

ASKAP: planned are $36 \times 12\text{-m}$ dishes

Phase 1 = BETA: $6 \times 12\text{-m}$ dishes



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WALLABY – the ASKAP HI All-Sky Survey

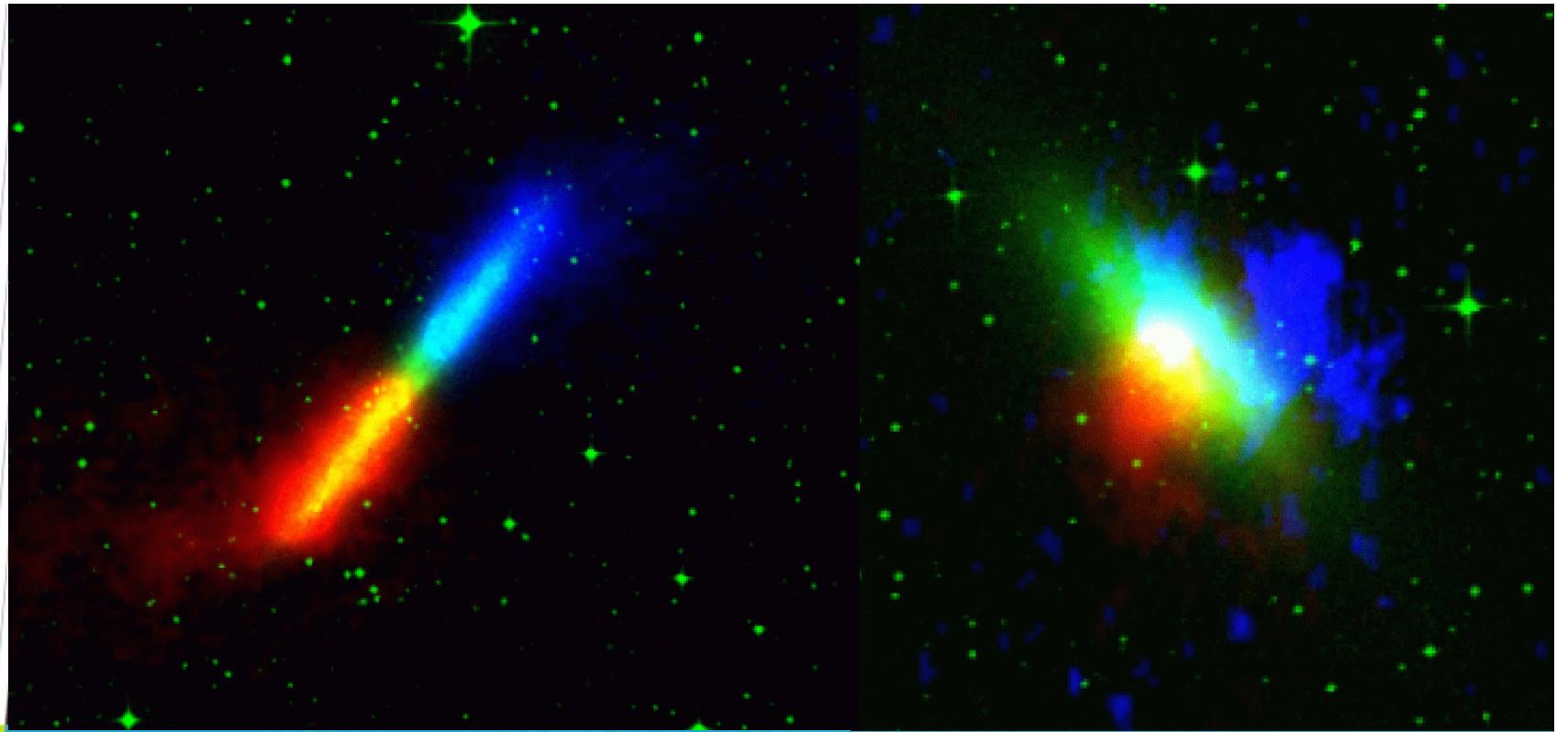
Bärbel Koribalski

CSIRO Astronomy and Space Science

Australia Telescope National Facility

“Gas in Galaxies” – Kloster Seeon – June 2011



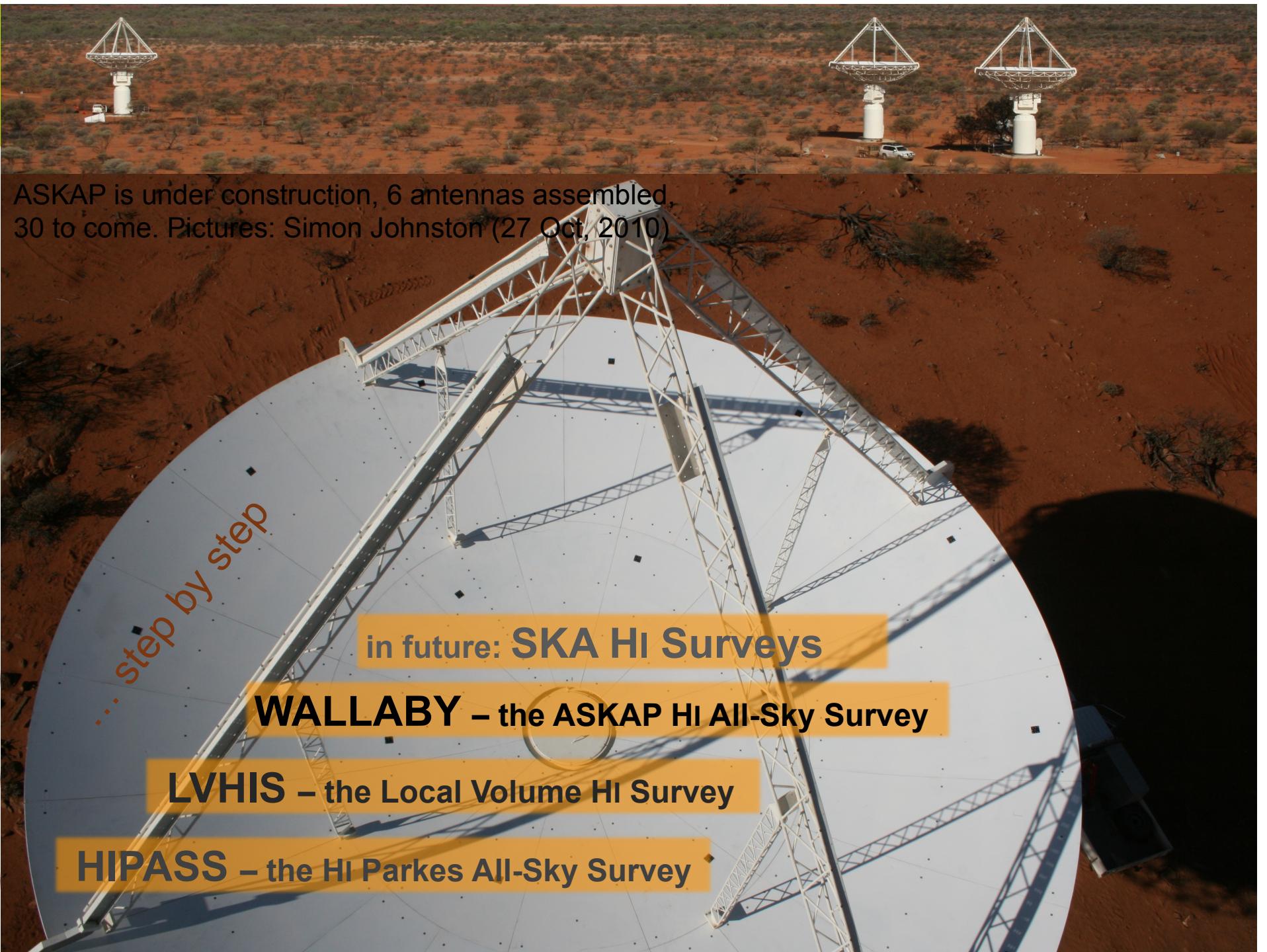


www.csiro.au

[www.atnf.csiro.au / research / WALLABY](http://www.atnf.csiro.au/research/WALLABY)

