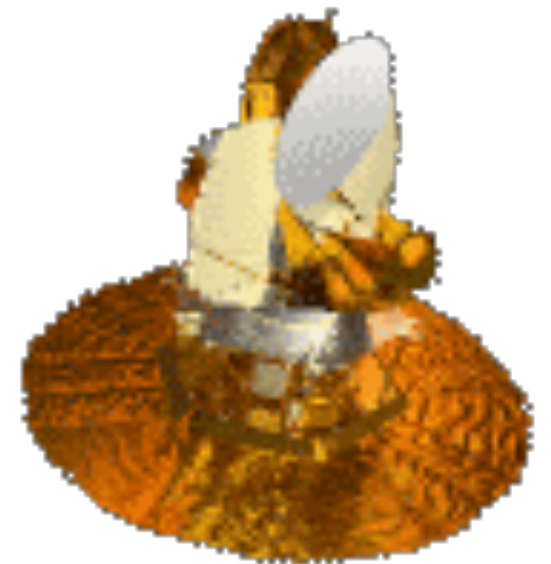


Critical Tests of Theory of the Early Universe using the Cosmic Microwave Background

Eiichiro Komatsu, Max-Planck-Institut für Astrophysik
Instituto de Física Corpuscular (IFIC), Valencia
September 14, 2017



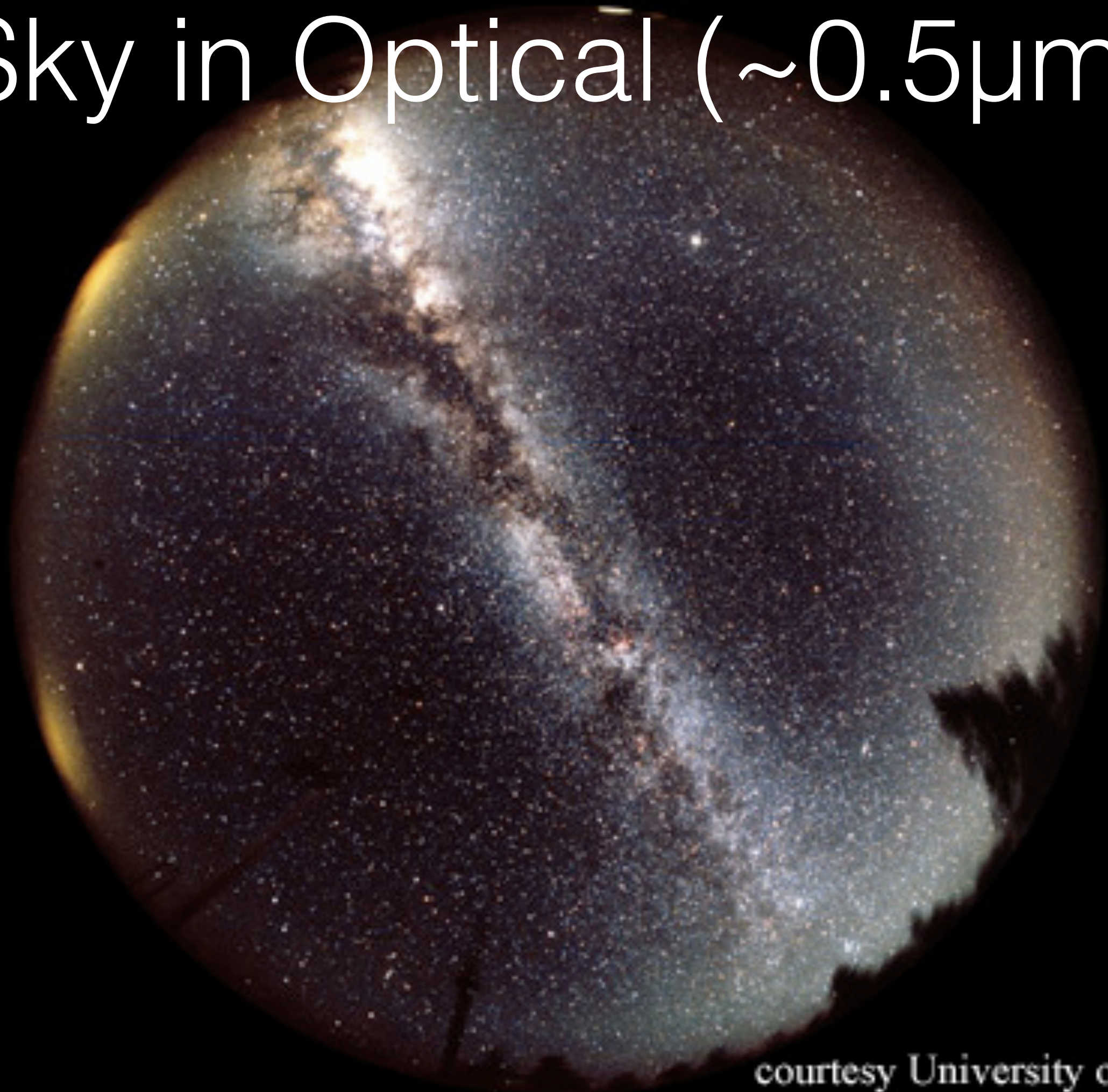
Breakthrough in Cosmological Research

- We can actually **see** the physical condition of the universe when it was very young



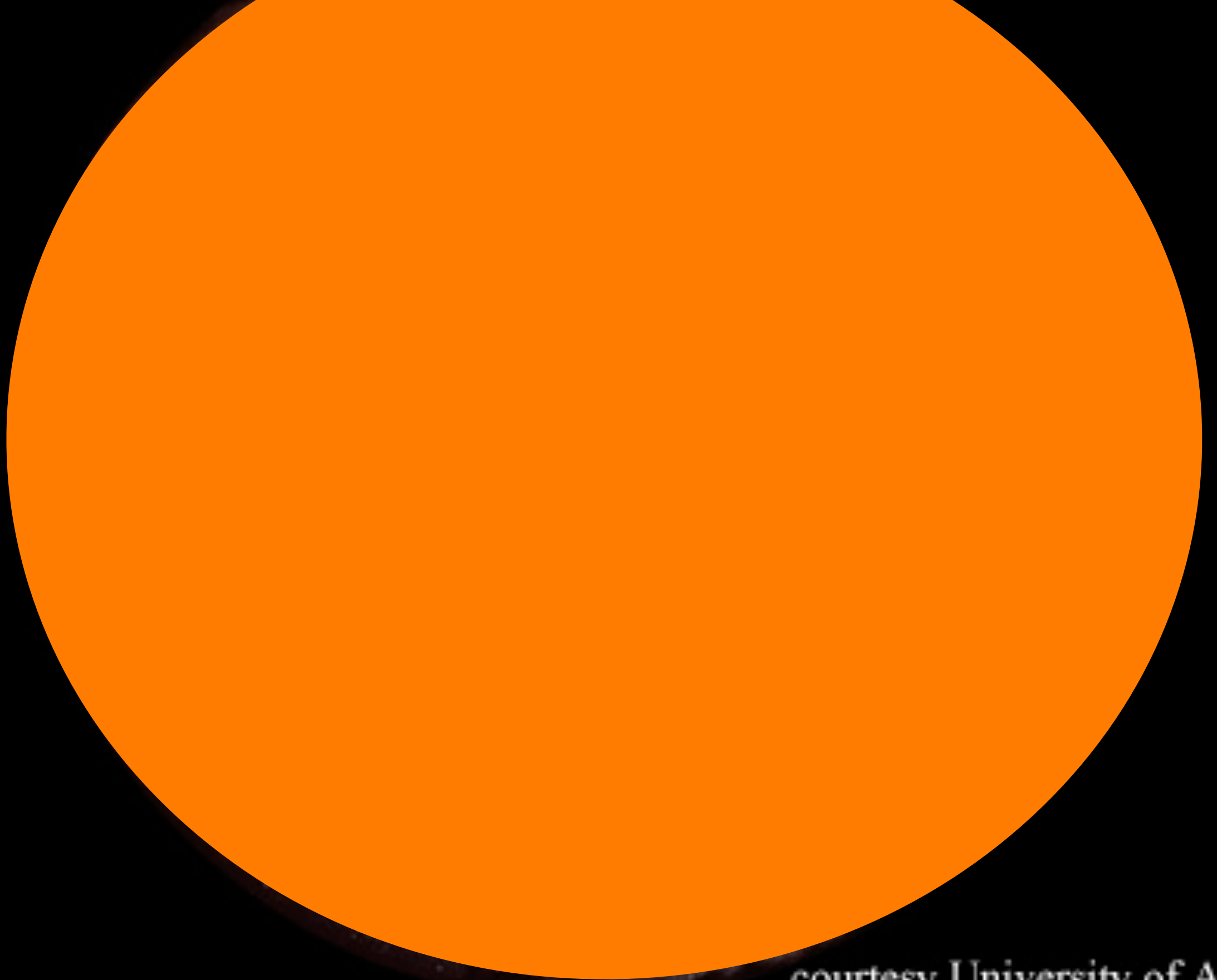
From "Cosmic Voyage"

Sky in Optical ($\sim 0.5\mu\text{m}$)



courtesy University of Arizona

Sky in Microwave ($\sim 1\text{mm}$)



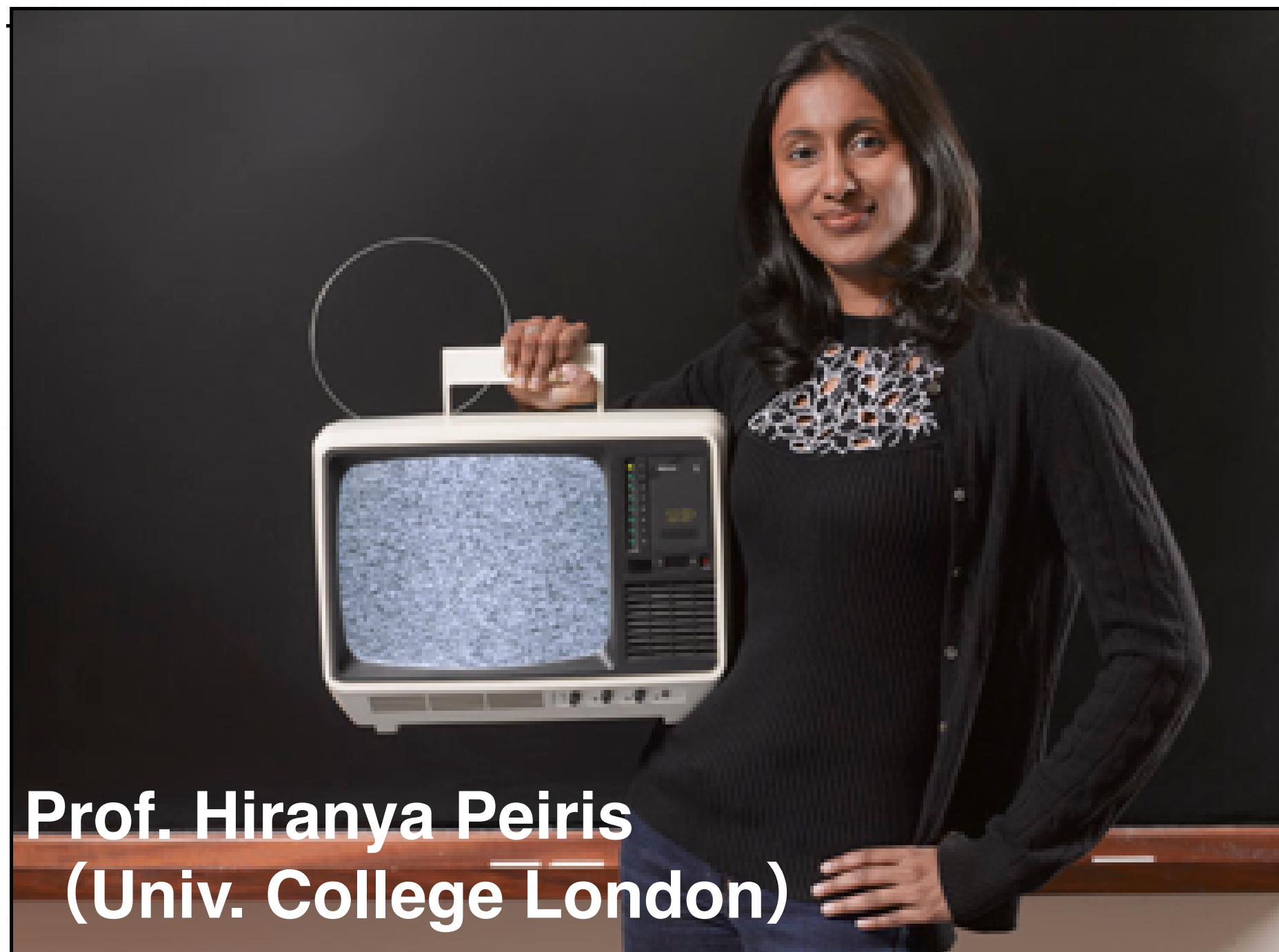
courtesy University of Arizona

Sky in Microwave ($\sim 1\text{mm}$)

*Light from the fireball Universe
filling our sky (2.7K)*

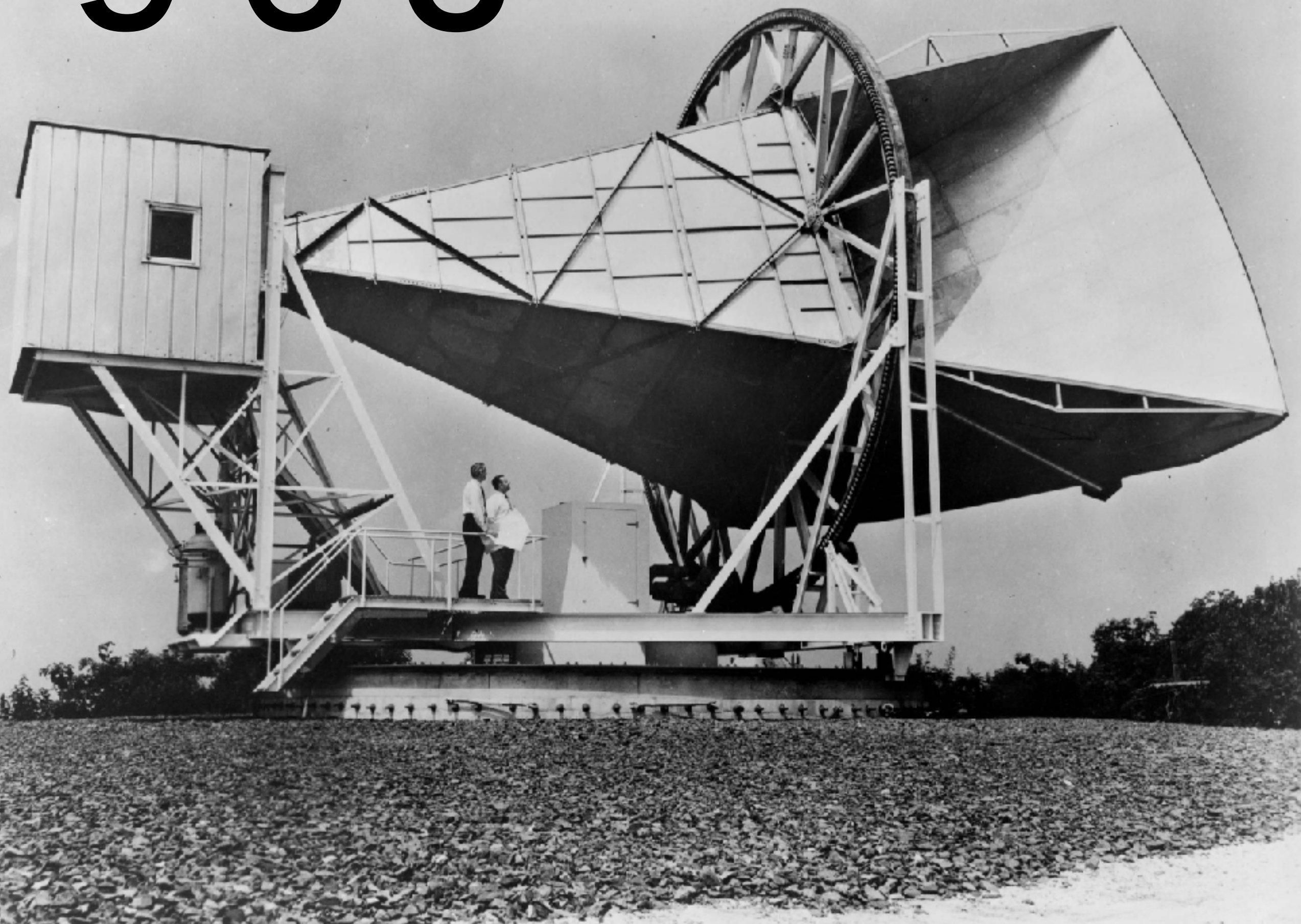
**The Cosmic Microwave
Background (CMB)**

410 photons
per
cubic centimeter!!



All you need to do is to detect radio waves. For example, 1% of noise on the TV is from the fireball Universe

1965



1:25 model of the antenna at Bell Lab
The 3rd floor of Deutsches Museum



The real detector system used by Penzias & Wilson

The 3rd floor of Deutsches Museum



**Donated by Dr. Penzias,
who was born in Munich**



Horn antenna

Calibrator, cooled
to 5K by liquid helium

Amplifier

Recorder

Hornantennenanschluss

Hohlleiterzug

V
Vergleichs-
quelle

R
Rauschquelle

F
Frequenzmischer
und Verstärker

M
MASER-Verstärker

Schreiber

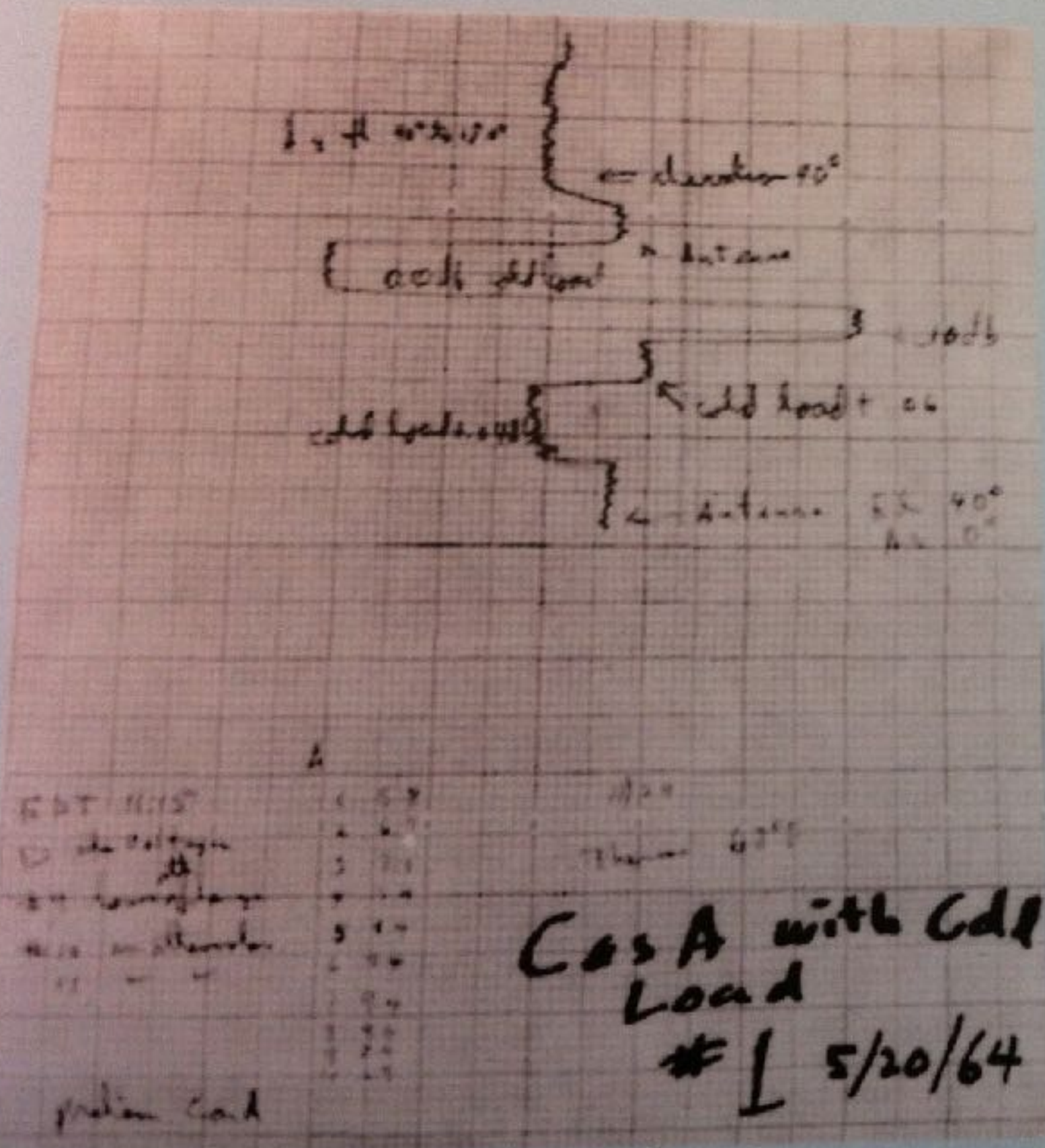
many
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May 20, 1964 CMB Discovered

