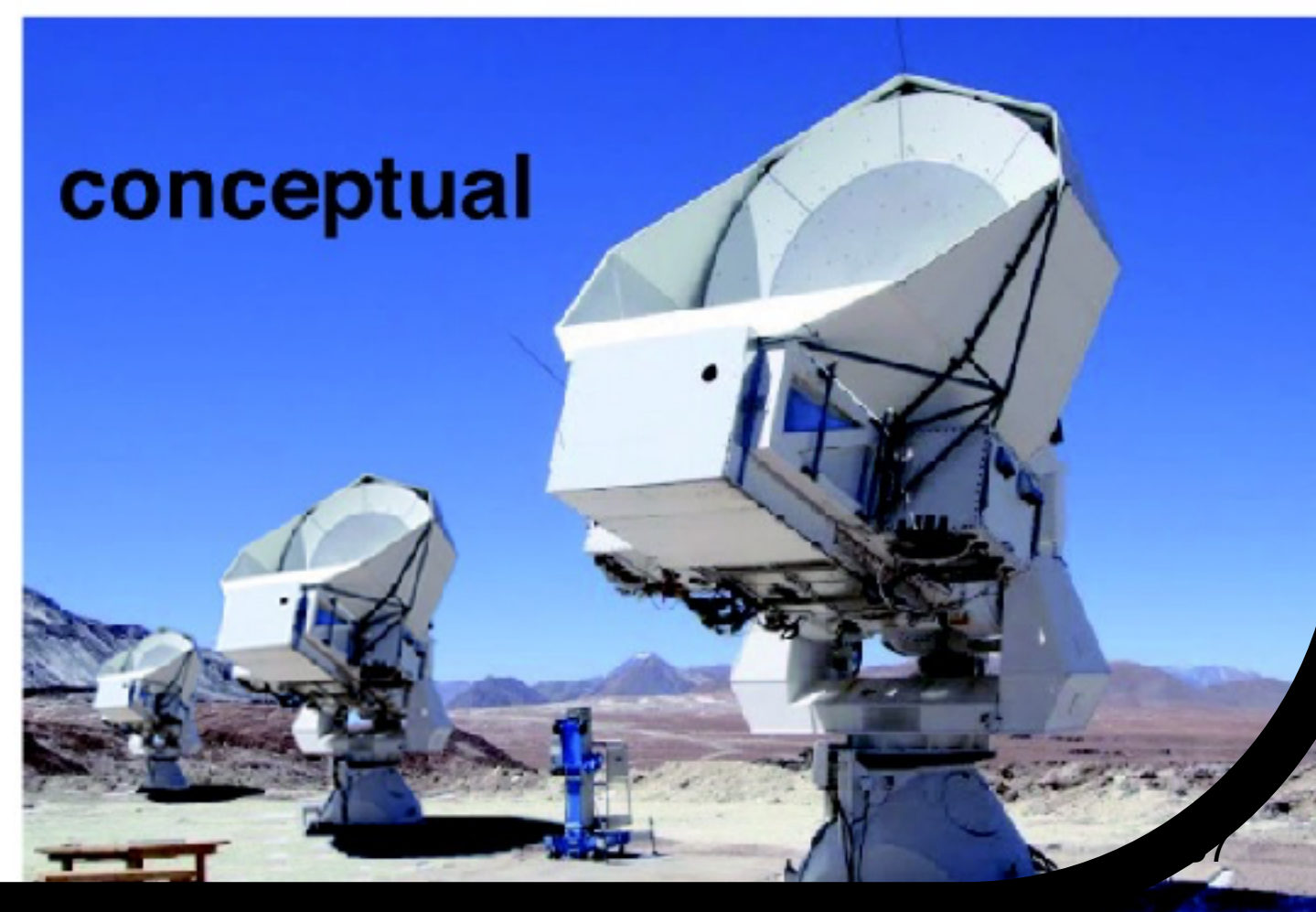


**Early 2020s**  
~\$100M

**Advanced Atacama  
Cosmology Telescope**