Project leaders:
Bärbel Koribalski &
Lister Staveley-Smith





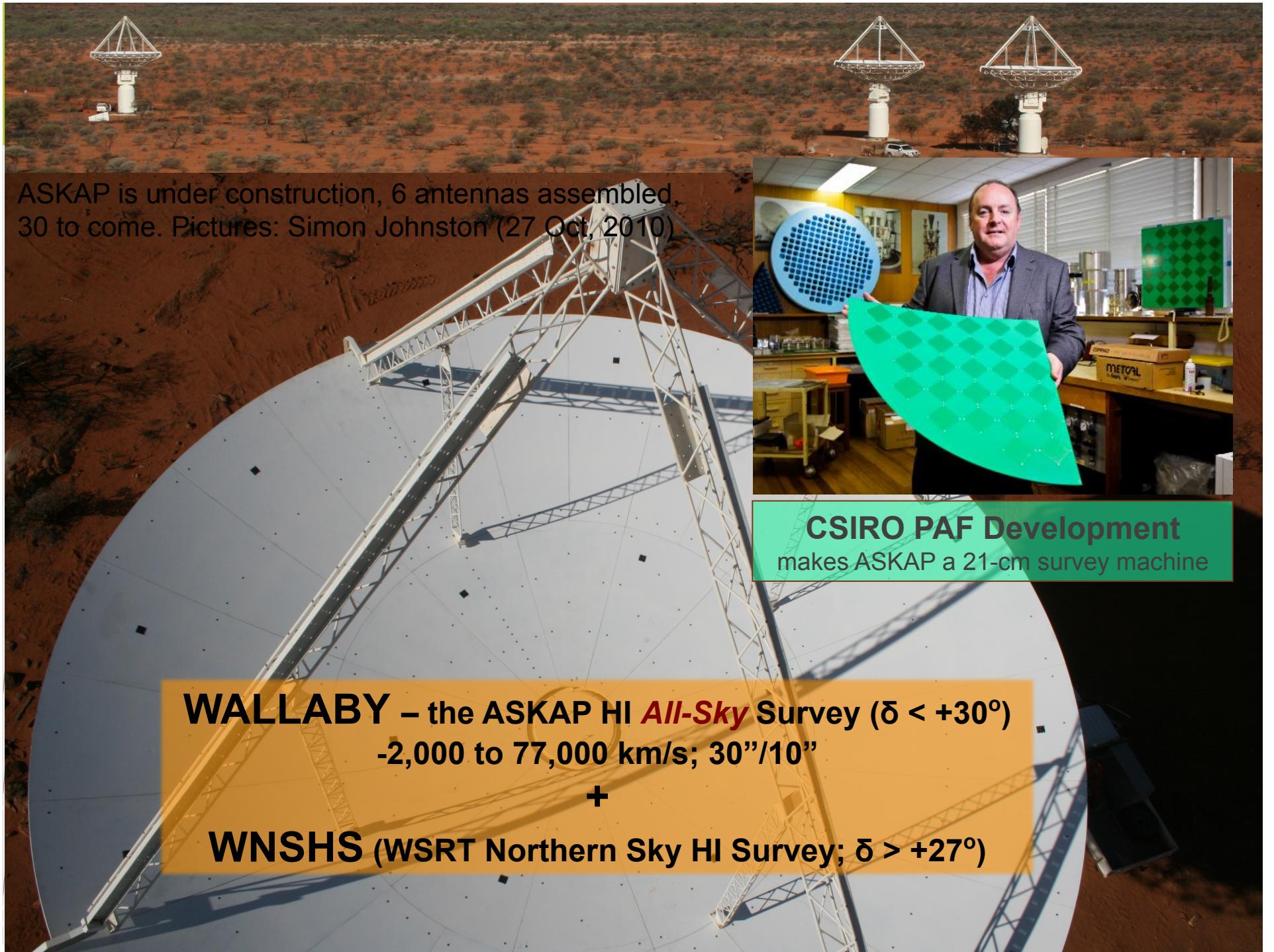
ASKAP is under construction, 6 antennas assembled,
30 to come. Pictures: Simon Johnston (27 Oct, 2010)

in future: **SKA HI Surveys**

WALLABY – the ASKAP HI All-Sky Survey

LVHIS – the Local Volume HI Survey

HIPASS – the HI Parkes All-Sky Survey



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CSIRO PAF Development
makes ASKAP a 21-cm survey machine

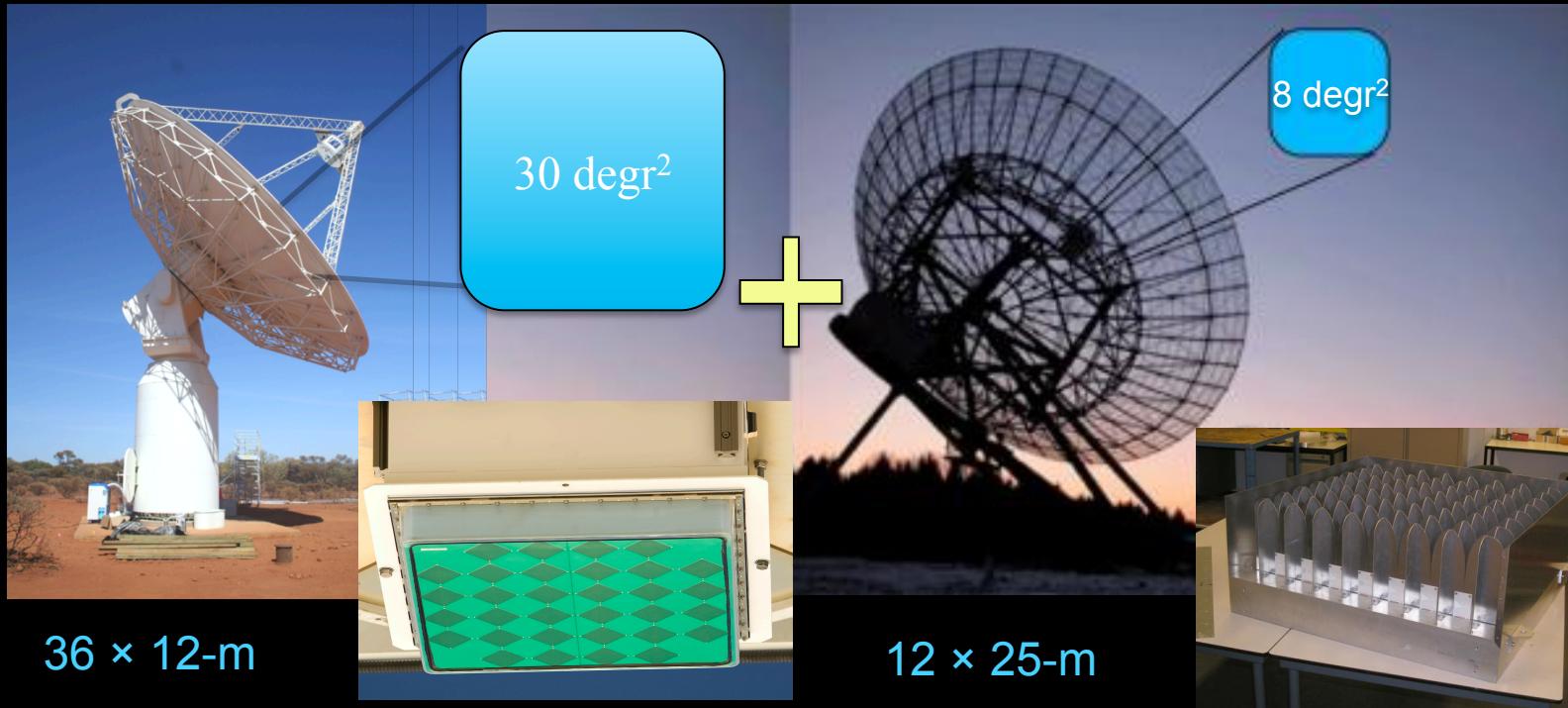
WALLABY – the ASKAP HI *All-Sky* Survey ($\delta < +30^\circ$)
-2,000 to 77,000 km/s; $30''/10''$

+

WNSHS (WSRT Northern Sky HI Survey; $\delta > +27^\circ$)

HI (21-cm) All-Sky Survey

ASKAP + WSRT with Apertif



To achieve all-sky coverage we require \approx 1200 ASKAP pointings (left), integrate 8h each ($\delta < 30$ deg r) + \approx 1300 WSRT/Apertif pointings (right), 4h each ($\delta > 30$ deg r) to achieve the same sensitivity and resolution.

