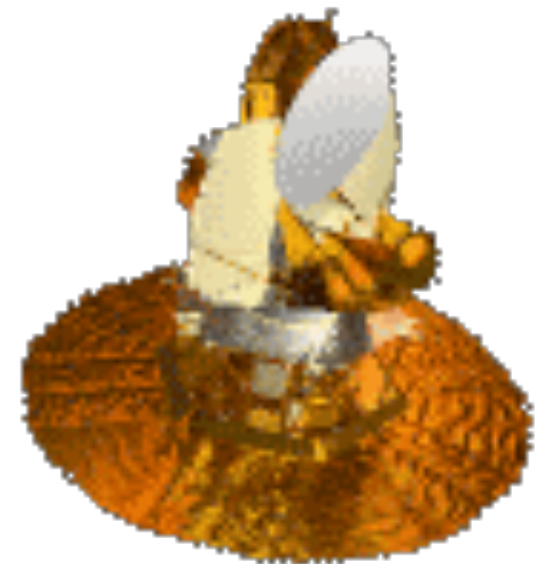


# Critical Tests of Theory of the Early Universe using the Cosmic Microwave Background

Eiichiro Komatsu, Max-Planck-Institut für Astrophysik  
Physikalisches Kolloquium, Universität Tübingen  
Mai 23, 2017



# Breakthrough in Cosmological Research

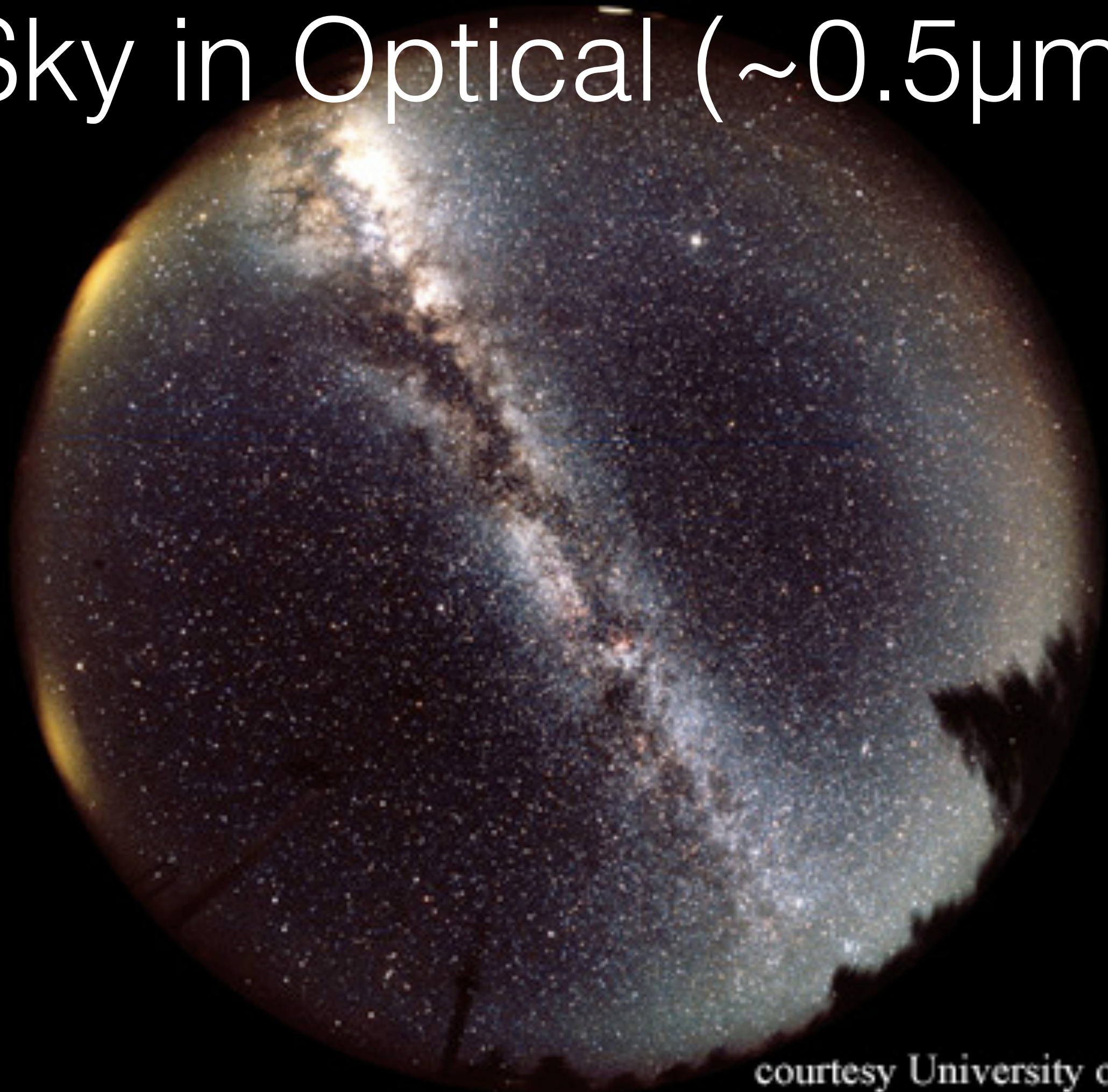
- We can actually **see** the physical condition of the universe when it was very young



*From "Cosmic Voyage"*



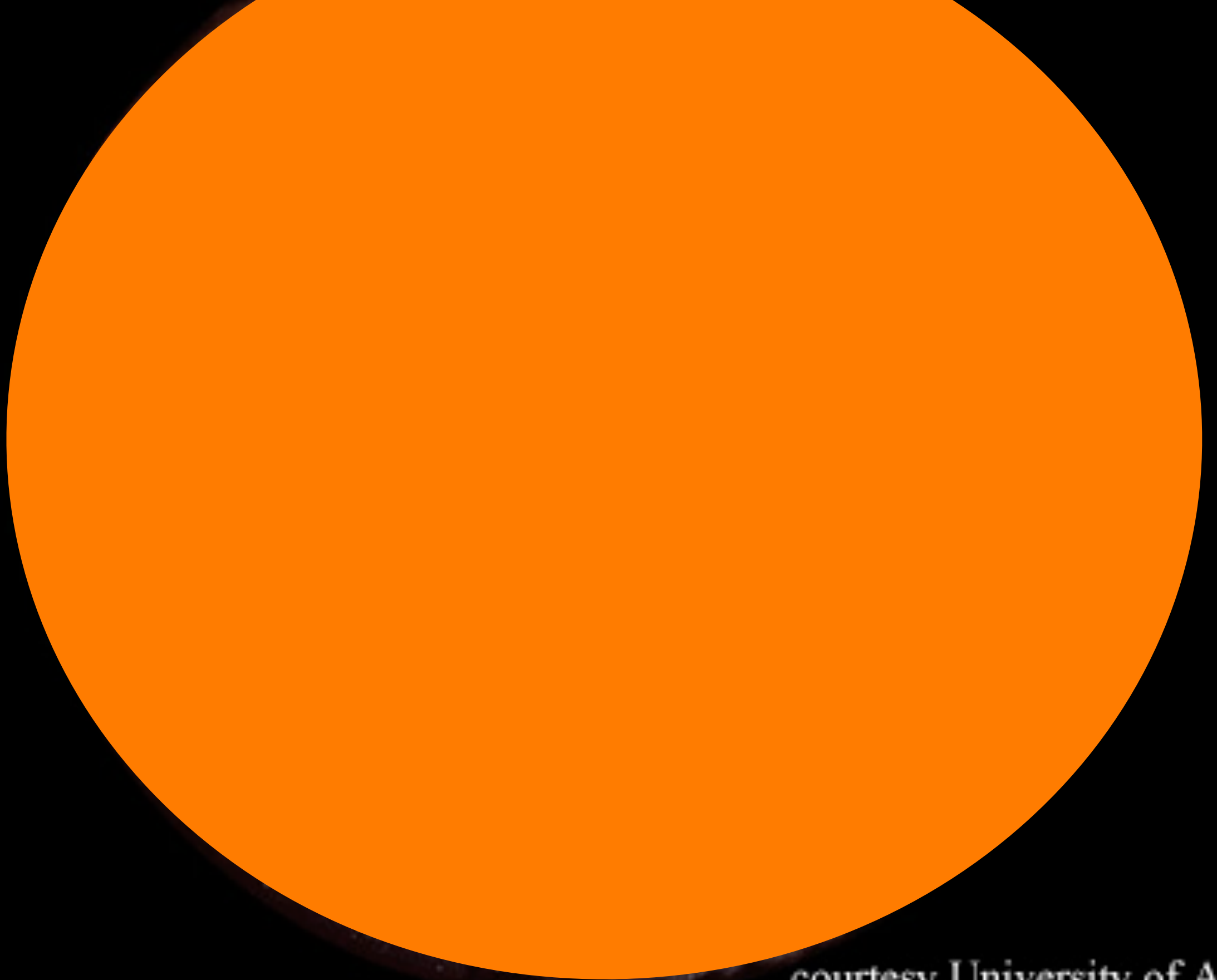
# Sky in Optical ( $\sim 0.5\mu\text{m}$ )



courtesy University of Arizona



# Sky in Microwave ( $\sim 1\text{mm}$ )



courtesy University of Arizona

# Sky in Microwave ( $\sim 1\text{mm}$ )

*Light from the fireball Universe  
filling our sky (2.7K)*

**The Cosmic Microwave  
Background (CMB)**

**410 photons**  
per  
cubic centimeter!!

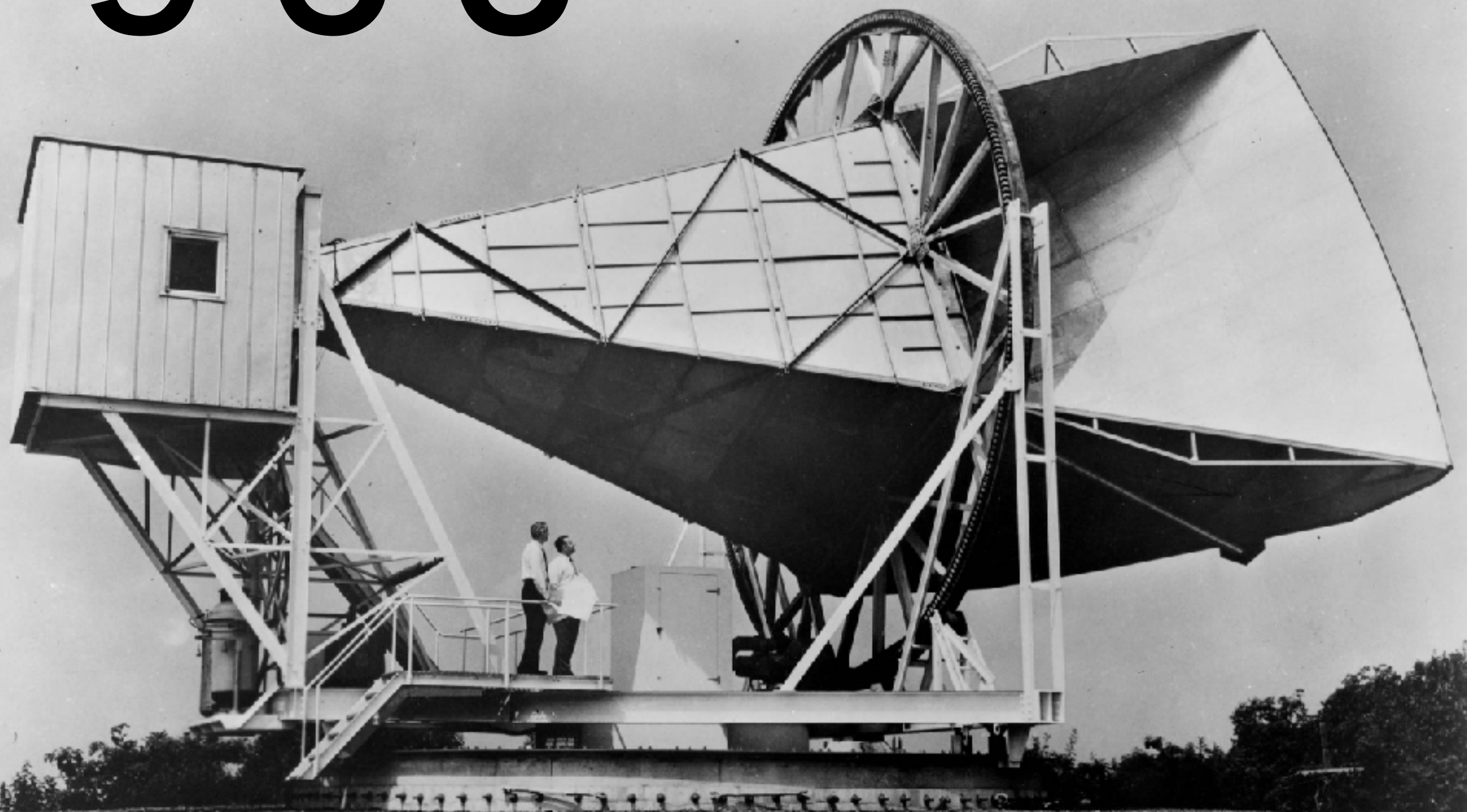


A photograph of Prof. Hiranya Peiris, a woman with long dark hair, wearing a black cardigan over a black top with a colorful patterned collar. She is holding a vintage, light-colored television set with a handle. The TV screen displays a blue, grainy, noisy pattern. The background is dark, and a wooden ledge is visible at the bottom.

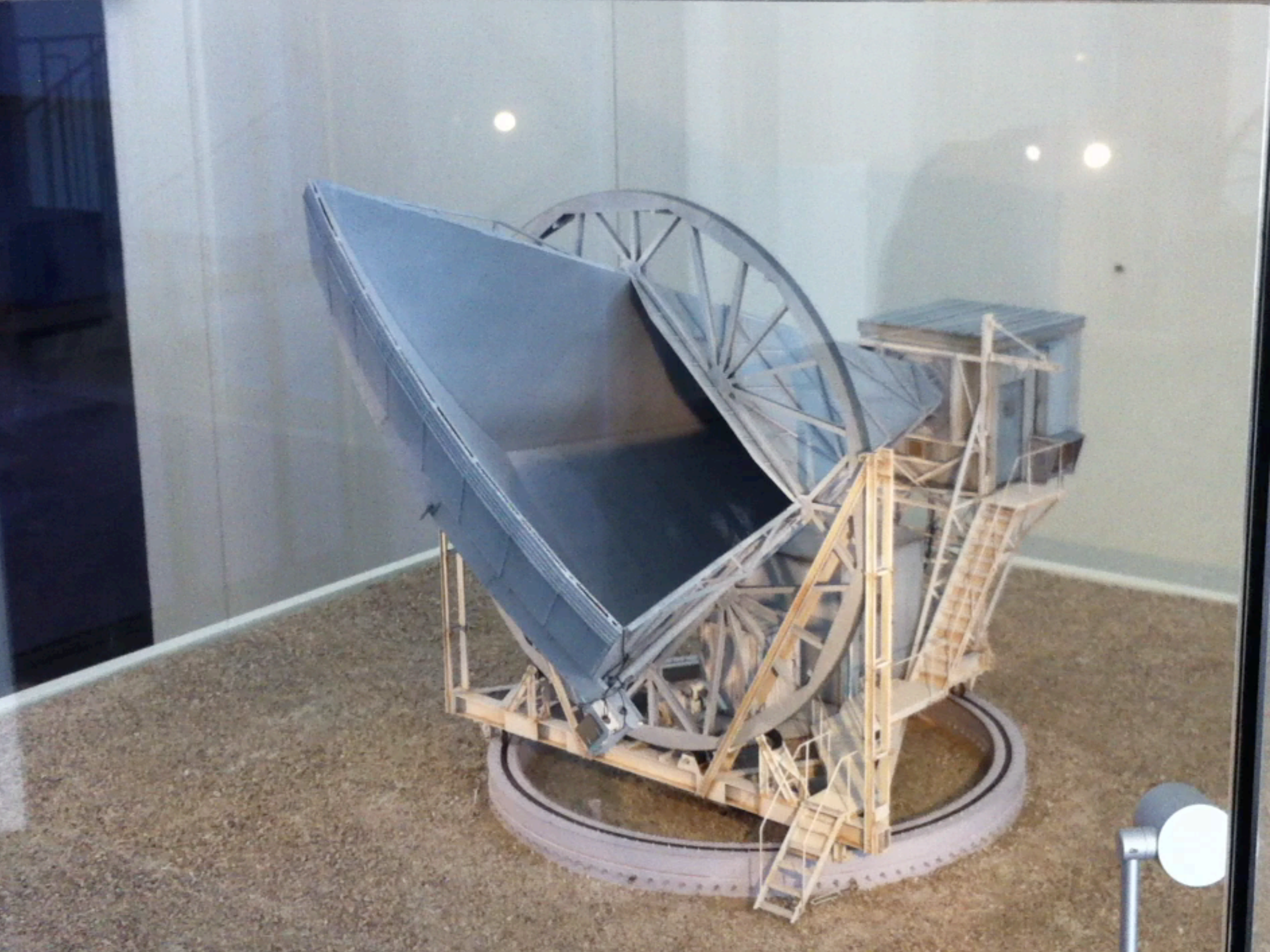
**Prof. Hiranya Peiris**  
**(Univ. College London)**

**All you need to do is to detect radio waves. For example, 1% of noise on the TV is from the fireball Universe**

1965









# The real detector system used by Penzias & Wilson

## The 3rd floor of Deutsches Museum



**Donated by Dr. Penzias,  
who was born in Munich**





Hornantennenanschluss

Horn antenna

Hohlleiterzug

V  
Vergleichs-  
quelle

M

Calibrator, cooled  
to 5K by liquid helium

Amplifier

Recorder

es

composed of many  
audible by a radio  
noise.  
on characteristic  
operation can be  
using the horn

collected by  
channel to the  
is brought  
or much like  
an electrical  
a recorder.

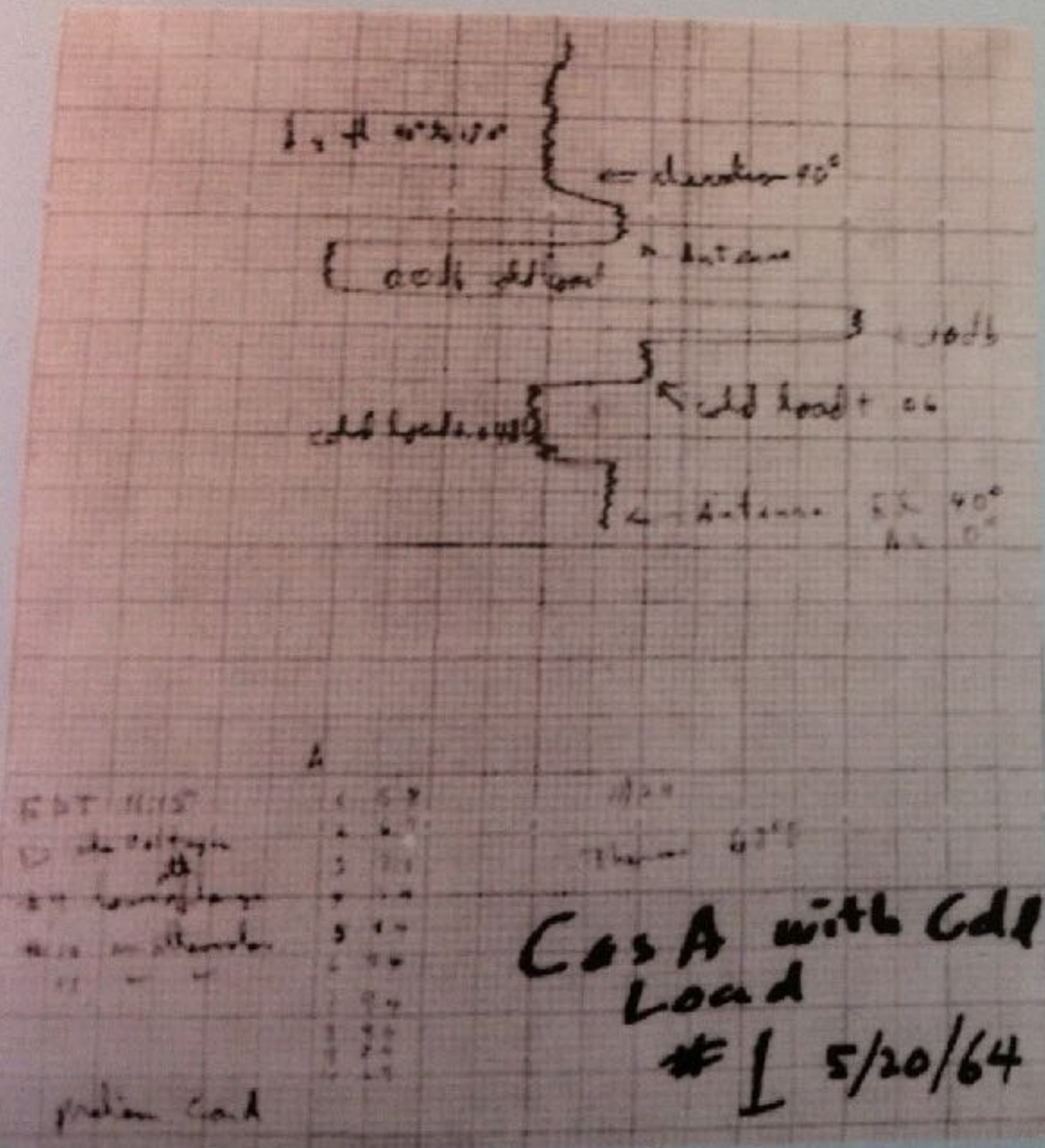
own  
th the

be  
with the



# May 20, 1964 CMB Discovered

$$6.7 - 2.3 - 0.8 - 0.1 \\ = 3.5 \pm 1.0 \text{ K}$$

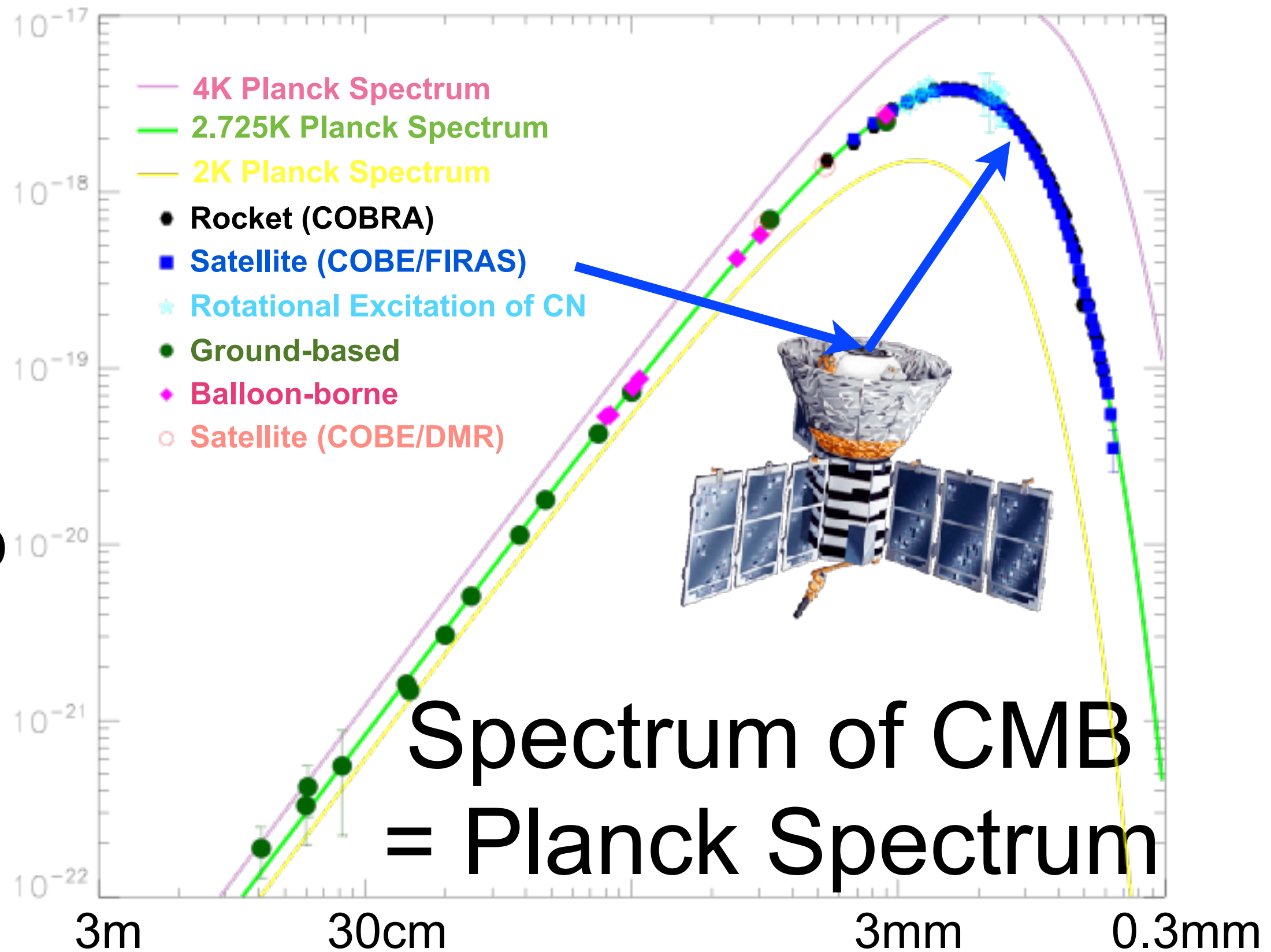


Schreiberaufzeichnung der ersten Messung des Mikrowellenhintergrundes am 20.5.1964

Recording of the first measurement of cosmic microwave background radiation taken on 5/20/1964.