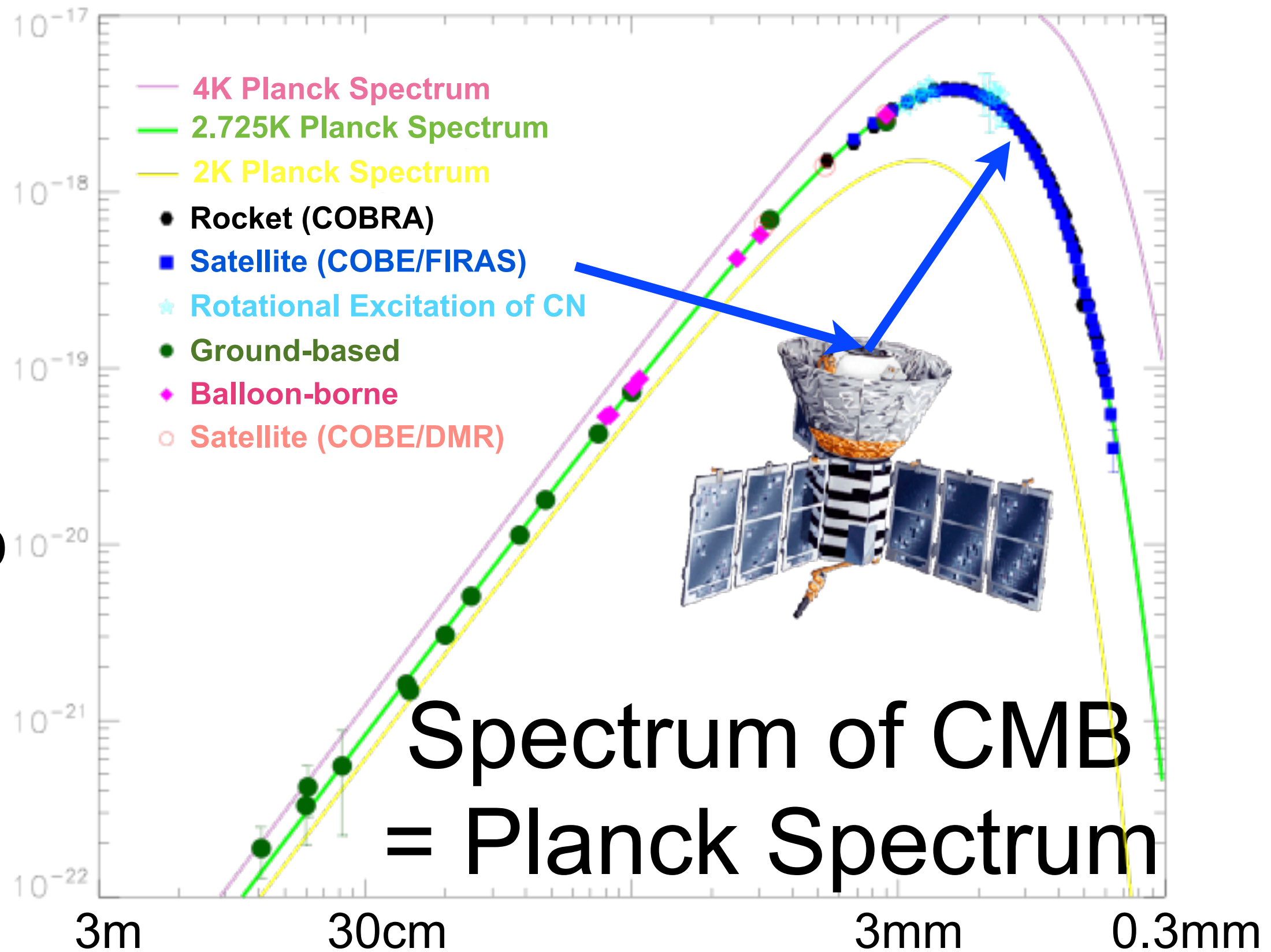
$$6.7 - 2.3 - 0.8 - 0.1 \\ = 3.5 \pm 1.0 \text{ K}$$



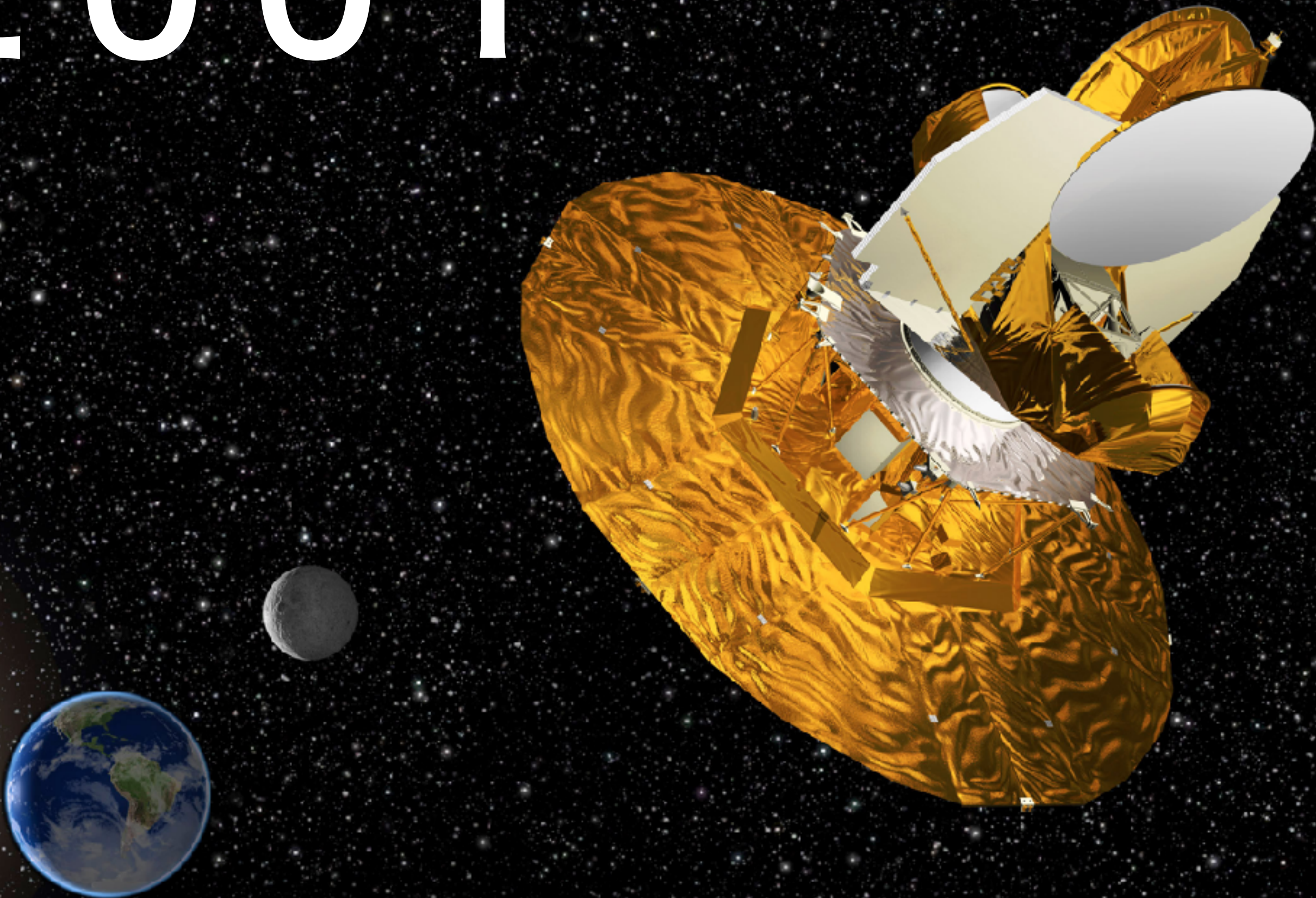
Schreiberaufzeichnung der ersten Messung des Mikrowellenhintergrundes am 20.5.1964

Recording of the first measurement of cosmic microwave background radiation taken on 5/20/1964.

Brightness



2001

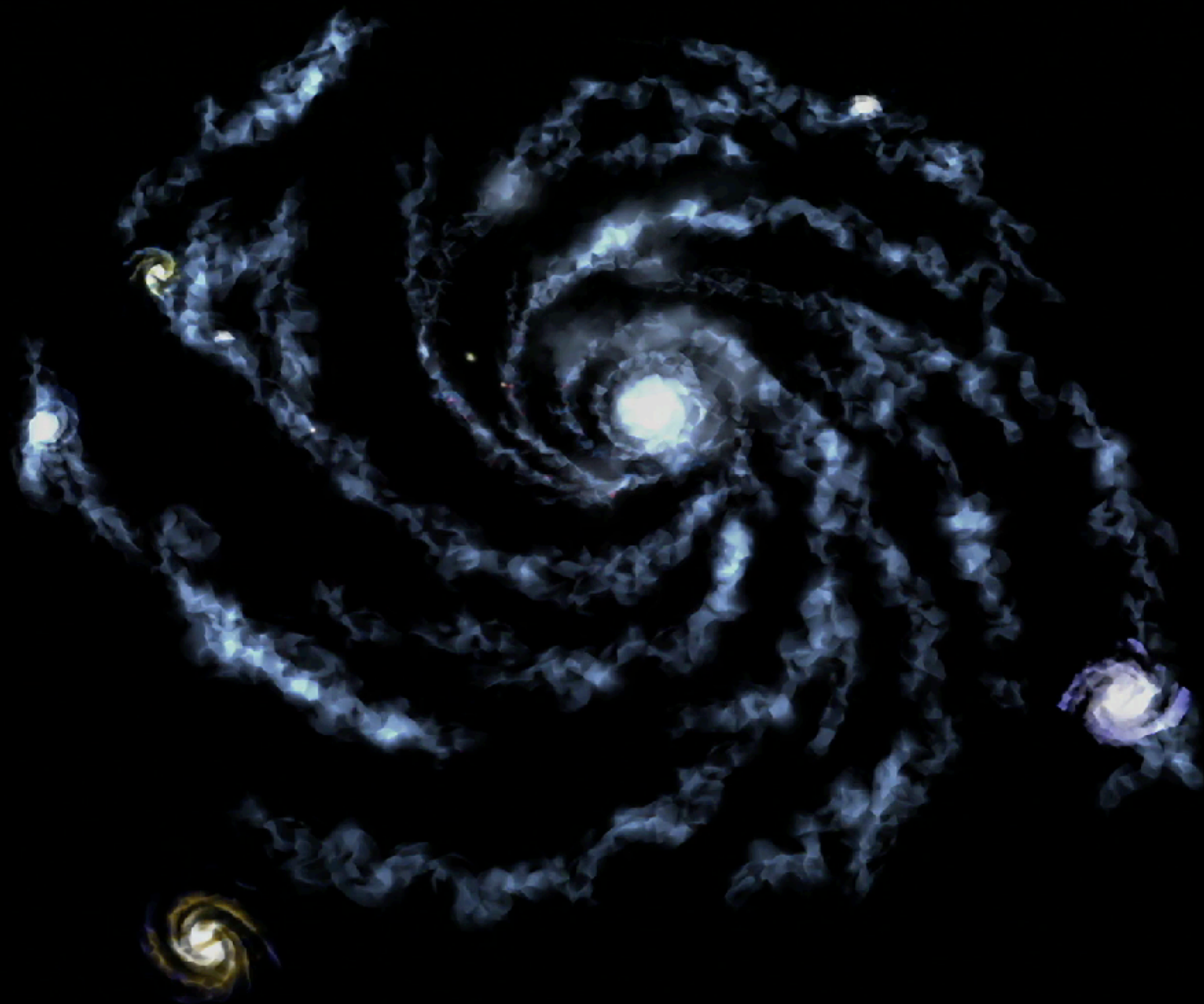


WMAP Science Team

July 19, 2002

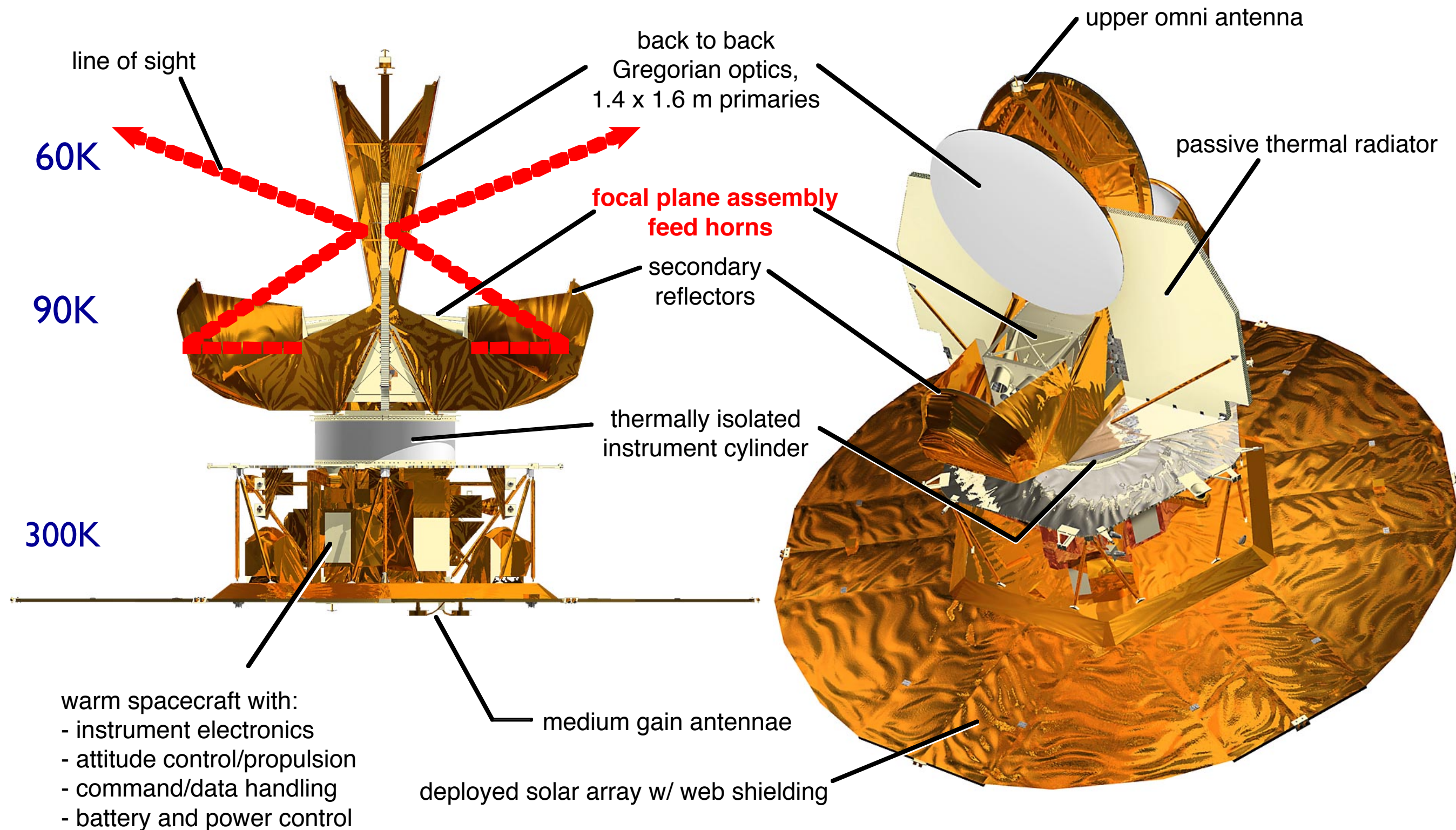


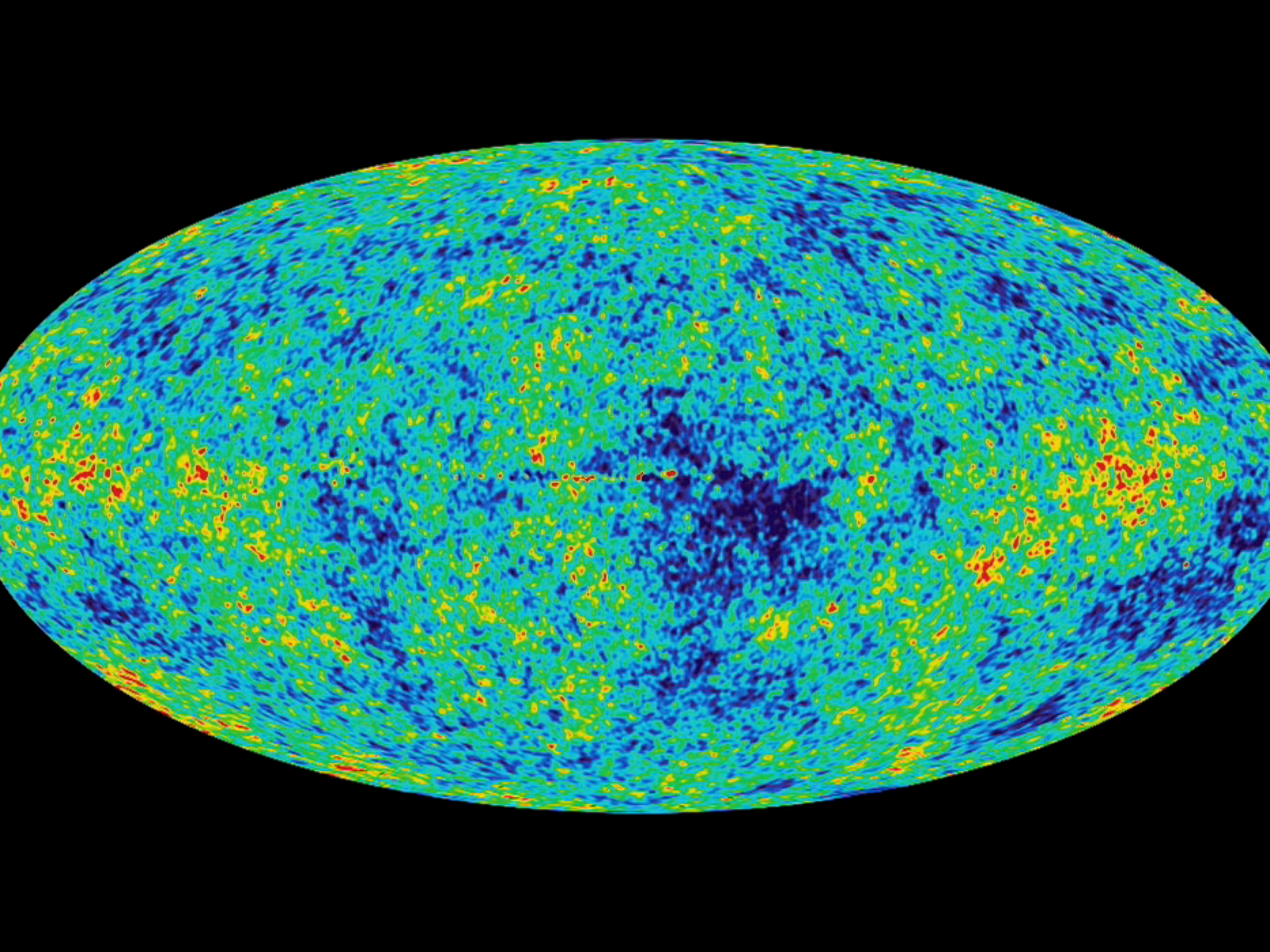
- WMAP was launched on June 30, 2001
- The WMAP mission ended after 9 years of operation