The ASKAP PAF – a new radio camera



April 2011: Front view & rear view of the ASKAP PAF.

- ◆ **PAF = Phased Array Feeds (checkerboard array: 188 elements)**
- ◆ **Beamformer: creates up to 36 beams, each 1.2 degr FWHM**
- ◆ **resulting field of view is 30 square degrees (5.5 deg × 5.5 deg)**

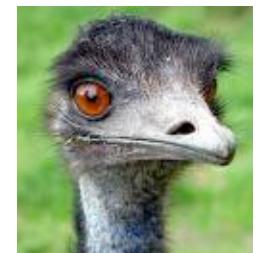
ASKAP SSPs



- WALLABY (Koribalski/Staveley-Smith) All sky $\text{H}\alpha$ survey to $z \sim 0.2$
- EMU (Norris) All sky continuum to 10 μJy rms

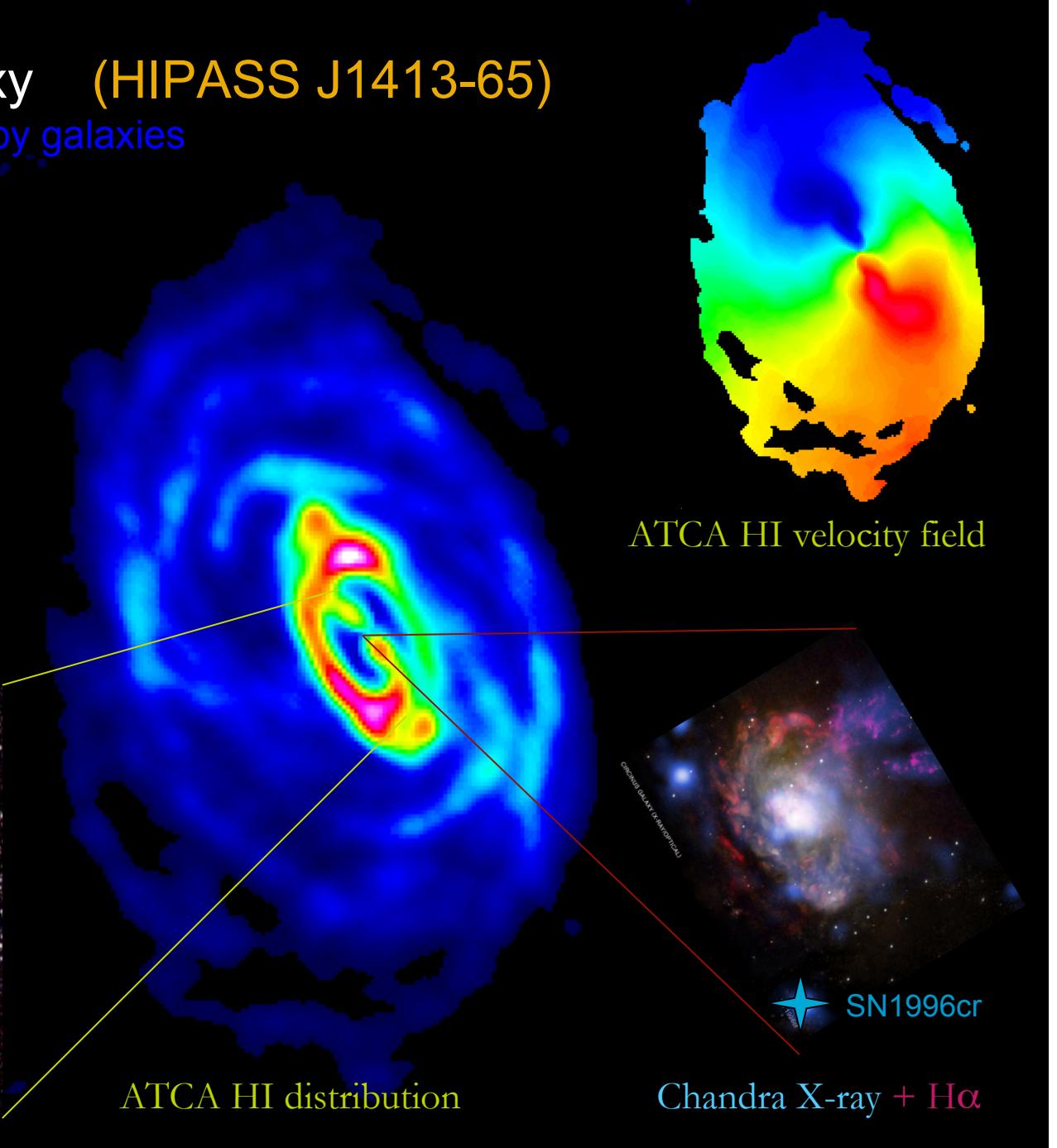
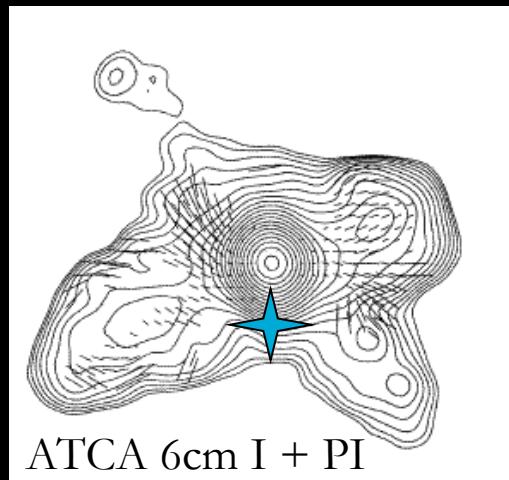
- POSSUM (Gaensler/Landecker/Taylor) Polarization / RM grid
- FLASH (Sadler) $\text{H}\alpha$ absorption to $z \sim 1$
- VAST (Murphy/Chatterjee) Transients and variables (>5 sec)
- CRAFT (Dodson/Macquart) Fast transients (<5 sec)
- GASKAP (Dickey) Galactic and Magellanic $\text{H}\alpha$ and OH
- DINGO (Meyer) Deep $\text{H}\alpha$ emission survey

- COAST (Stairs) Pulsar timing and searching
- VLBI (Tingay) ASKAP as part of the LBA



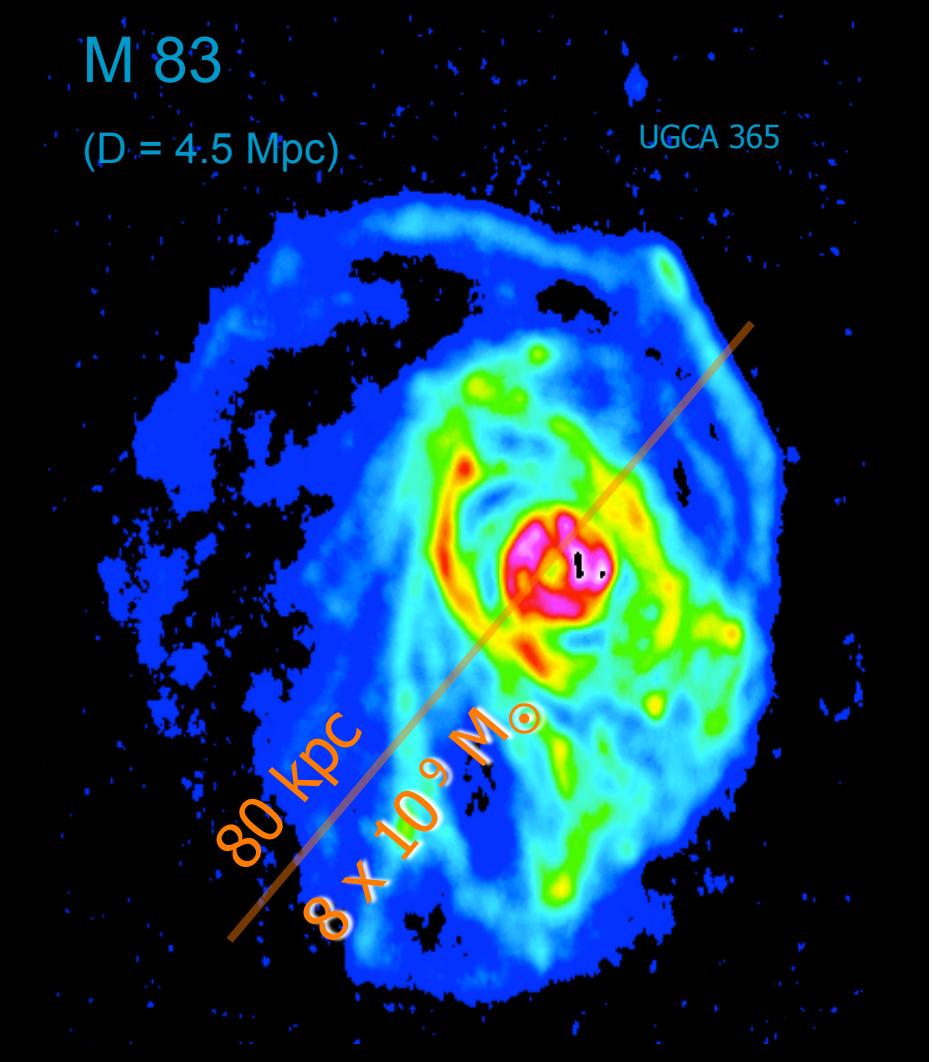
The Circinus Galaxy (HIPASS J1413-65)