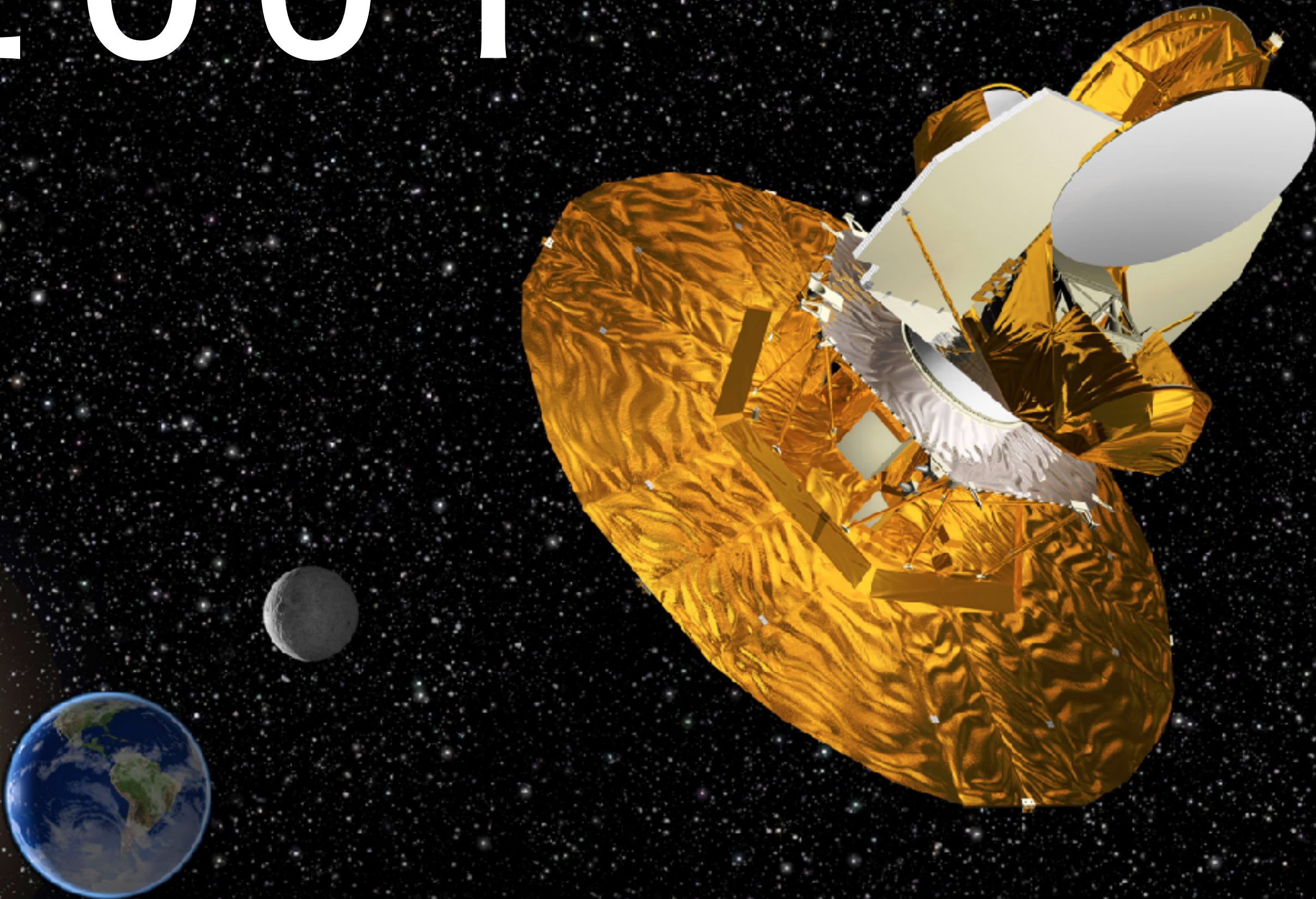


Brightness





2001





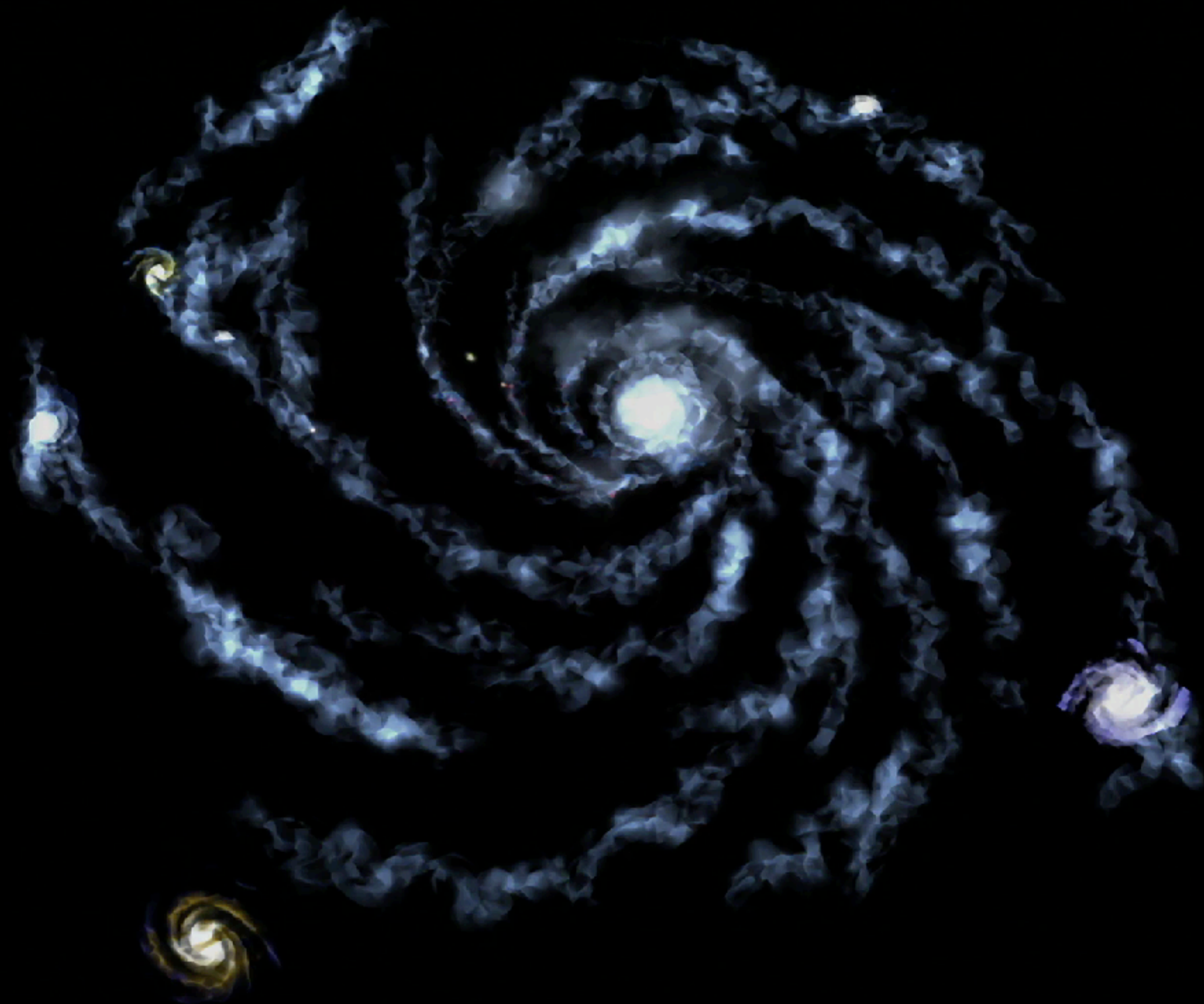
# WMAP Science Team

## July 19, 2002



- WMAP was launched on June 30, 2001
- The WMAP mission ended after 9 years of operation

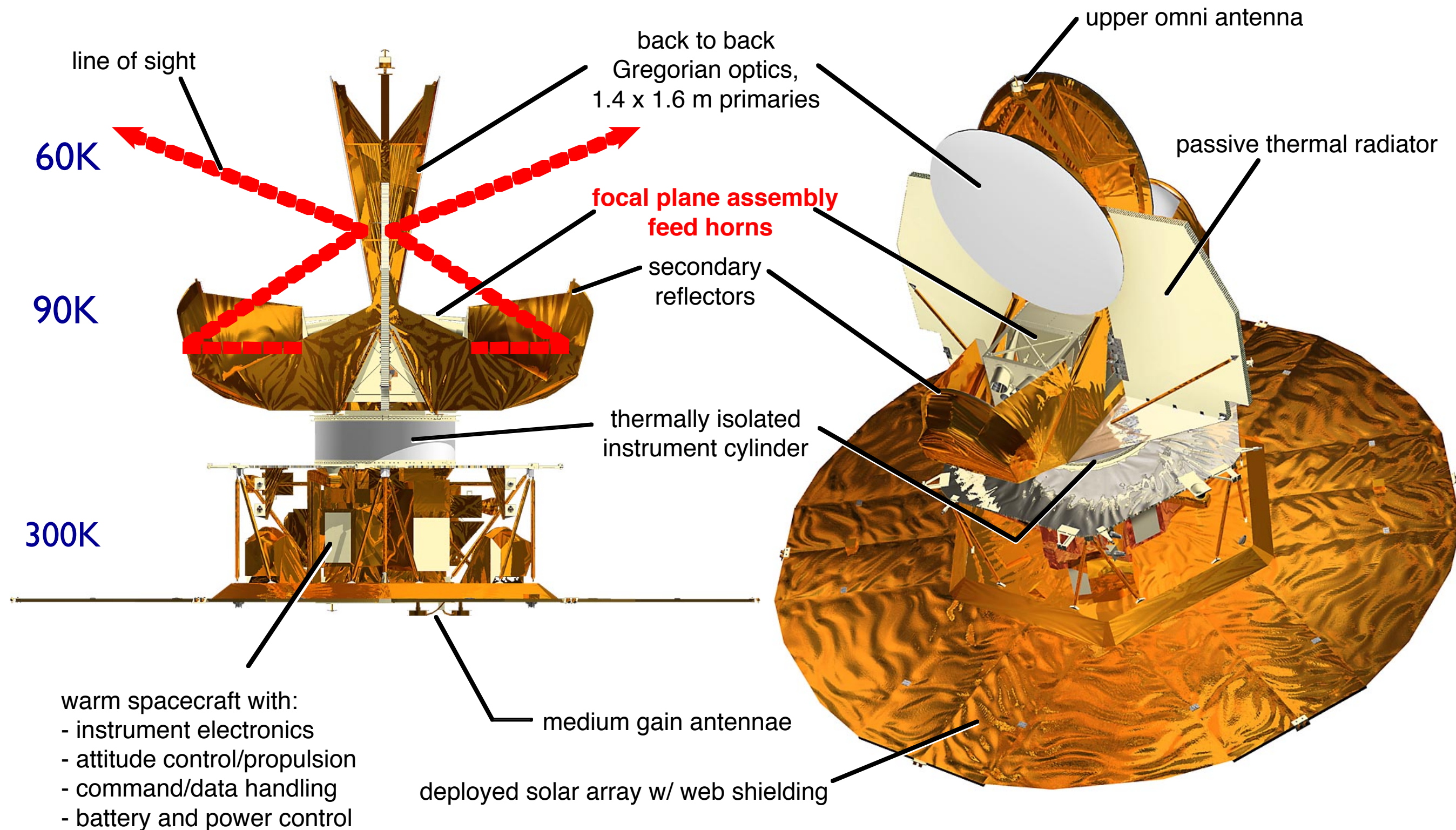




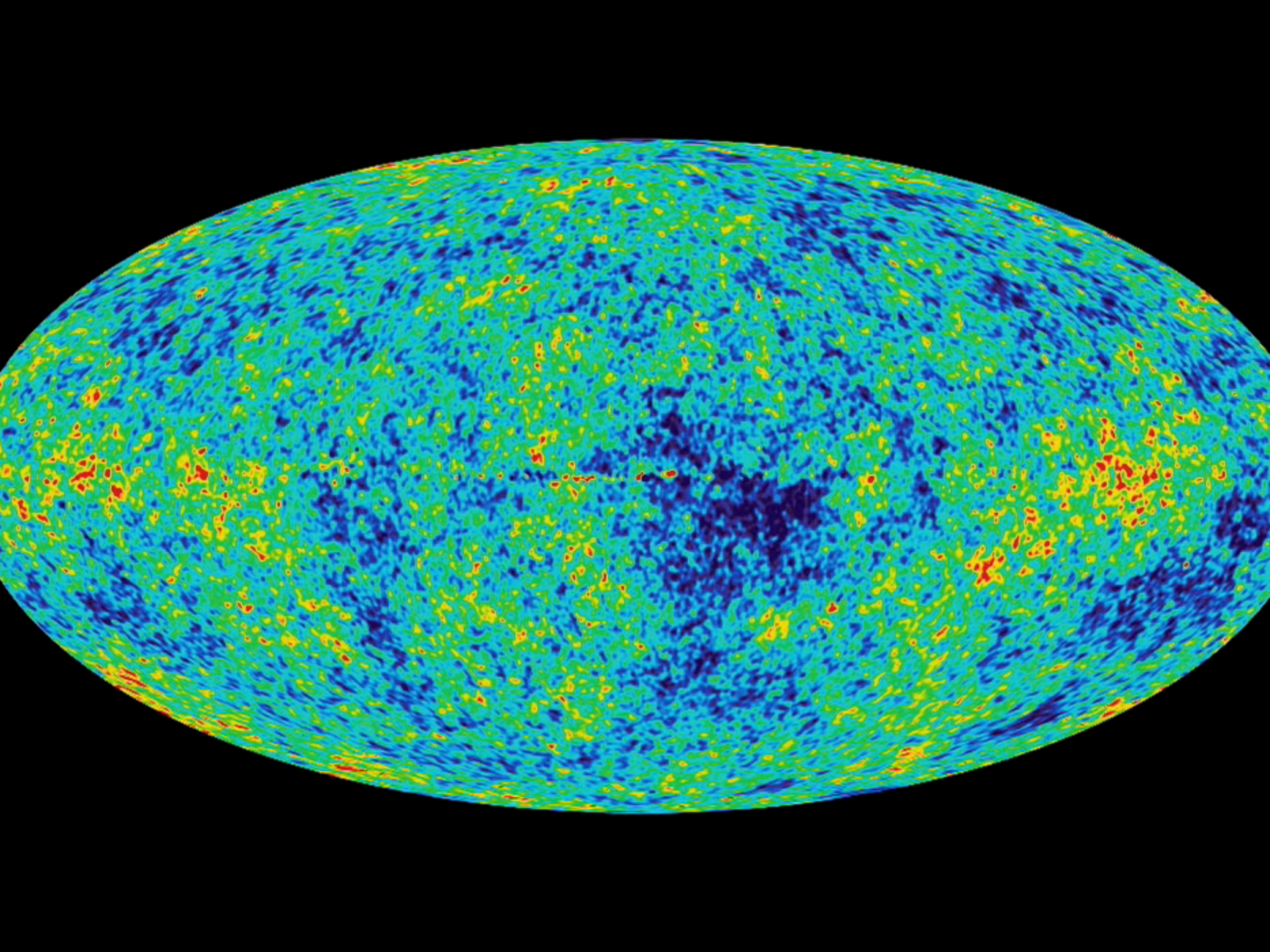


# WMAP Spacecraft

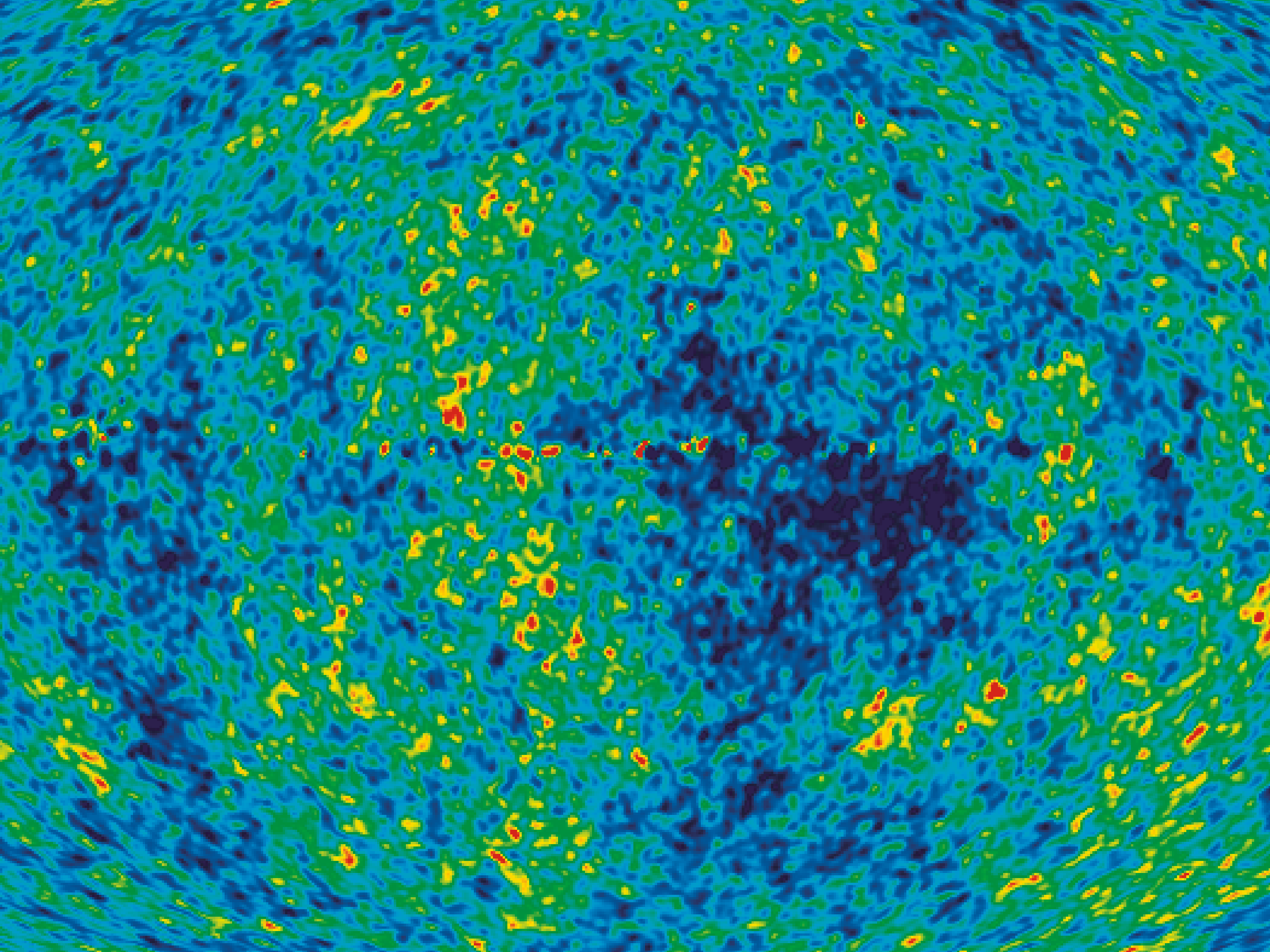
## No cryogenic components











# Our Origin

- WMAP taught us that **galaxies, stars, planets, and ourselves originated from tiny fluctuations in the early Universe**