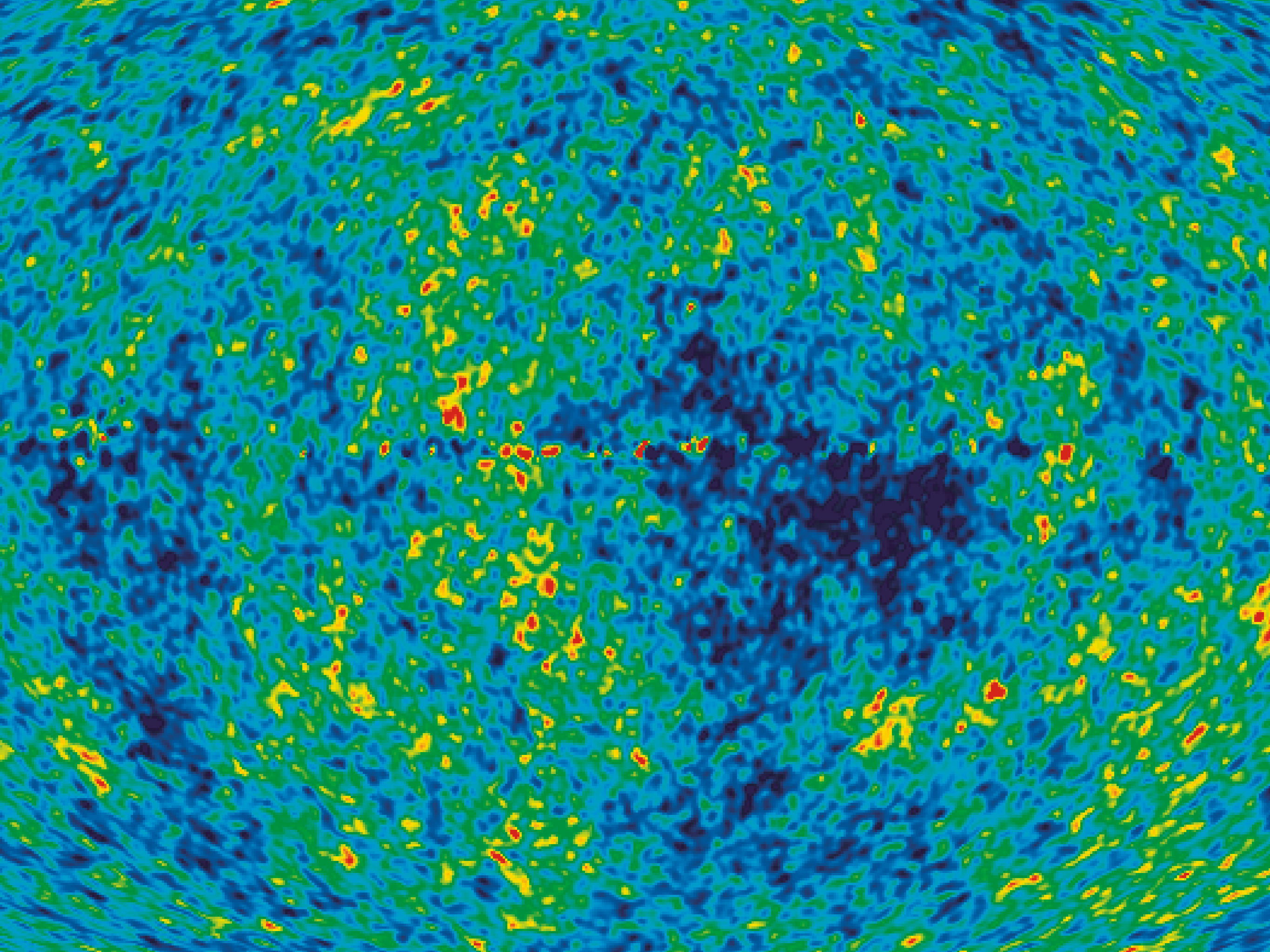


WMAP Spacecraft

No cryogenic components







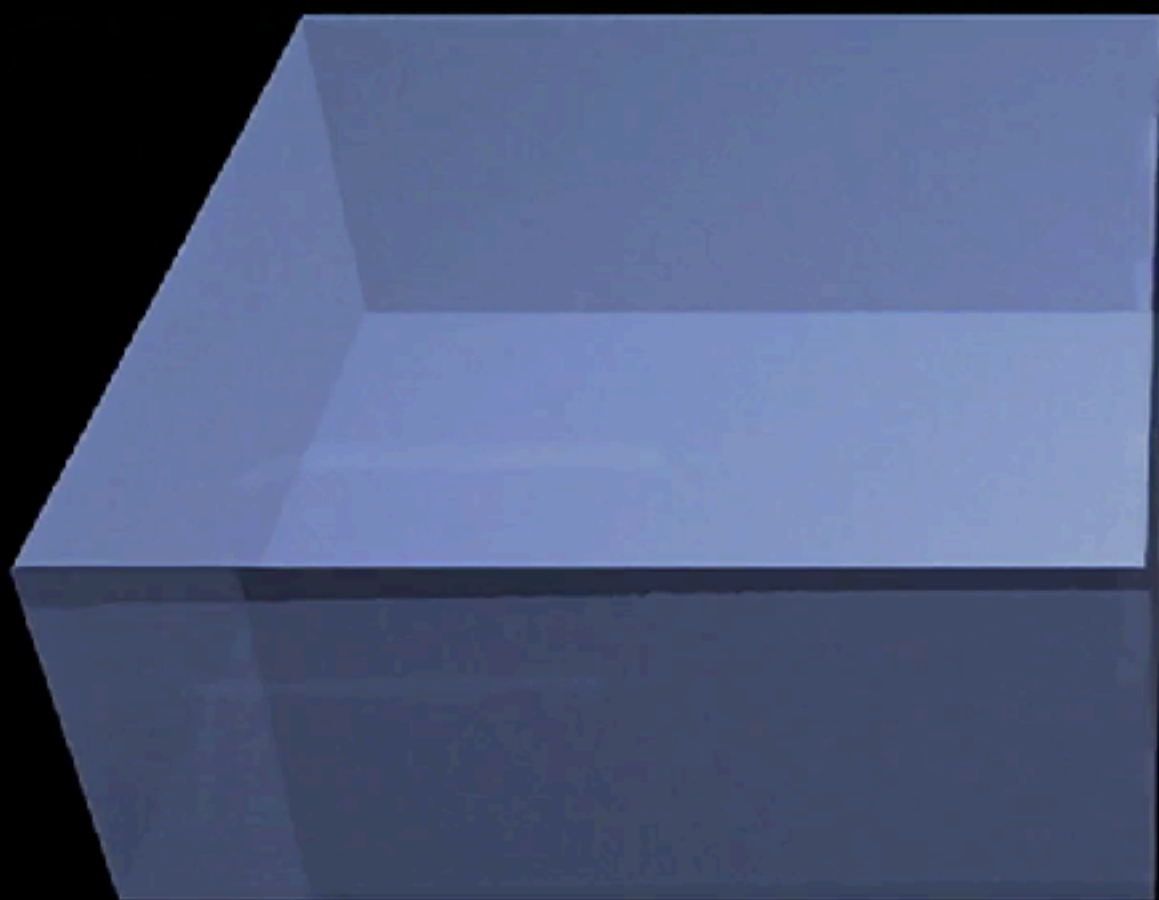
Our Origin

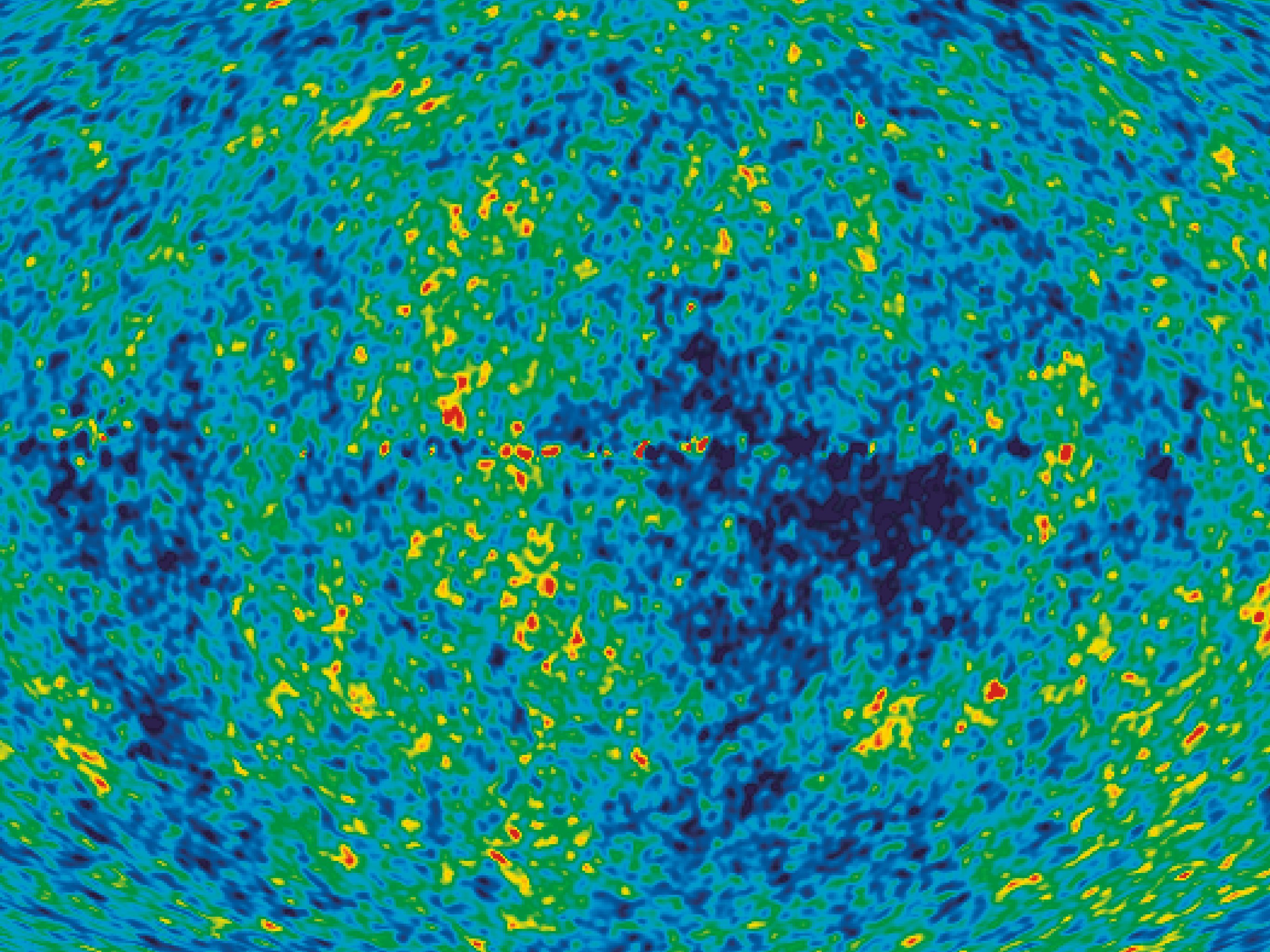
- WMAP taught us that **galaxies, stars, planets, and ourselves originated from tiny fluctuations in the early Universe**



Sopa de Miso Cósmica

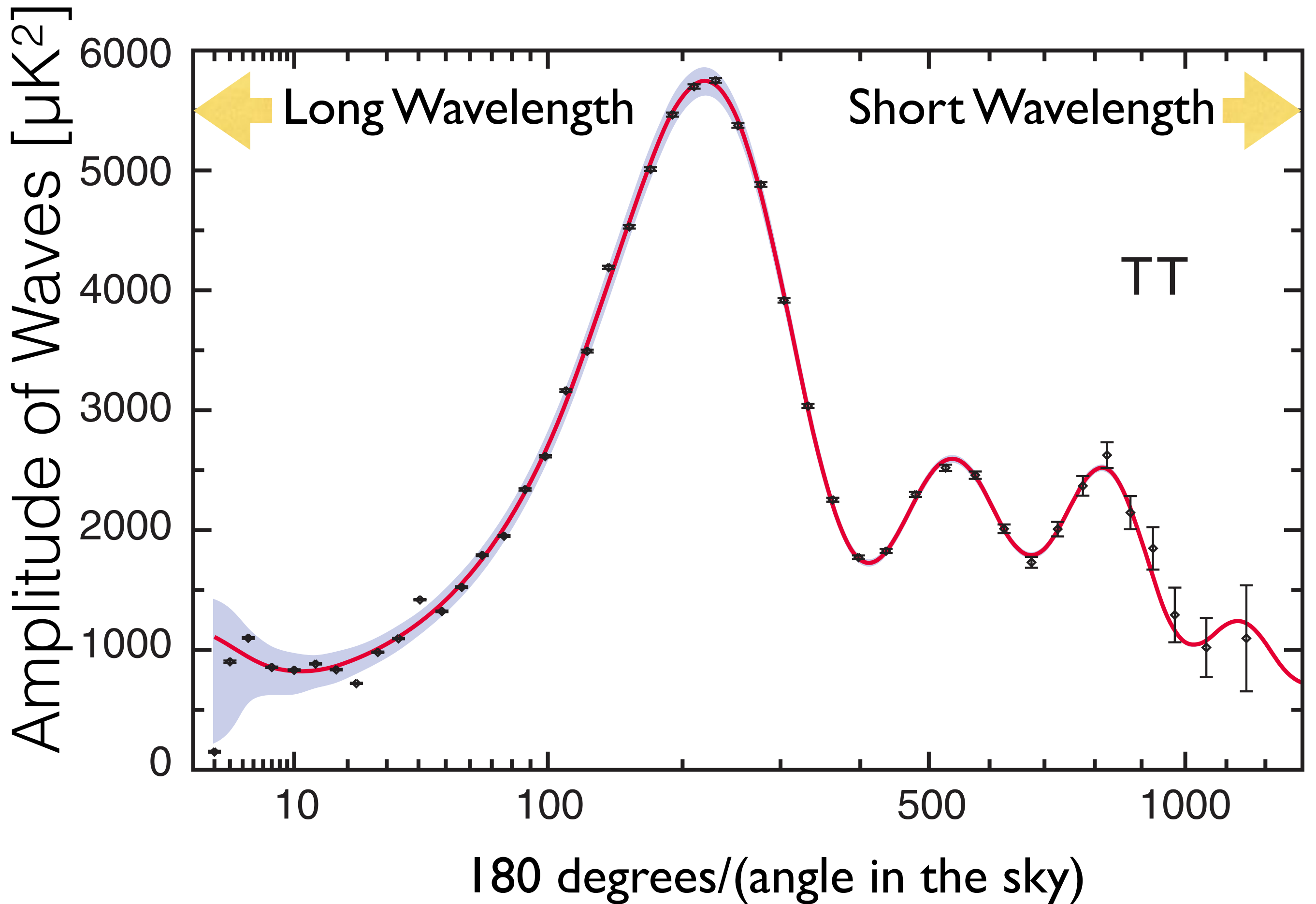
- When matter and radiation were hotter than 3000 K, matter was completely ionised. The Universe was filled with plasma, which behaves just like a soup
- Think about a Miso soup (if you know what it is). Imagine throwing Tofus into a Miso soup, while changing the density of Miso
- And imagine watching how ripples are created and propagate throughout the soup

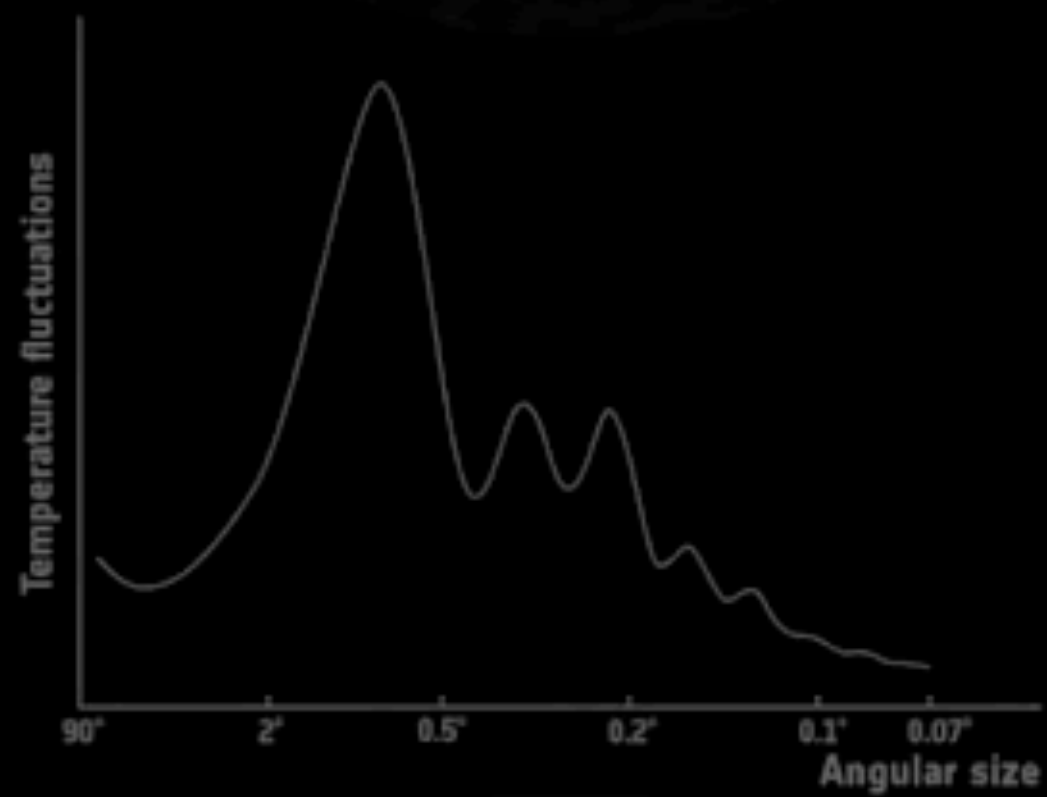




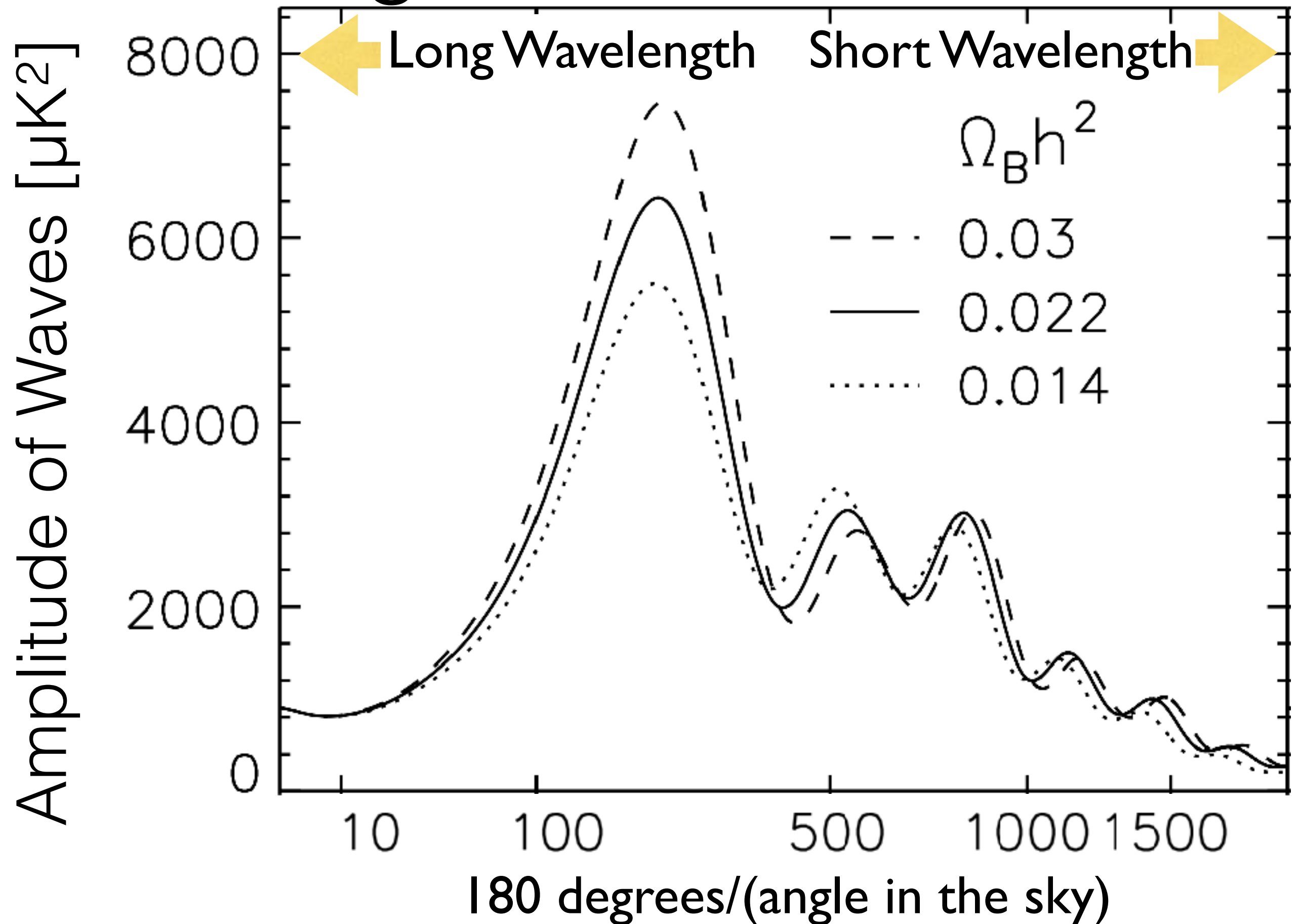
Data Analysis

- Decompose temperature fluctuations in the sky into a set of waves with various wavelengths
- Make a diagram showing the strength of each wavelength

