Workshop on the progenitor-supernova-remnant connection @Ringberg Castle, Germany 2017/07/25

Rotation aided Neutrino-driven Explosion of Core-Collapse Supernovae

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Explosion Mechanism of Core-Collapse Supernovae



- 1. Methods and initial models
- 2. Results
 - 1. Shock revival helped by a rotational instability
 - 2. How the instability grows
 - 3. How the instability helps the explosion
 - 4. Features of neutrino and gravitational waves
- 3. Discussion on progenitors
- 4. Summary

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Method in this presentation (now updated)

Hydro: 3DnSNe

- Geometry: Spherical coordinate, PWL interp.
- HLL scheme

Neutrino Radiation Transport: IDSA+Leakage

- For ν_e and $\overline{
 u}_e$, flux-IDSA (Takiwaki+2014)
- For ν_X , Leakage Scheme (Rosswog & Liebendoerfer 2003)
- Minimum set of the reactions considered (Takiwaki+14) Gravity:
- Newtonian Monopole approximation

While our method is not state-of-the art, that does NOT require high CPU resources.

We can perform many models.

Initial setups and Results



Initial setups and Results



 For non-rotating 3D simulations, Bernhard Mueller will talk. I'll skip it.

Initial setups and Results



 I focus on most interesting models. Rotation plays an essential role in the explosion.

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Evolution of the Shock and Explosion Energy



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Growth of Spiral Mode





Rotational energy(T)/gravitational energy(W) reach some criteria => Spiral mode arises In the rigid ball: 14% Ott+ 2009 In SNe case: ~ 6% (Called low-T/W instability)



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Energy Transport by spiral mode



Spiral mode transport energy from center to outer region and helps explosion.

Key aspects of Neutrino Mechanism





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Neutrino signals from rotating model



Anisotropy of neutrino emission:



Takiwaki+ in prep Like Light house, periodic signal of the neutrino will come to an observer.

Neutrino signals from rotating model



The time variability is visible at the direction of equator by the observation of IceCube (as well as Hyper Kamiokande). From the pole, the variability disappers.

Neutrino signals from rotating model



Takiwaki+ in prep



Takiwaki+ in prep

Bounce signal can be observed from side view. Non-axisymmetric motion emits GWs at later phase.

Feature of GWs from Rotational Explosion Takiwaki+ in prep Viewing from side direction 10⁻²¹ 800 Bounce 600 Only visible from Equator requency [Hz] 10^{-22} 400 Spiral mode ~220Hz 10⁻²³ 200 10⁻²⁴ 100 50 100 0 150 200 Time after bounce [ms] In addition to g-mode signal, GW from spiral mode arises and spin period of PNS surface can be extracted. 24



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Progenitor?

- Rapid rotation of $\Omega = 2.0 \text{rad/s}$ is assumed in our model.
- In solar metallicity, such a rotation is difficult due to large mass loss and large angular momentum transfer via magnetic fields.
- If the metallicity become lower, Chemically homogeneous evolution might be possible and large angular momentum could remain in the core.
- In that sense, we expect the star like our model would live in LMC and SMC (however, a few such progenitors are found in our Galaxy. Martins 2013).

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Summary

- 1. We found a new model of ν -driven explosion that is helped by the rotation.
- 2. The rotation induced low-T/|W| instability that makes a spiral mode.
- 3. Pushed by the spiral mode, the shock expands oblately.
- 4. The neutrino and GW signal oscillate with their characteristic frequency that are related to the angular velocity of the neutrino sphere.

Method in the latest version

Hydro: 3DnSNe

- Geometry: Spherical coordinate, PWL interp. (Mignone 2014)
- HLL scheme

Neutrino Radiation Transport: IDSA+Leakage

- flux-IDSA for all flavor (Takiwaki+2014)
- Standard set of the reactions considered (Liebendoerfer+ 2005)

Gravity:

• Effective GR Monopole approximation(Marek+ 2006)