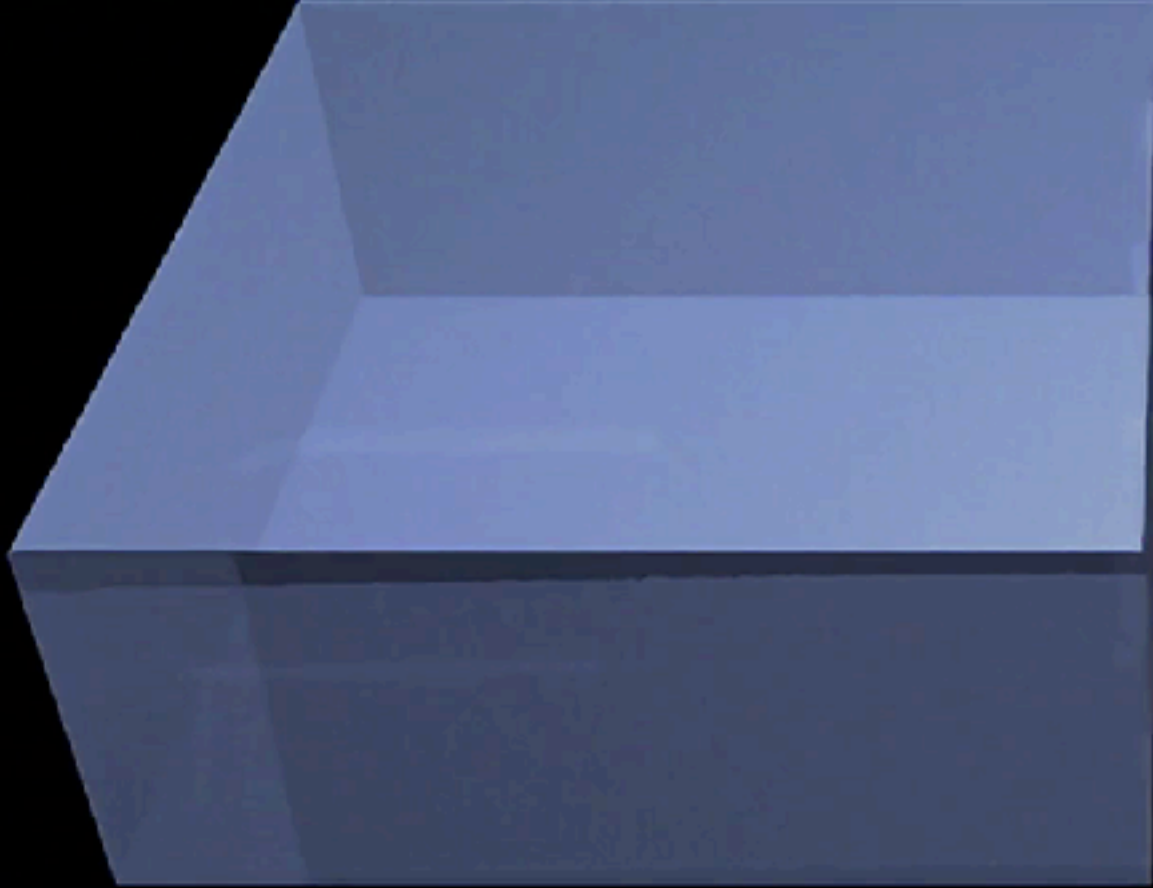


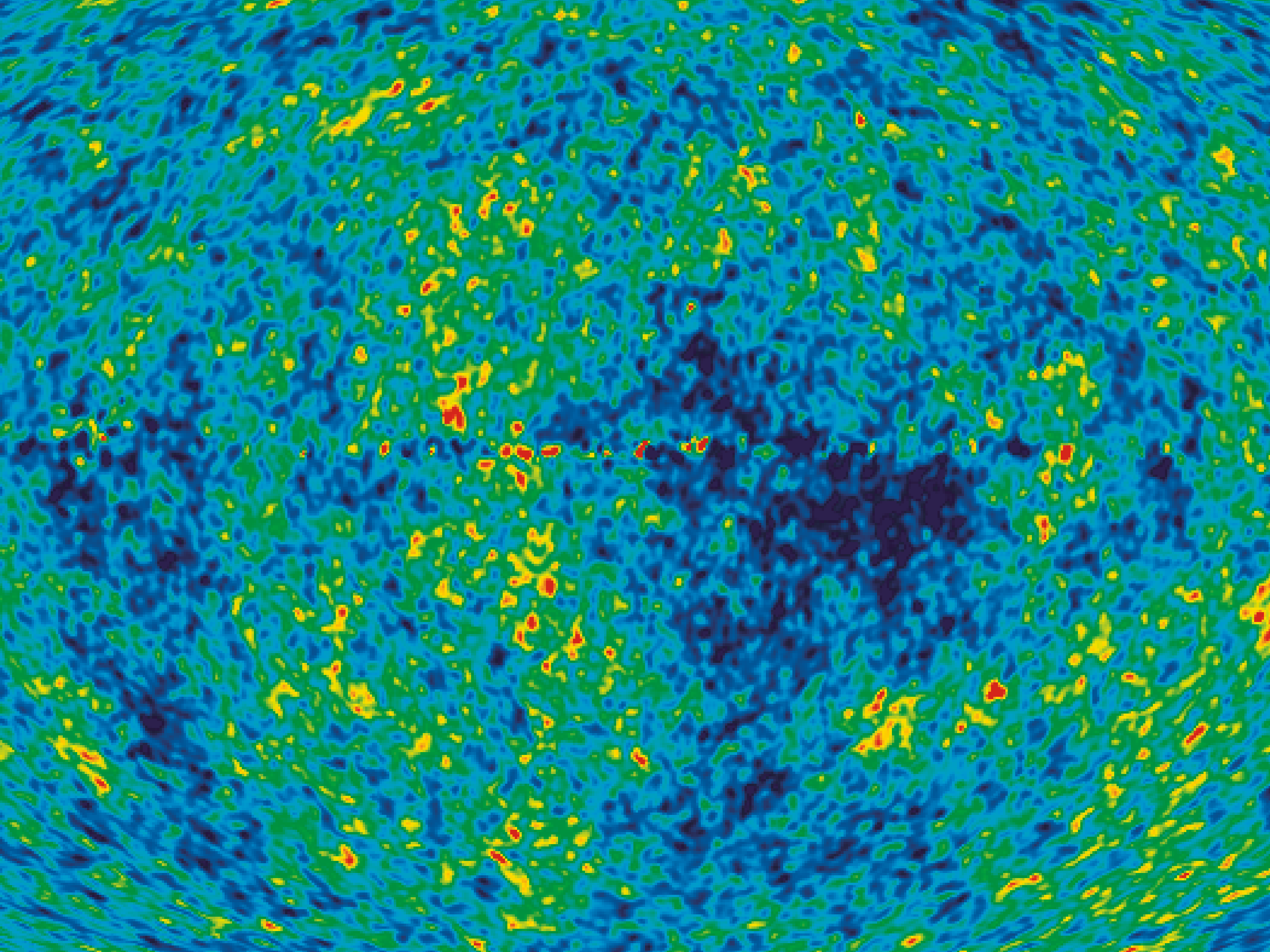


# Kosmische Miso Suppe

- When matter and radiation were hotter than 3000 K, matter was completely ionised. The Universe was filled with plasma, which behaves just like a soup
- Think about a Miso soup (if you know what it is). Imagine throwing Tofus into a Miso soup, while changing the density of Miso
- And imagine watching how ripples are created and propagate throughout the soup



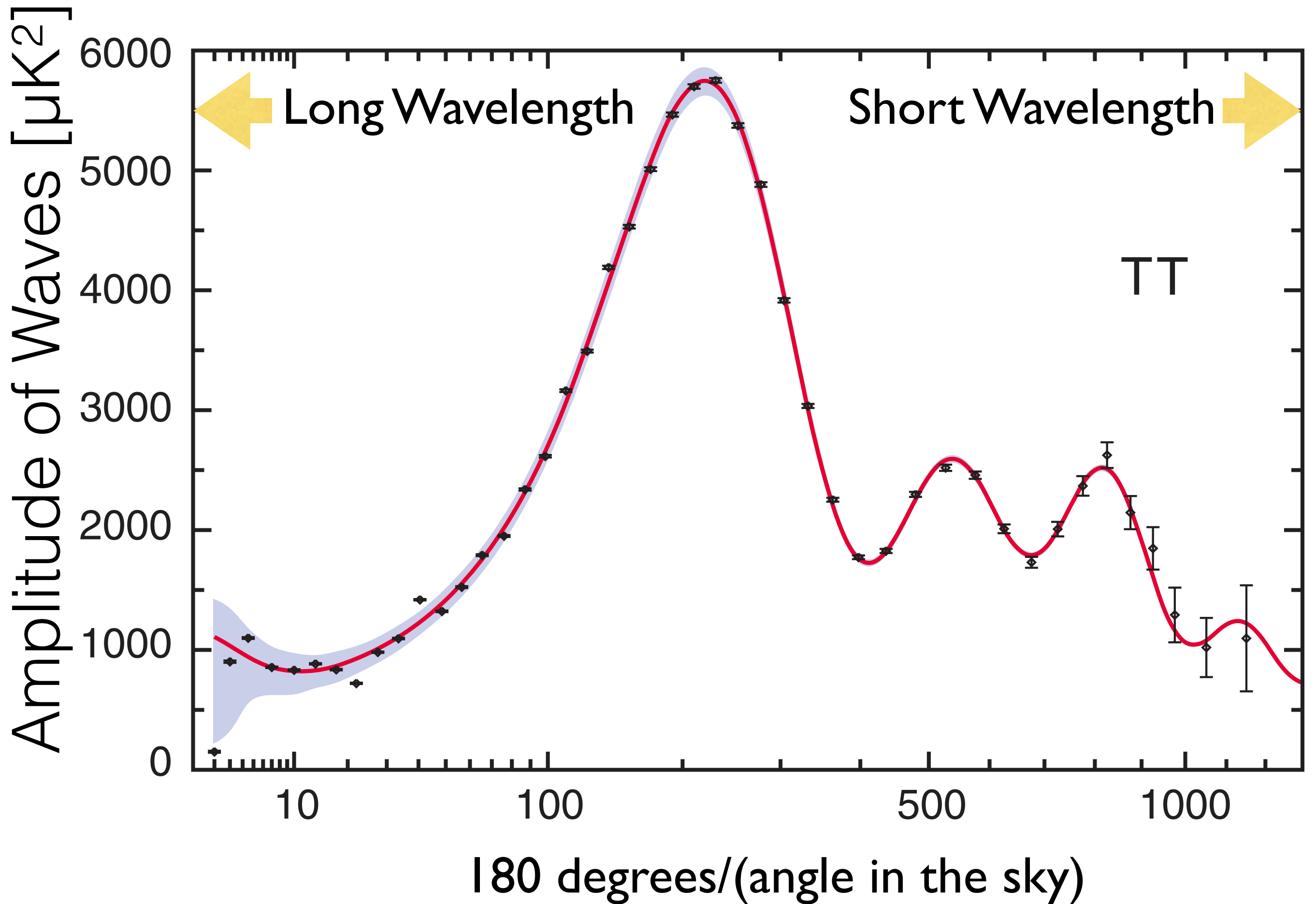




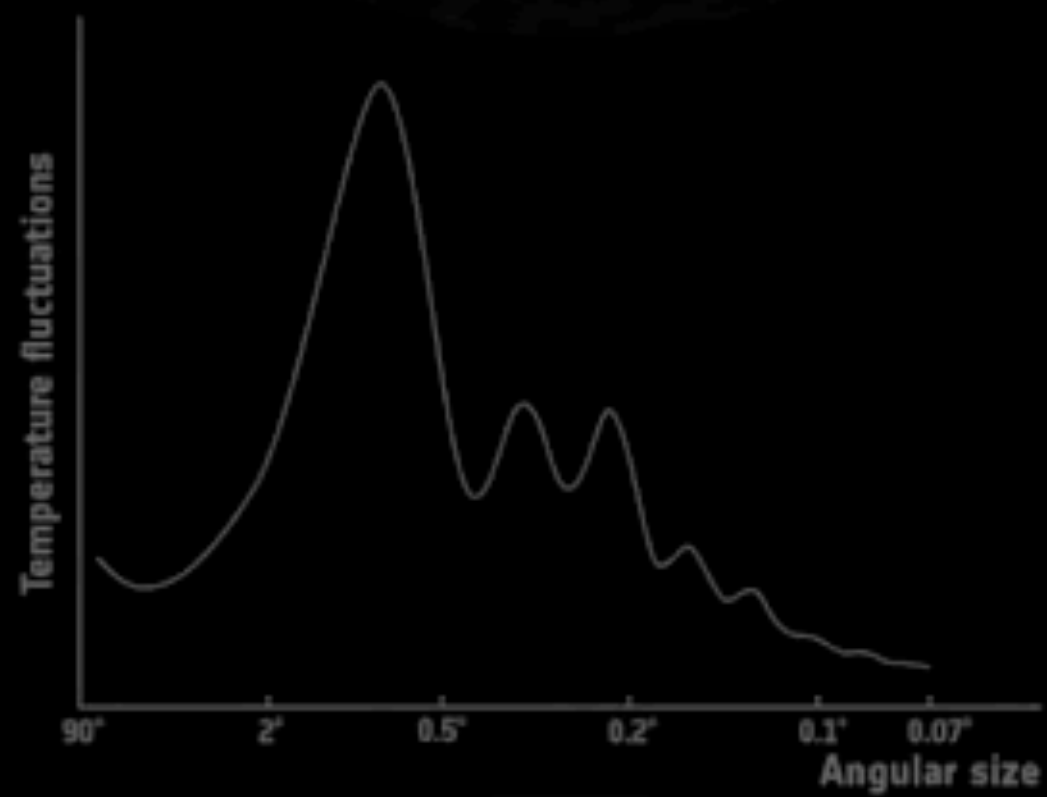


# Data Analysis

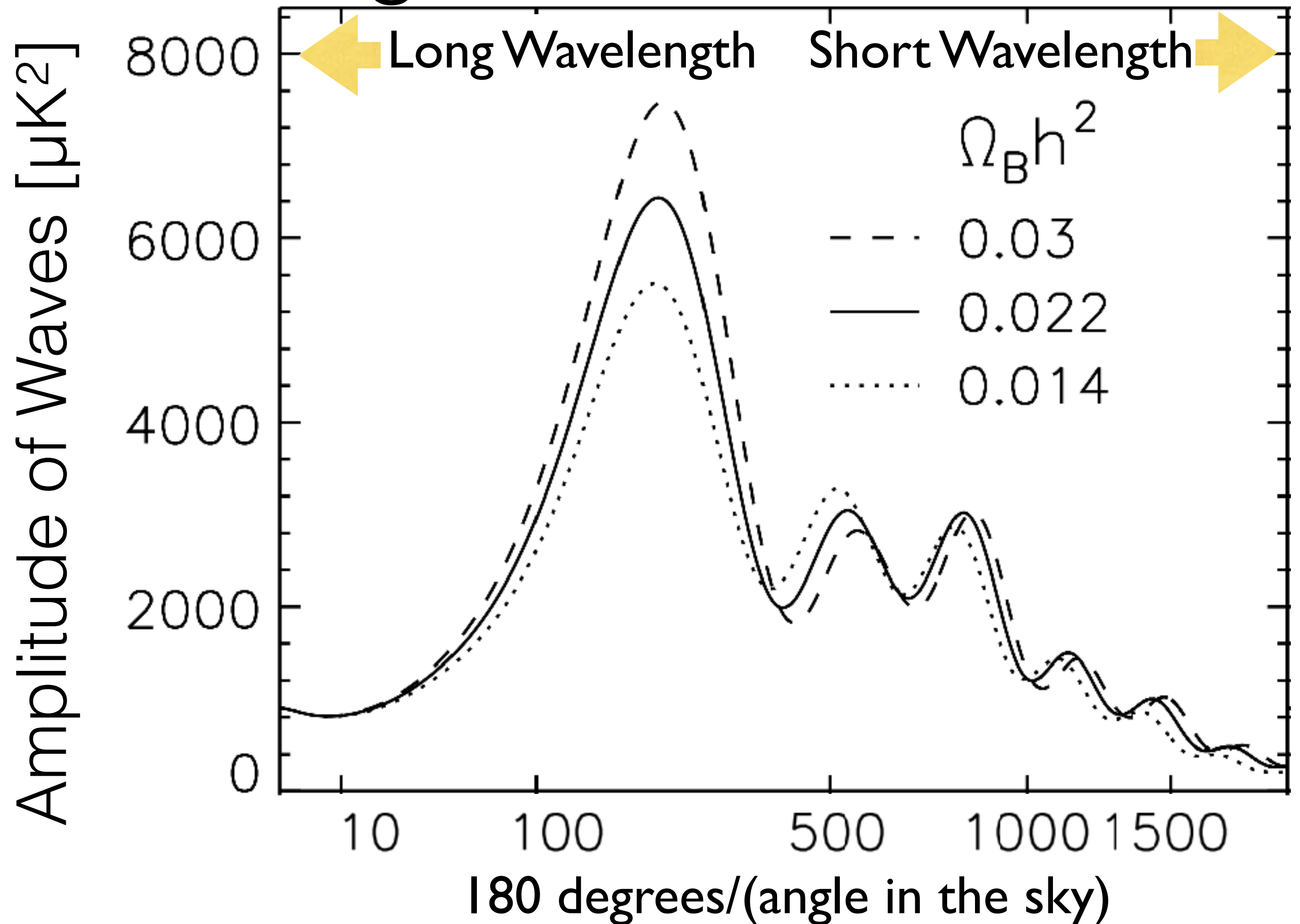
- Decompose temperature fluctuations in the sky into a set of waves with various wavelengths
- Make a diagram showing the strength of each wavelength