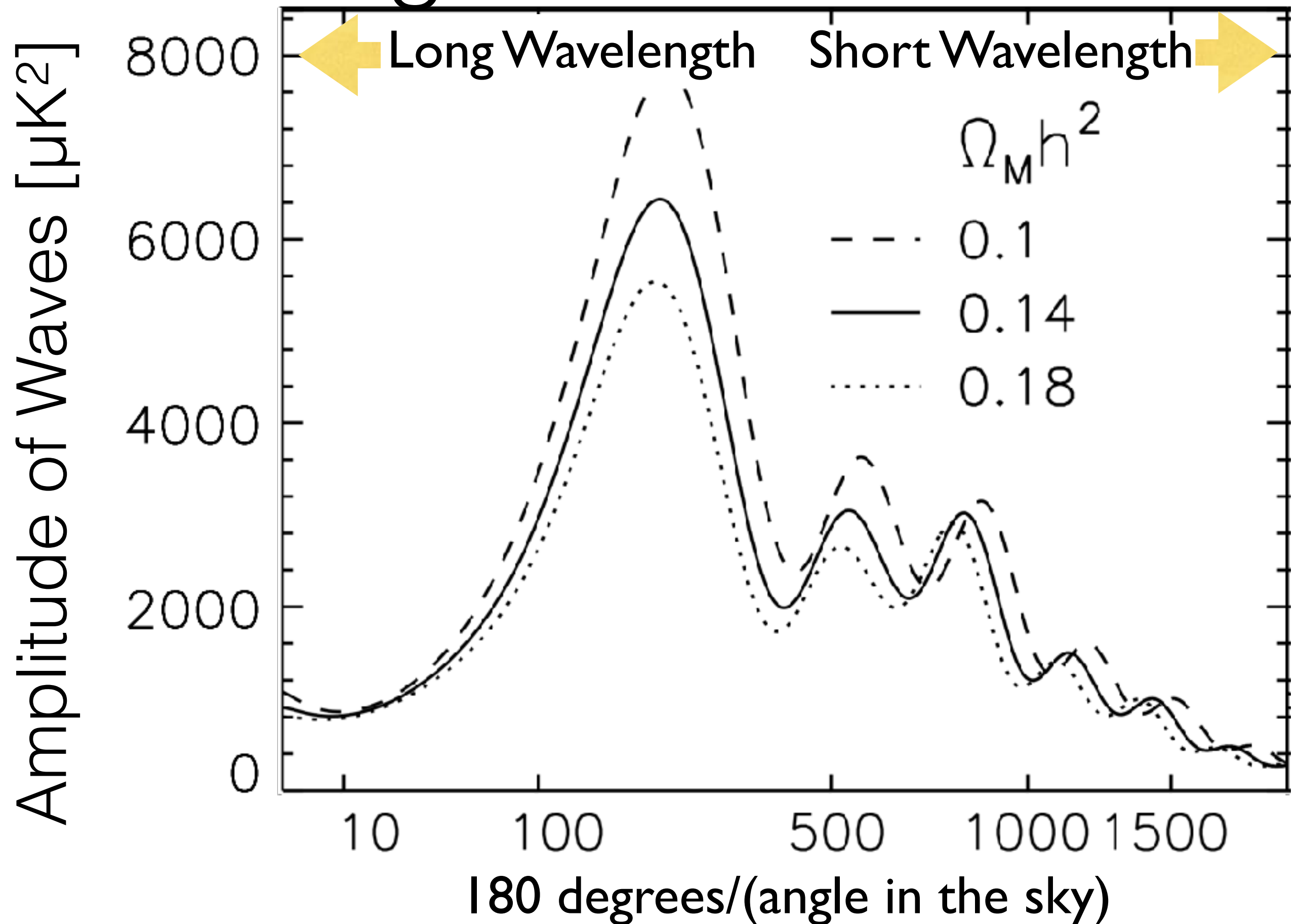


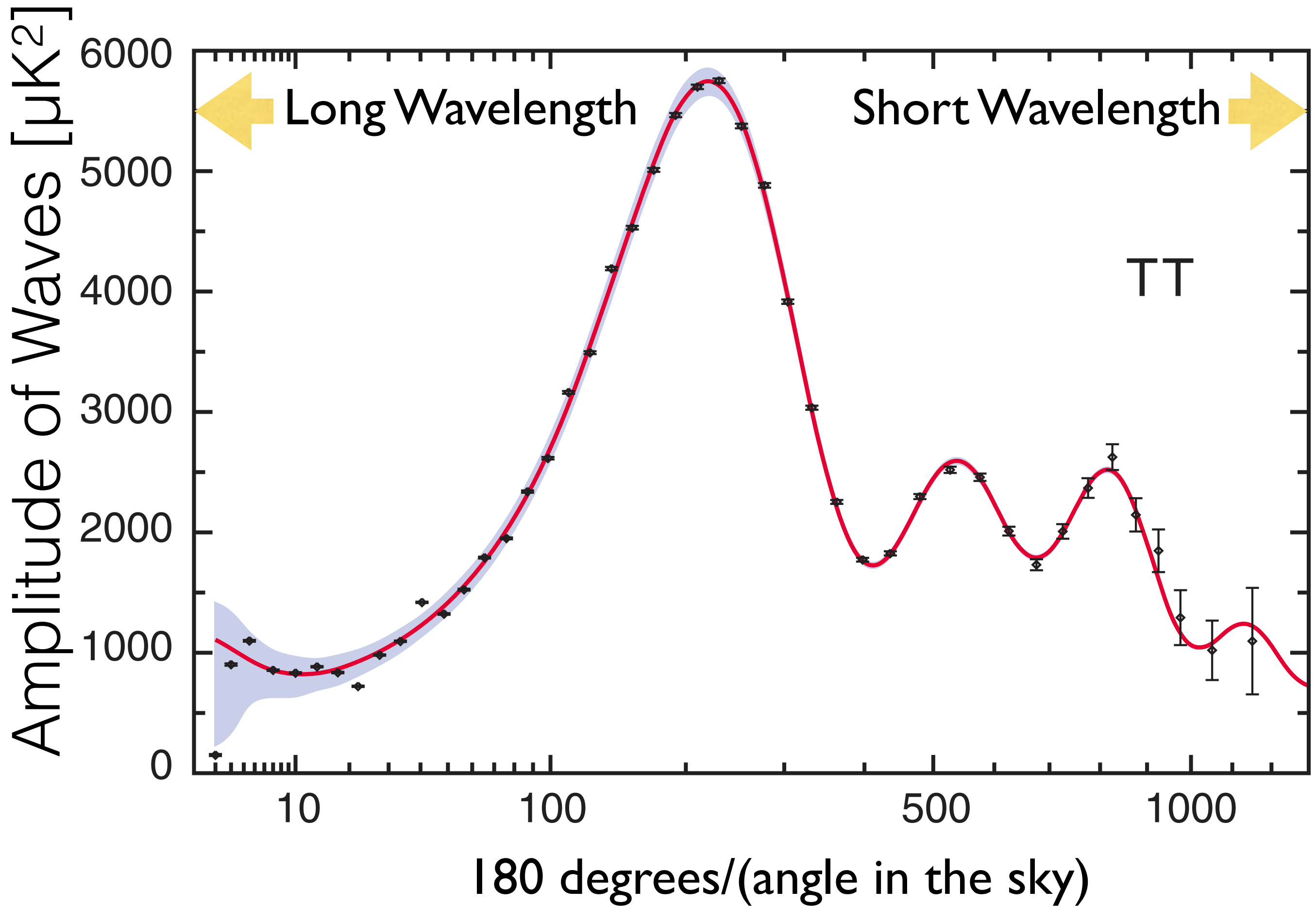


Measuring Abundance of H&He

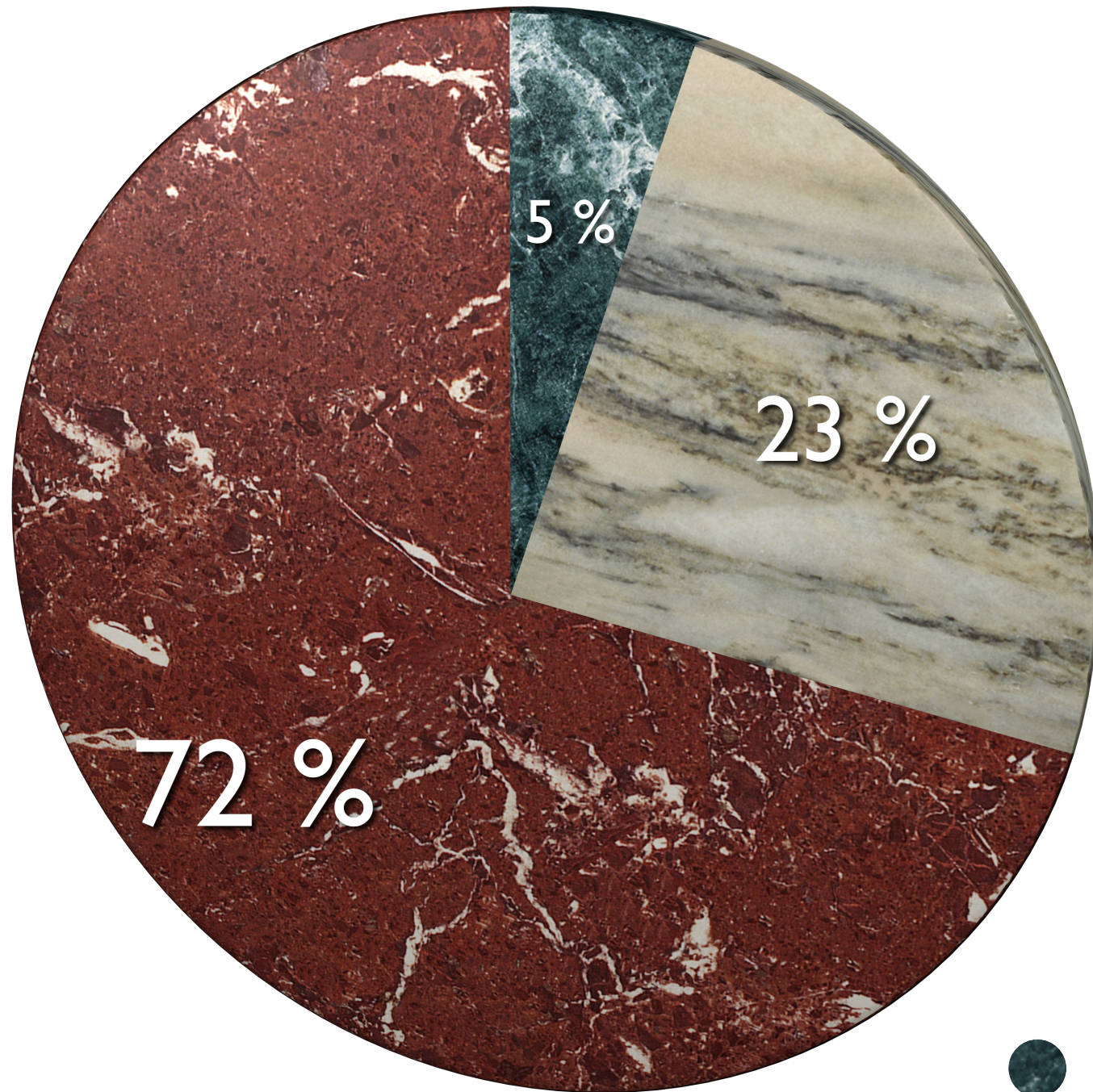


Measuring Total Matter Density





Cosmic Pie Chart

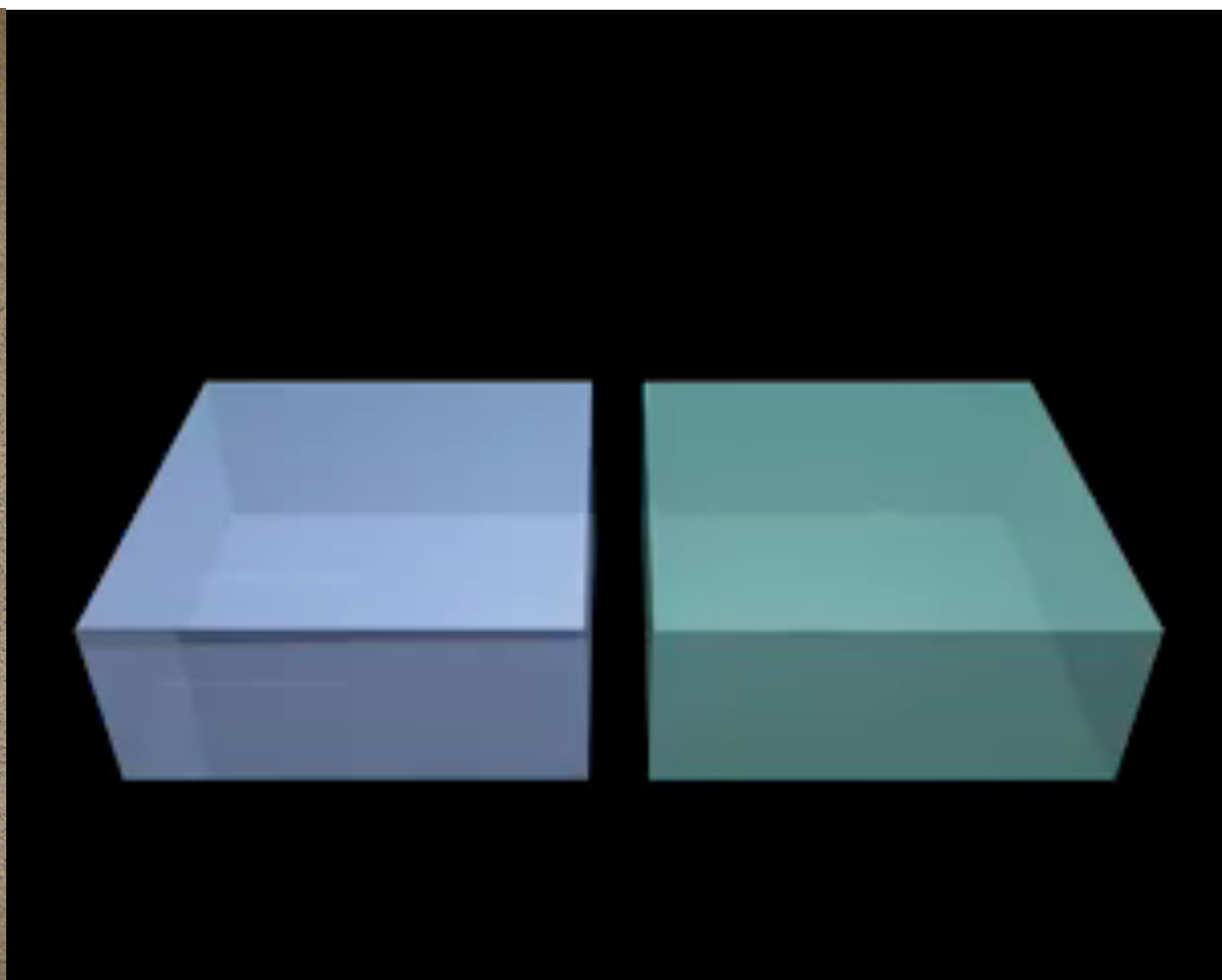


- WMAP determined the abundance of various components in the Universe
- As a result, **we came to realise that we do not understand 95% of our Universe...**



Origin of Fluctuations

- Who dropped those Tofus into the cosmic Miso soup?



Leading Idea

- **Quantum Mechanics at work in the early Universe**
- Uncertainty Principle:
 - [Energy you can borrow] x [Time you borrow] $\sim h$
 - Time was very short in the early Universe = You could borrow a lot of energy
- **Those energies became the origin of fluctuations**
- How did quantum fluctuations on the microscopic scales become macroscopic fluctuations over cosmological sizes?

Cosmic Inflation

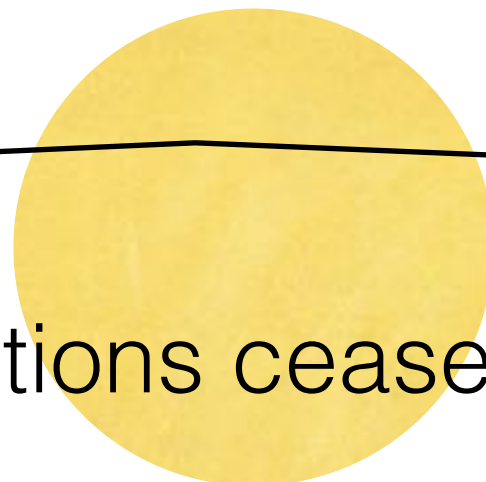
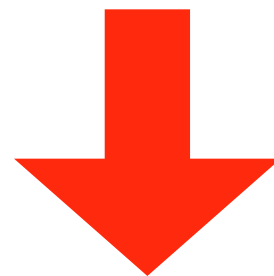
- In a tiny fraction of a second, the size of an atomic nucleus became the size of the Solar System
- In 10^{-36} second, space was stretched by at least a factor of 10^{26}

Stretching Micro to Macro

Quantum fluctuations on
microscopic scales



Inflation!



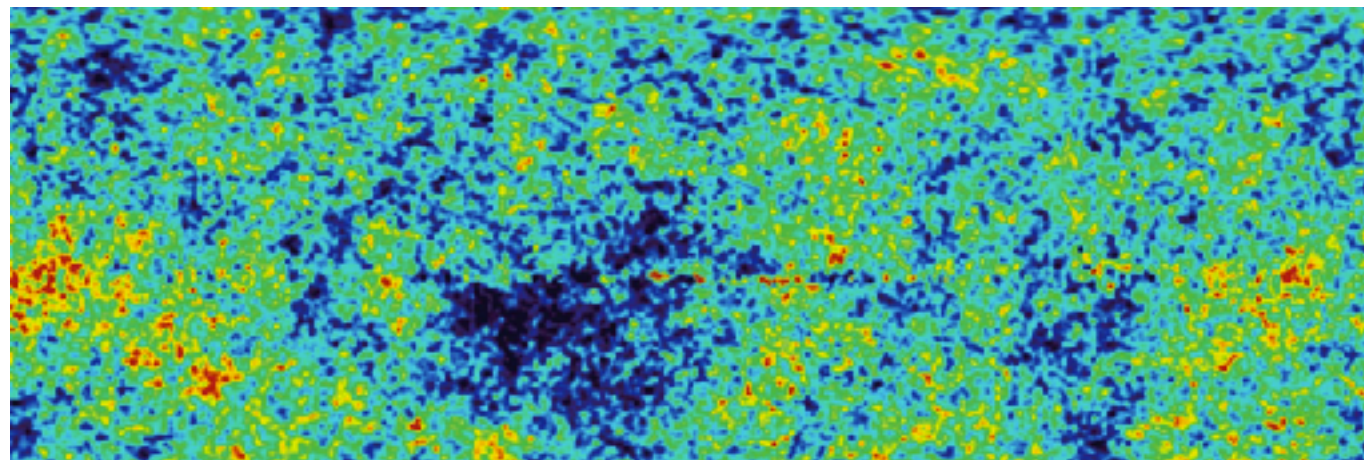
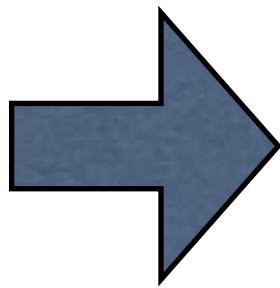
- Quantum fluctuations cease to be quantum
- Become macroscopic, classical fluctuations

Key Predictions of Inflation

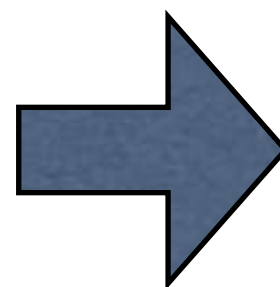
 ζ

scalar
mode

- Fluctuations we observe today in CMB and the matter distribution originate from quantum fluctuations generated during inflation

 h_{ij}

tensor
mode



We measure distortions in space

- A distance between two points in space

$$d\ell^2 = a^2(t)[1 + 2\zeta(\mathbf{x}, t)][\delta_{ij} + h_{ij}(\mathbf{x}, t)]dx^i dx^j$$

- ζ : “curvature perturbation” (scalar mode)
 - Perturbation to the determinant of the spatial metric
- h_{ij} : “gravitational waves” (tensor mode)
 - Perturbation that does not change the determinant (area)



$$\sum_i h_{ii} = 0$$

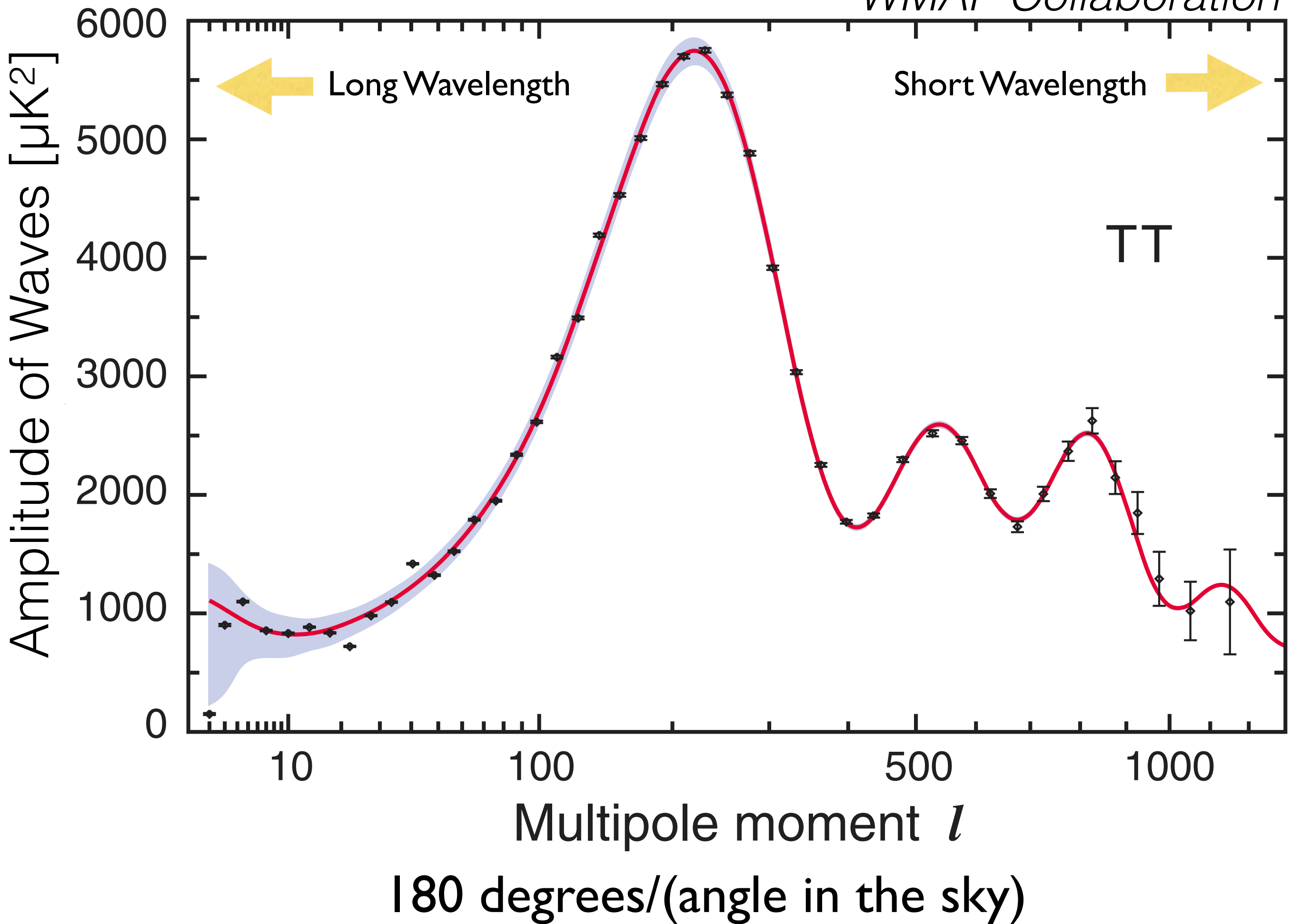
Heisenberg's Uncertainty Principle

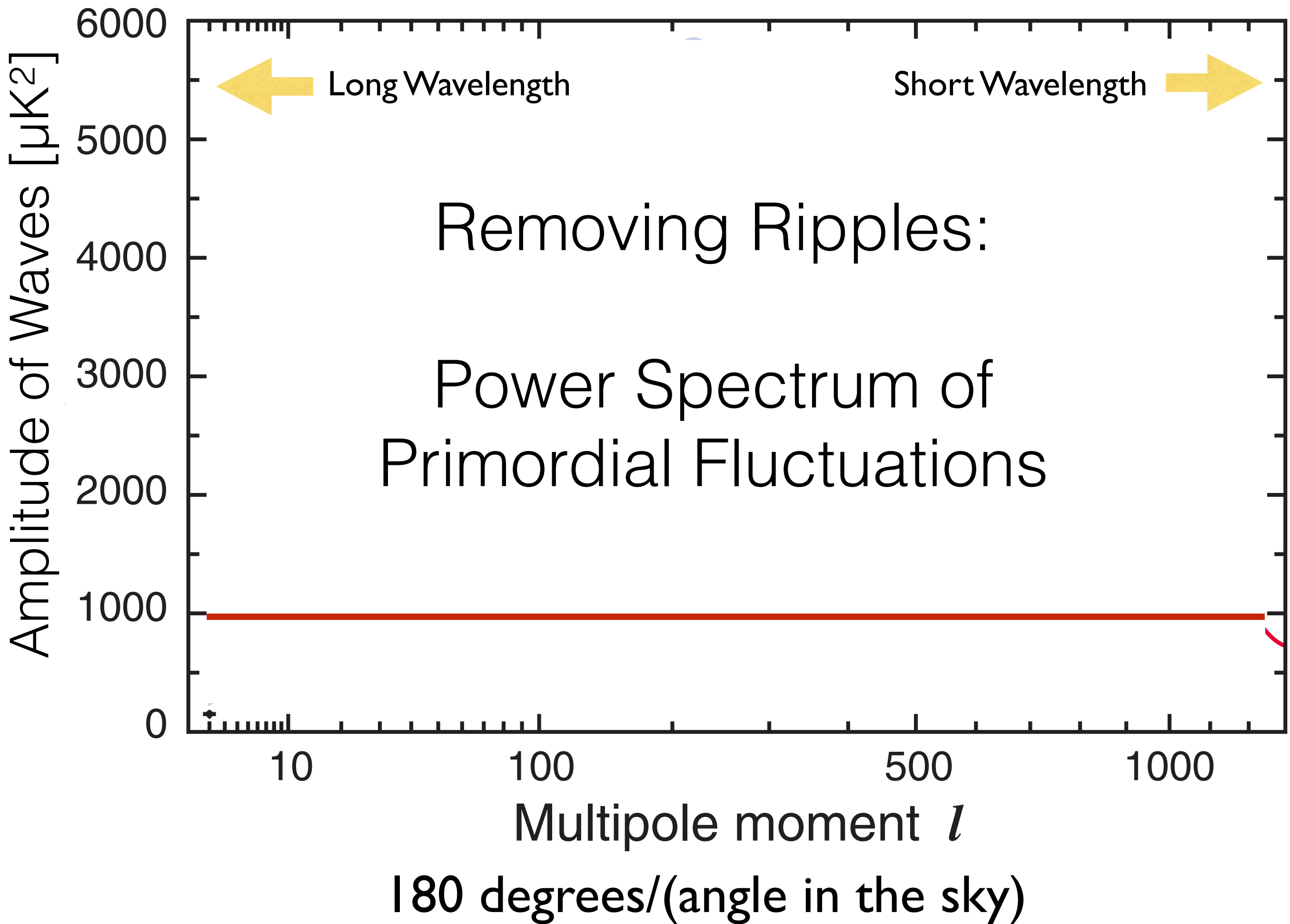
- [Energy you can borrow] x [Time you borrow] = constant
- Suppose that the distance between two points increases in proportion to **$a(t)$** [which is called the scale factor] by the expansion of the universe
- Define the “expansion rate of the universe” as

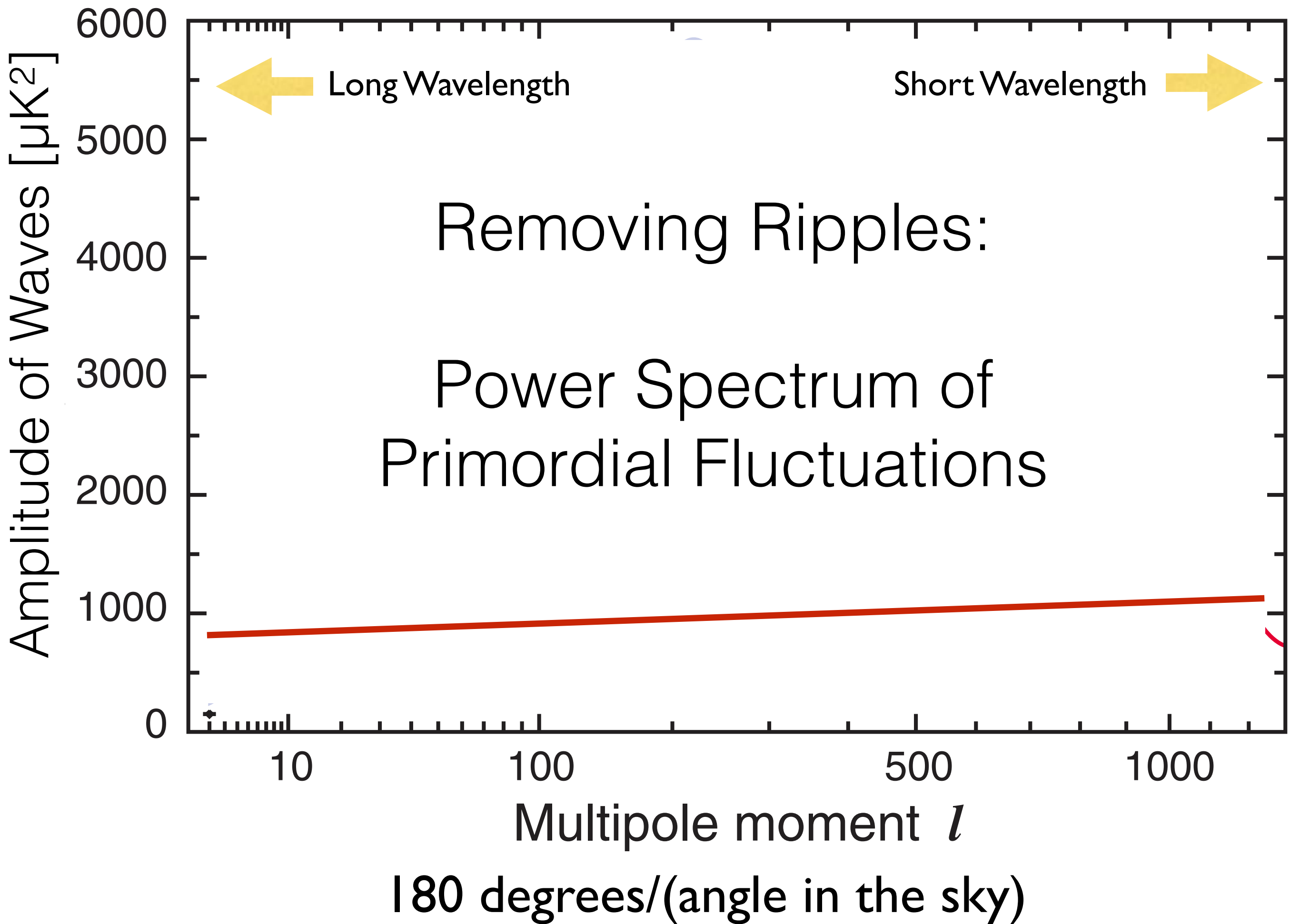
$$H \equiv \frac{\dot{a}}{a} \quad [\text{This has units of } 1/\text{time}]$$