= one of the largest nearby galaxies



WALLABY – sensitivity & scales

- e.g., 8h integration time
 - ➔ 5σ M_{HI} limit
 - = $5 \times 10^6 M_{\odot}$ ($D = 10 \text{ Mpc}$)
 - = $5 \times 10^8 M_{\odot}$ ($D = 100 \text{ Mpc}$)
 - = $3 \times 10^{10} M_{\odot}$ ($D = 800 \text{ Mpc}$)
- 10" beam
 - = 0.5 kpc ($D = 10 \text{ Mpc}$)
 - = 5 kpc ($D = 100 \text{ Mpc}$)
 - = 39 kpc ($D = 800 \text{ Mpc}$)
- 30" beam
 - = 1.5 kpc ($D = 10 \text{ Mpc}$)
 - = 15 kpc ($D = 100 \text{ Mpc}$)
 - = 116 kpc ($D = 800 \text{ Mpc}$)



WALLABY will explore

- the *Gaseous* Universe
 - the *Dynamic* Universe
 - the *3D* Universe
 - the *Dark* Universe
- and
- the *Unknown* Universe

... probing the Gaseous Universe

WALLABY will discover

- new dwarf galaxies in the *Local Group*
- hundreds of dwarfs in the *Local Volume* ($5 \times 10^6 M_{\odot}$ out to 10 Mpc)
- diffuse HI clouds, tails and filaments connecting galaxies
- $10^8 M_{\odot}$ out to 60 Mpc

M 81 dwarf A

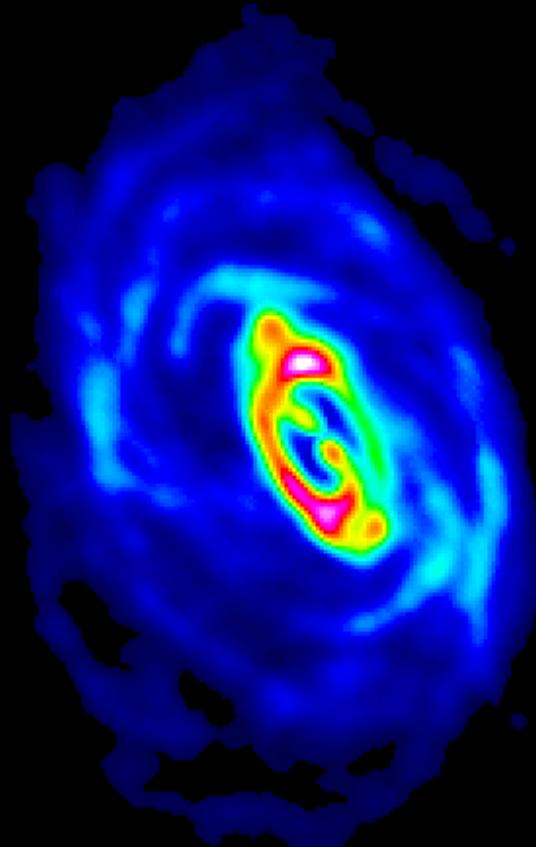
Holmberg I

Holmberg II

IC 2574

VLA HI + Spitzer MIR + Galex UV; Walter et al. (2009)

... probing the Gaseous Universe

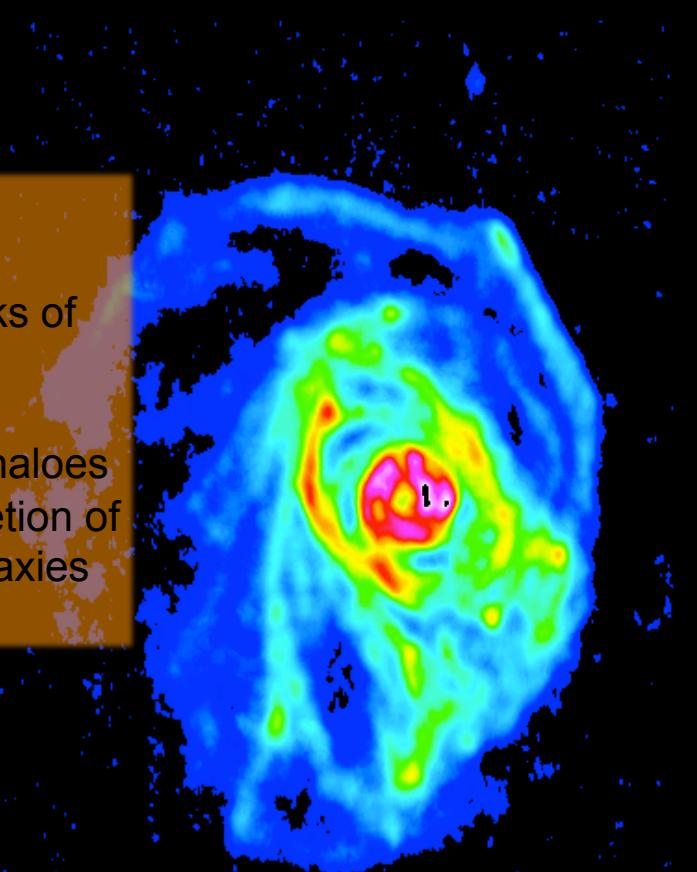


The Circinus
Galaxy

M 83 + UGC 365

WALLABY will reveal

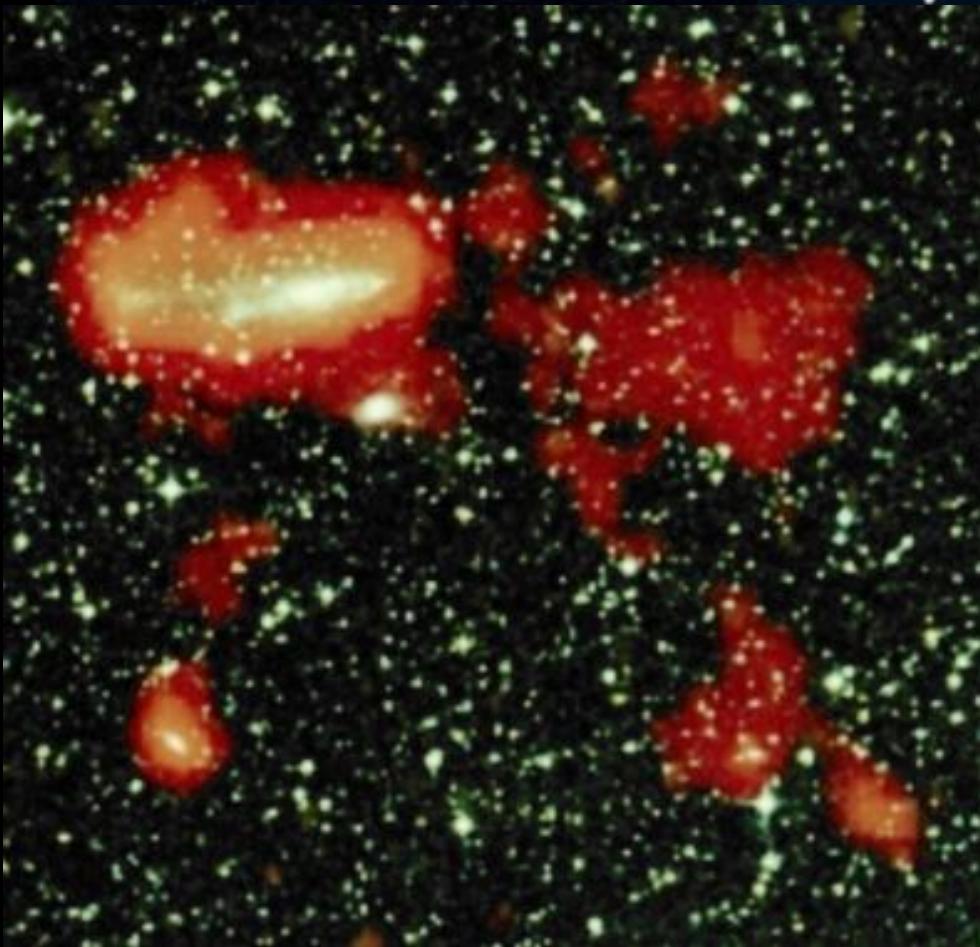
- the large-scale HI disks of spiral galaxies
- HI streams in galaxy haloes resulting from the accretion of neighbouring dwarf galaxies



BK et al. 2011, in prep.