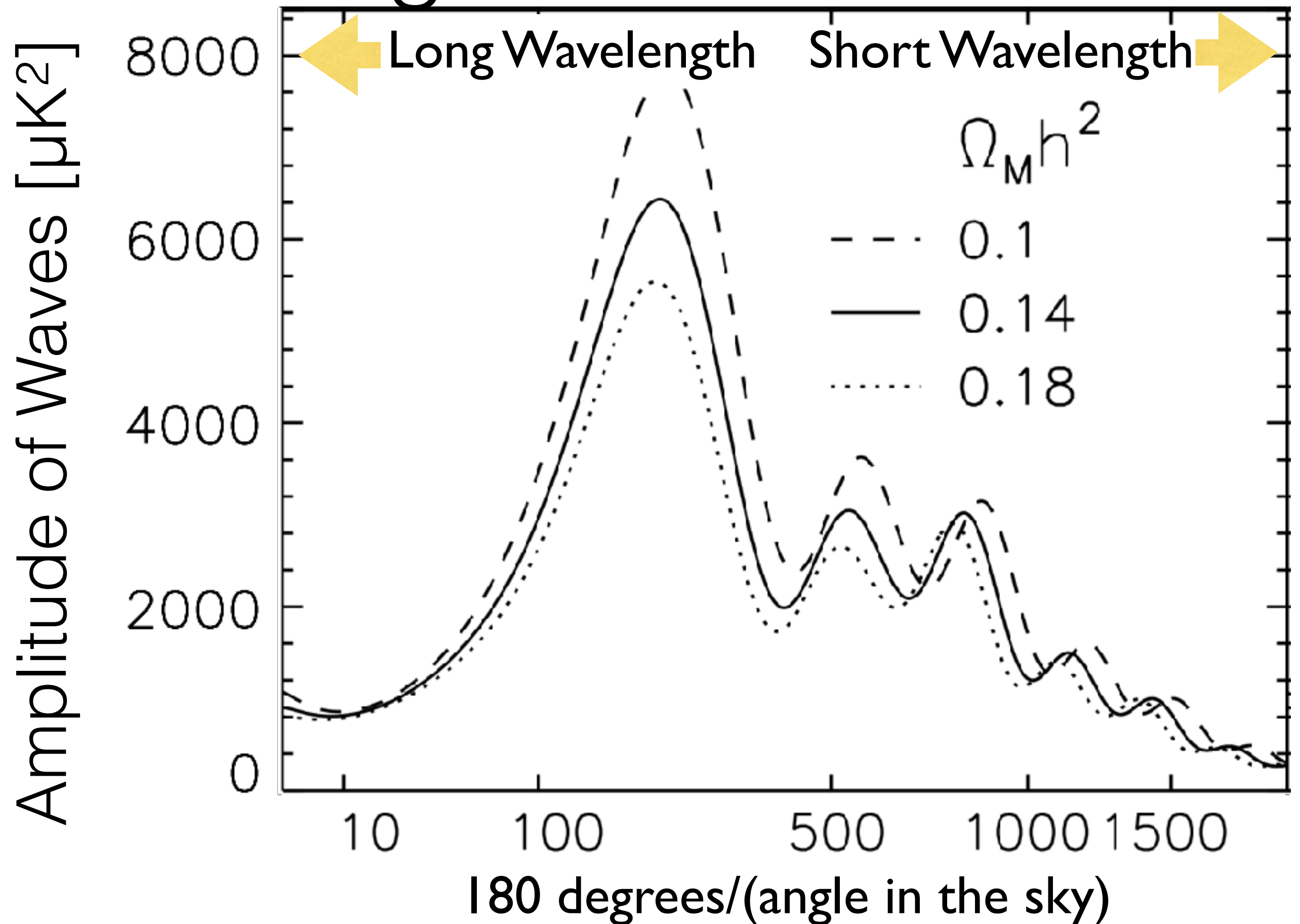


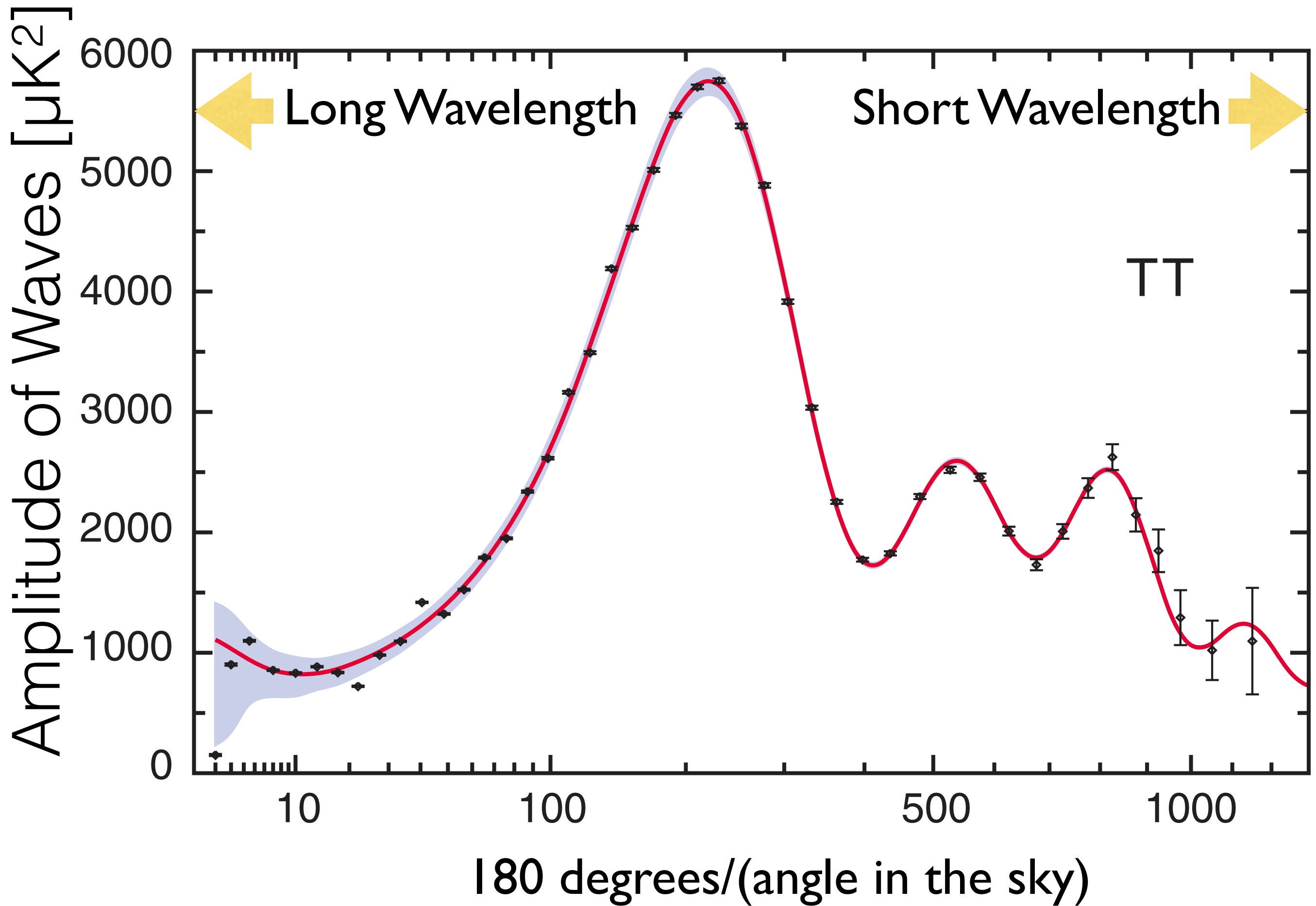
# Measuring Abundance of H&He





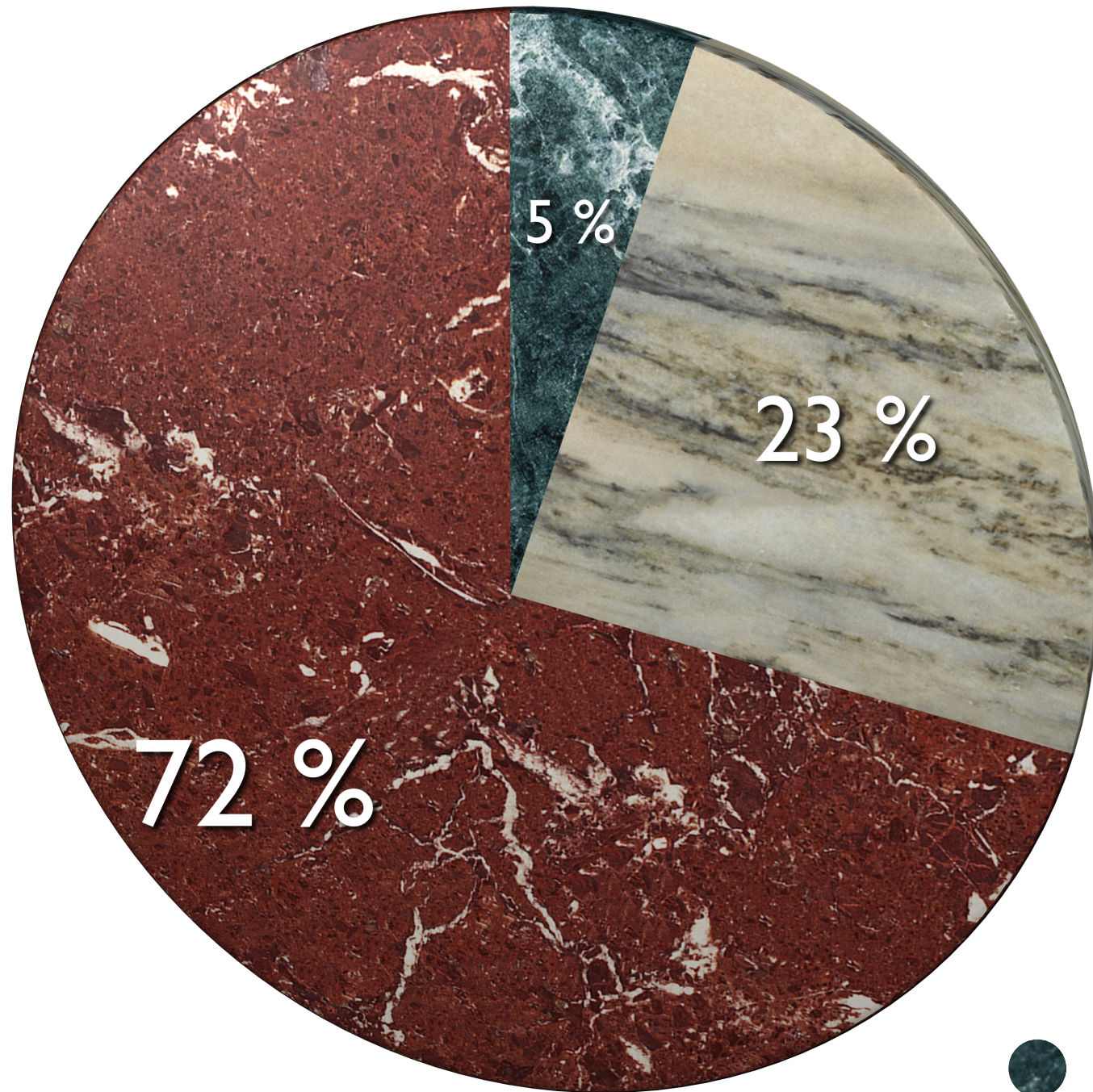
# Measuring Total Matter Density







# Cosmic Pie Chart



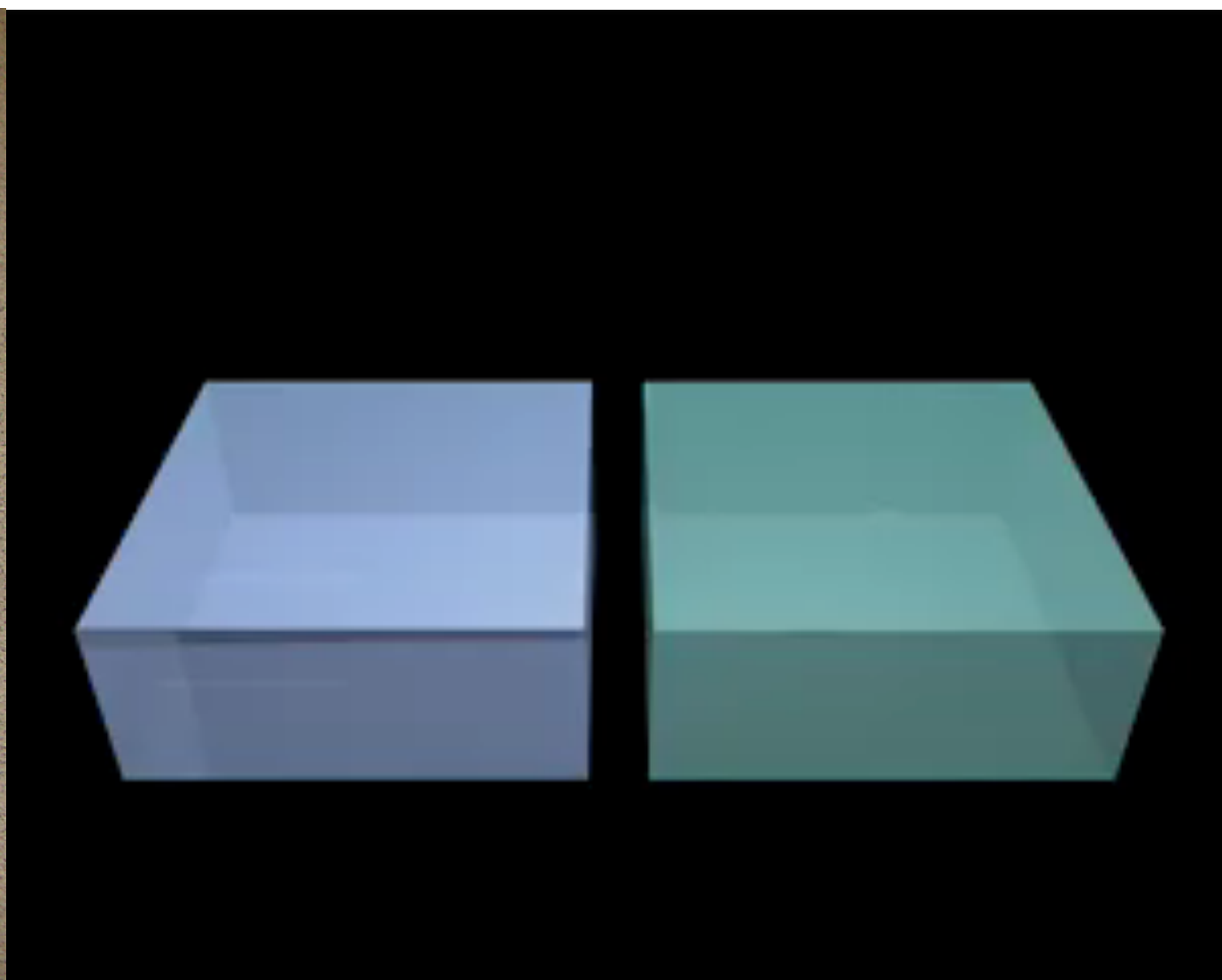
- WMAP determined the abundance of various components in the Universe
- As a result, **we came to realise that we do not understand 95% of our Universe...**





# Origin of Fluctuations

- Who dropped those Tofus into the cosmic Miso soup?





*Mukhanov & Chibisov (1981); Guth & Pi (1982); Hawking (1982); Starobinsky (1982);  
Bardeen, Turner & Steinhardt (1983)*

# Leading Idea

- **Quantum Mechanics at work in the early Universe**
- Uncertainty Principle:
  - $[\text{Energy you can borrow}] \times [\text{Time you borrow}] \sim h$
  - Time was very short in the early Universe = You could borrow a lot of energy
- **Those energies became the origin of fluctuations**
- How did quantum fluctuations on the microscopic scales become macroscopic fluctuations over cosmological sizes?

# Cosmic Inflation

- In a tiny fraction of a second, the size of an atomic nucleus became the size of the Solar System
- In  $10^{-36}$  second, space was stretched by at least a factor of  $10^{26}$

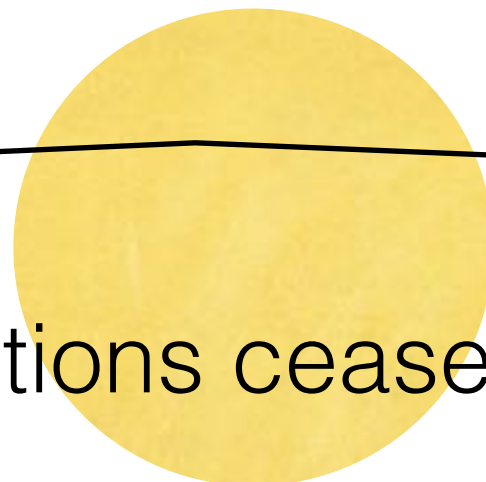
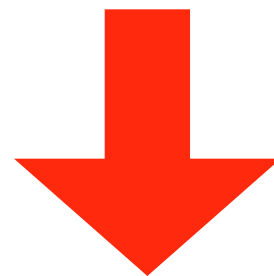


# Stretching Micro to Macro

Quantum fluctuations on  
microscopic scales



# Inflation!



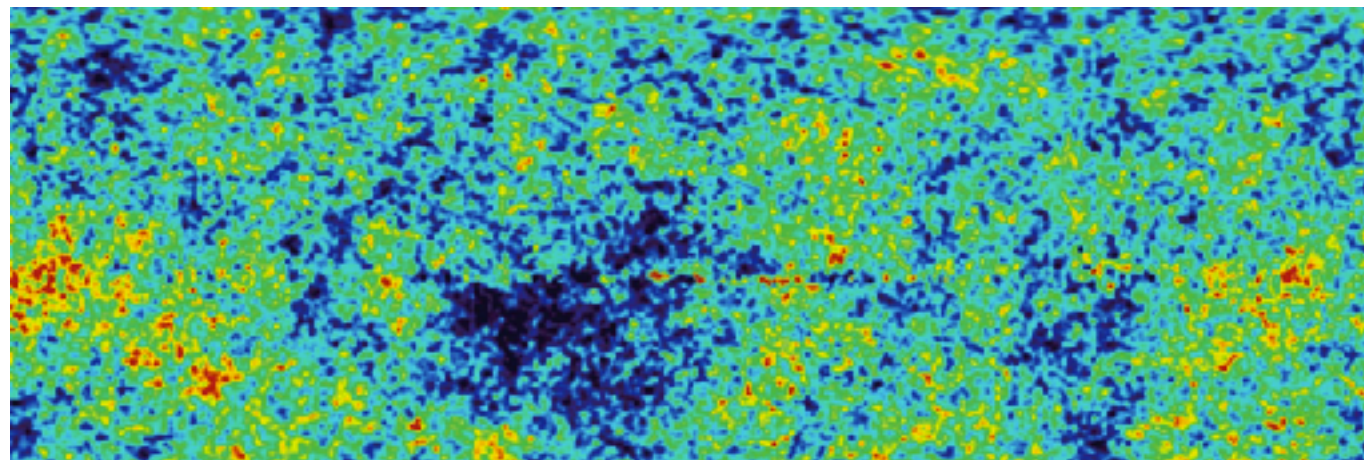
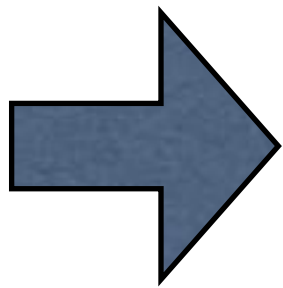
- Quantum fluctuations cease to be quantum
- Become macroscopic, classical fluctuations

# Key Predictions of Inflation

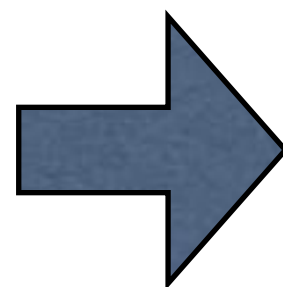
 $\zeta$ 

scalar  
mode

- Fluctuations we observe today in CMB and the matter distribution originate from quantum fluctuations generated during inflation

 $h_{ij}$ 

tensor  
mode





# We measure distortions in space

- A distance between two points in space

$$d\ell^2 = a^2(t)[1 + 2\zeta(\mathbf{x}, t)][\delta_{ij} + h_{ij}(\mathbf{x}, t)]dx^i dx^j$$

- $\zeta$ : “curvature perturbation” (scalar mode)
  - Perturbation to the determinant of the spatial metric
- $h_{ij}$ : “gravitational waves” (tensor mode)
  - Perturbation that does not change the determinant (area)



$$\sum_i h_{ii} = 0$$

# Heisenberg's Uncertainty Principle

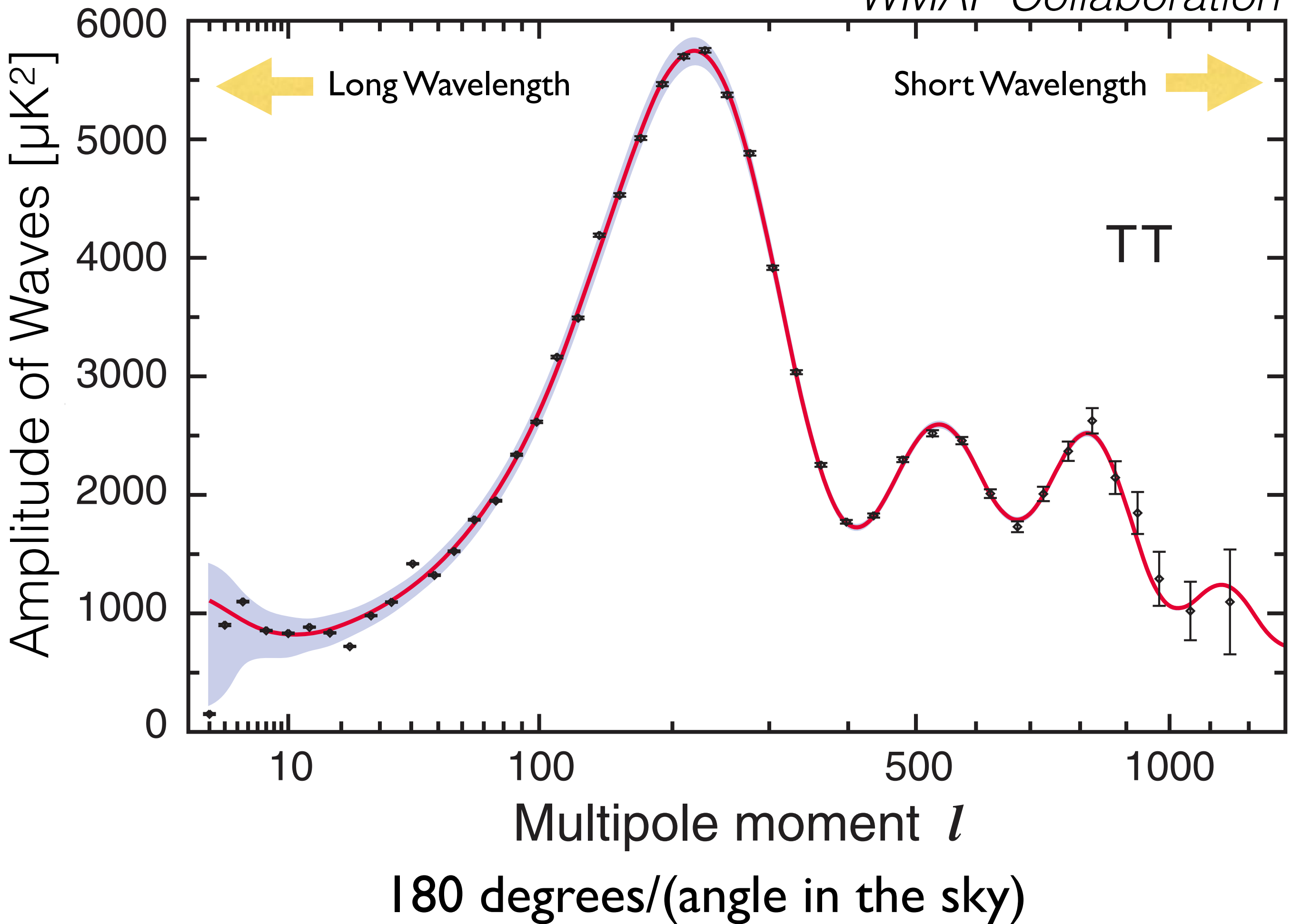
- [Energy you can borrow] x [Time you borrow] = constant
- Suppose that the distance between two points increases in proportion to  **$a(t)$**  [which is called the scale factor] by the expansion of the universe
- Define the “expansion rate of the universe” as

$$H \equiv \frac{\dot{a}}{a} \quad [\text{This has units of } 1/\text{time}]$$

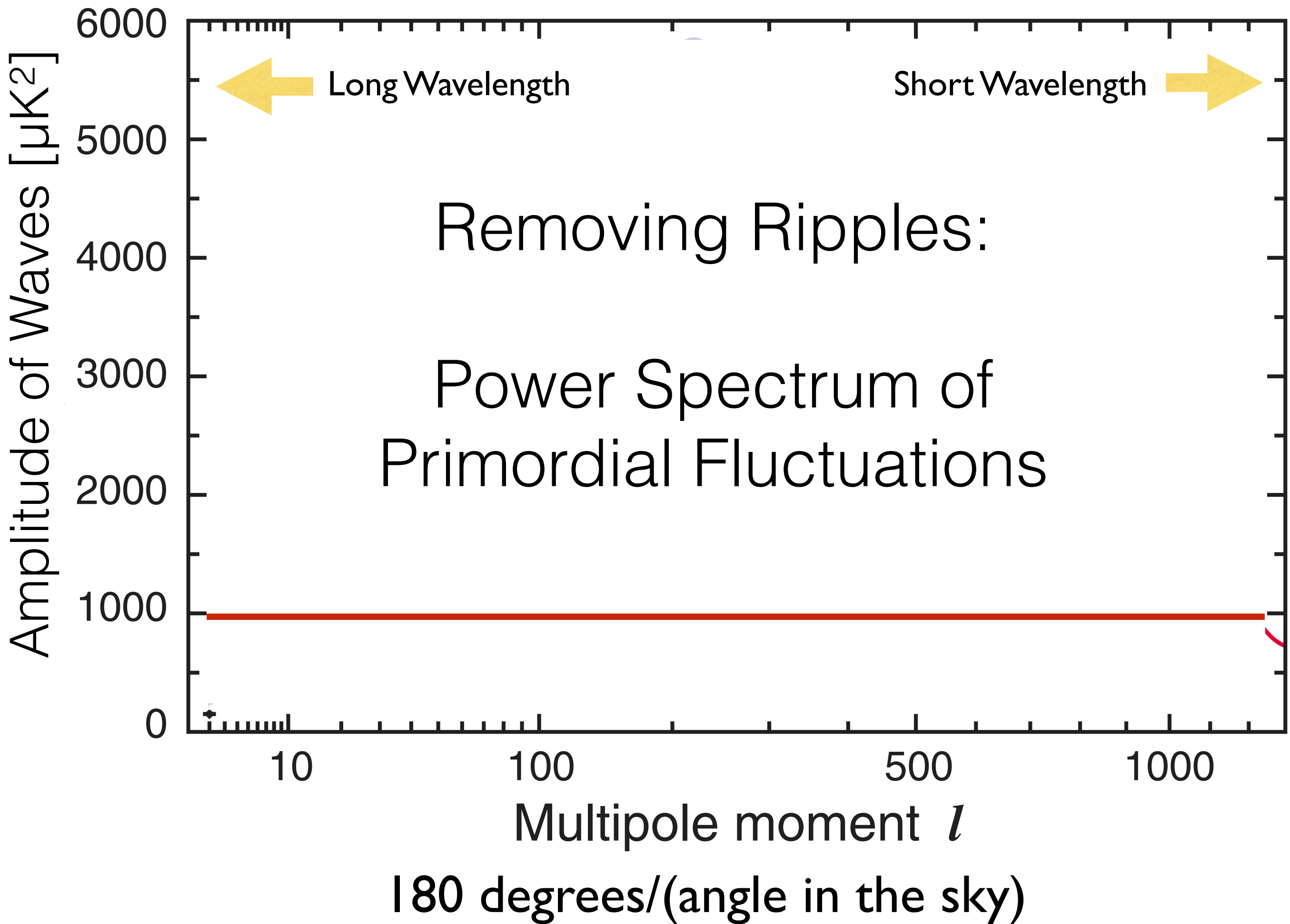


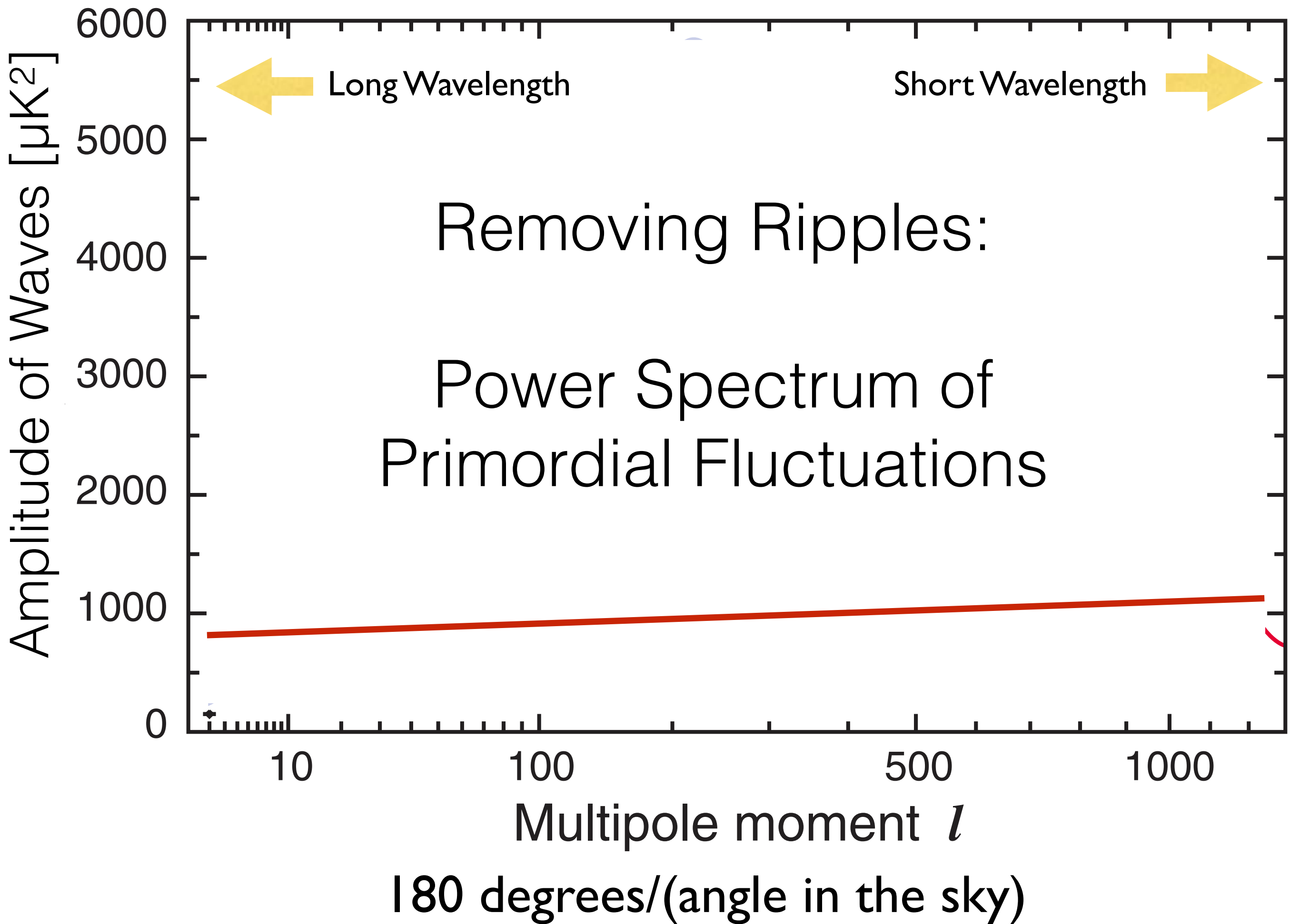
# Fluctuations are proportional to $H$

- [Energy you can borrow] x [Time you borrow] = constant
- $H \equiv \frac{\dot{a}}{a}$  [This has units of 1/time]
- Then, **both  $\zeta$  and  $h_{ij}$  are proportional to  $H$**
- Inflation occurs in  $10^{-36}$  second - this is such a short period of time that you can borrow a lot of energy!  
 **$H$  during inflation in energy units is  $10^{14}$  GeV**