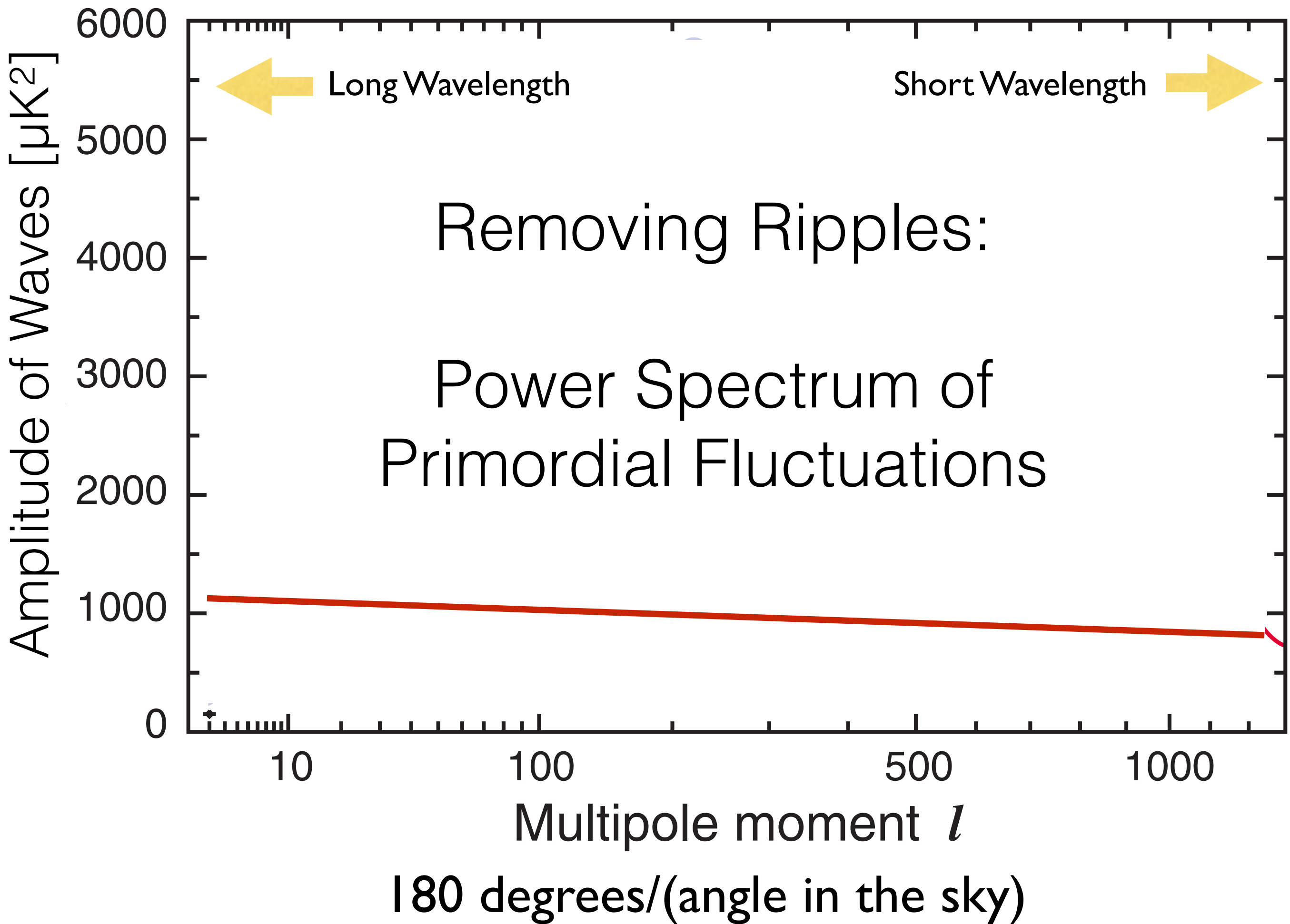
Fluctuations are proportional to H

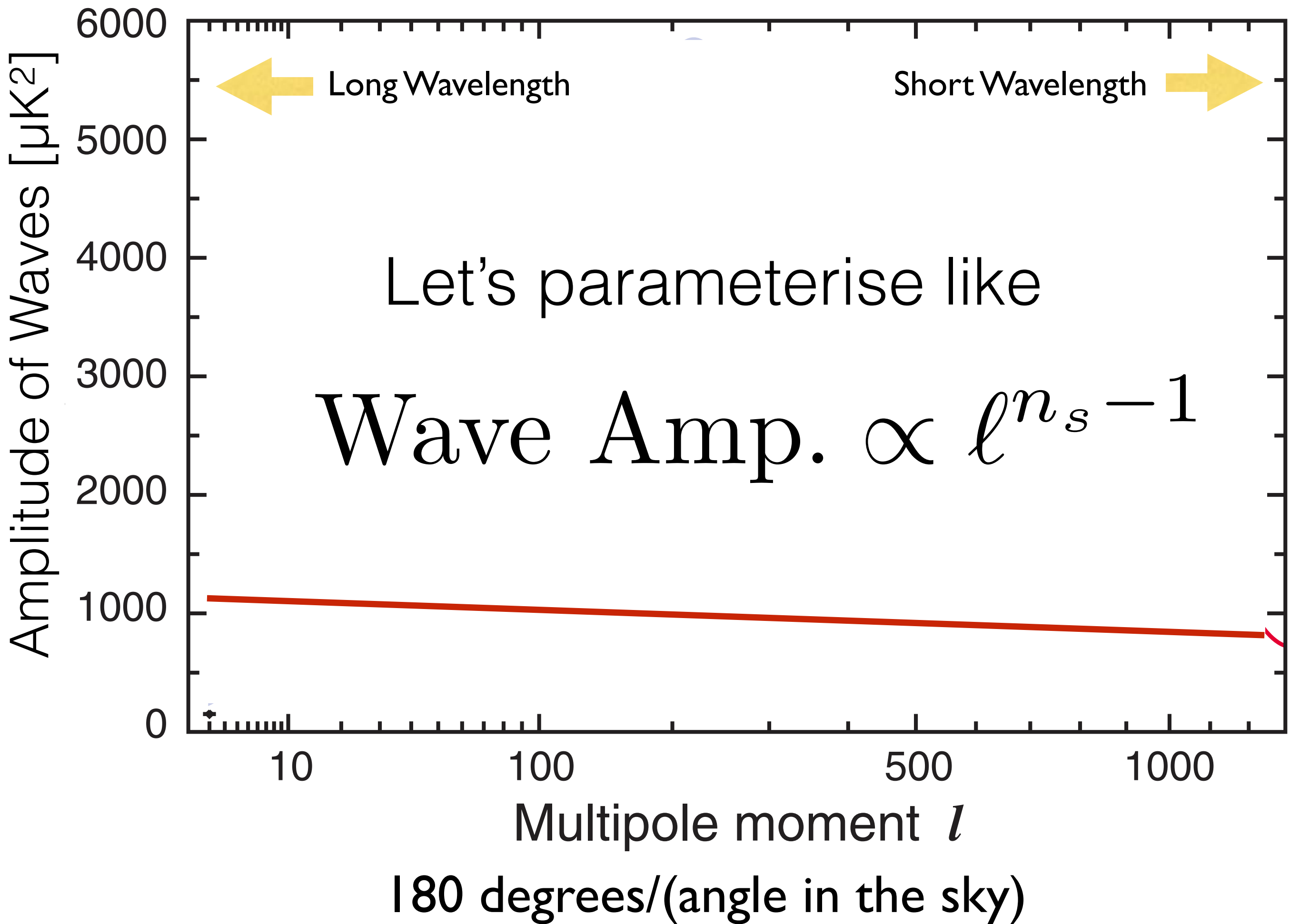
- [Energy you can borrow] x [Time you borrow] = constant
- $H \equiv \frac{\dot{a}}{a}$ [This has units of 1/time]
- Then, **both ζ and h_{ij} are proportional to H**
- Inflation occurs in 10^{-36} second - this is such a short period of time that you can borrow a lot of energy!
 H during inflation in energy units is 10^{14} GeV