... probing the Gaseous Universe

NGC 3263 group – English, BK et al. (2010)



WALLABY will discover

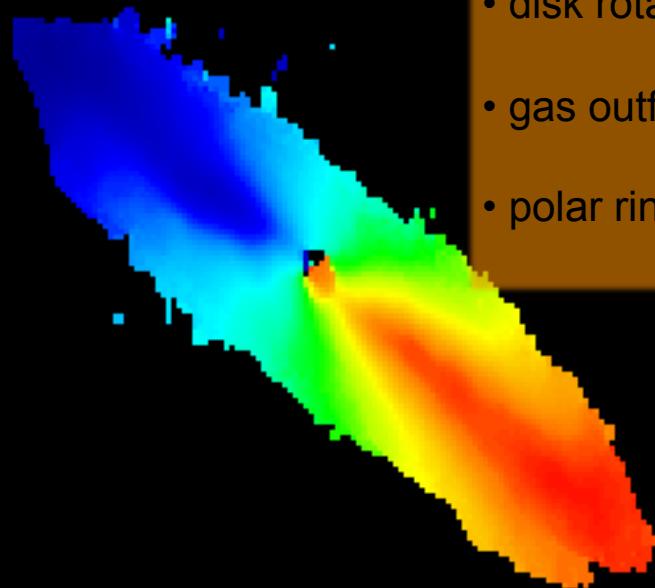
- tidal tails and HI plumes as the signatures of interacting galaxies
- distant high-density HI clouds and filaments, most likely in groups & clusters

IC 2554 – BK, Gordon & Jones (2003)

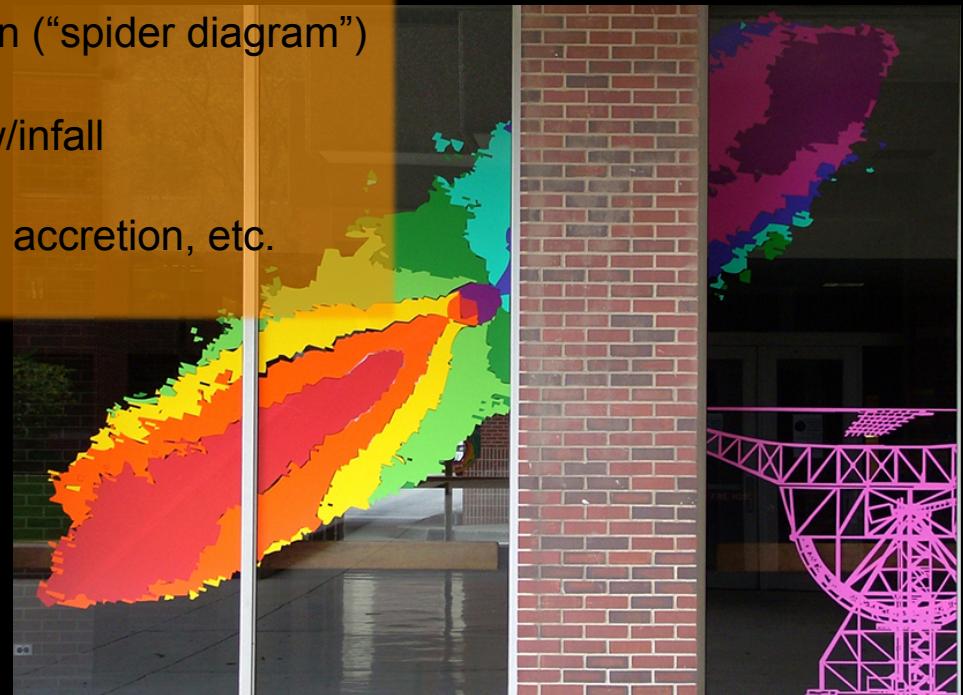
... exploring the Dynamic Universe

WALLABY will reveal

- the large-scale HI velocity fields of spiral galaxies
- disk rotation (“spider diagram”)
- gas outflow/infall
- polar rings, accretion, etc.



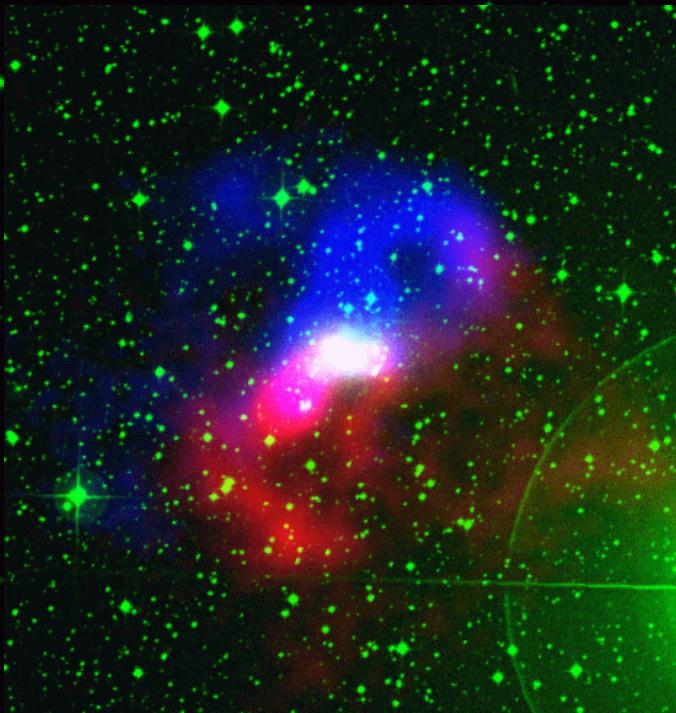
The starburst
galaxy NGC 253



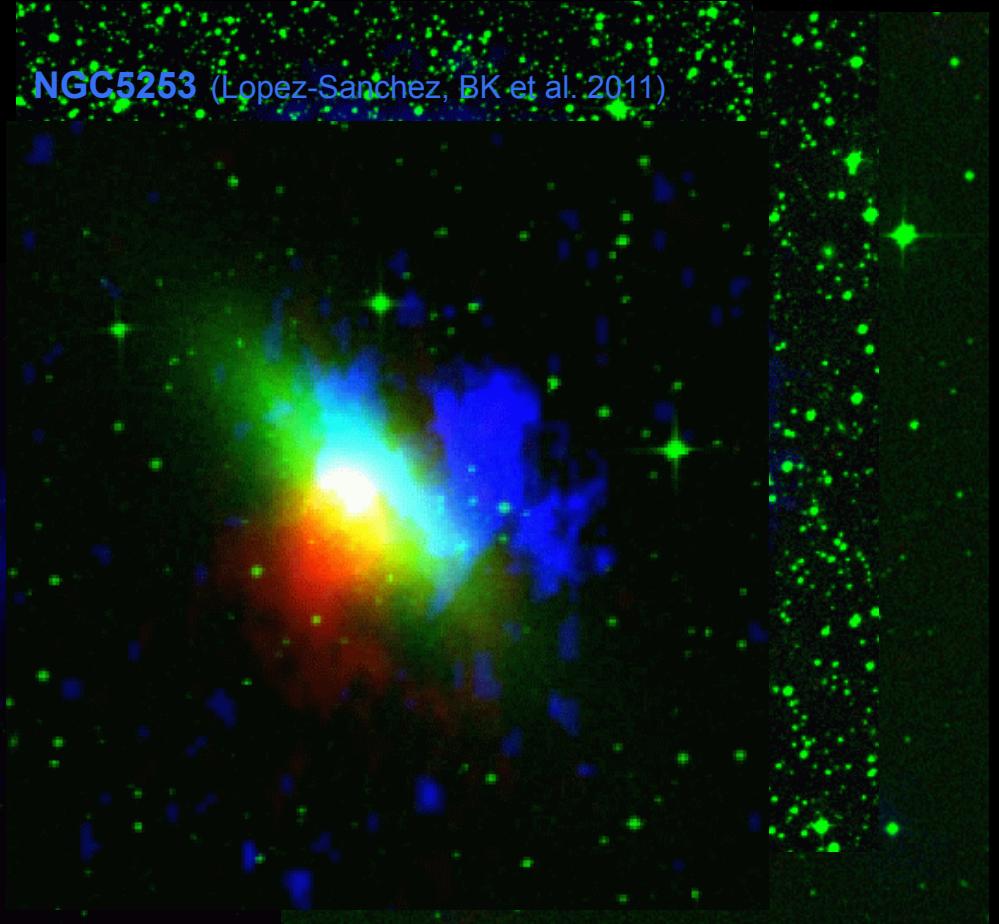
Dennison Mural @ U Michigan, IYA2009:
The Universe – Yours to Discover