**The Simons Array**



**South Pole Telescope "3G"**



**The South Pole  
Observatory**

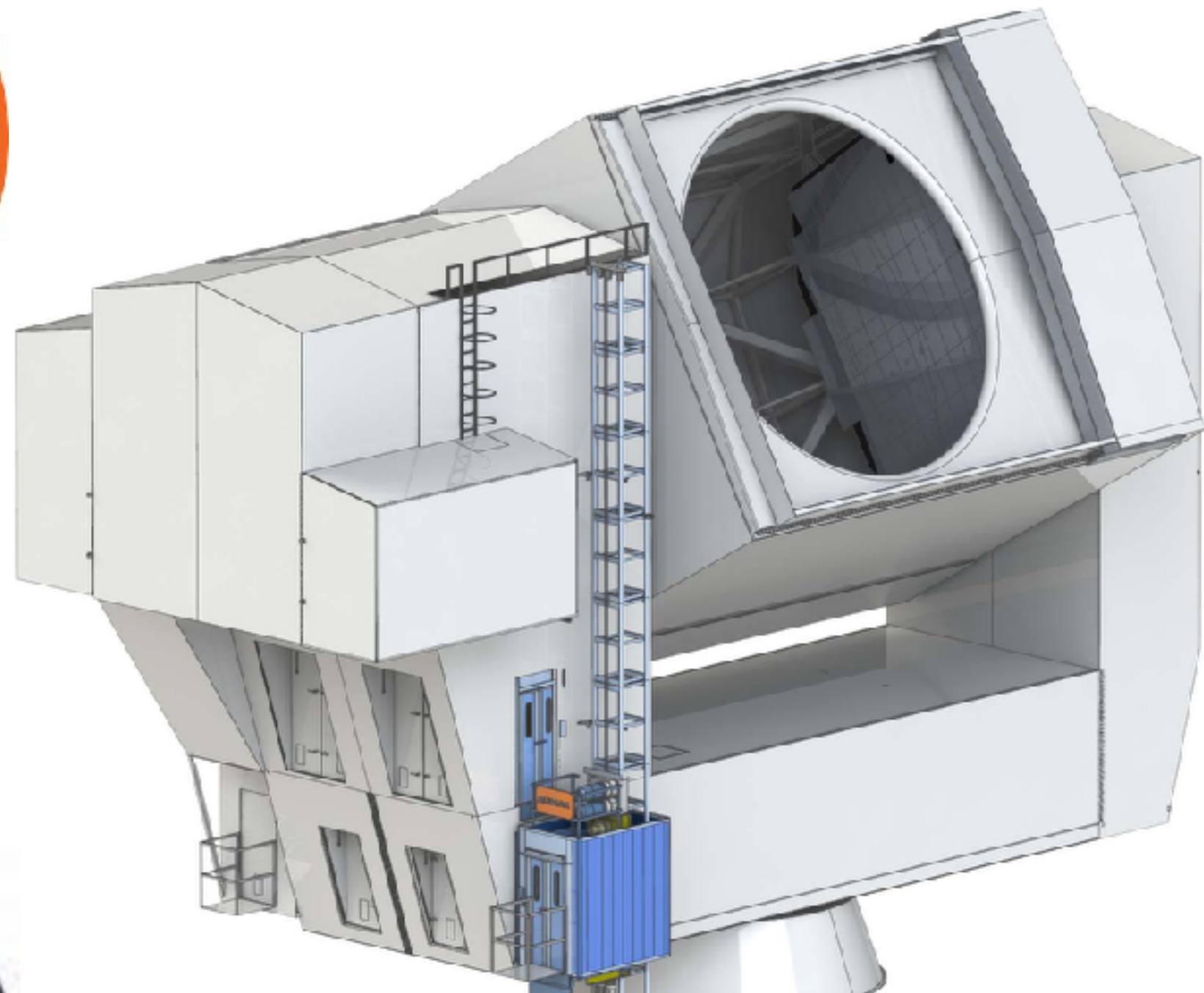