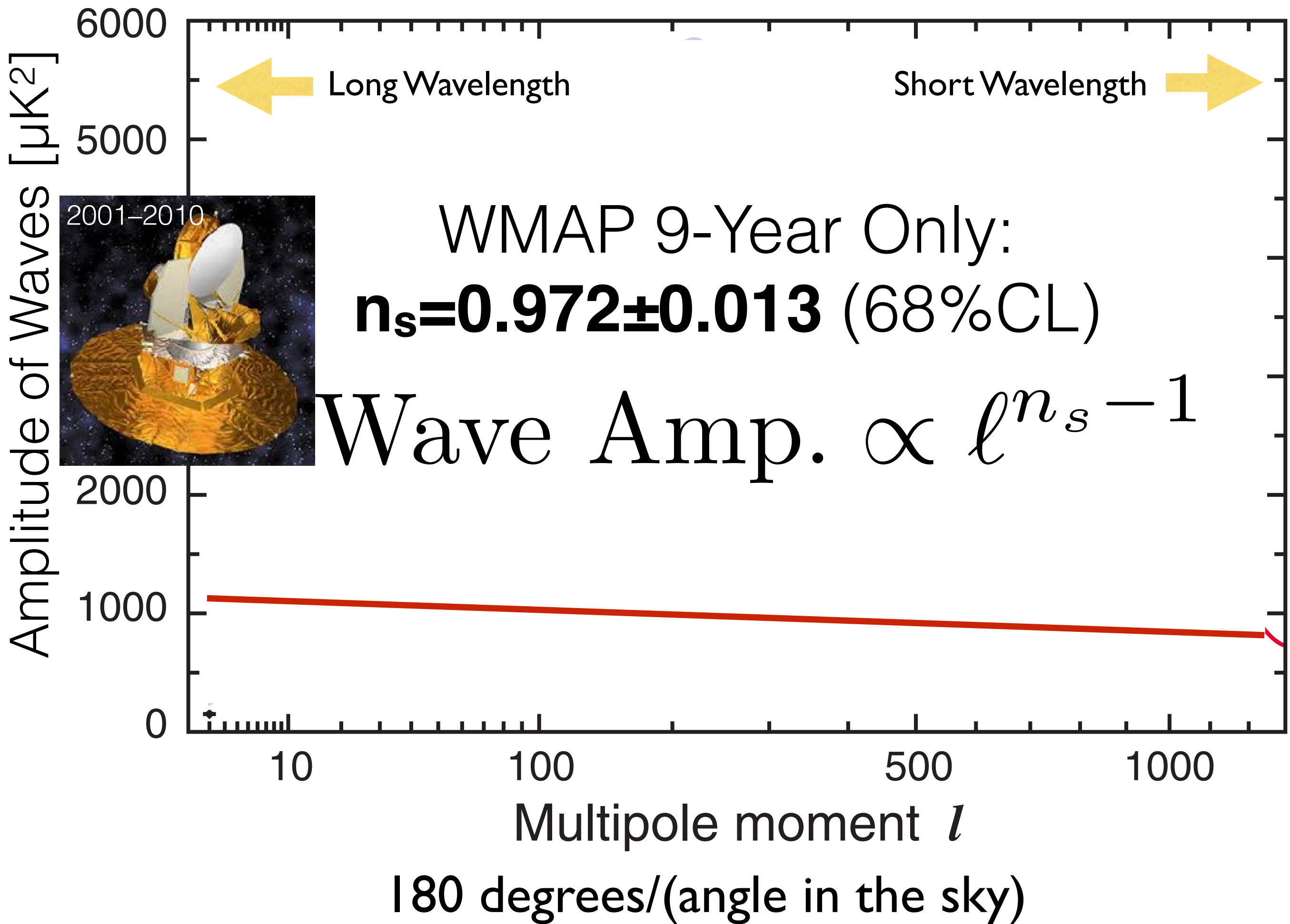


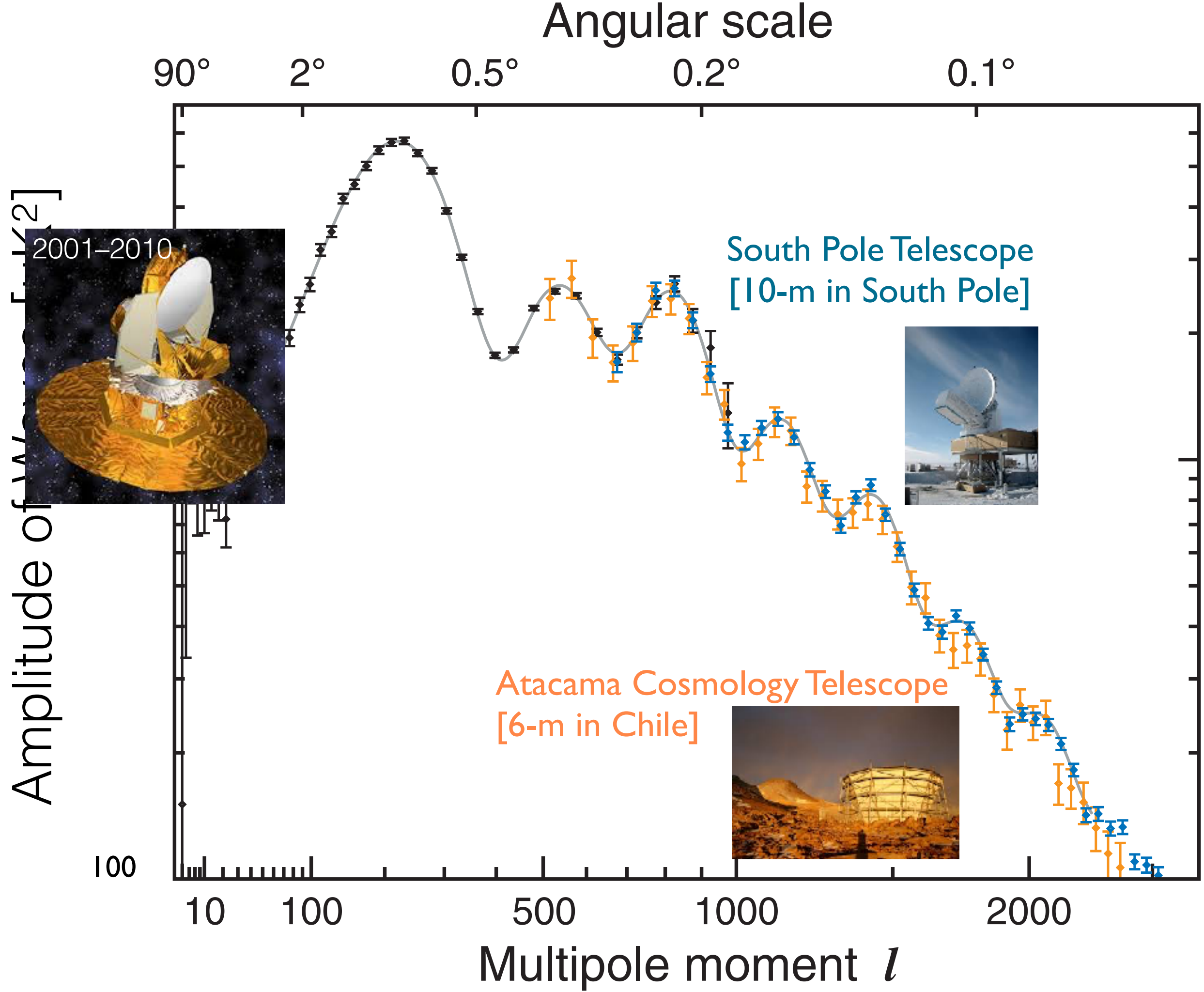


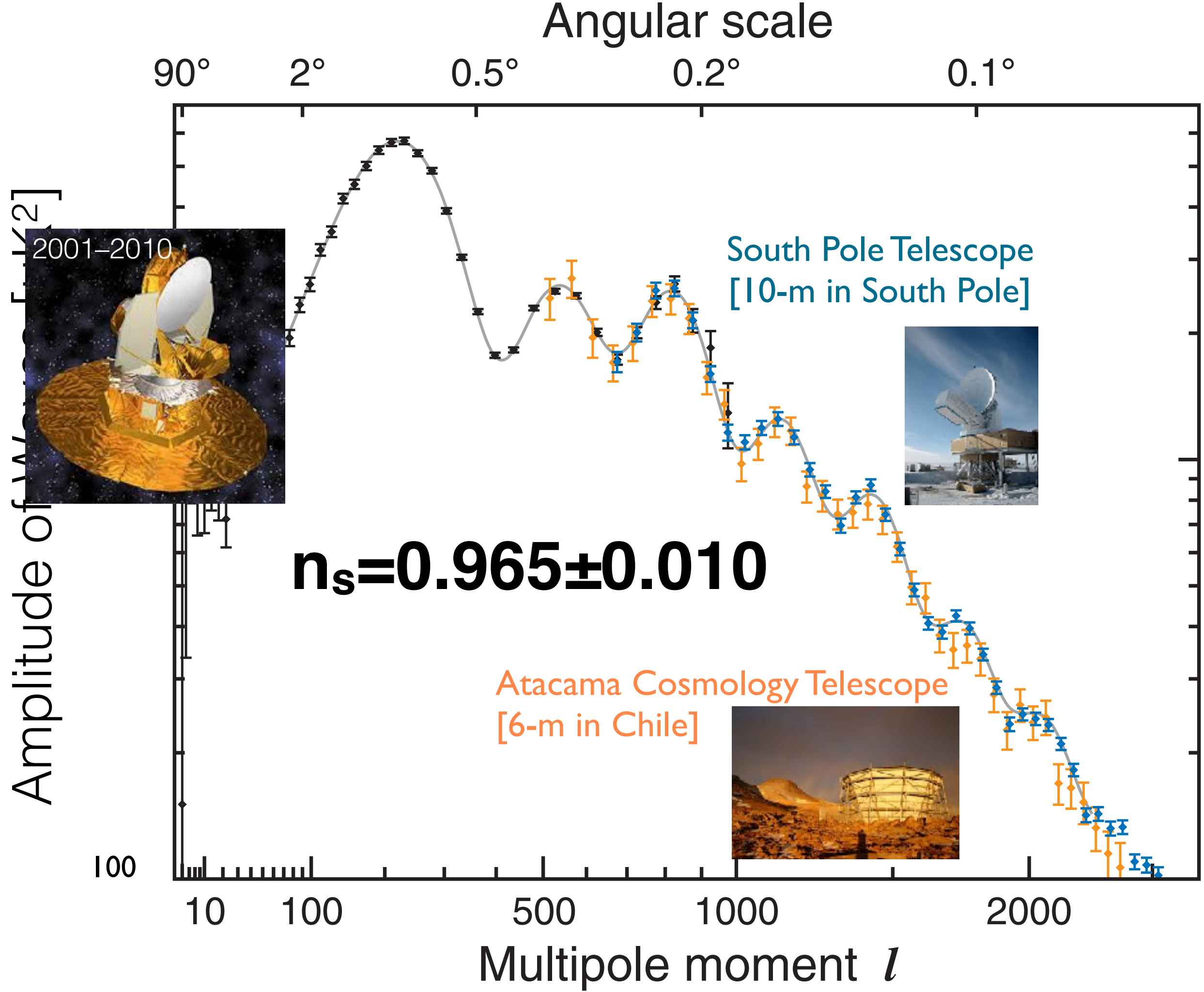


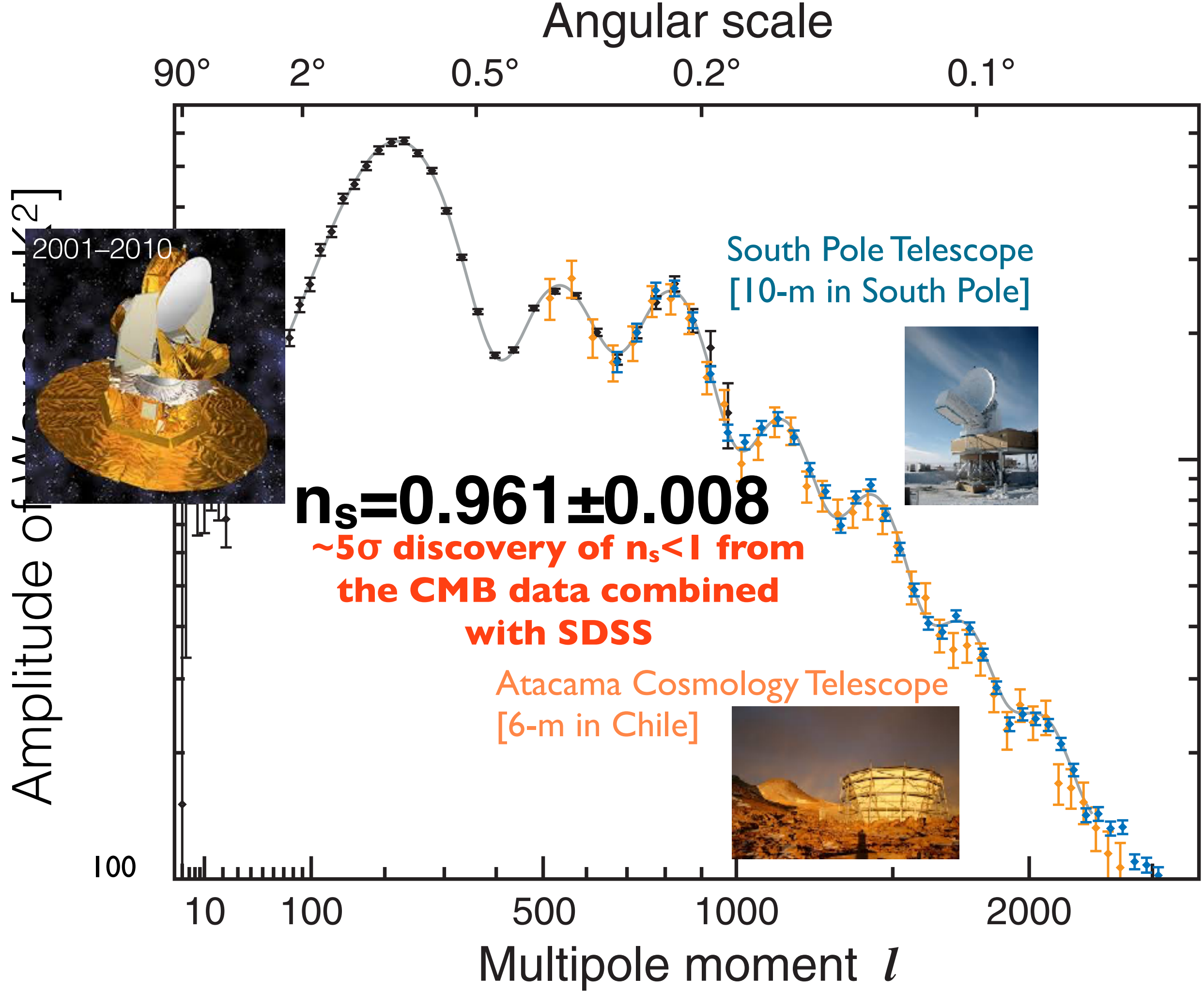








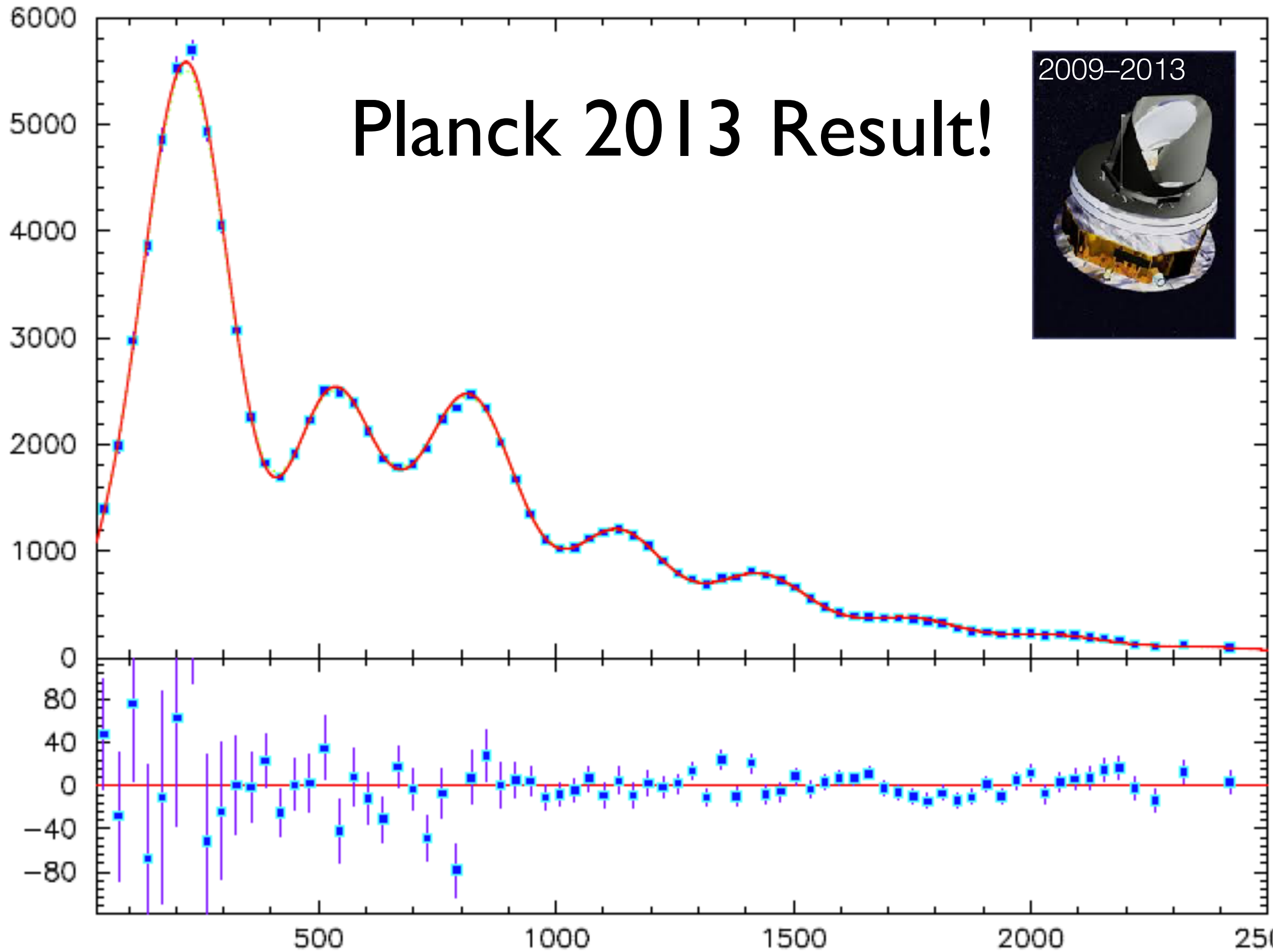
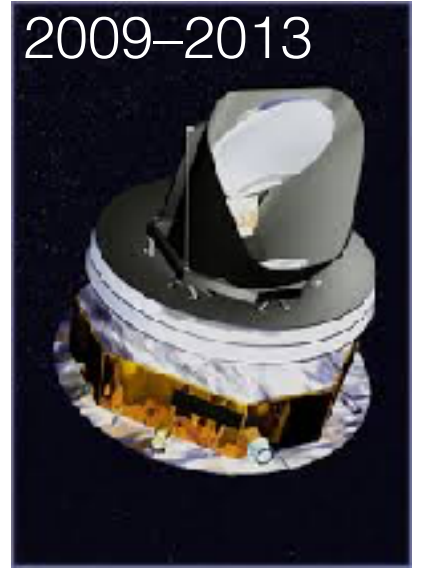




Residual Amplitude of Waves [μK^2]

Planck 2013 Result!

2009–2013



l 80 degrees/(angle in the sky)

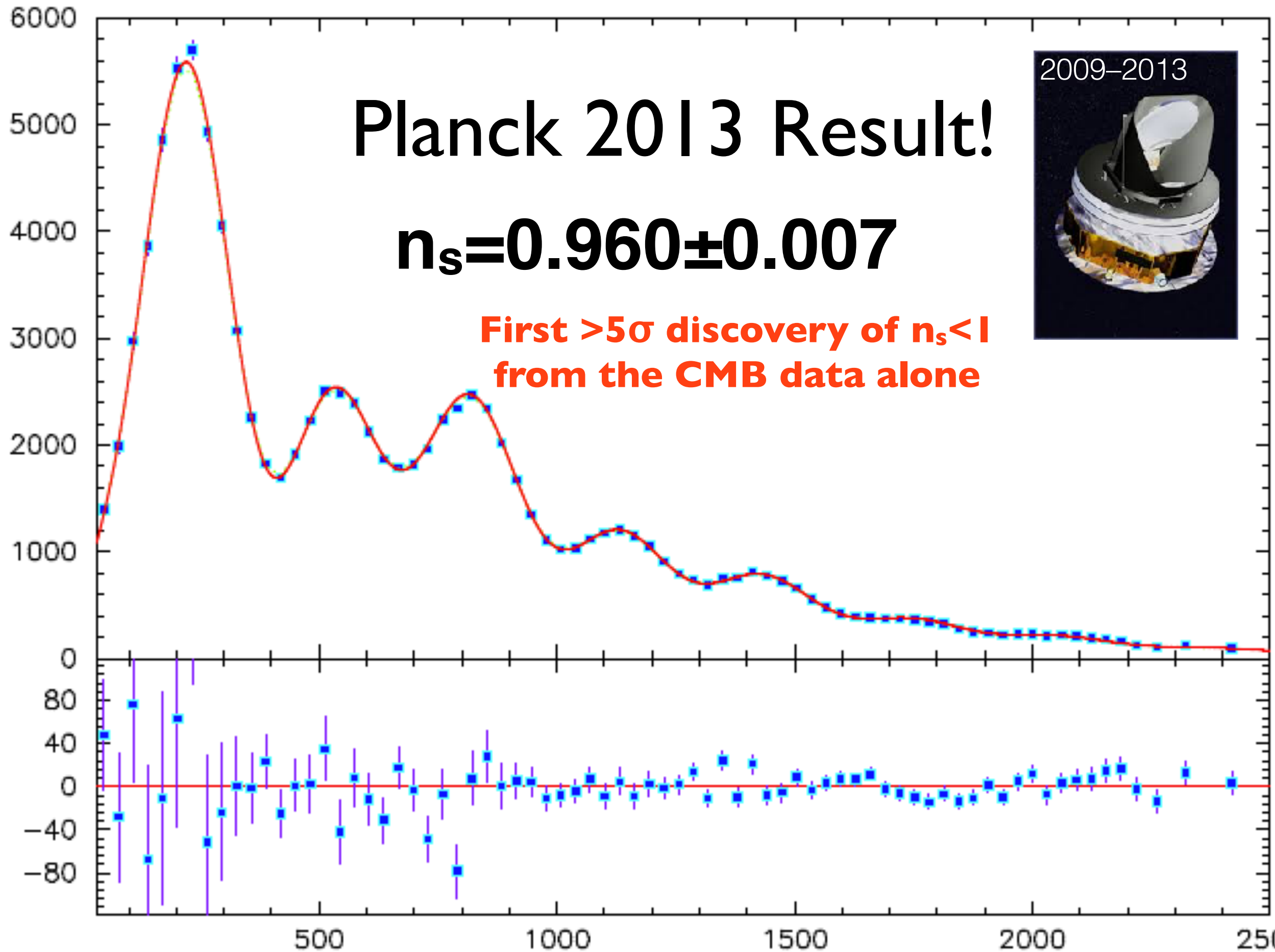
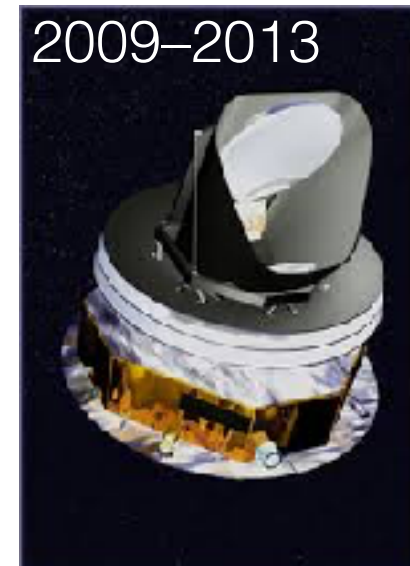
Residual Amplitude of Waves [μK^2]

Planck 2013 Result!

$$n_s = 0.960 \pm 0.007$$

**First $>5\sigma$ discovery of $n_s < 1$
from the CMB data alone**

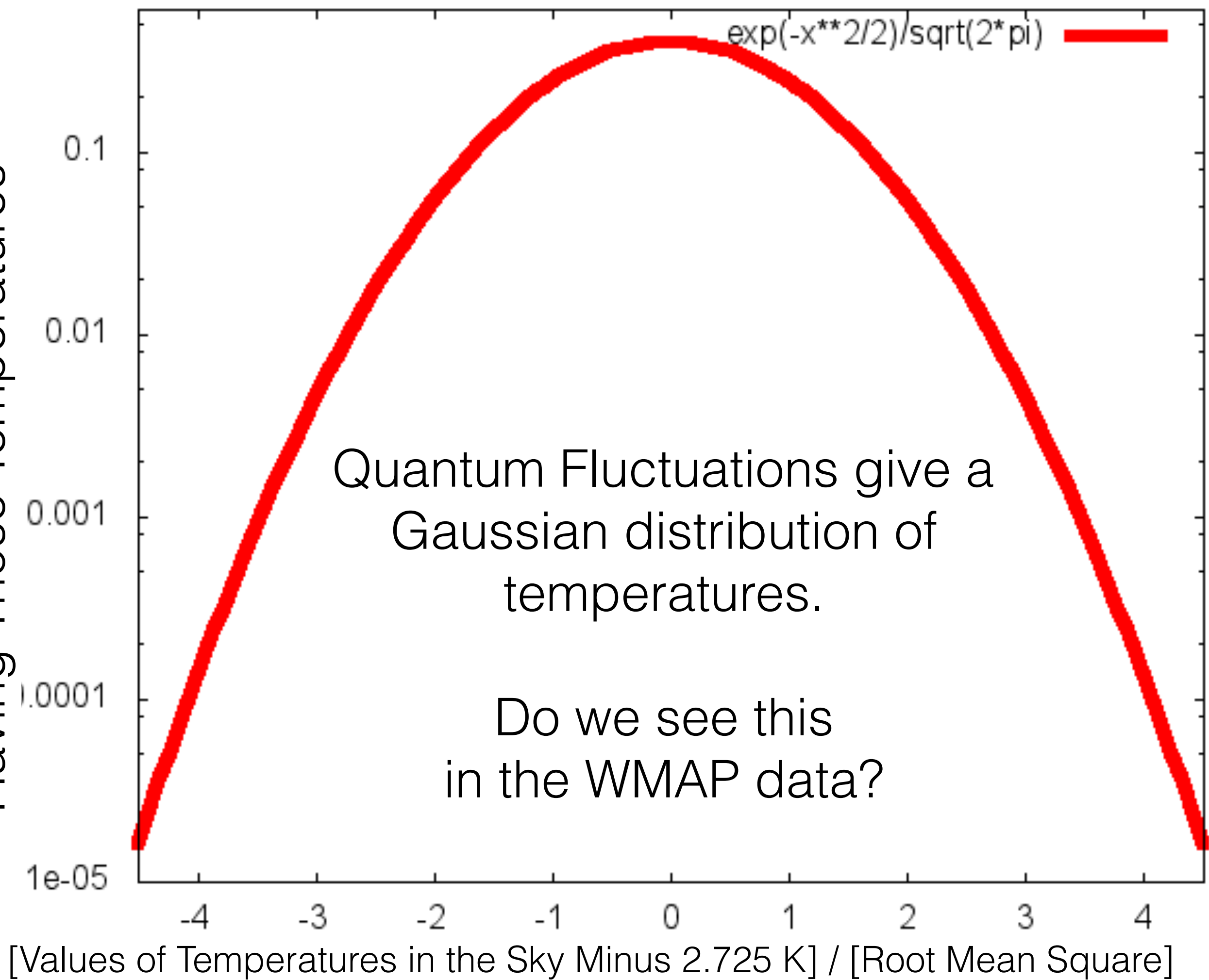
2009–2013



l 80 degrees/(angle in the sky)

How do we know that
primordial fluctuations were of
quantum mechanical origin?

Fraction of the Number of Pixels
Having Those Temperatures



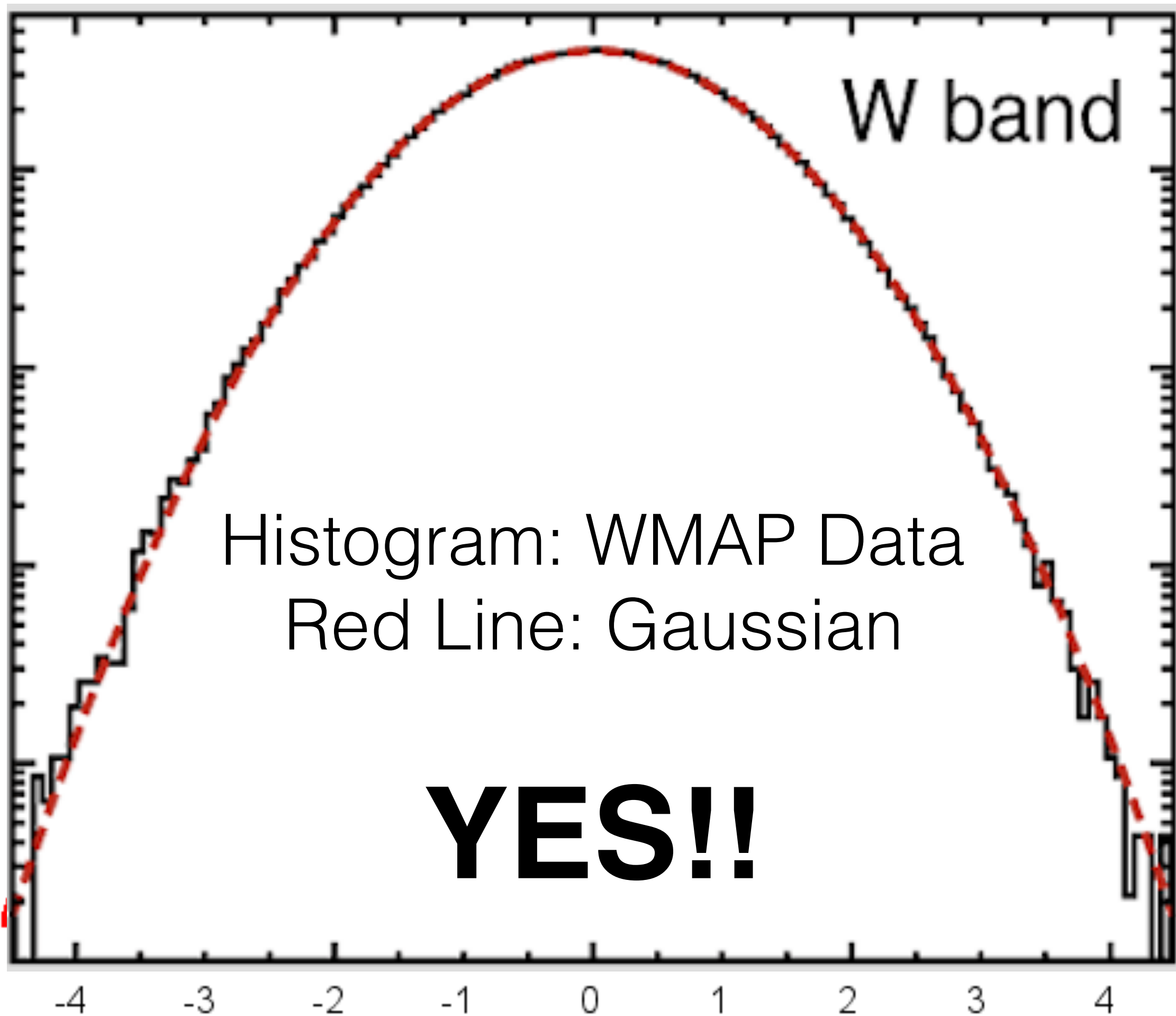
Fraction of the Number of Pixels
Having Those Temperatures

W band

Histogram: WMAP Data
Red Line: Gaussian

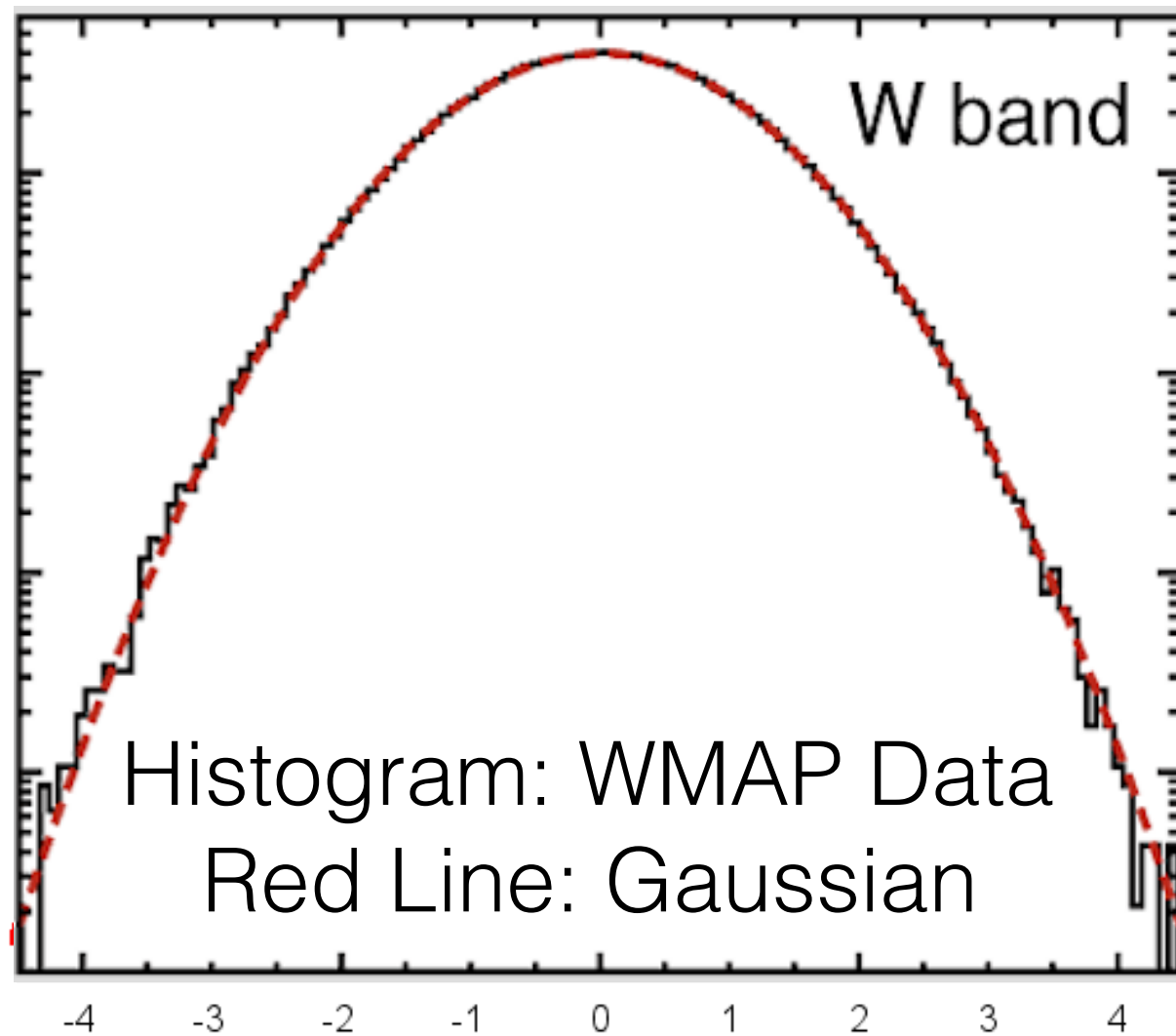
YES!!