... exploring the Dynamic Universe

IC 4662 (van Eymeren, BK et al. 2009)

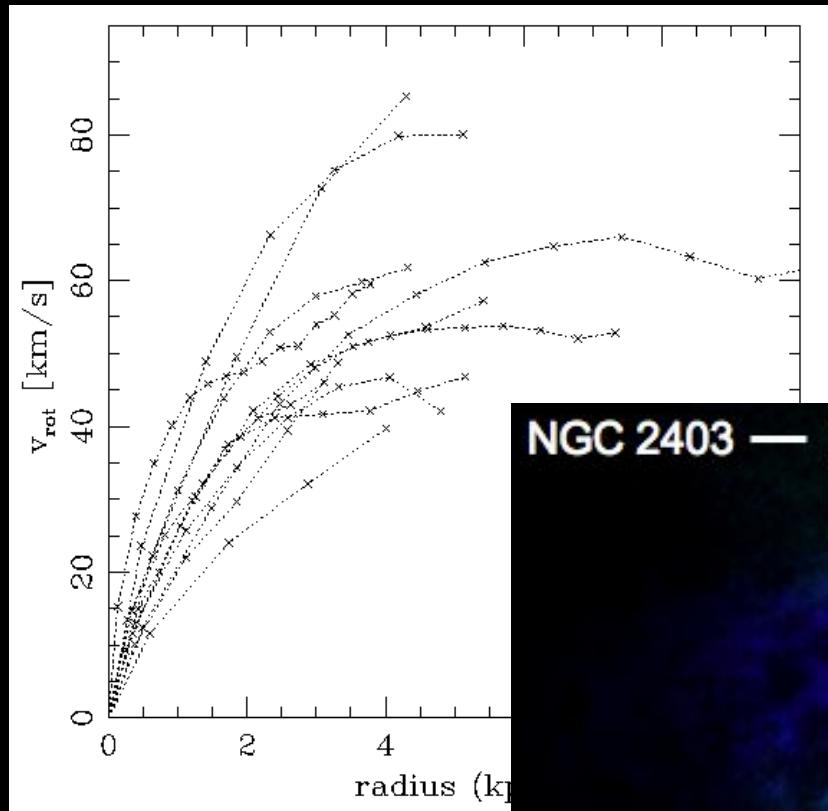


NGC5253 (Lopez-Sanchez, BK et al. 2011)



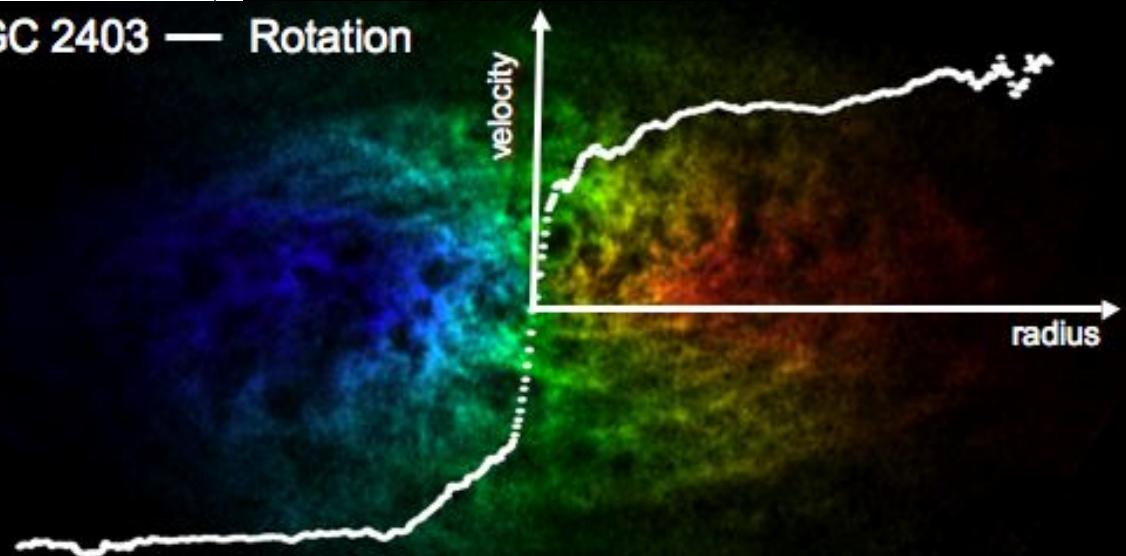
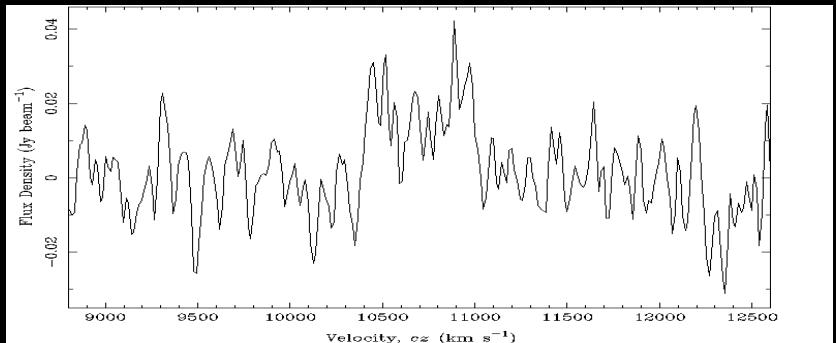
LVHIS galaxies (gas + stars)

... uncover **the Dark Universe**



NGC 2403 — Rotation

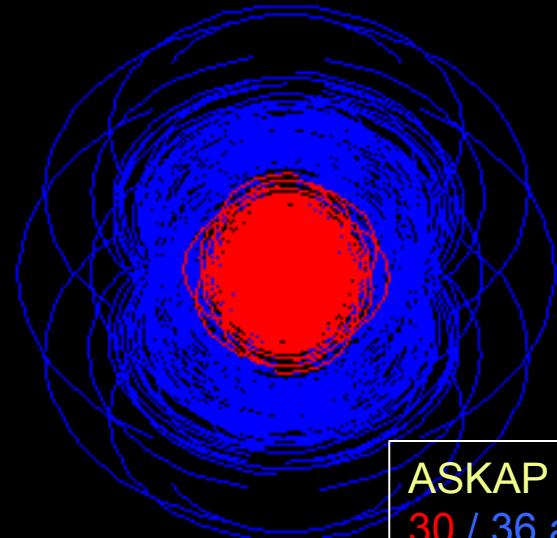
LVHIS: Kirby, BK et al. (2010)