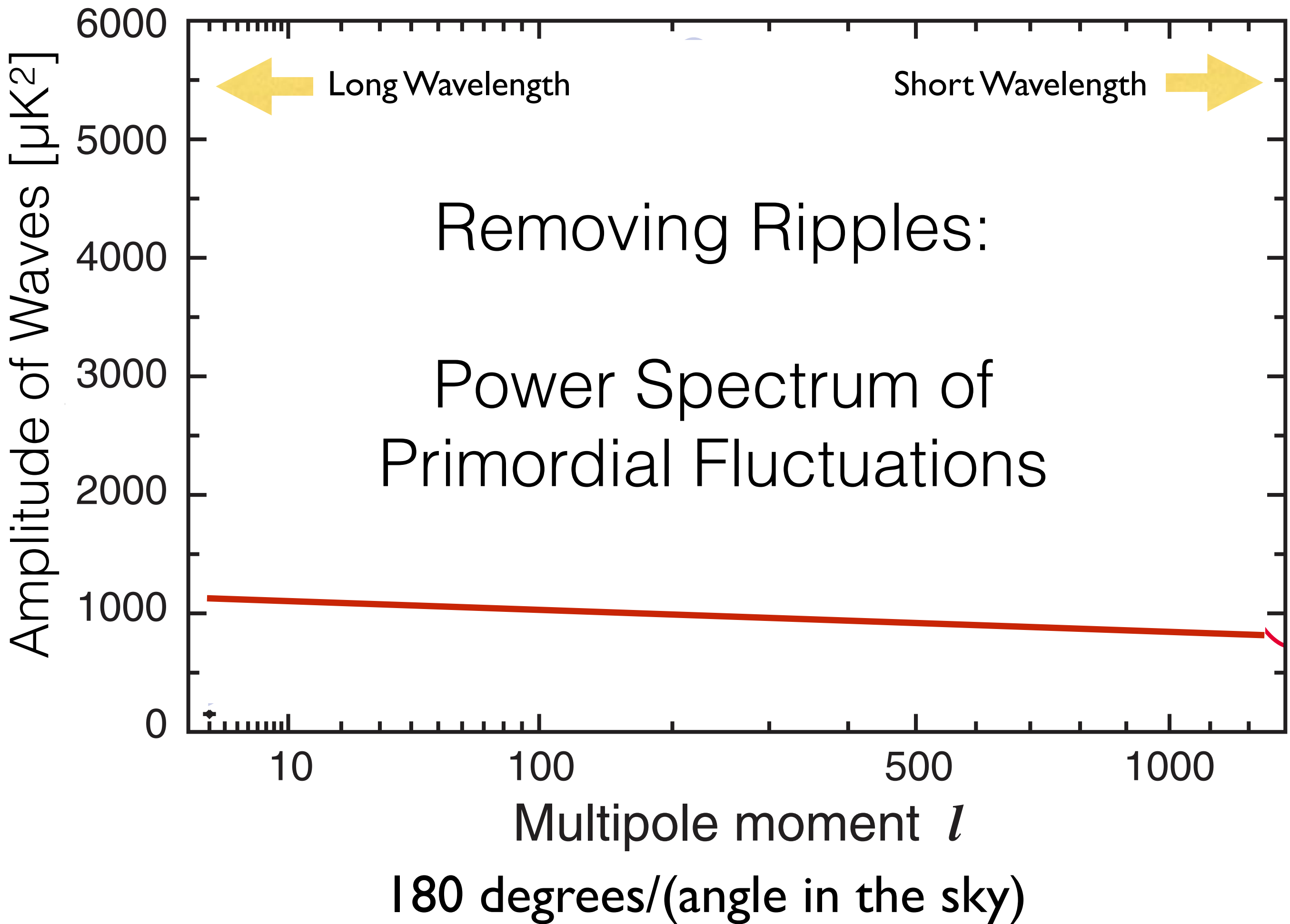


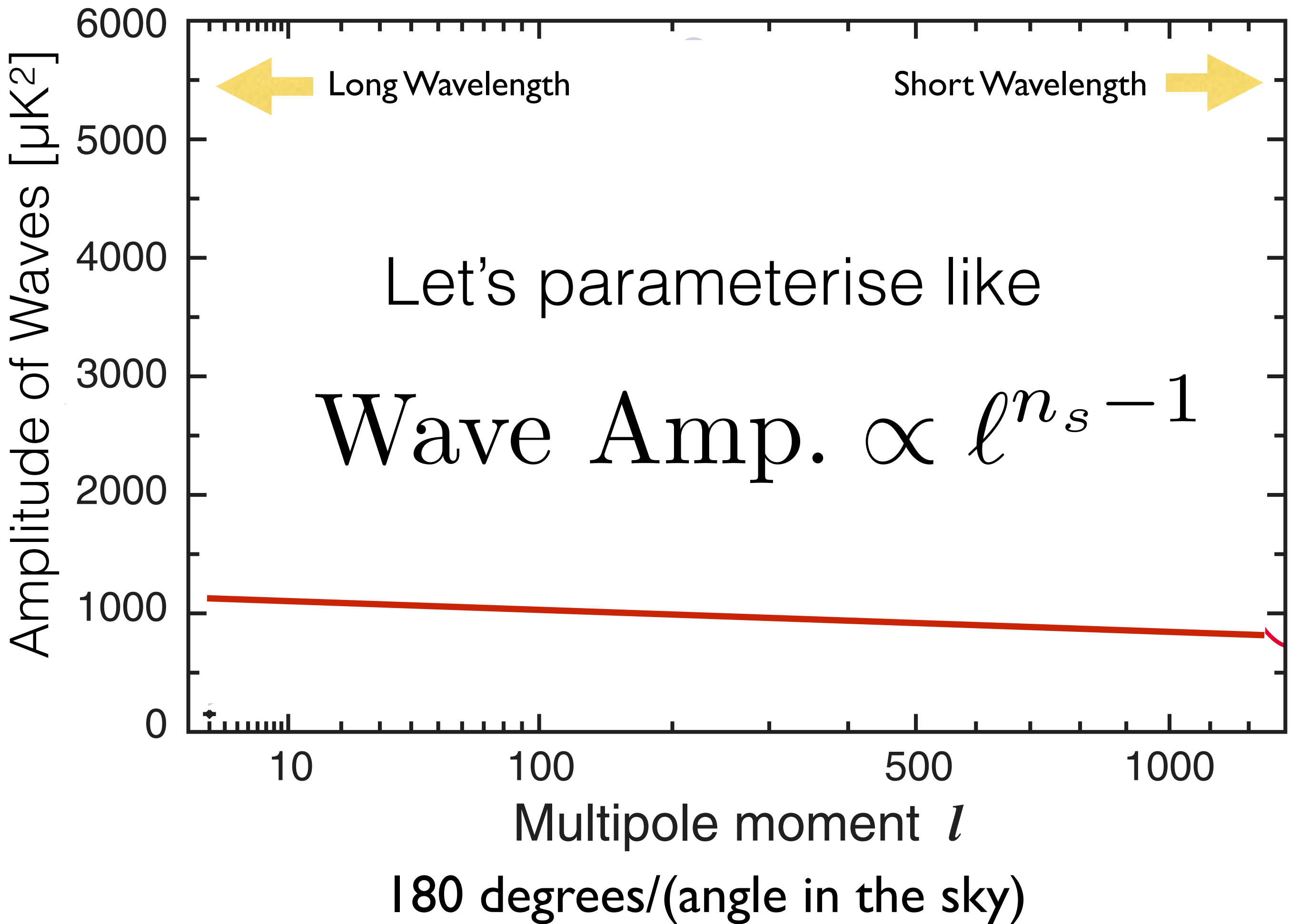


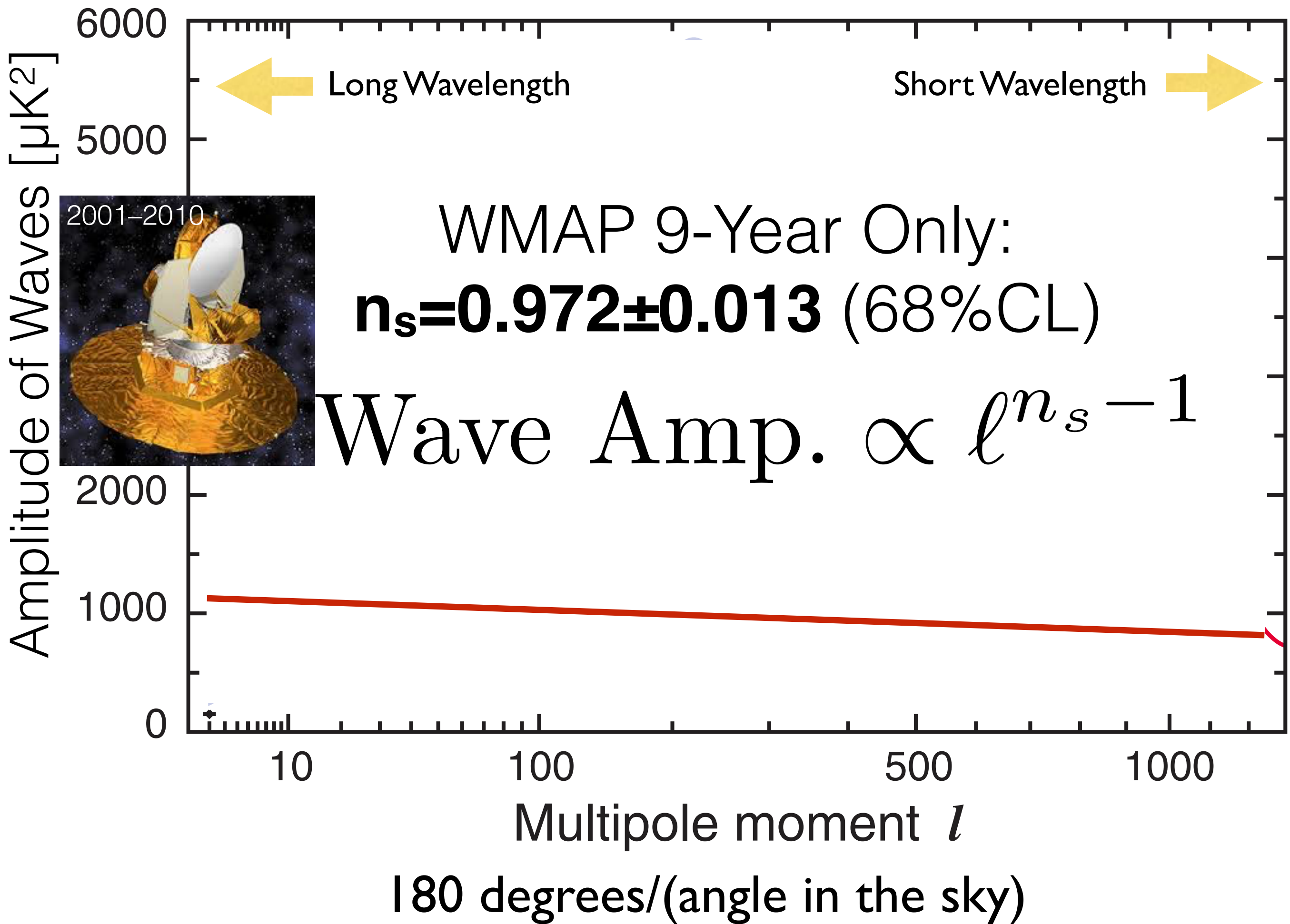




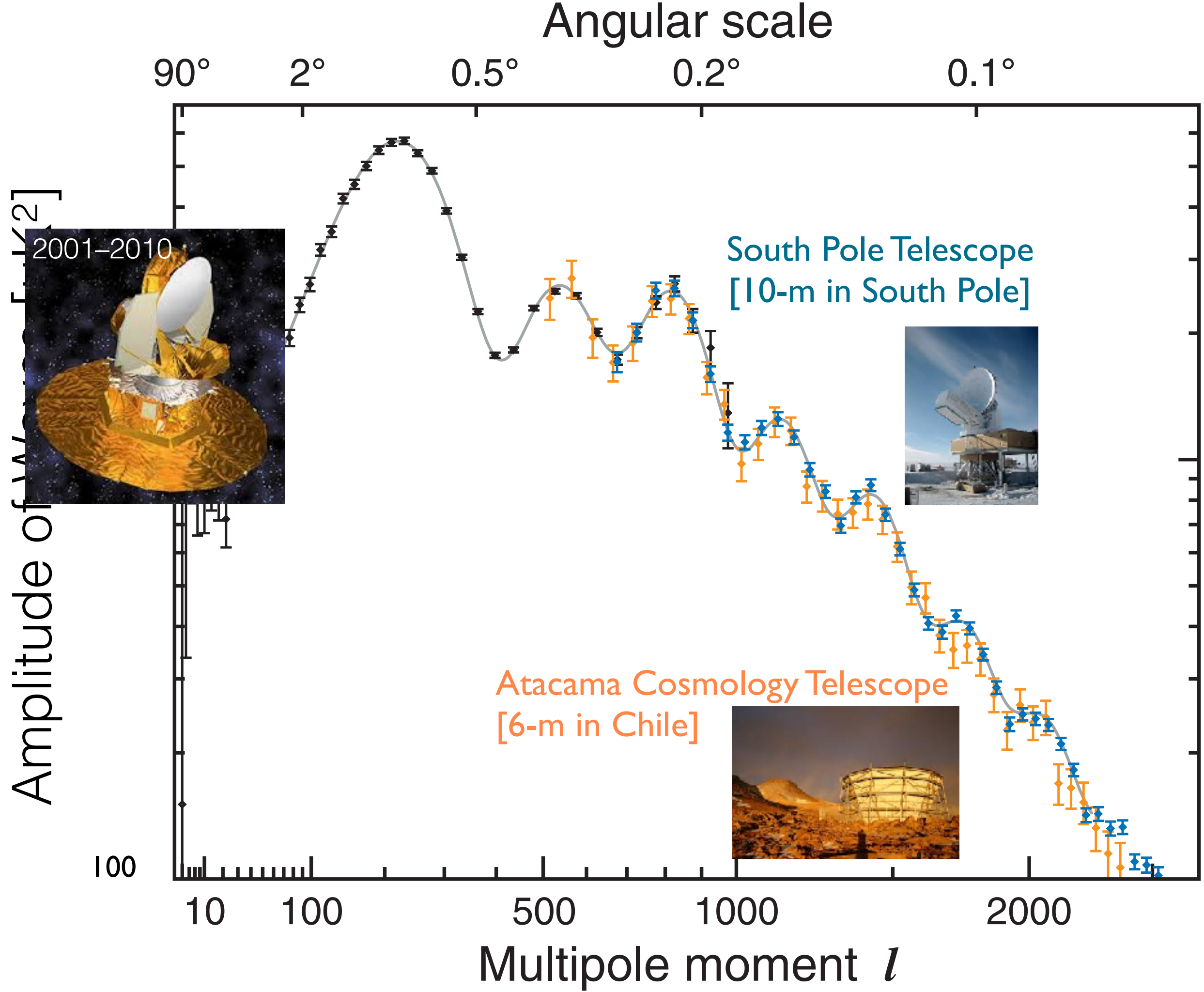


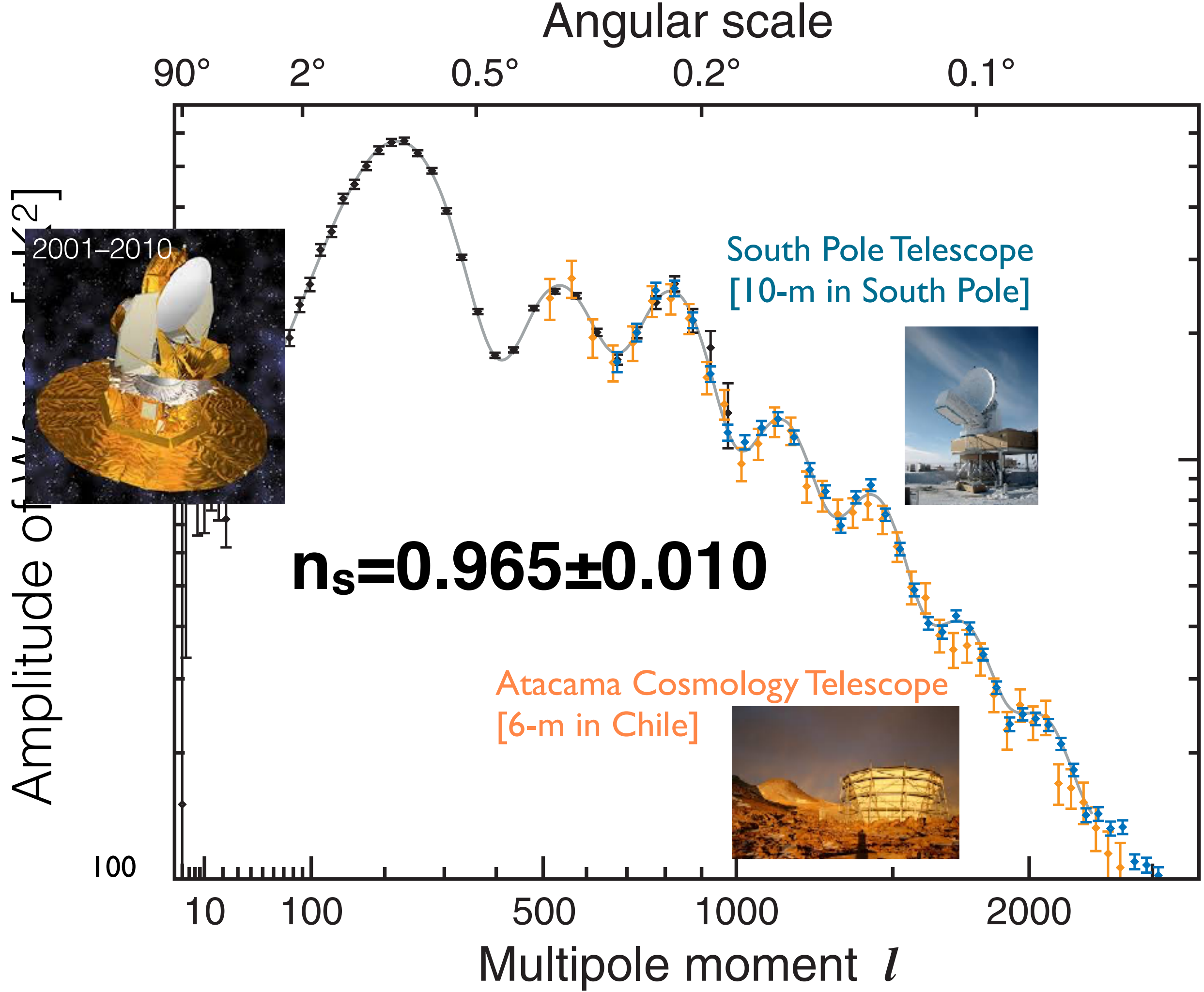


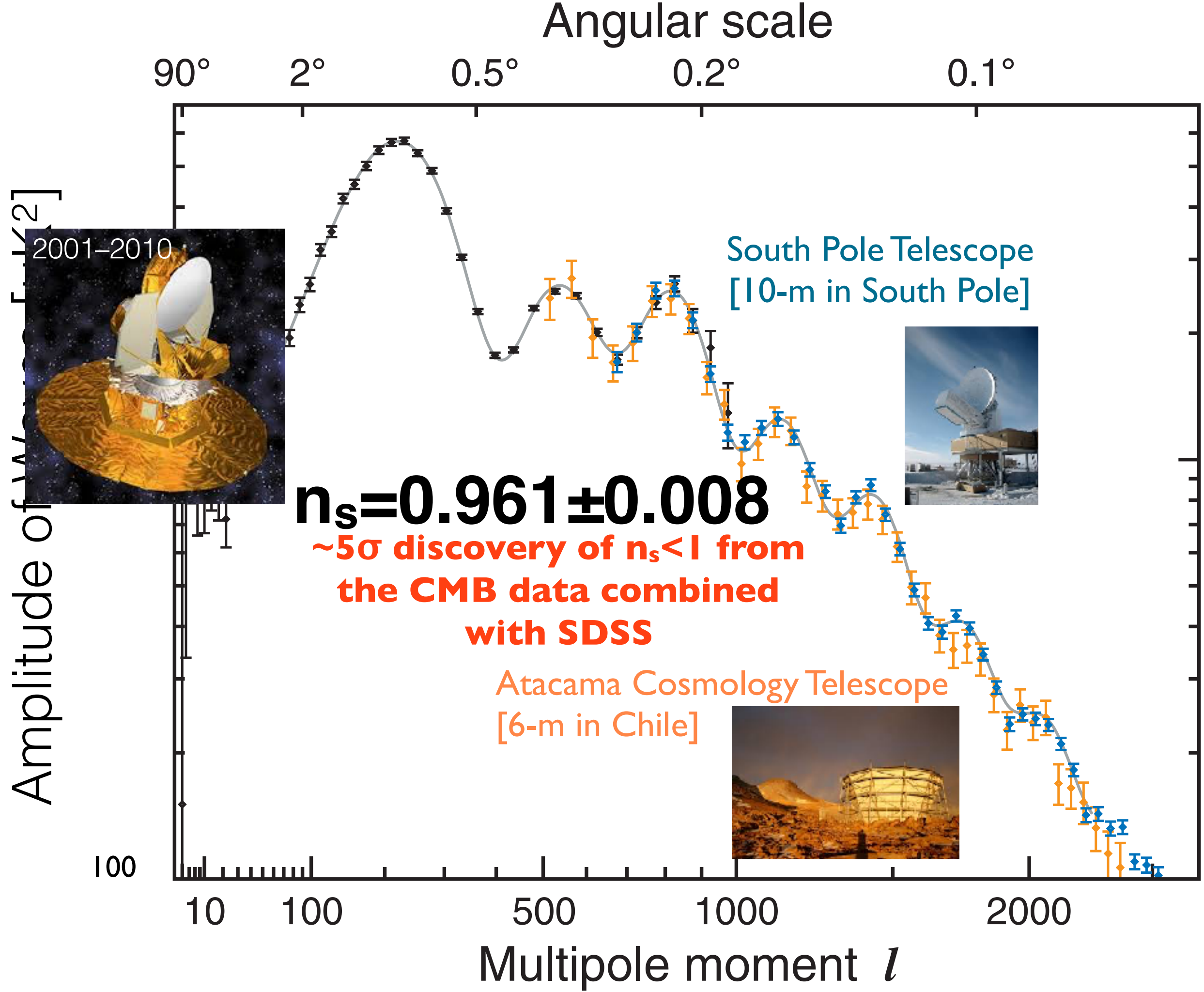










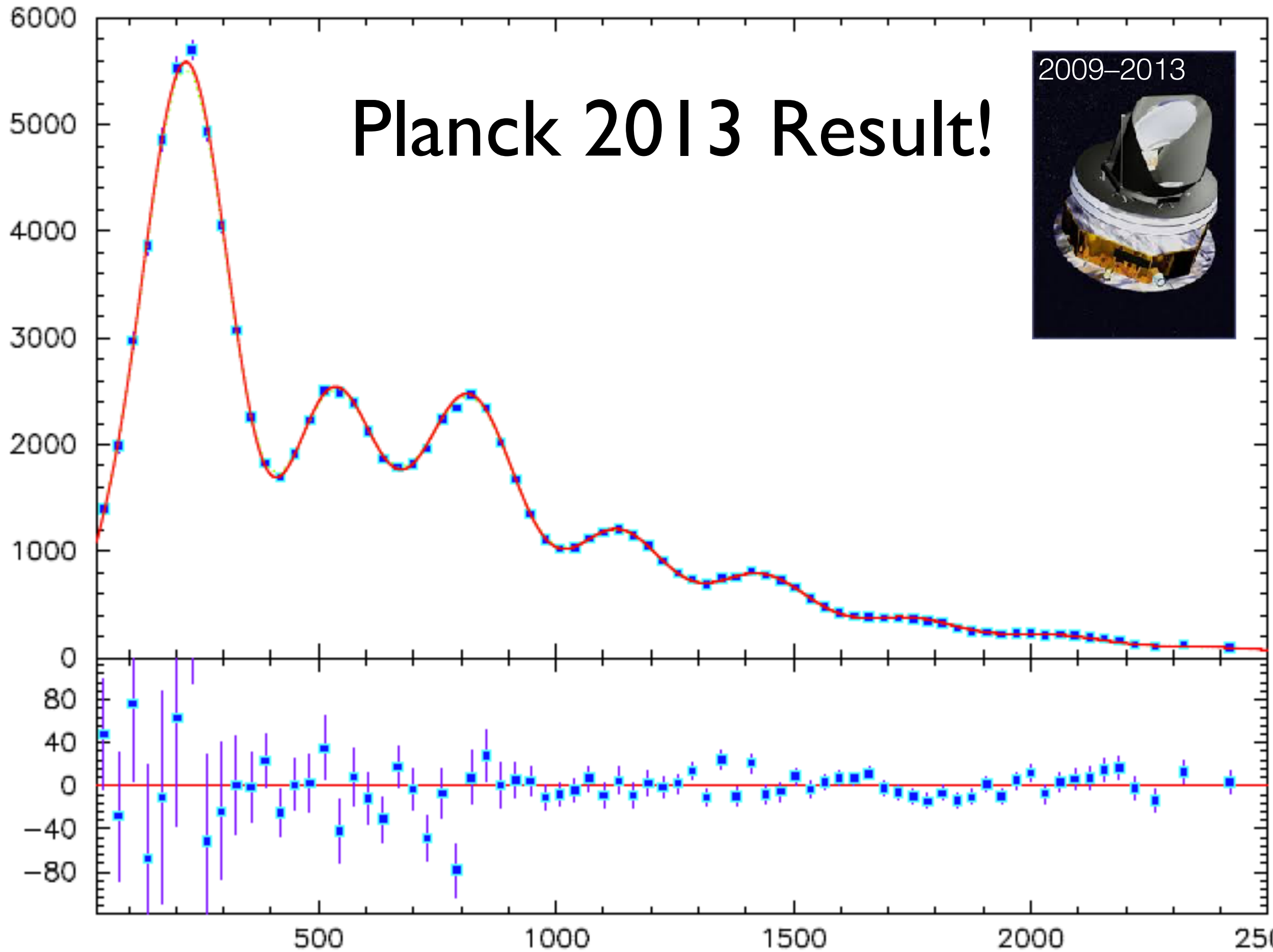
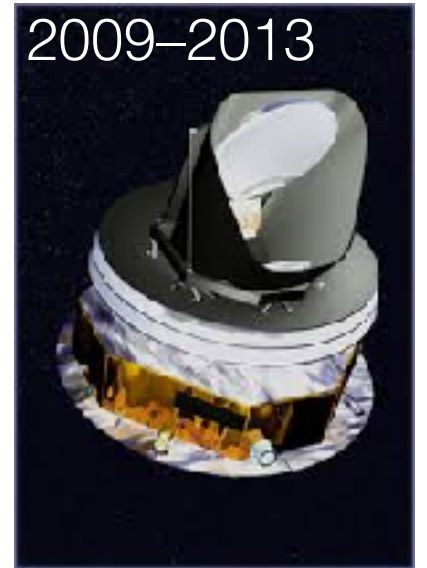




Residual Amplitude of Waves [ $\mu\text{K}^2$ ]

# Planck 2013 Result!

2009–2013



$l$  80 degrees/(angle in the sky)

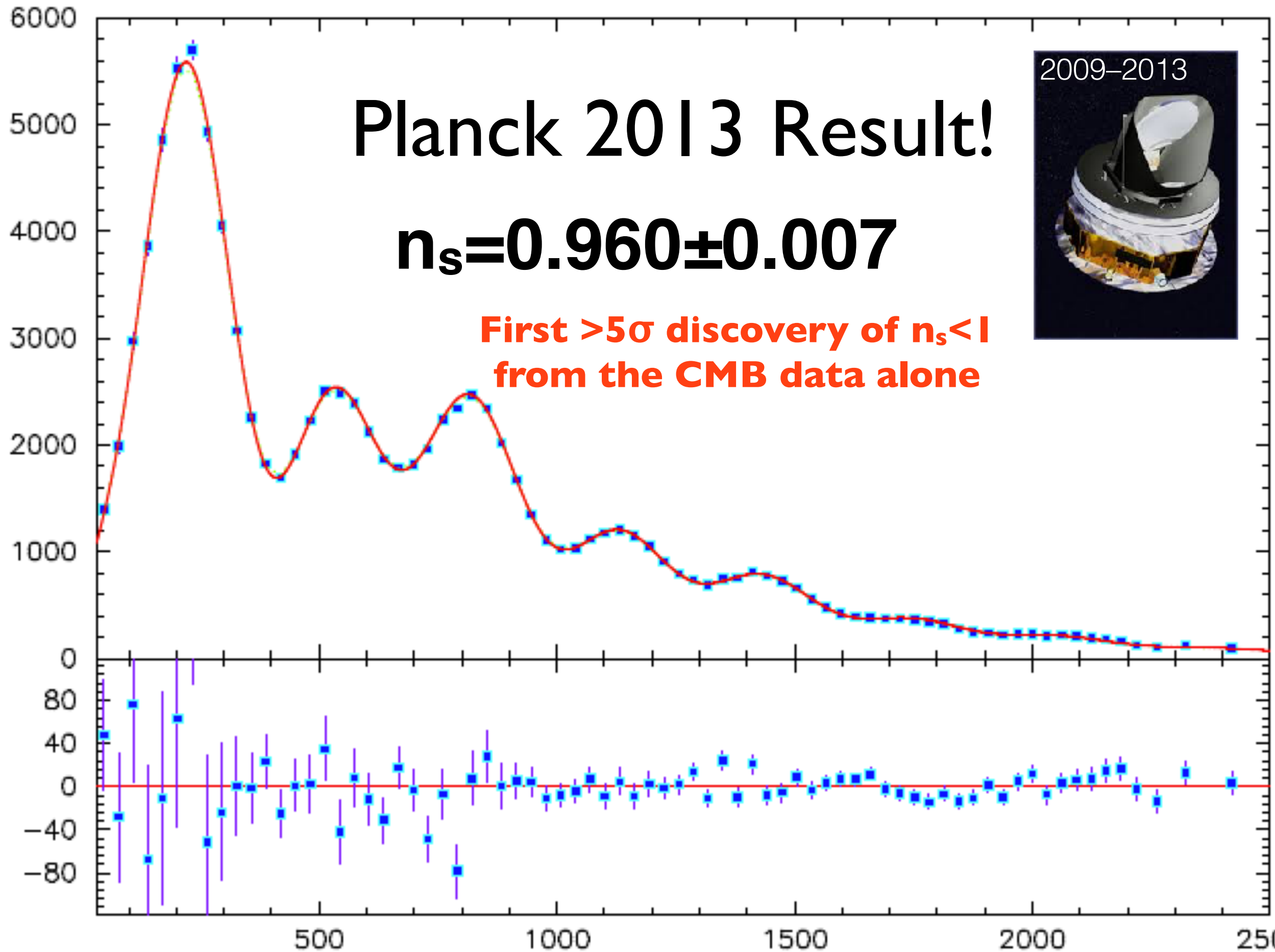
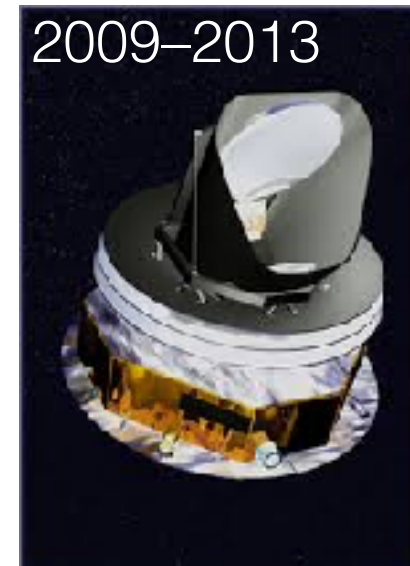
Residual Amplitude of Waves [ $\mu\text{K}^2$ ]

# Planck 2013 Result!

$$n_s = 0.960 \pm 0.007$$

**First  $>5\sigma$  discovery of  $n_s < 1$   
from the CMB data alone**

2009–2013



$l$  80 degrees/(angle in the sky)

Predicted in 1981.  
Finally discovered in 2013  
by WMAP and Planck