$[\text{Values of Temperatures in the Sky Minus } 2.725 \text{ K}] / [\text{Root Mean Square}]$



Testing Gaussianity

Fraction of the Number of Pixels
Having Those Temperatures



[Values of Temperatures in the Sky Minus
2.725 K]/ [Root Mean Square]

Since a Gauss distribution is symmetric, it must yield a vanishing **3-point function**

$$\langle \delta T^3 \rangle \equiv \int_{-\infty}^{\infty} d\delta T \, P(\delta T) \delta T^3$$

More specifically, we measure this using temperatures at three different locations and average:

$$\langle \delta T(\hat{n}_1) \delta T(\hat{n}_2) \delta T(\hat{n}_3) \rangle$$

Lack of non-Gaussianity

- The WMAP data show that the distribution of temperature fluctuations of CMB is very precisely Gaussian
 - with an upper bound on a deviation of **0.2%** (95%CL)

$$\zeta(\mathbf{x}) = \zeta_{\text{gaus}}(\mathbf{x}) + \frac{3}{5} f_{\text{NL}} \zeta_{\text{gaus}}^2(\mathbf{x}) \text{ with } f_{\text{NL}} = 37 \pm 20 \text{ (68\% CL)}$$

WMAP 9-year Result

- The Planck data improved the upper bound by an order of magnitude: deviation is **<0.03%** (95%CL)

$$f_{\text{NL}} = 0.8 \pm 5.0 \text{ (68\% CL)}$$

Planck 2015 Result

CMB Research: Next Frontier

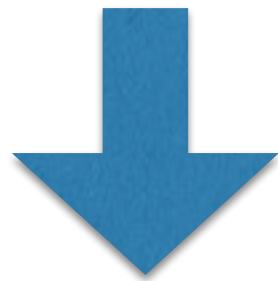
Primordial Gravitational Waves

*Extraordinary claims require extraordinary evidence.
The same quantum fluctuations could also generate
gravitational waves, and we wish to find them*

Measuring GW

- GW changes the distances between two points

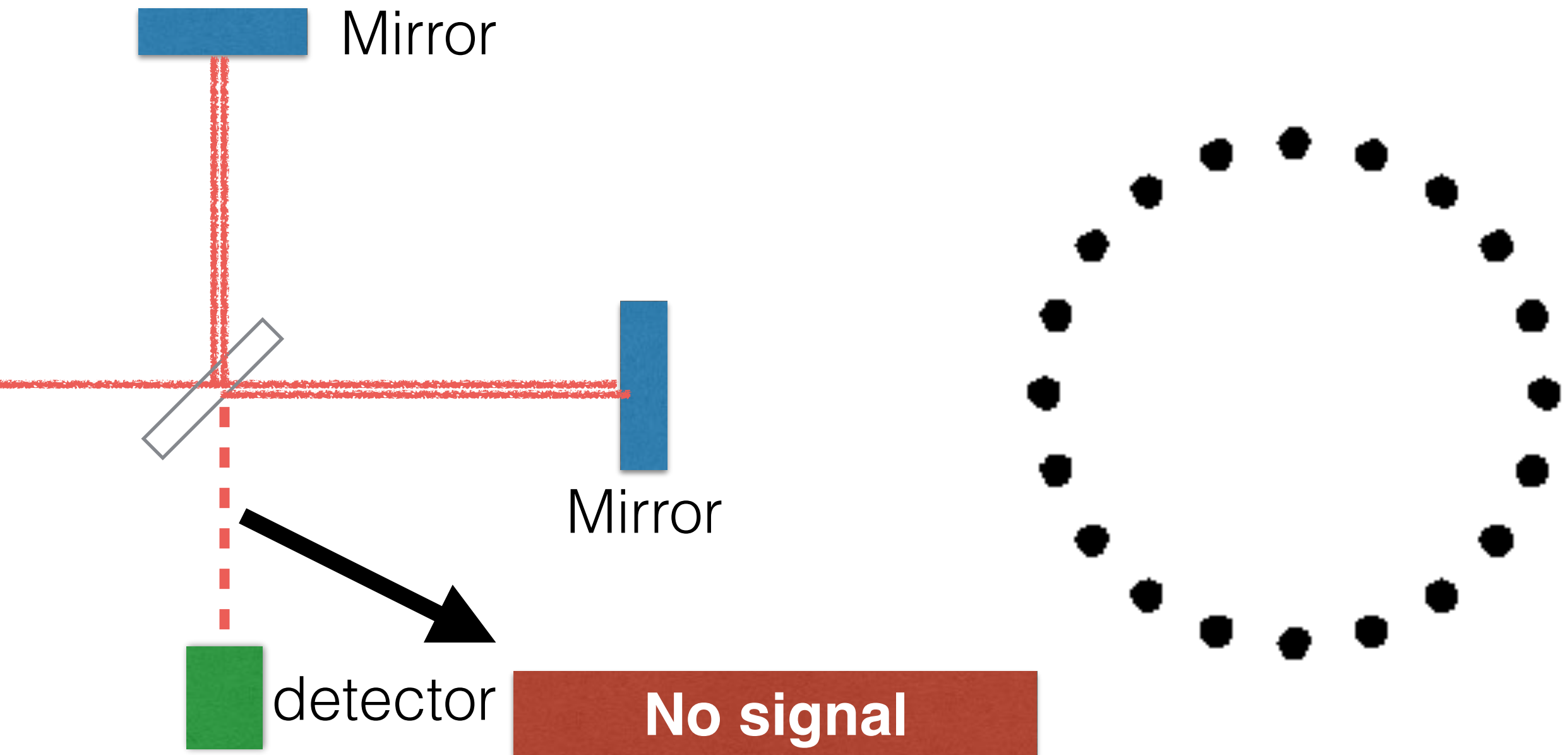
$$d\ell^2 = d\mathbf{x}^2 = \sum_{ij} \delta_{ij} dx^i dx^j$$



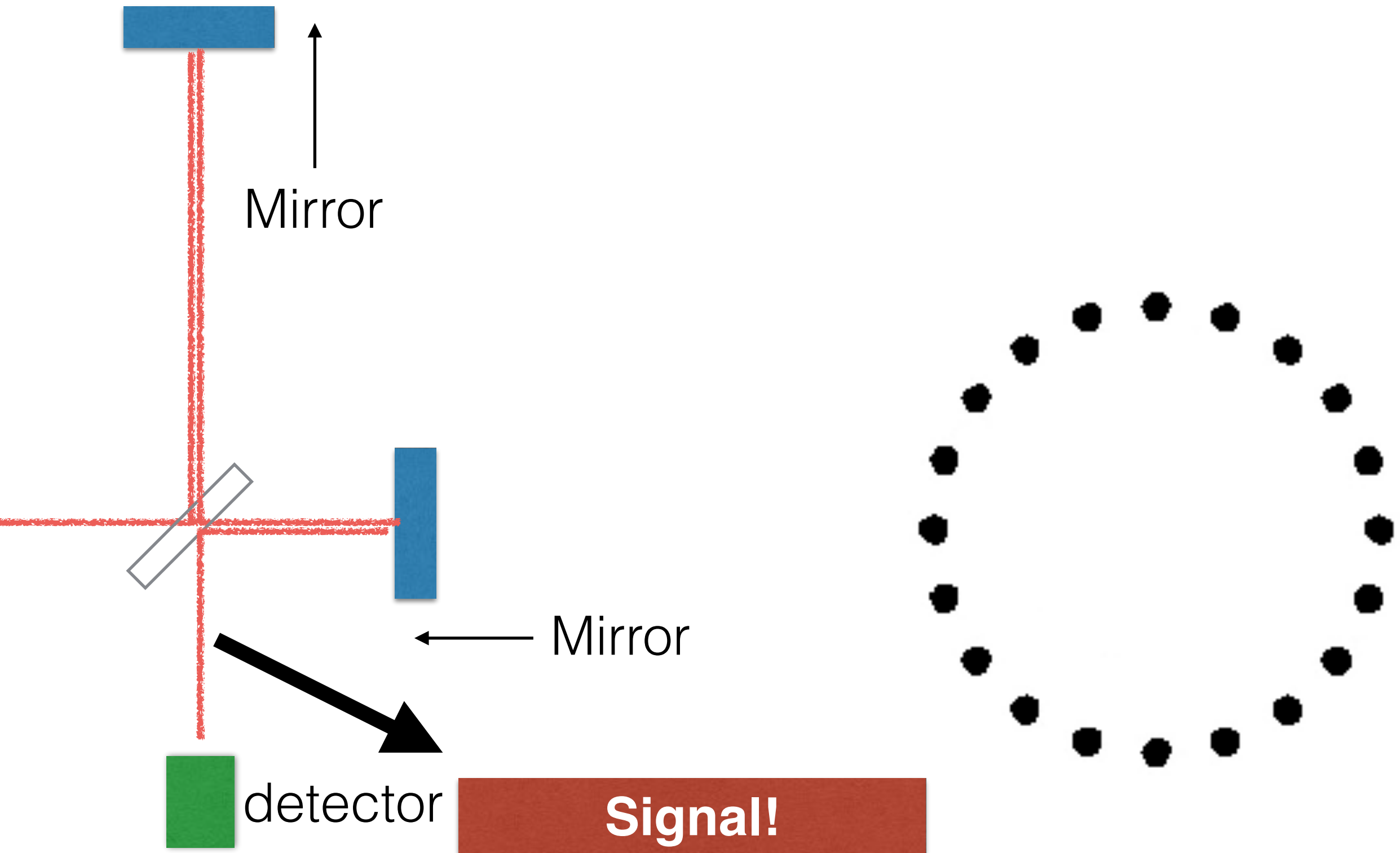
$$d\ell^2 = \sum_{ij} (\delta_{ij} + \textcolor{red}{h_{ij}}) dx^i dx^j$$



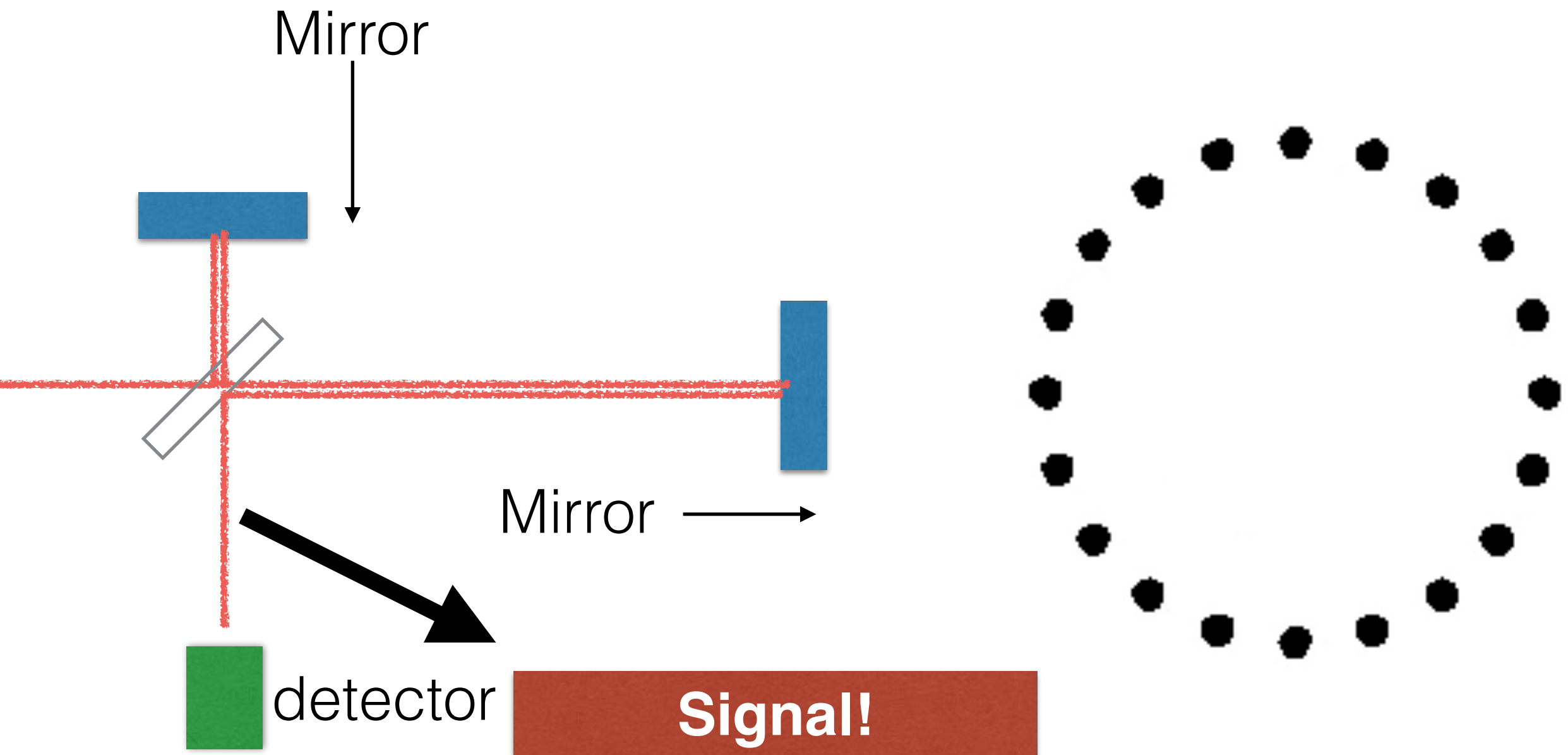
Laser Interferometer



Laser Interferometer



Laser Interferometer



LIGO detected GW from binary blackholes, with the wavelength of thousands of kilometres

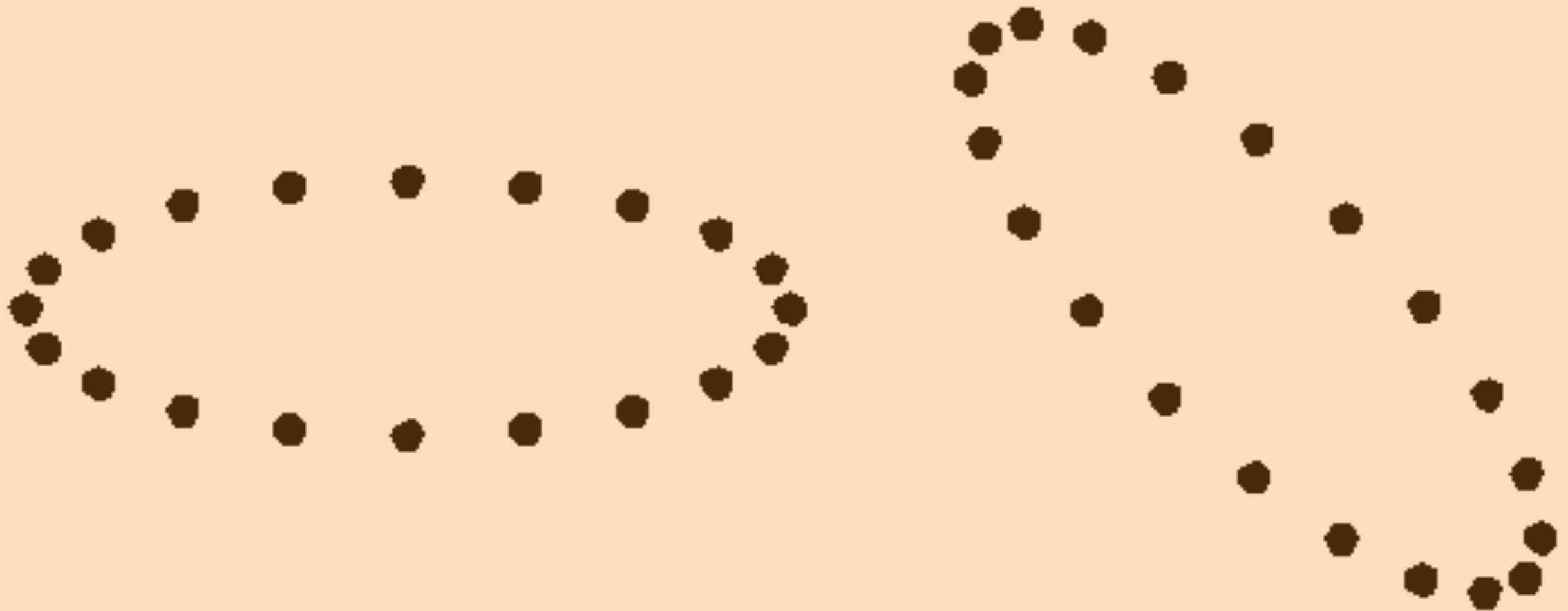
But, the primordial GW affecting the CMB has a wavelength of **billions of light-years!!** How do we find it?

Detecting GW by CMB

Isotropic electro-magnetic fields

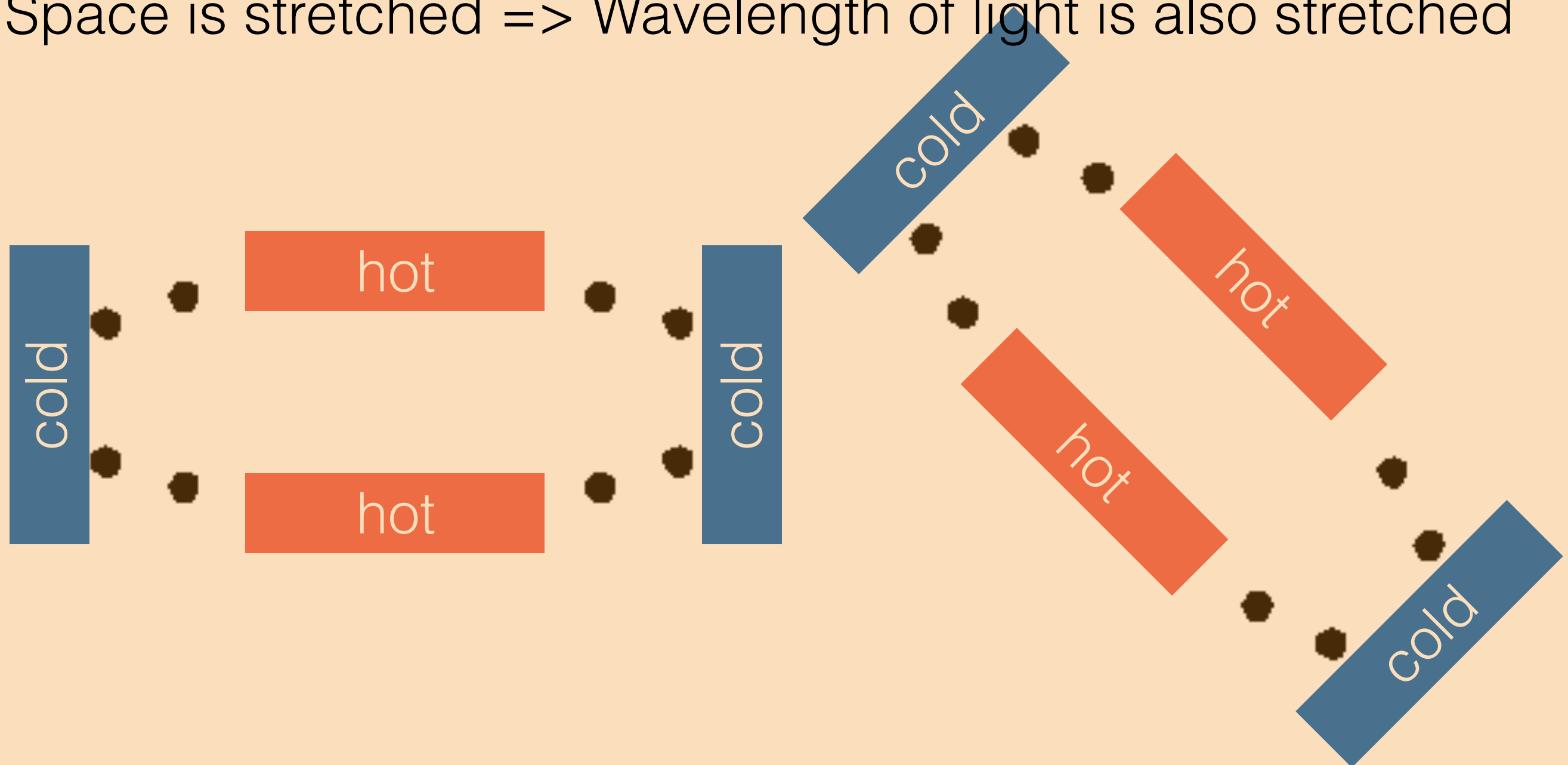
Detecting GW by CMB

GW propagating in isotropic electro-magnetic fields



Detecting GW by CMB

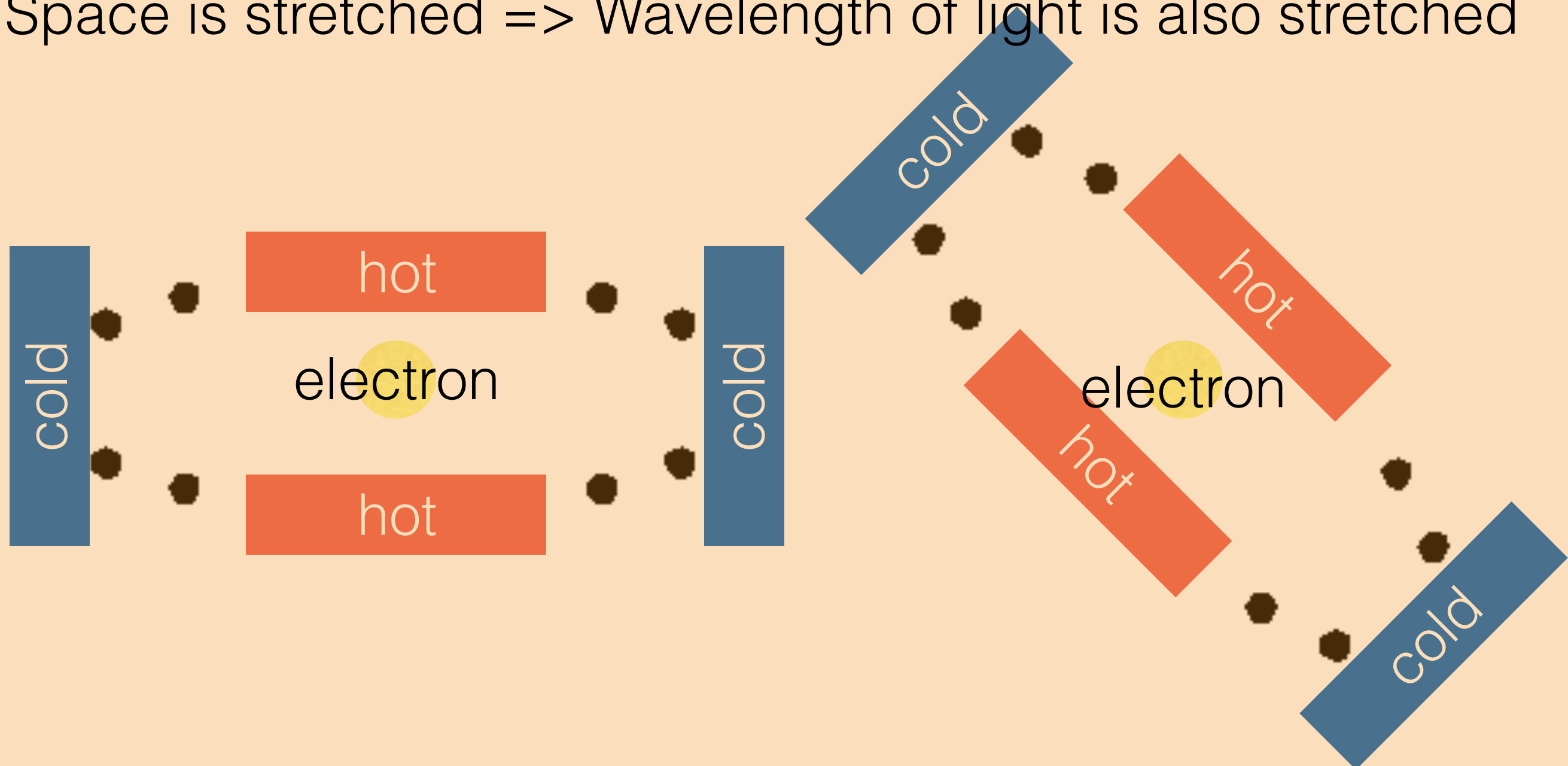
Space is stretched => Wavelength of light is also stretched