**BICEP/Keck Array**



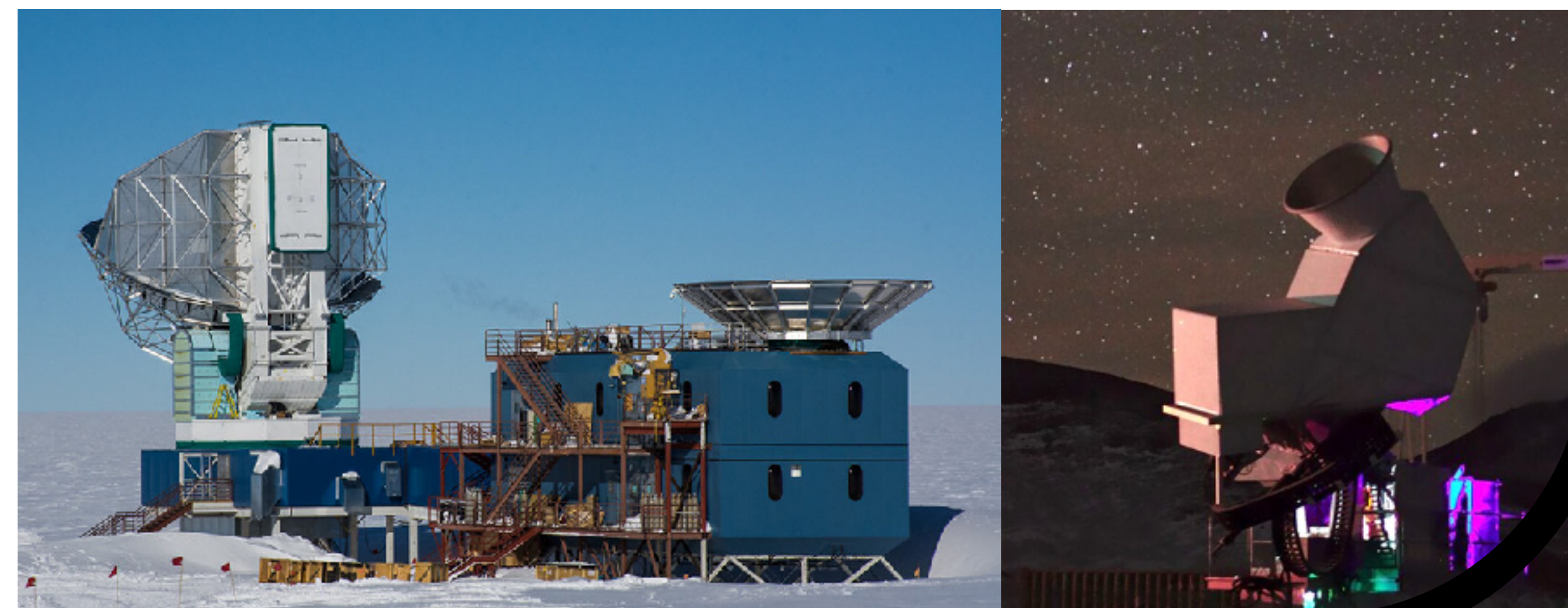
**CLASS**







Bringing all together:  
US-led CMB Stage IV  
Late 2020s (~\$600M)