- Inflation must end
- Inflation predicts  $n_s \sim 1$ , but not exactly equal to 1. Usually  $n_s < 1$  is expected
- **The discovery of  $n_s < 1$  has been the dream of cosmologists since 1992,** when the CMB anisotropy was first discovered and  $n_s \sim 1$  (to within 30%) was indicated

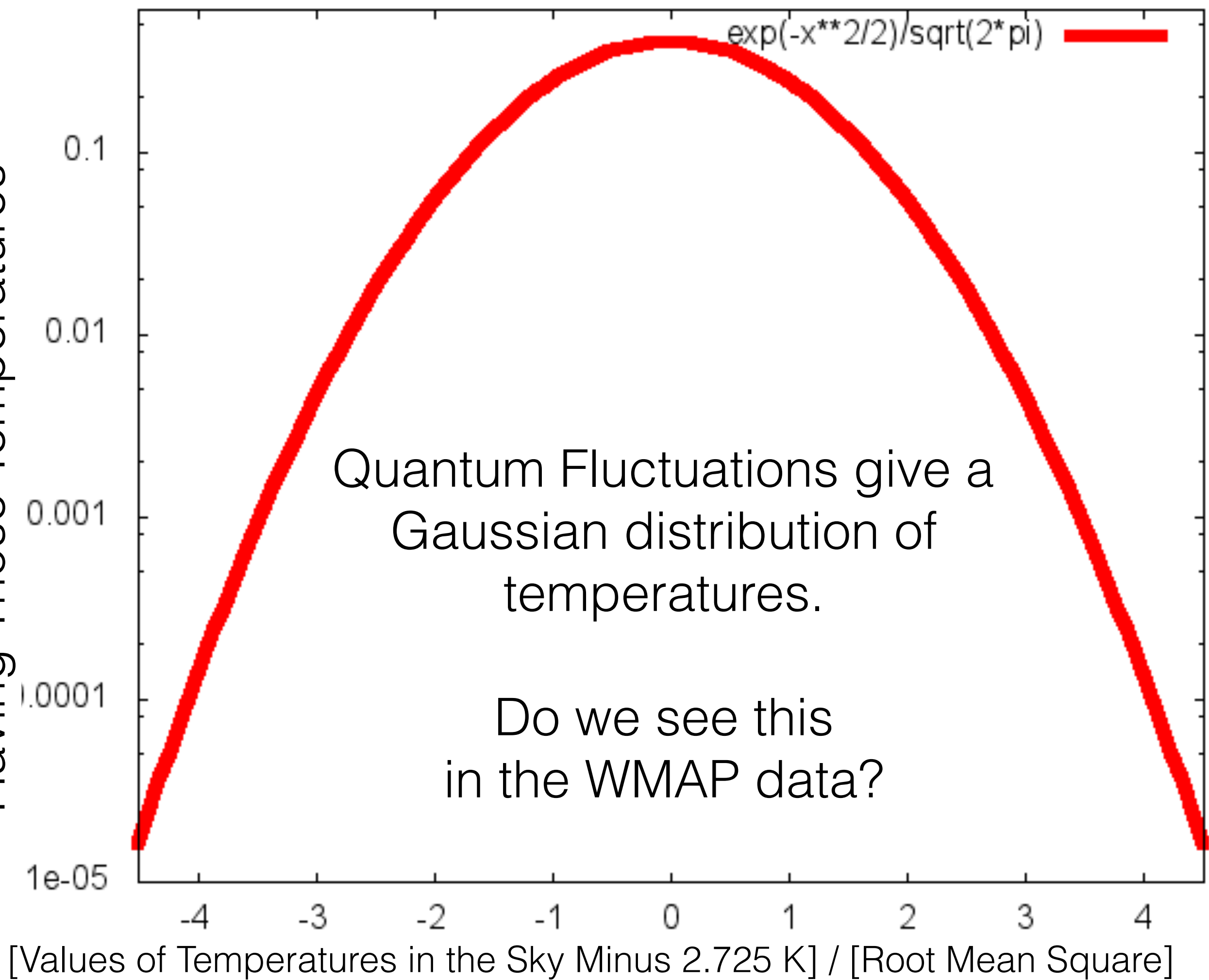


*Slava Mukhanov said in his 1981 paper that  $n_s$  should be less than 1*



How do we know that  
primordial fluctuations were of  
*quantum mechanical origin?*

Fraction of the Number of Pixels  
Having Those Temperatures



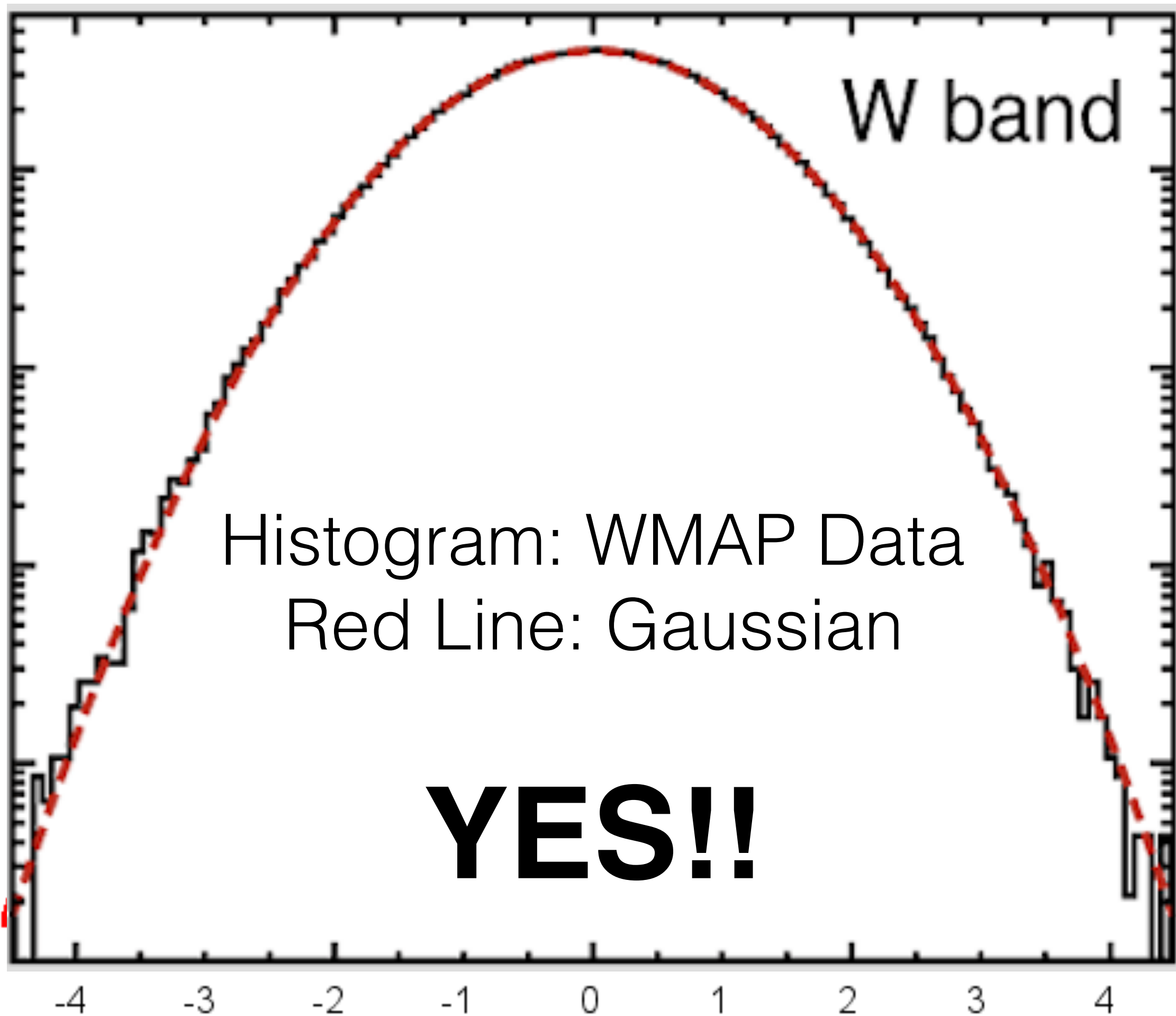
Fraction of the Number of Pixels  
Having Those Temperatures

W band

Histogram: WMAP Data  
Red Line: Gaussian

**YES!!**

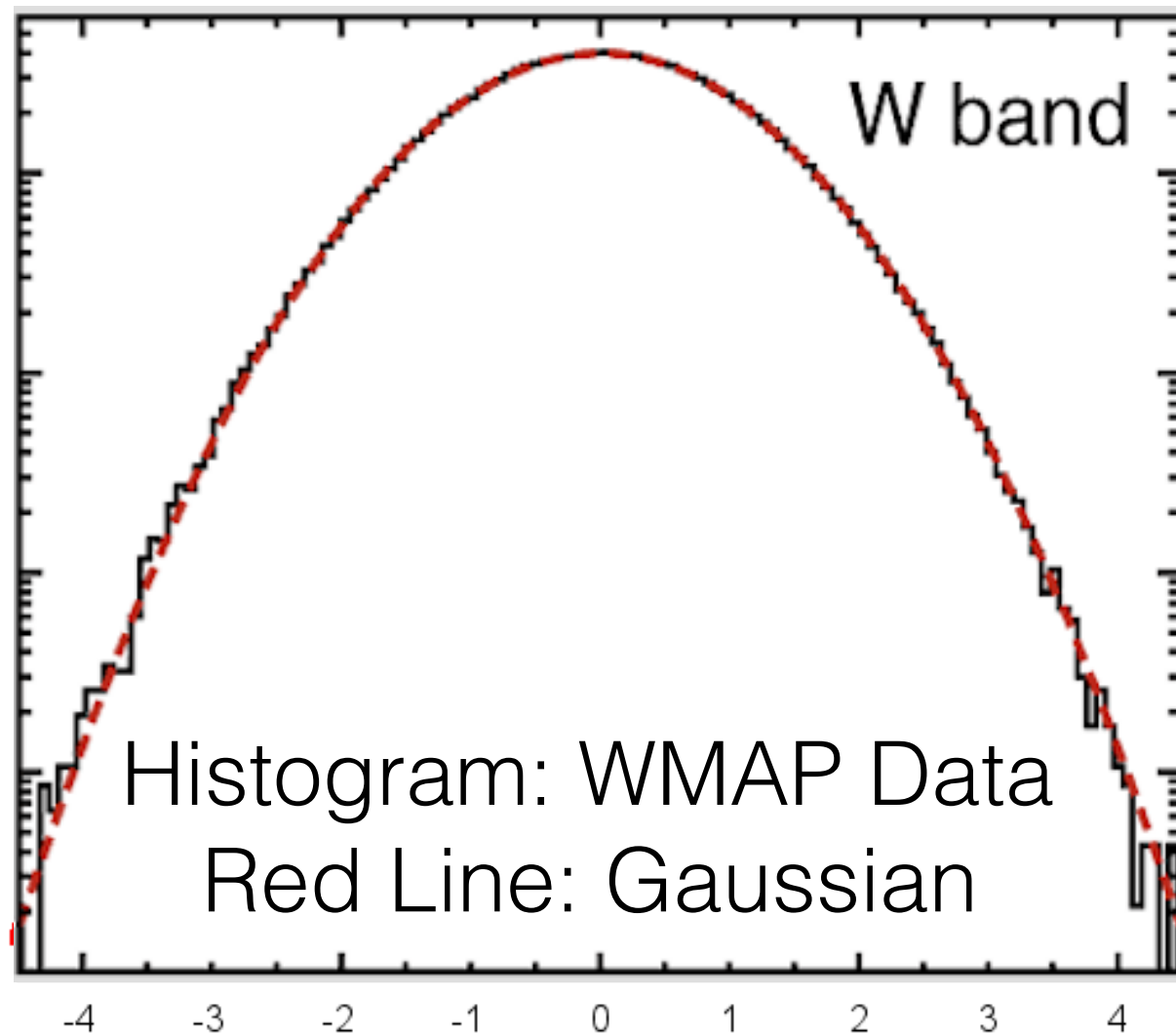
[Values of Temperatures in the Sky Minus 2.725 K] / [Root Mean Square]