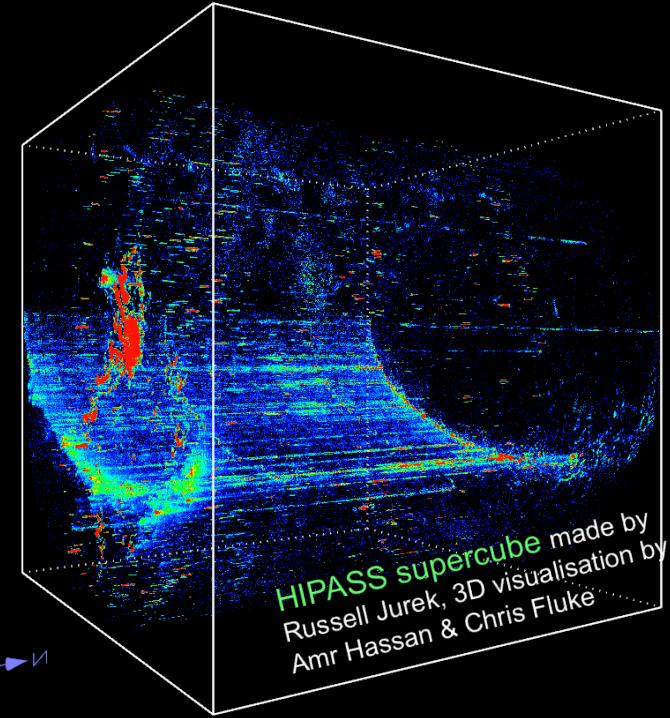
$$M_{\text{dyn}} = 2.31 \times 10^5 R_{\text{kpc}} v_{\text{rot}}^2$$

THINGS: Walter et al. (2009)

... exploring the 3D Universe

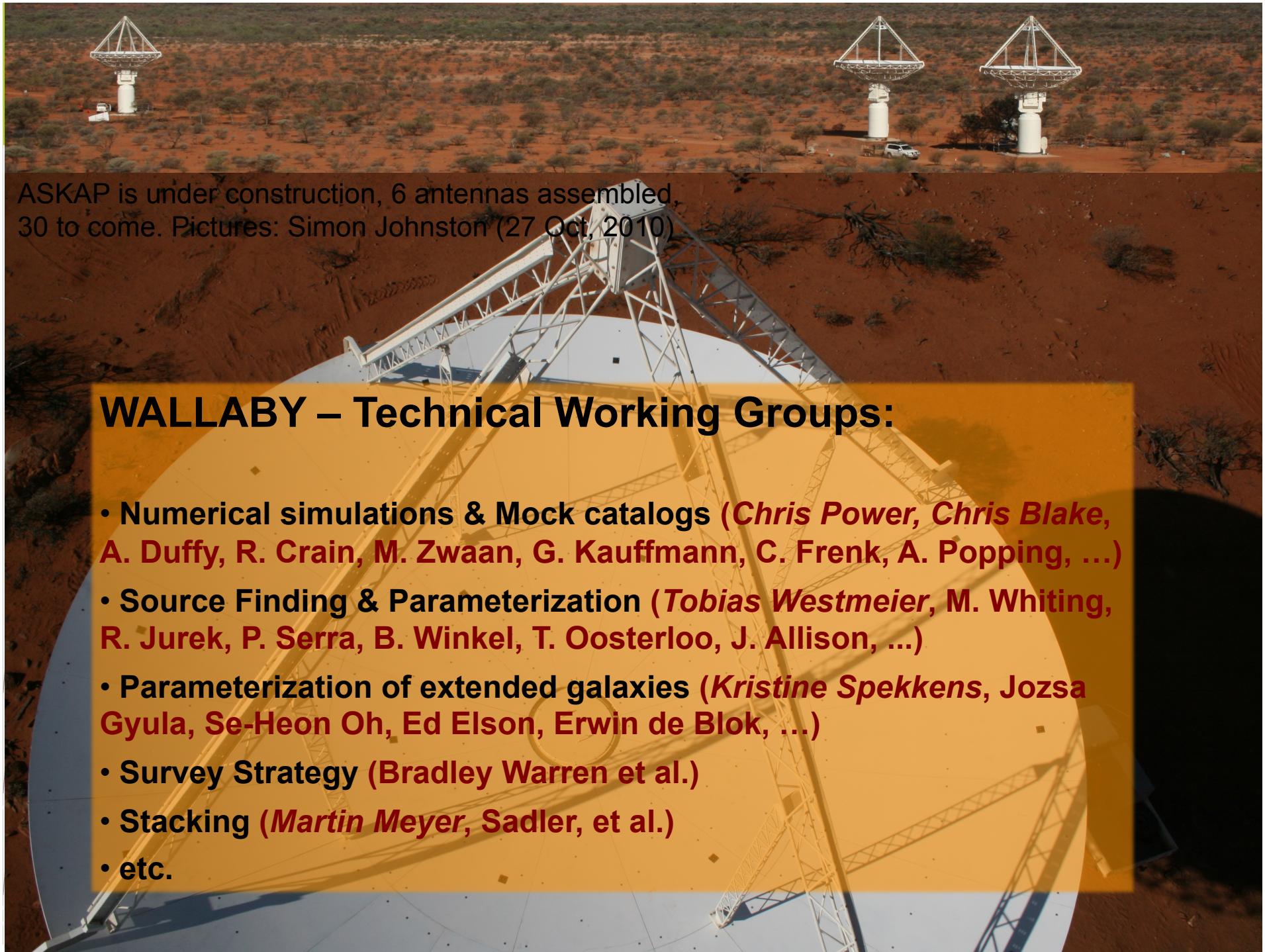


ASKAP uv-coverage:
30 / 36 antennas



1200 ×





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30 to come. Pictures: Simon Johnston (27 Oct, 2010)

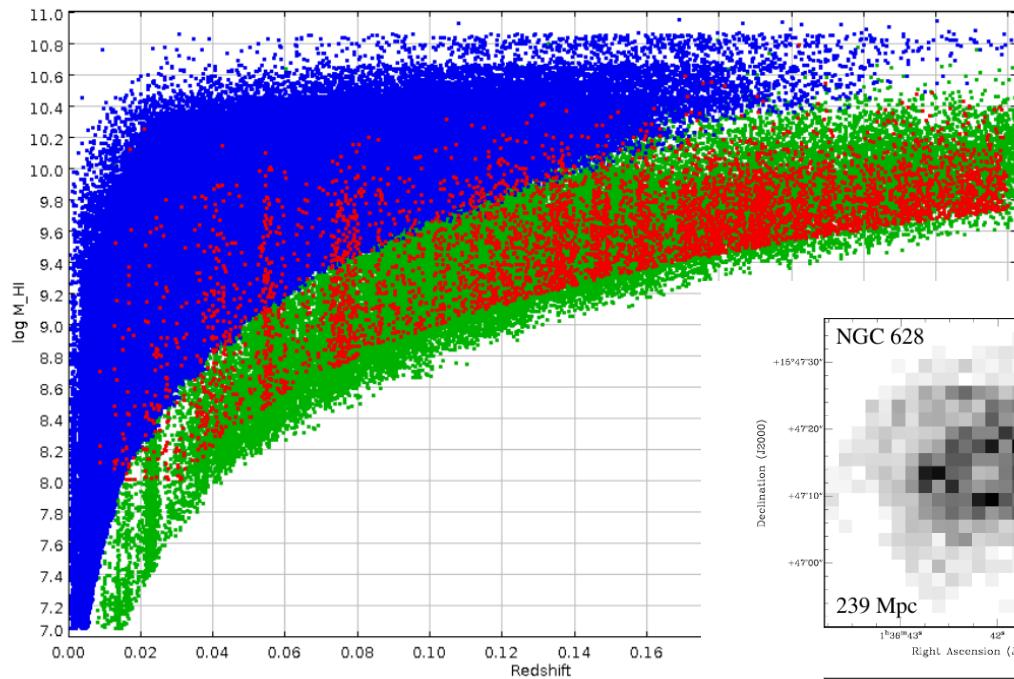
WALLABY – Technical Working Groups:

- Numerical simulations & Mock catalogs (*Chris Power, Chris Blake, A. Duffy, R. Crain, M. Zwaan, G. Kauffmann, C. Frenk, A. Popping, ...*)
- Source Finding & Parameterization (*Tobias Westmeier, M. Whiting, R. Jurek, P. Serra, B. Winkel, T. Oosterloo, J. Allison, ...*)
- Parameterization of extended galaxies (*Kristine Spekkens, Jozsa Gyula, Se-Heon Oh, Ed Elson, Erwin de Blok, ...*)
- Survey Strategy (*Bradley Warren et al.*)
- Stacking (*Martin Meyer, Sadler, et al.*)
- etc.