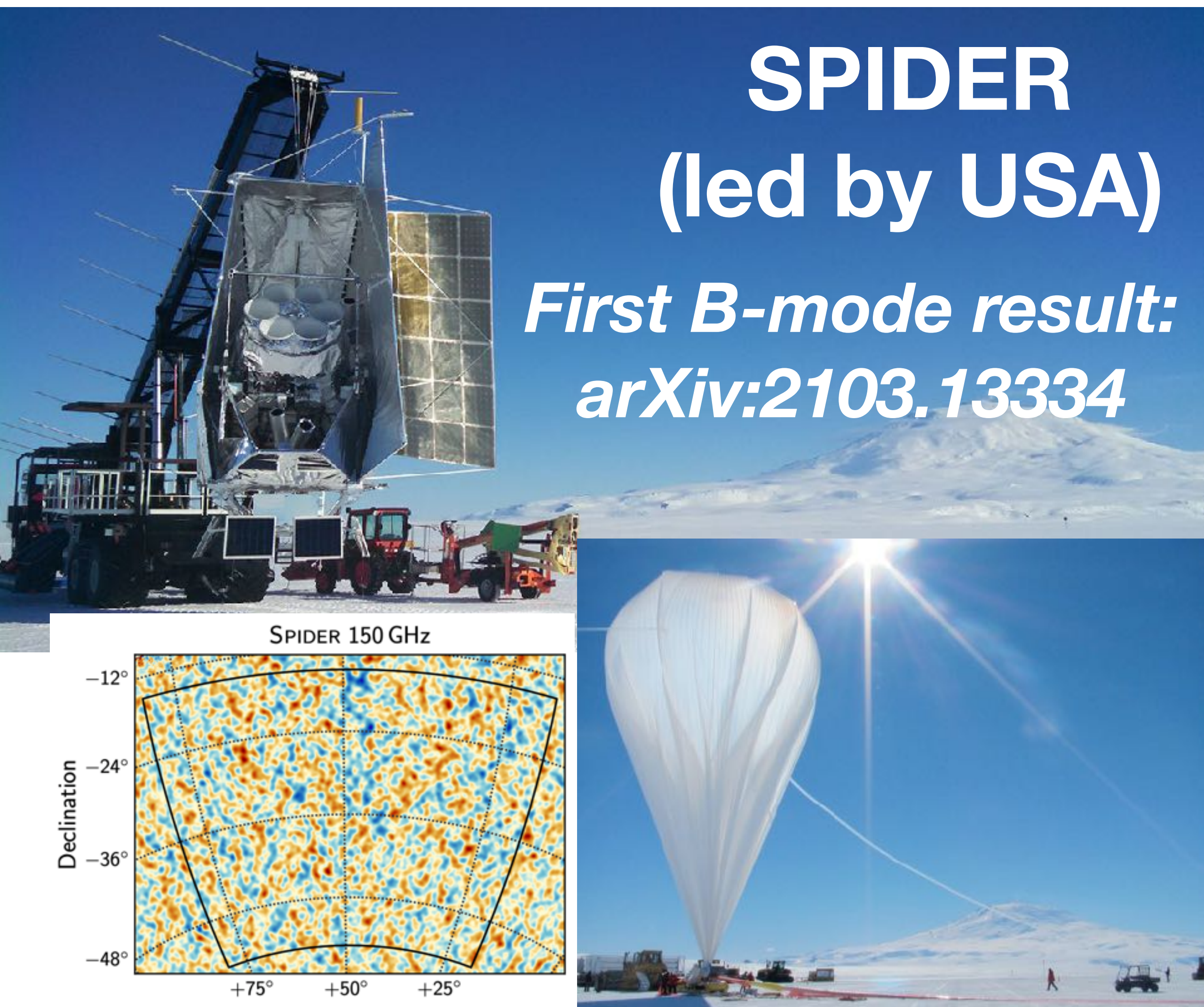


# Balloons!

“Almost space”