# Testing Gaussianity

Fraction of the Number of Pixels  
Having Those Temperatures



[Values of Temperatures in the Sky Minus  
2.725 K]/ [Root Mean Square]

Since a Gauss distribution is symmetric, it must yield a vanishing **3-point function**

$$\langle \delta T^3 \rangle \equiv \int_{-\infty}^{\infty} d\delta T P(\delta T) \delta T^3$$

More specifically, we measure this using temperatures at three different locations and average:

$$\langle \delta T(\hat{n}_1) \delta T(\hat{n}_2) \delta T(\hat{n}_3) \rangle$$

# *Non-Gaussianity:*

A Powerful Test of Quantum Fluctuations

- The WMAP data show that the distribution of temperature fluctuations of CMB is **very precisely** Gaussian
  - with an upper bound on a deviation of **0.2%**
- With improved data provided by the Planck mission, the upper bound is now **0.03%**

# CMB Research: Next Frontier

## Primordial Gravitational Waves

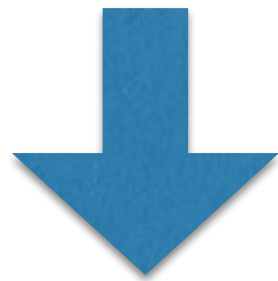
*Extraordinary claims require extraordinary evidence.  
The same quantum fluctuations could also generate  
gravitational waves, and we wish to find them*



# Measuring GW

- GW changes the distances between two points

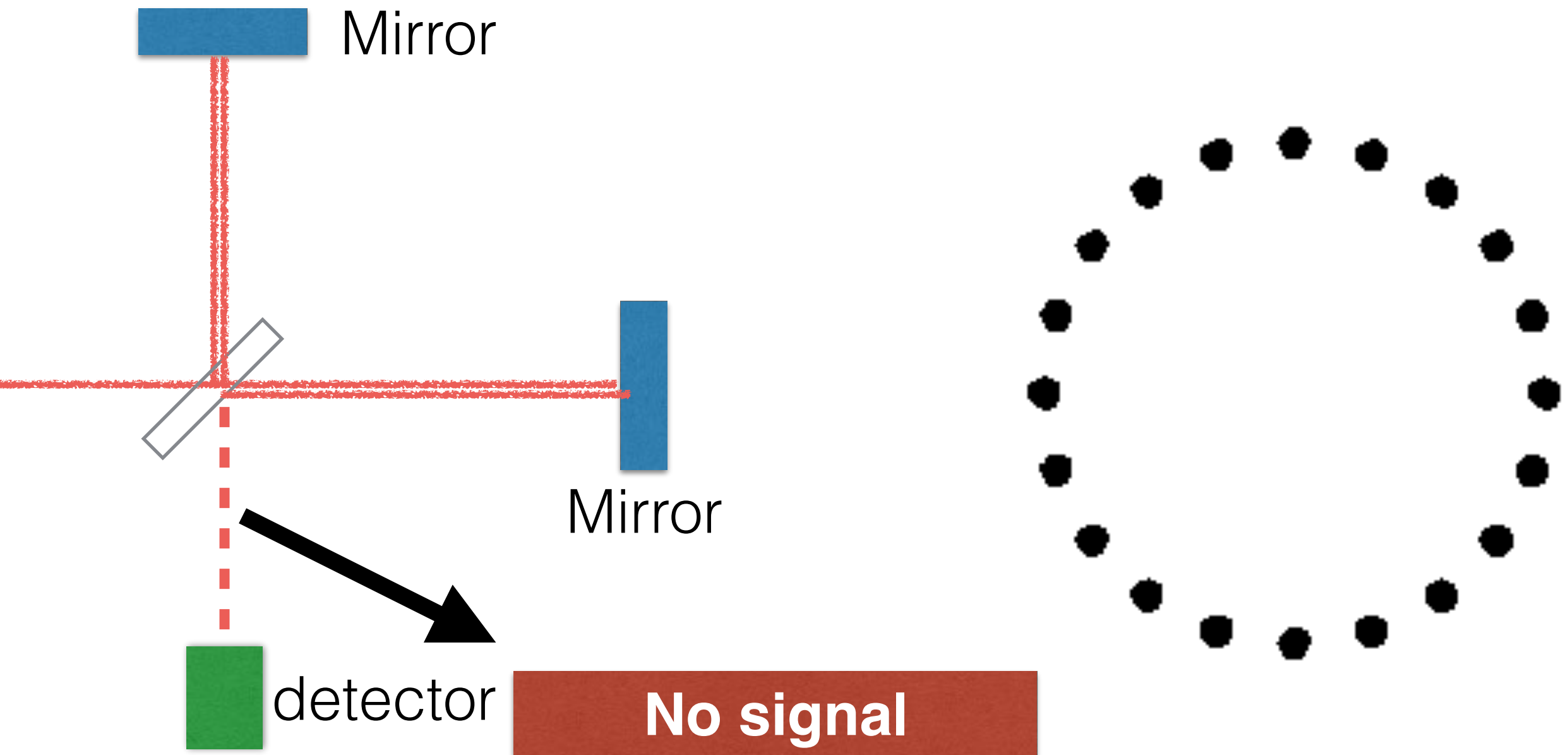
$$d\ell^2 = d\mathbf{x}^2 = \sum_{ij} \delta_{ij} dx^i dx^j$$



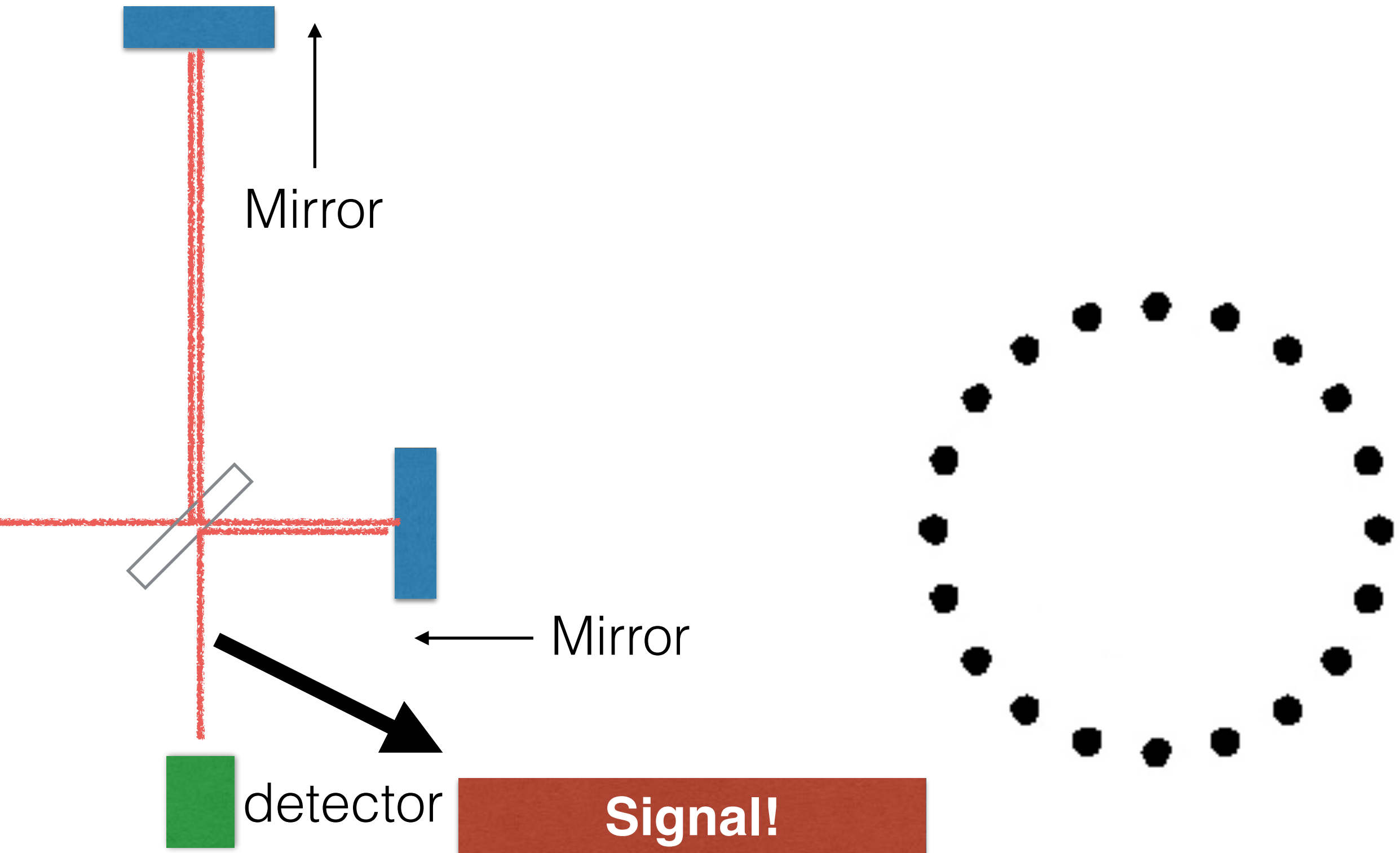
$$d\ell^2 = \sum_{ij} (\delta_{ij} + \textcolor{red}{h_{ij}}) dx^i dx^j$$



# Laser Interferometer

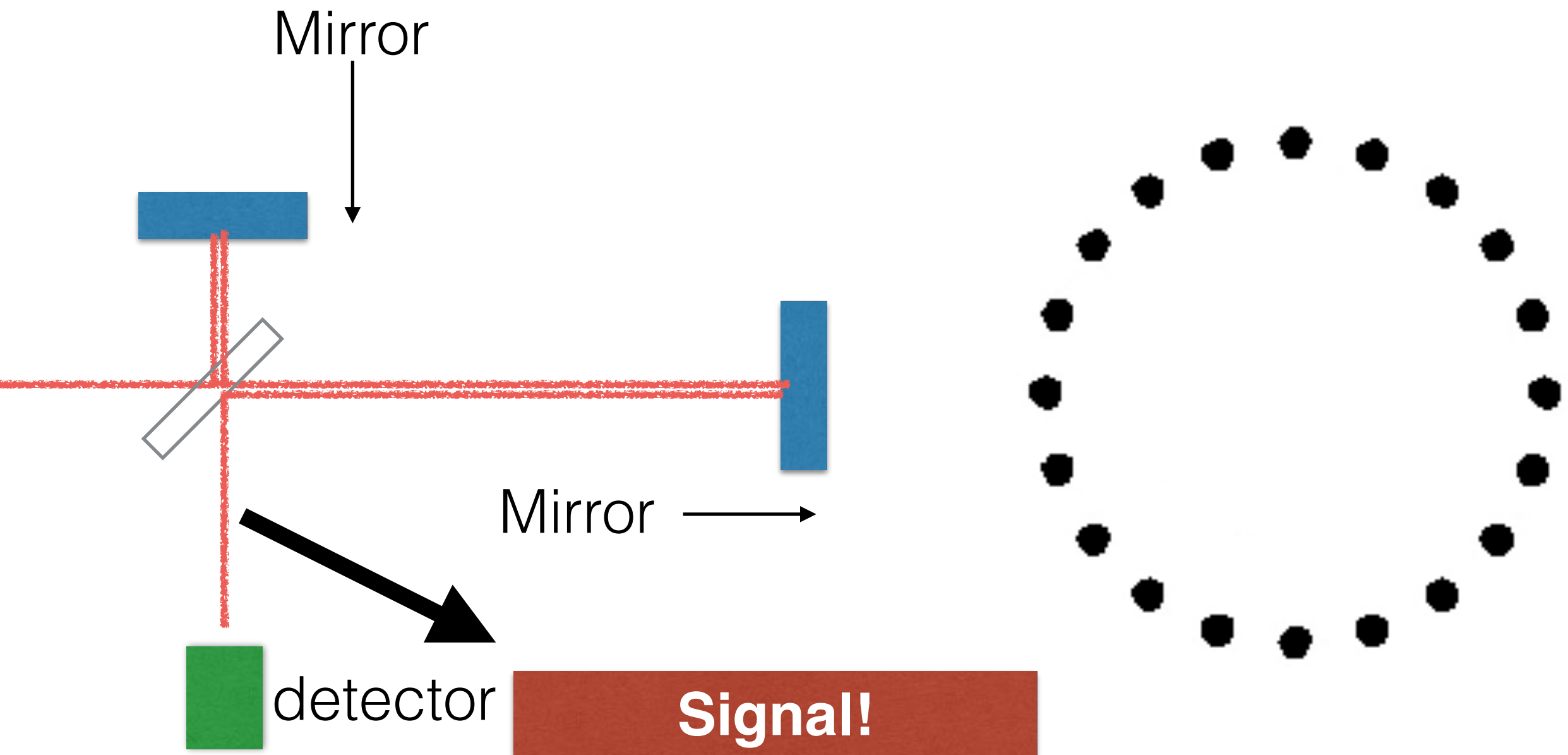


# Laser Interferometer





# Laser Interferometer



LIGO detected GW from binary blackholes, with the wavelength of thousands of kilometres

But, the primordial GW affecting the CMB has a wavelength of **billions of light-years!!** How do we find it?

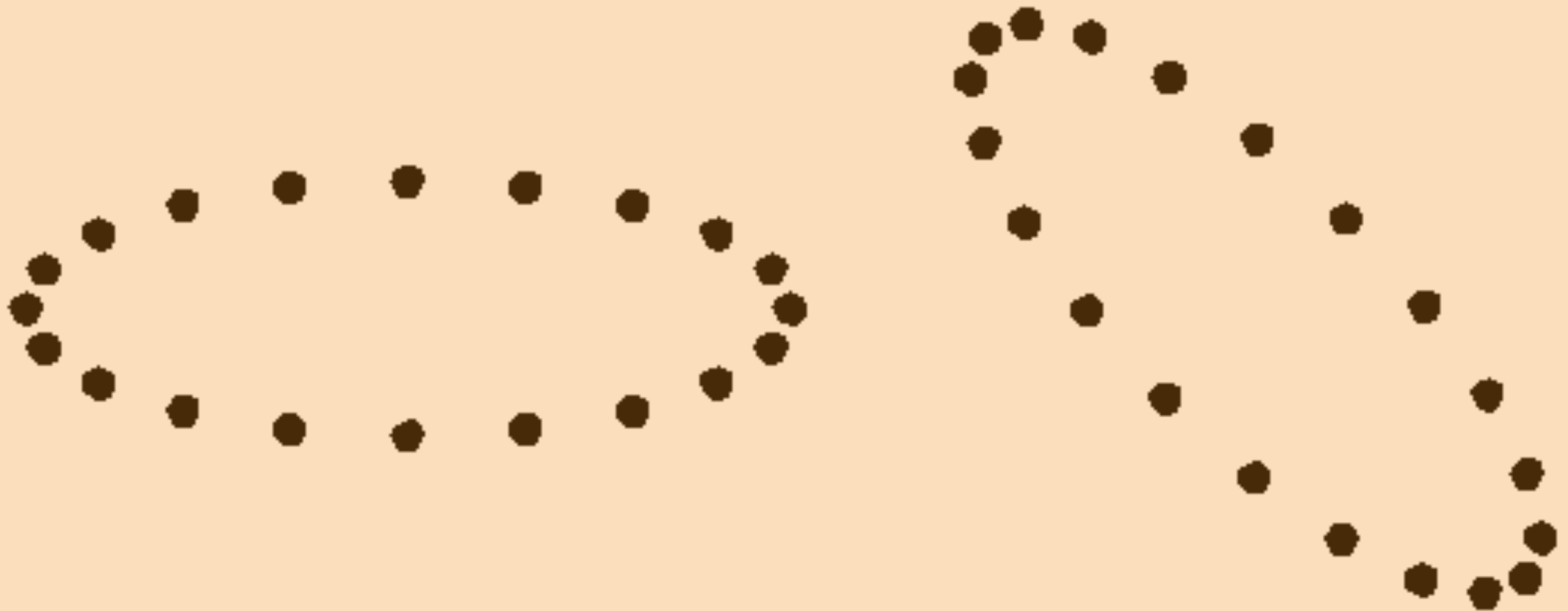
# Detecting GW by CMB

Isotropic electro-magnetic fields



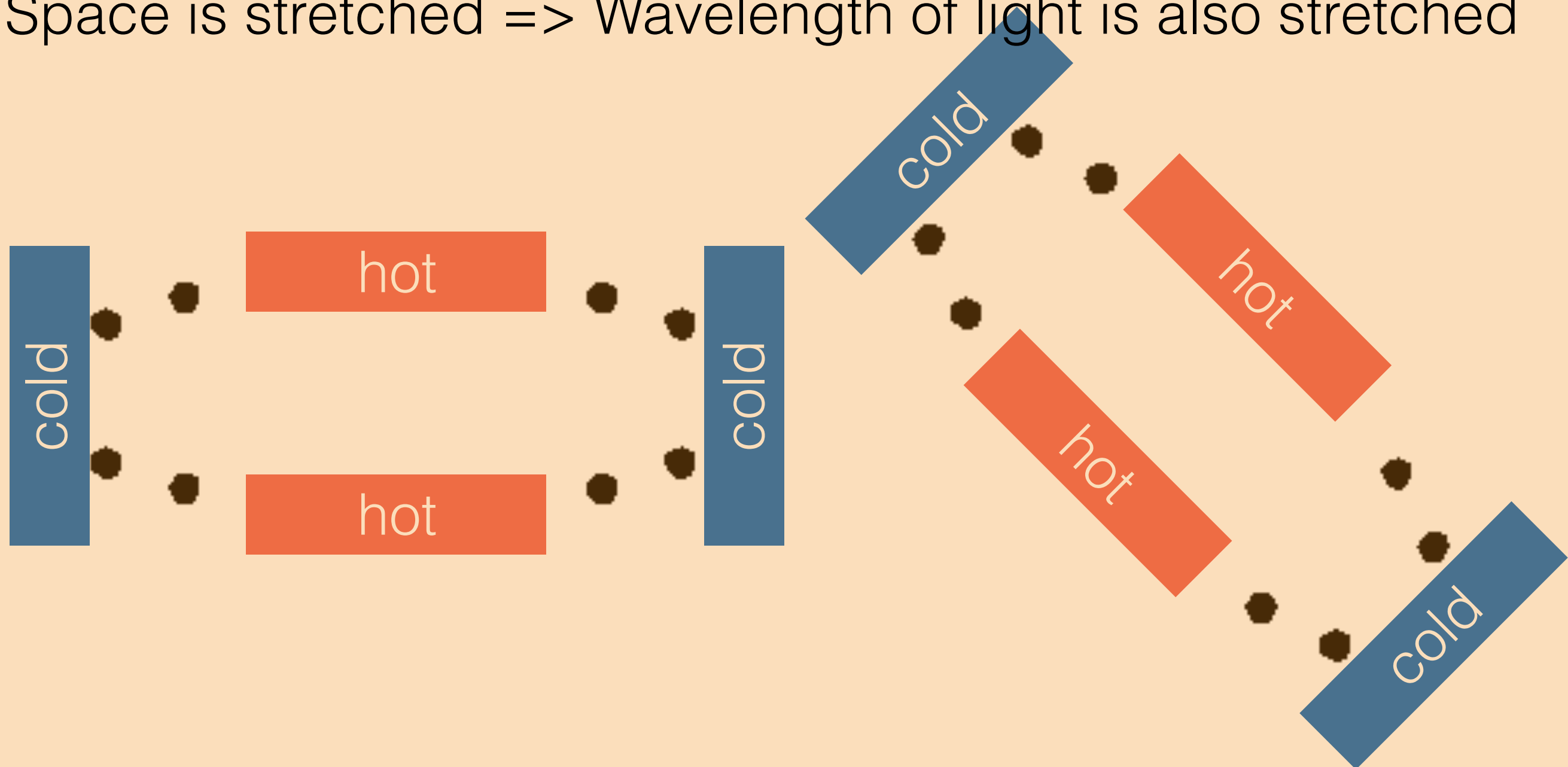
# Detecting GW by CMB

GW propagating in isotropic electro-magnetic fields



# Detecting GW by CMB

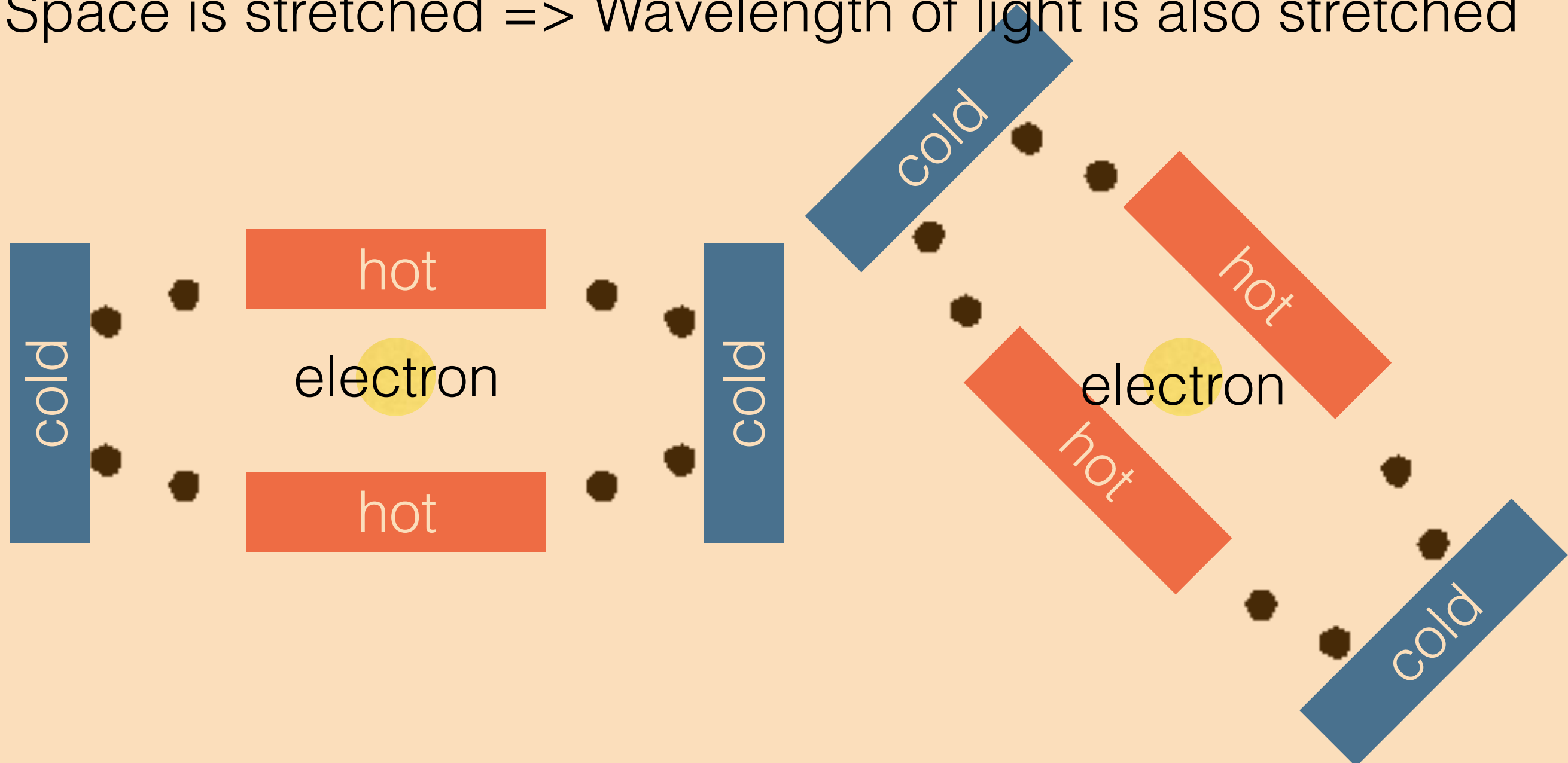
Space is stretched => Wavelength of light is also stretched