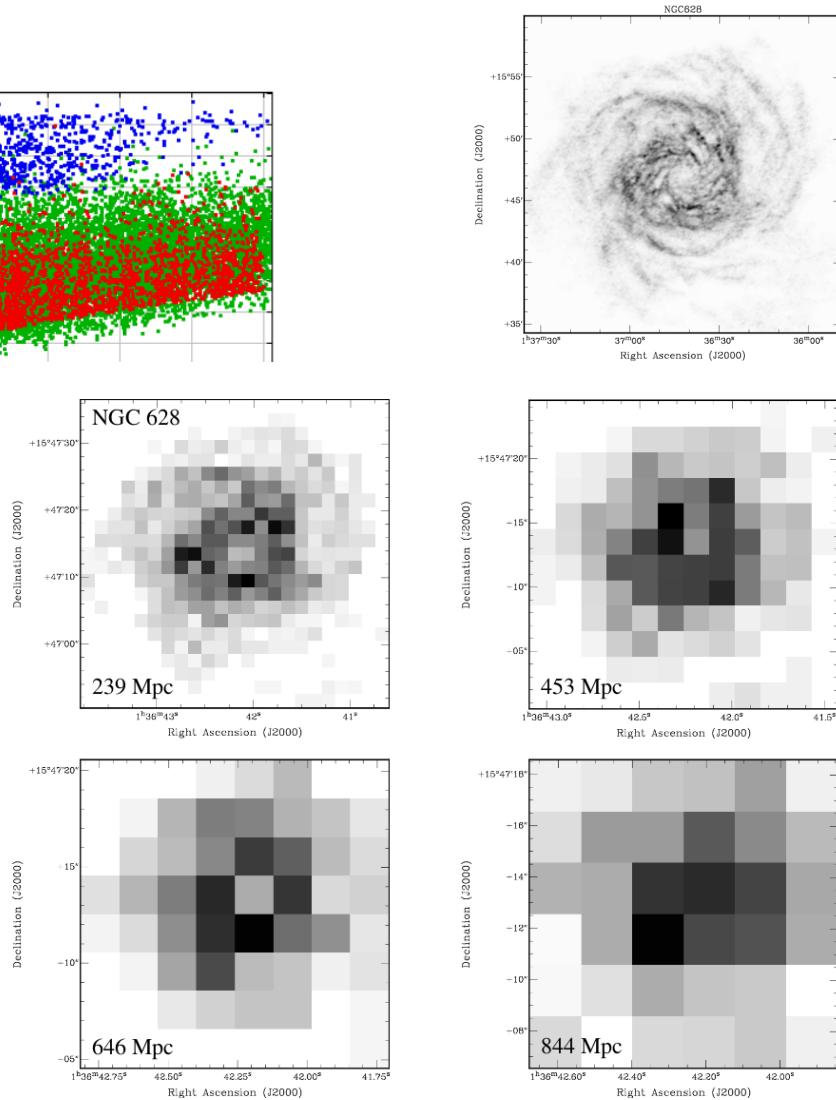
ASKAP simulations – input galaxies



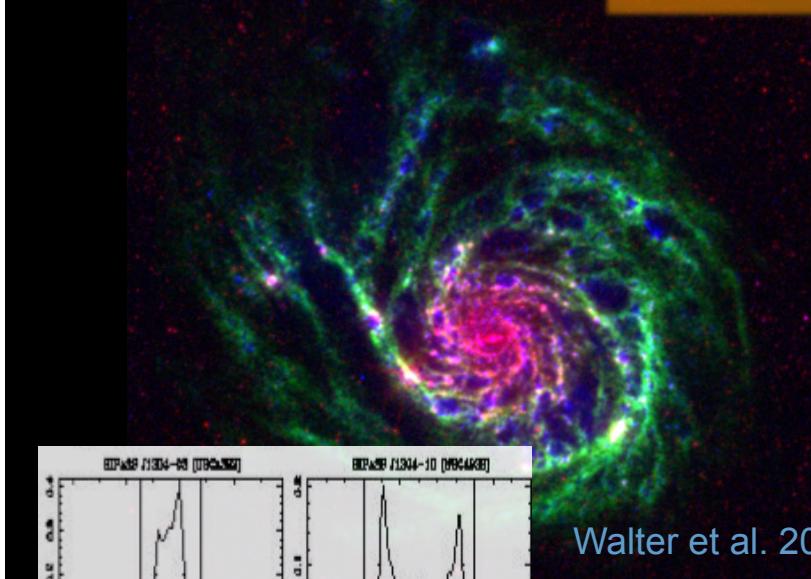
HI mass vs. redshift for **WALLABY**, **DINGO deep**, and **simulation**



300 HI data sets for nearby galaxies;
red-shifted to several distances and
distributed on cosmic web simulation.

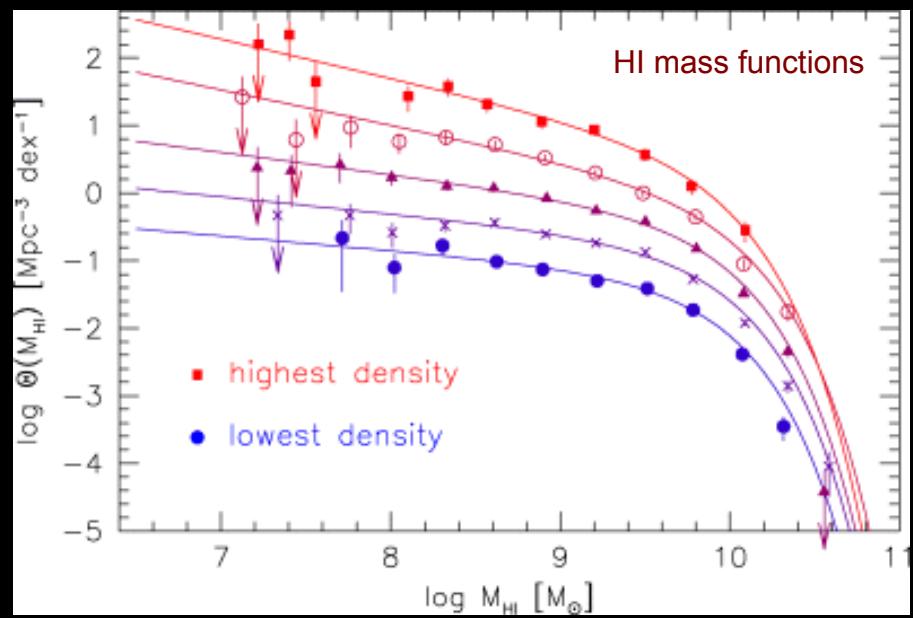
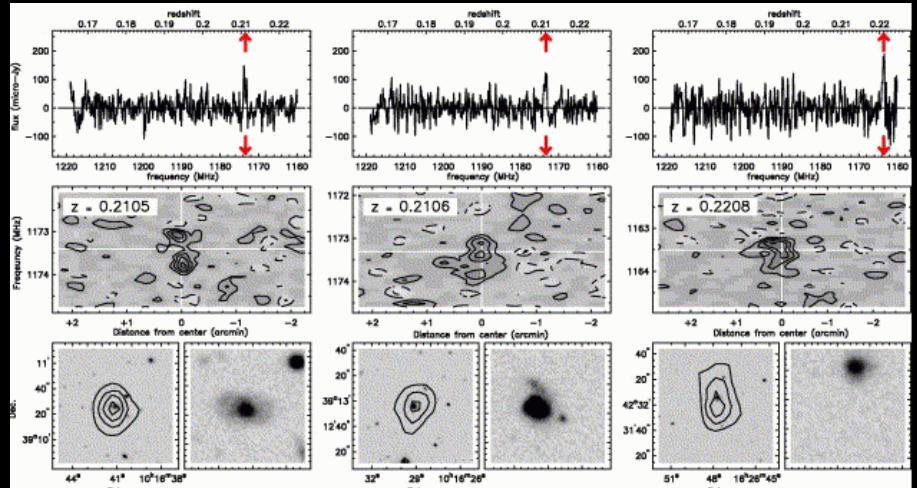


WALLABY in summary



500 000 HI detections,
ie. 500 000 redshifts, HI
& dynamical masses, ...

Koribalski et al. 2004



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Phase 1 = BETA: $6 \times 12\text{-m}$ dishes



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CSIRO Astronomy and Space Science
Australia Telescope National Facility
“Gas in Galaxies” – Kloster Seeon – June 2011