## SPIDER (led by USA)

*First B-mode result:  
arXiv:2103.13334*



## LSPE/SWIPE (led by Italy)





# 2029– LiteBIRD



**JAXA  
+ NASA  
+ CSA  
+ Europe**

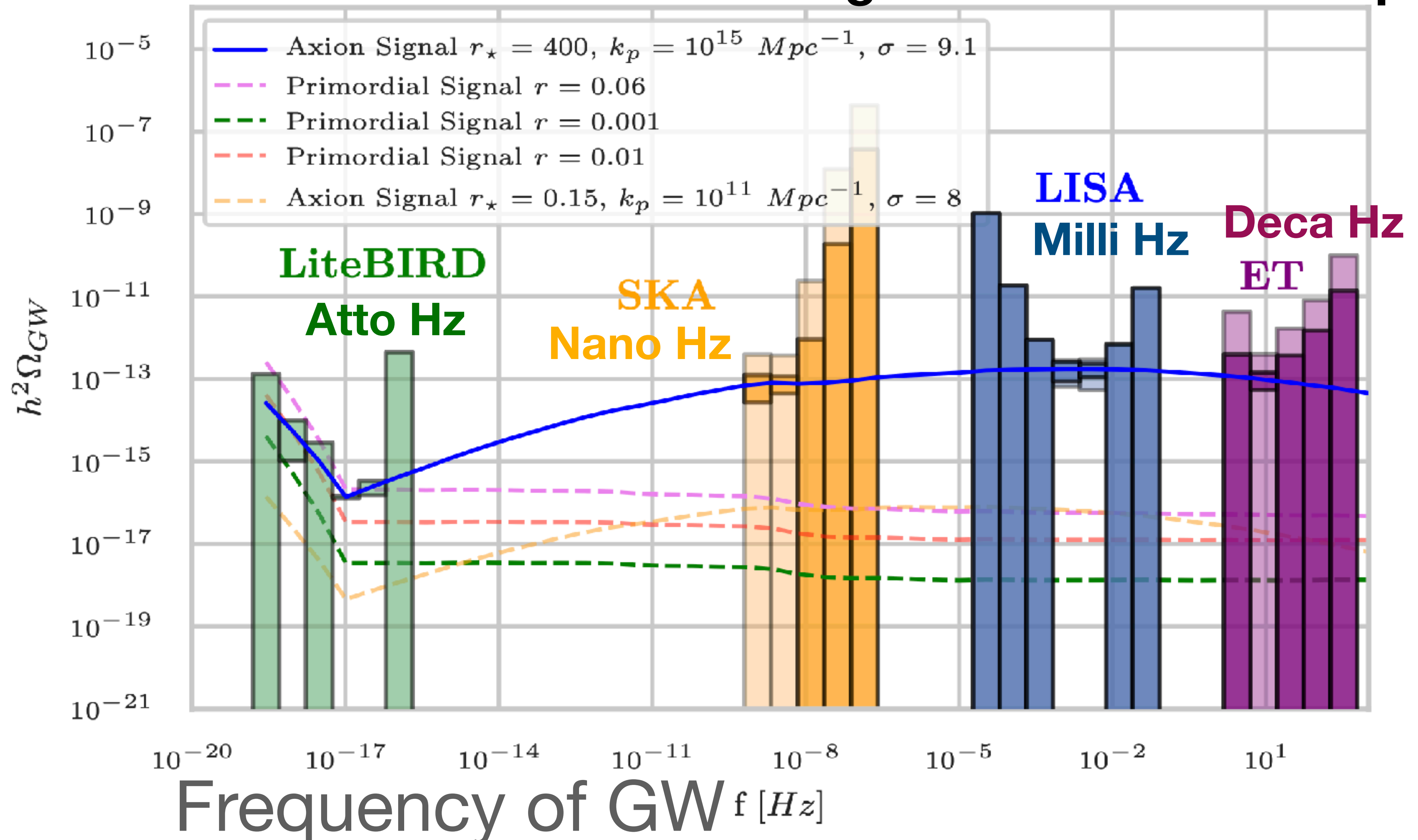
**A few thousand super-conducting  
microwave sensors in space.  
Selected by JAXA to fly to L2!**



# But let's recall again: not just CMB!

We can measure it across 21 orders of magnitude in the GW frequency

Energy Density of GW  
today





# Summary

## Towards finding our origins

- **The Quest So Far:**

- There is very good evidence that we all came from the quantum fluctuation in the early Universe, generated during the period of **cosmic inflation**.

- **The New Quest:**

- Discovery of the primordial gravitational wave with the wavelength of billions of light years gives **definitive evidence for inflation**.
- Hoping to find the first evidence from ground-based and balloon-borne experiments within the next 10 years.
- Then, the definitive measurement will come from **LiteBIRD** in early 2030s.