# Detecting GW by CMB

## Polarisation

Space is stretched => Wavelength of light is also stretched





# Detecting GW by CMB

## Polarisation

Space is stretched => Wavelength of light is also stretched

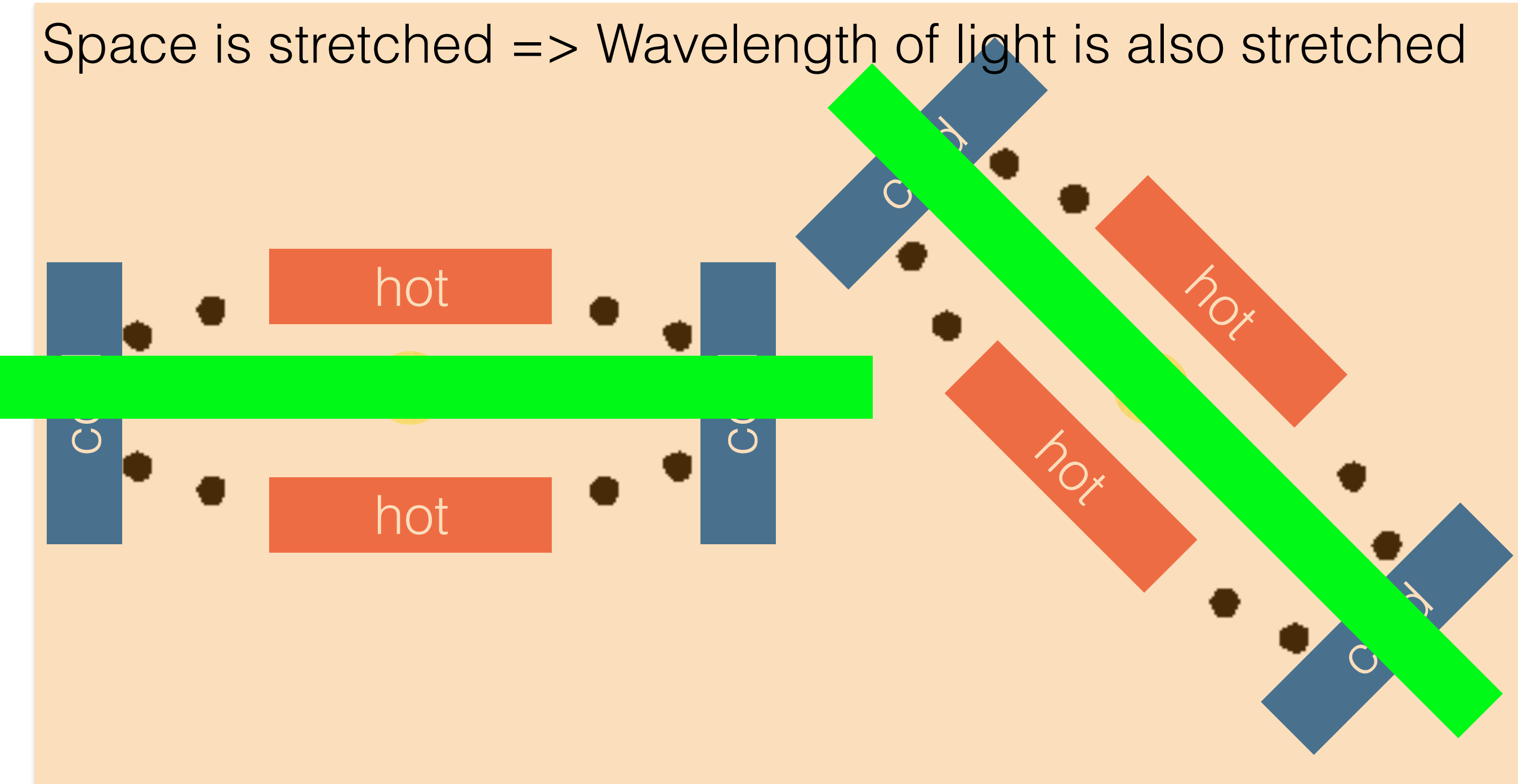


Photo Credit: TALEX



horizontally polarised



Photo Credit: TALEX





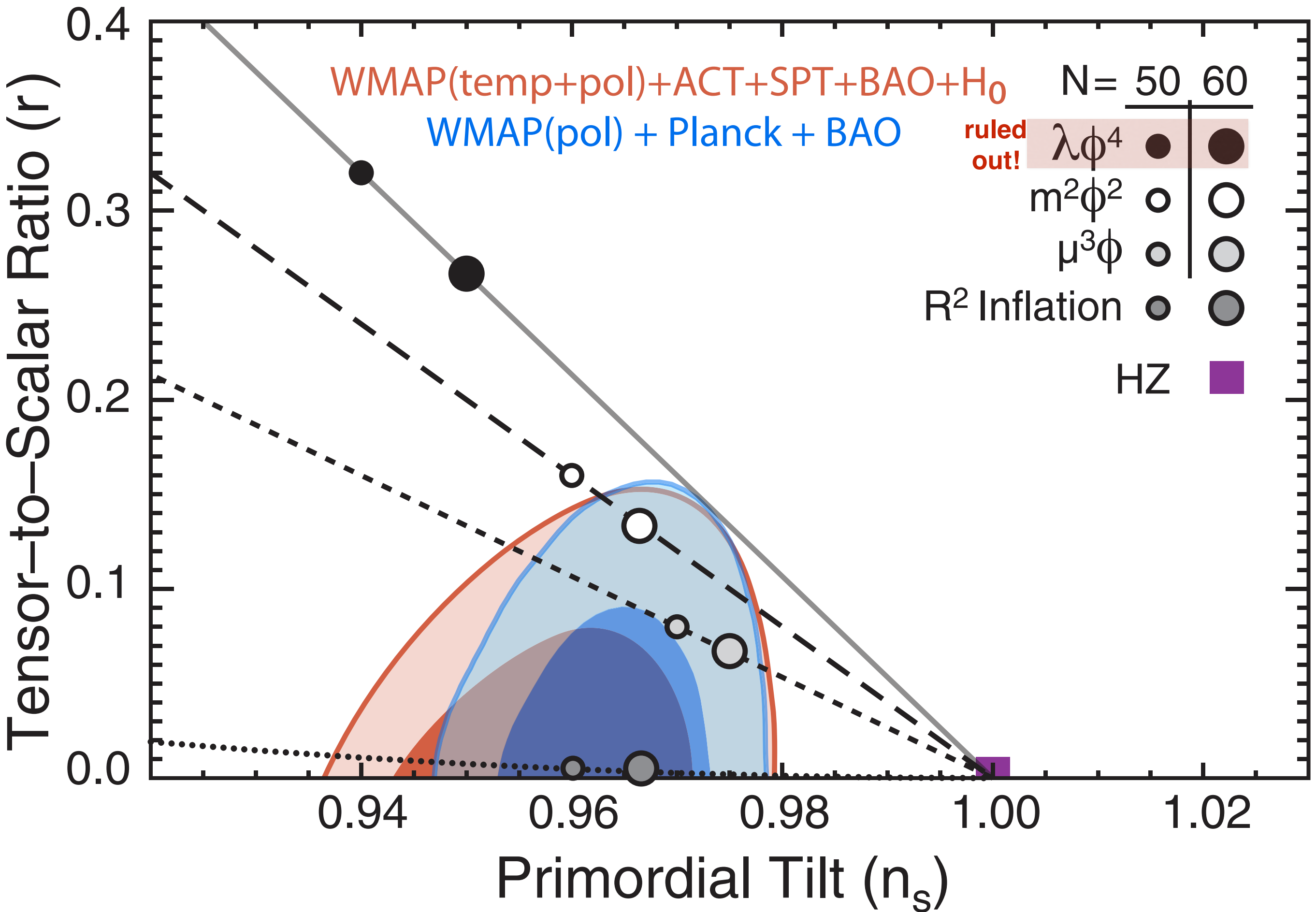
# If polarisation from GW is found...

- Then what?
- The next step is to nail the specific model of inflation

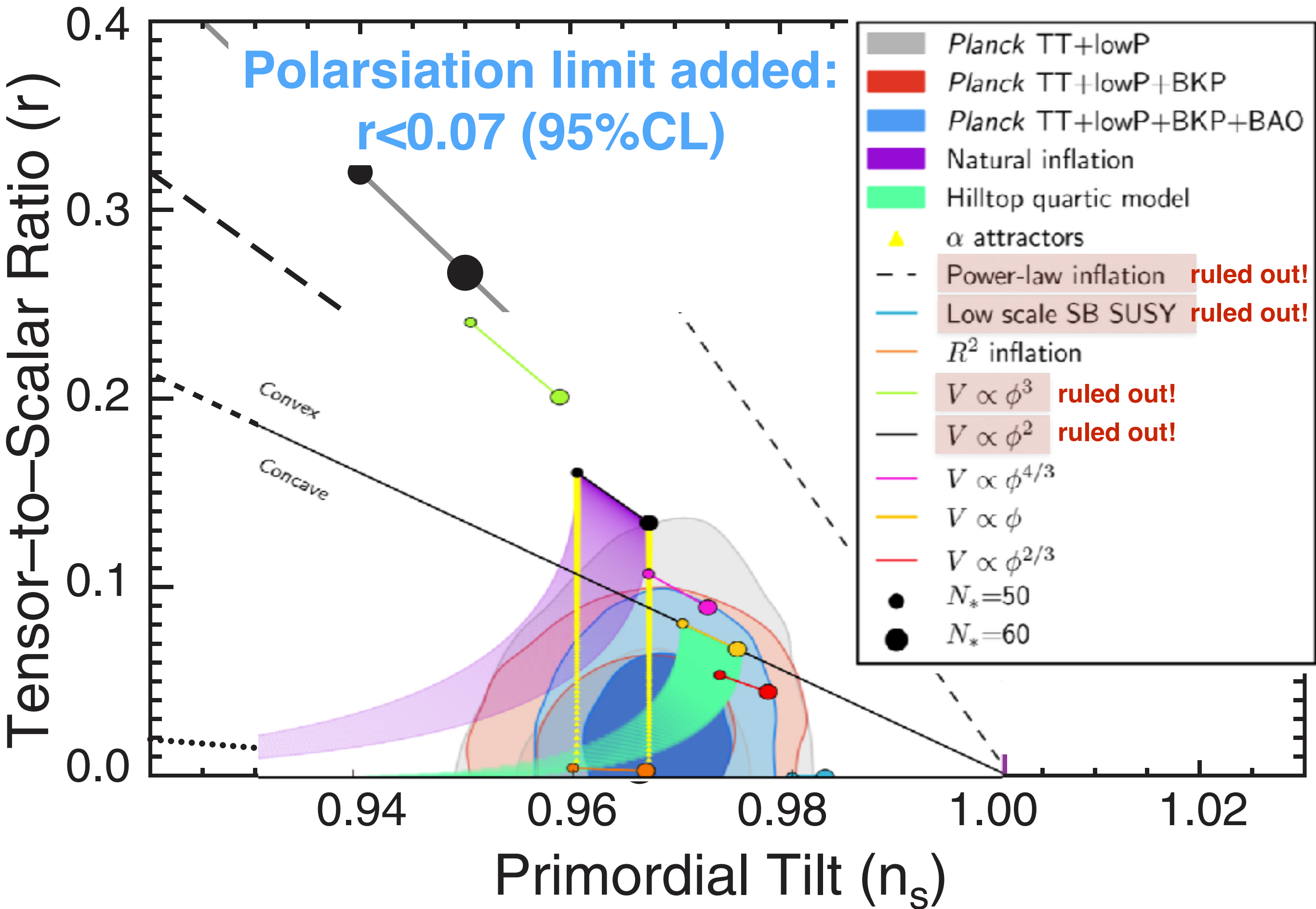
# Tensor-to-scalar Ratio

$$r \equiv \frac{\langle h_{ij} h^{ij} \rangle}{\langle \zeta^2 \rangle}$$

- We really want to find this quantity!  
**The current upper bound:  $r < 0.07$**







# March 17, 2014

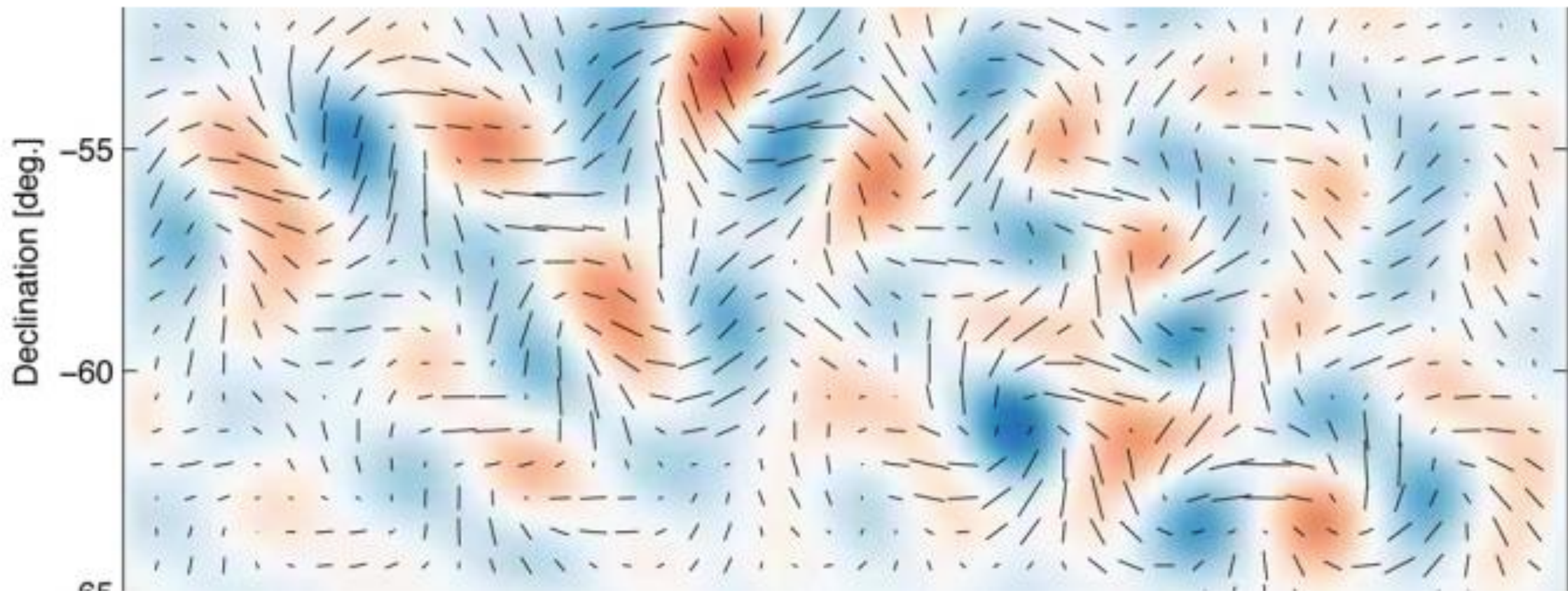
BICEP2's announcement



## First Direct Evidence of Cosmic Inflation

**Release No.:** 2014-05

**For Release:** Monday, March 17, 2014 - 10:45am



**Cambridge, MA -** Almost 14 billion years ago, the universe we inhabit burst into existence in an extraordinary event that initiated the Big Bang. In the first fleeting fraction of a second, the universe expanded exponentially, stretching far beyond the view of our best telescopes. All this, of course, was just theory.



SPACE & COSMOS

The New York Times

# Space Ripples Reveal Big Bang's Smoking Gun

By DENNIS OVERBYE MARCH 17, 2014

BBC

News

Sport

Weather

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17 March 2014 Last updated at 14:46 GMT

Share



## Cosmic inflation: 'Spectacular' discovery hailed

By Jonathan Amos

Science correspondent, BBC News



Cambridge, MA - Almost 14 billion years ago, a flash of light that initiated the Big Bang. In the far beyond the view of our best tel

Süddeutsche.de

Wissen

Politik Panorama Kultur Wirtschaft Sport München Bayern Digital Auto Reise Video

Home > Wissen > Urknall > Urknall - Gravitationswellen belegen inflationäres Universum

Süddeutsche.de als Startseite einrichten

Hir

17. März 2014, 17:34 Gravitationswellen

## Signale aus der Geburtsstunde des Universums

Von Patrick Illinger

# January 30, 2015

Joint Analysis of BICEP2 data and Planck data

# Speck of Interstellar Dust Obscures Glimpse of Big Bang

By DENNIS OVERBYE JAN. 30, 2015

**BBC**

News Sport Weather Earth Future Shop

**NEWS** SCIENCE & ENVIRONMENT

Home UK Africa Asia Australia Europe Latin America Mid-East US & Canada Business Health S

30 January 2015 Last updated at 20:54 GMT



## Cosmic inflation: New study says BICEP claim was wrong

By Jonathan Amos  
Science correspondent, BBC News

Süddeutsche.de

Wissen

Politik Panorama Kultur Wirtschaft Sport München Bayern Digital Auto Reise Video

Home > Wissen > Kosmologie - Urknall-Forscher gestehen Irrtum ein

[Süddeutsche.de als Startseite einrichten](#)

Hlr

1. Februar 2015, 22:19 Kosmologie

## Urknall-Forscher gestehen Irrtum ein

Von Marlene Weiß



# Current Situation

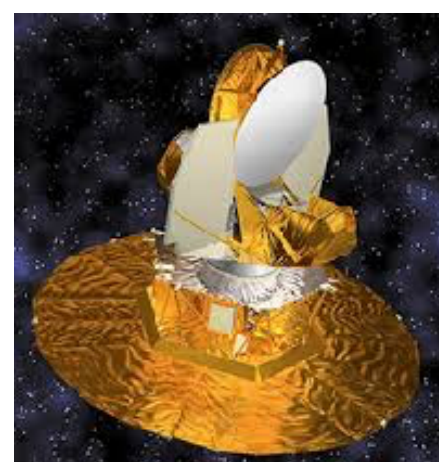
- Planck shows the evidence that the detected signal is not cosmological, but is due to dust
- No strong evidence that the detected signal is cosmological



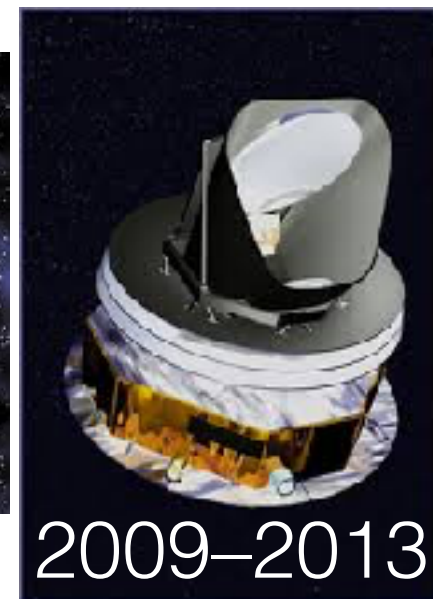
**The search continues!!**



1989–1993



2001–2010



2009–2013



202X–



# JAXA

+ possibly NASA

## LiteBIRD

2025– [proposed]

Target  $1\sigma$  uncertainty:  
 $\delta r = 10^{-3}$

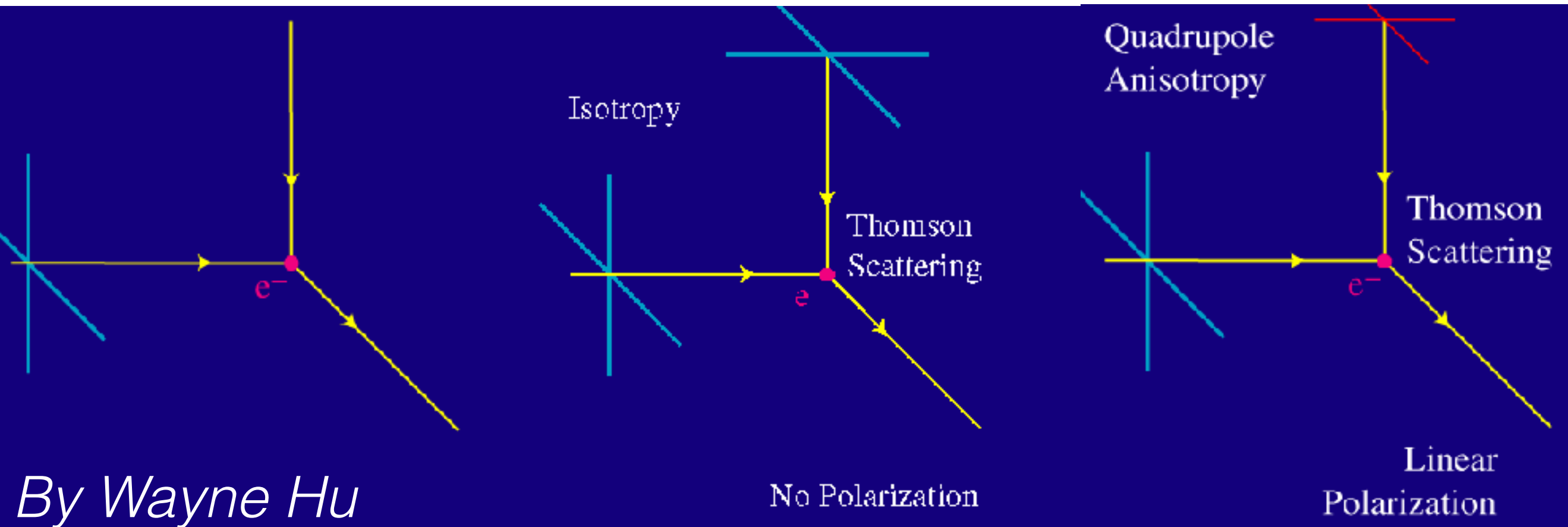


# Conclusion

- The WMAP and Planck's temperature data provide **strong evidence for the quantum origin of structures in the universe**
- The next goal: unambiguous measurement of polarisation from gravitational waves
- **LiteBIRD** proposal: a CMB polarisation satellite in 2025



# Physics of CMB Polarisation



- Necessary and sufficient conditions for generating polarisation in CMB:
  - Thomson scattering
  - Quadrupolar temperature anisotropy around an electron