Detecting GW by CMB

Polarisation

Space is stretched => Wavelength of light is also stretched



Detecting GW by CMB

Polarisation

Space is stretched => Wavelength of light is also stretched

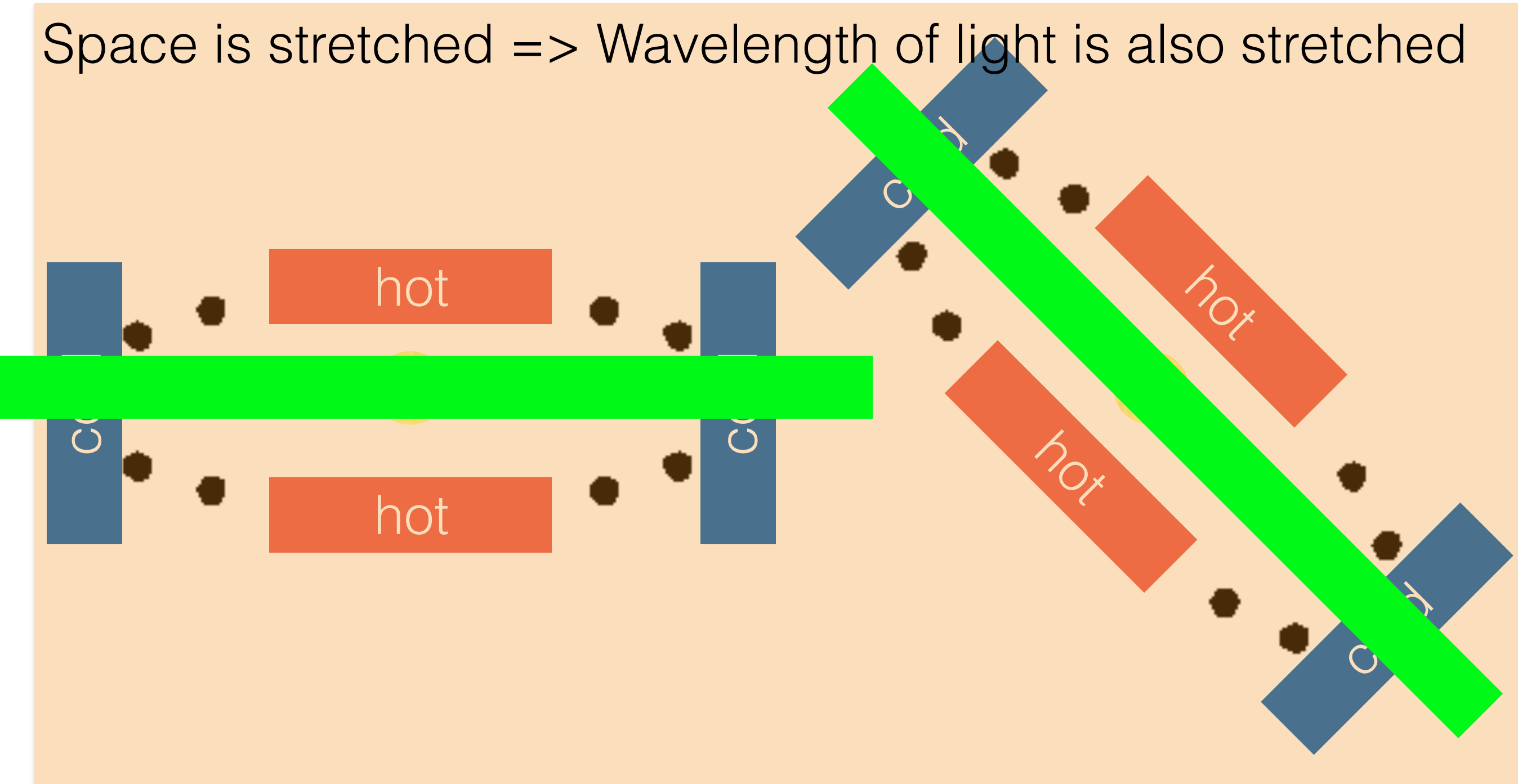


Photo Credit: TALEX



horizontally polarised

Photo Credit: TALEX



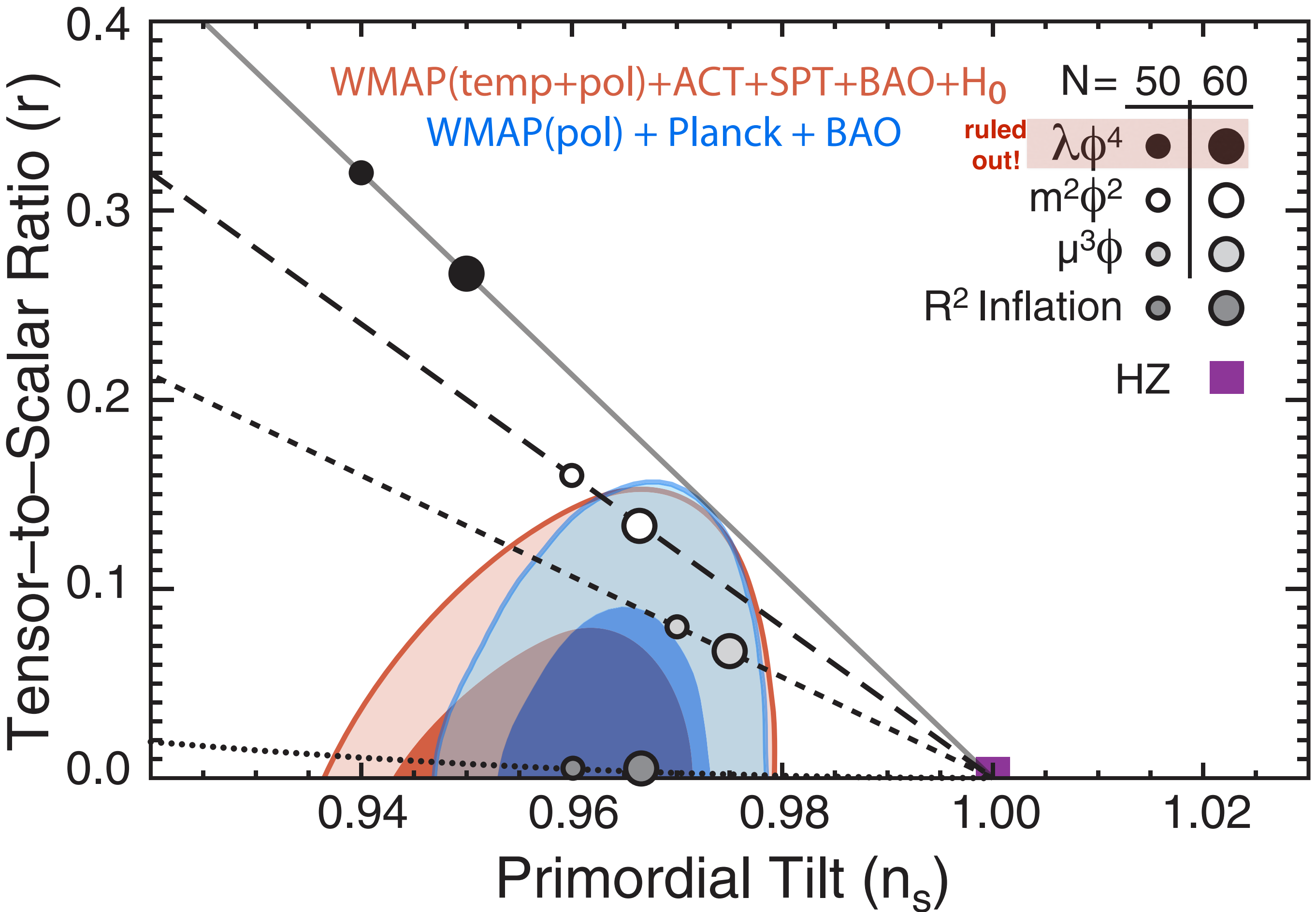
If polarisation from GW is found...

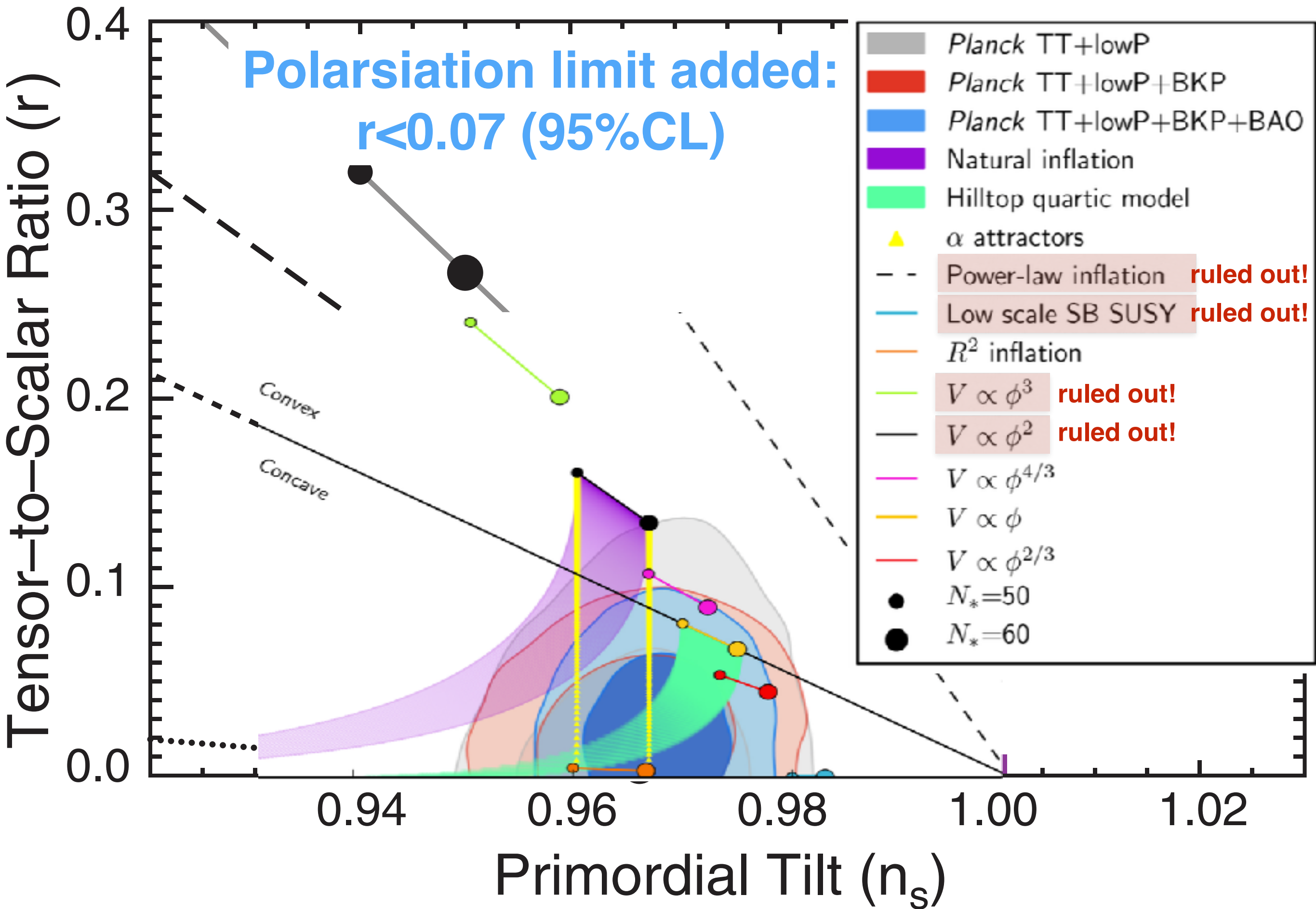
- Then what?
- The next step is to nail the specific model of inflation

Tensor-to-scalar Ratio

$$r \equiv \frac{\langle h_{ij} h^{ij} \rangle}{\langle \zeta^2 \rangle}$$

- We really want to find this quantity!
The current upper bound: $r < 0.07$





March 17, 2014

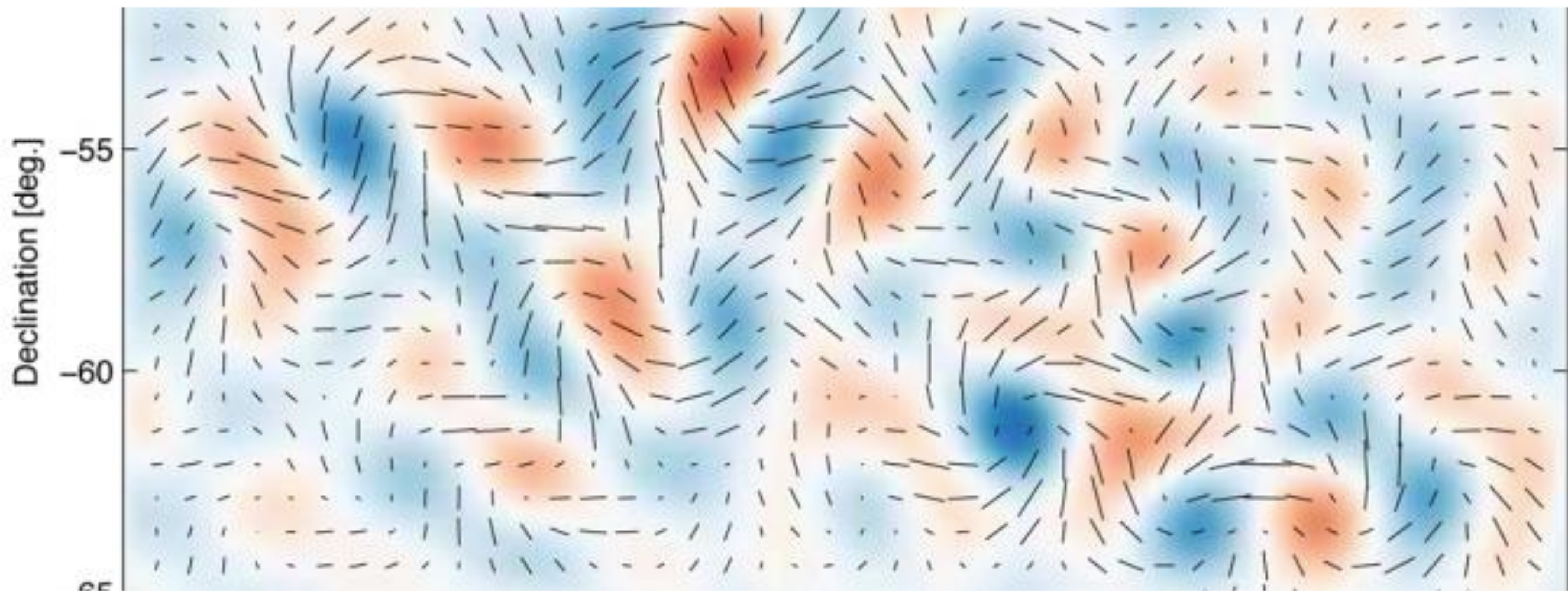
BICEP2's announcement



First Direct Evidence of Cosmic Inflation

Release No.: 2014-05

For Release: Monday, March 17, 2014 - 10:45am



Cambridge, MA - Almost 14 billion years ago, the universe we inhabit burst into existence in an extraordinary event that initiated the Big Bang. In the first fleeting fraction of a second, the universe expanded exponentially, stretching far beyond the view of our best telescopes. All this, of course, was just theory.

SPACE & COSMOS

The New York Times

Space Ripples Reveal Big Bang's Smoking Gun

By DENNIS OVERBYE MARCH 17, 2014

BBC

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17 March 2014 Last updated at 14:46 GMT

Share



Cosmic inflation: 'Spectacular' discovery hailed

By Jonathan Amos

Science correspondent, BBC News



Cambridge, MA - Almost 14 billion years ago, a flash of light and energy that initiated the Big Bang. In the far beyond the view of our best tel

Süddeutsche.de

Wissen

Politik Panorama Kultur Wirtschaft Sport München Bayern Digital Auto Reise Video

Home > Wissen > Urknall > Urknall - Gravitationswellen belegen inflationäres Universum

Süddeutsche.de als Startseite einrichten

Hir

17. März 2014, 17:34 Gravitationswellen

Signale aus der Geburtsstunde des Universums

Von Patrick Illinger

January 30, 2015

Joint Analysis of BICEP2 data and Planck data

Speck of Interstellar Dust Obscures Glimpse of Big Bang

By DENNIS OVERBYE JAN. 30, 2015

BBC

News Sport Weather Earth Future Shop

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30 January 2015 Last updated at 20:54 GMT



Cosmic inflation: New study says BICEP claim was wrong

By Jonathan Amos
Science correspondent, BBC News

Süddeutsche.de

Wissen

Politik Panorama Kultur Wirtschaft Sport München Bayern Digital Auto Reise Video

Home > Wissen > Kosmologie - Urknall-Forscher gestehen Irrtum ein

[Süddeutsche.de als Startseite einrichten](#)

Hir

1. Februar 2015, 22:19 Kosmologie

Urknall-Forscher gestehen Irrtum ein

Von Marlene Weiß

Current Situation

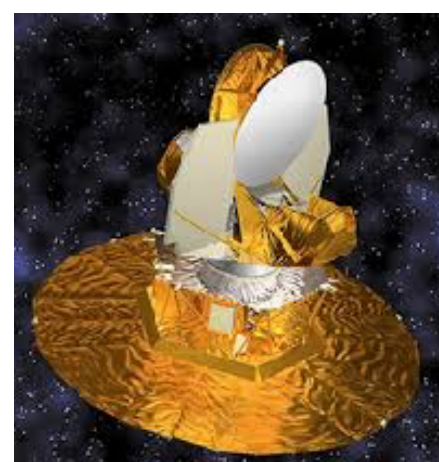
- Planck shows the evidence that the detected signal is not cosmological, but is due to dust
- No strong evidence that the detected signal is cosmological



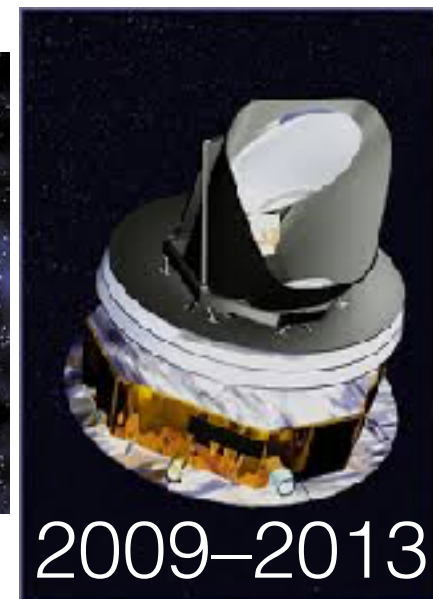
The search continues!!



1989–1993



2001–2010



2009–2013



202X–

JAXA

+ possible participations
from USA, Canada,
Europe

LiteBIRD

2025– [proposed]



Target: $\delta r < 0.001$

JAXA

+ possible participations
from USA, Canada,
Europe

LiteBIRD

2025– [proposed]



**Polarisation satellite dedicated to
measure CMB polarisation from
primordial GW, with a few thousand
super-conducting detectors in space**

JAXA

+ possible participations
from USA, Canada,
Europe

LiteBIRD

2025– [proposed]

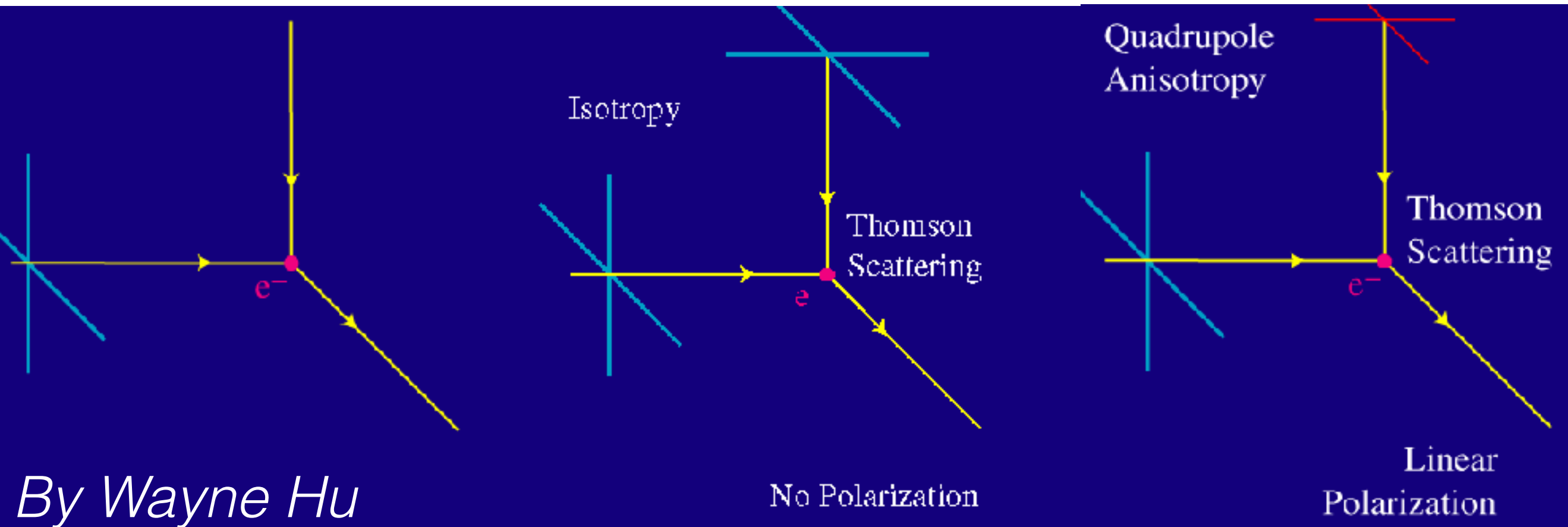


Down-selected by JAXA as
one of the two missions
competing for a launch in mid 2020's

Conclusion

- The WMAP and Planck's temperature data provide **strong evidence for the quantum origin of structures in the universe**
- The next goal: unambiguous measurement of polarisation from gravitational waves
- **LiteBIRD** proposal: a CMB polarisation satellite in mid 2020's

Physics of CMB Polarisation



- Necessary and sufficient conditions for generating polarisation in CMB:
 - Thomson scattering
 - Quadrupolar temperature anisotropy around an electron