

Parity Violation in Cosmology In search of new physics for the Universe The lecture slides are available at https://wwwmpa.mpa-garching.mpg.de/~komatsu/ lectures--reviews.html

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Topics From the syllabus

- 1. What is parity symmetry?
- **2. Chern-Simons interaction**
- 3. Parity violation 1: Cosmic inflation
- 4. Parity violation 2: Dark matter
- 5. Parity violation 3: Dark energy
- 6. Light propagation: birefringence
- 7. Physics of polarization of the cosmic microwave background
- 8. Recent observational results, their implications, and future prospects





2.1 Parity Symmetry in Electromagnetism (EM)

Maxwell's Equations In Heaviside units and c=1

$$abla \cdot \mathbf{E} =
ho \,, \quad -\dot{\mathbf{E}}$$
 $abla \cdot \mathbf{B} = 0 \,, \quad \dot{\mathbf{B}}$

- These equations are invariant under both spatial translation and rotation. They are also invariant under parity transformation, if E and j are vectors, ρ is
- a scalar, and **B** is a pseudovector.

$+ \nabla \times \mathbf{B} = \mathbf{j}$ $+ \nabla \times \mathbf{E} = 0$

Parity-flipping Maxwell's Equations In Heaviside units and c=1

$$(-
abla) \cdot (-\mathbf{E}) =
ho \,, \quad -(-(-
abla)) \cdot \mathbf{B} = 0 \,,$$

- a scalar, and **B** is a pseudovector.

$-\mathbf{E}) + (-\nabla) \times \mathbf{B} = (-\mathbf{j})$ $\dot{\mathbf{B}} + (-\nabla) \times (-\mathbf{E}) = 0$

 These equations are invariant under both spatial translation and rotation. • They are also invariant under parity transformation, if **E** and **j** are vectors, ρ is

Parity-flipping Maxwell's Equations In Heaviside units and c=1

$$(-\nabla) \cdot (-\mathbf{E}) = \rho, \quad -(-\nabla) \cdot \mathbf{B} = 0,$$

If there is a minimum time to be

 They are also invariant under parity a scalar, and B is a pseudovector.

$-\dot{\mathbf{E}}) + (-\nabla) \times \mathbf{B} = (-\mathbf{j})$ $\dot{\mathbf{B}} + (-\nabla) \times (-\mathbf{E}) = 0$

nagnetic monopole, a pseudoscalar!

They are also invariant under parity transformation, if E and j are vectors, ρ is

Simplifying Maxwell's Equations Let's go 4D.







Antisymmetric Field Strength Tensor, F^{µv} $F^{\mu\nu} = -F^{\nu\mu}$



• Equivalently,

$$F^{0i} = E_i$$
$$F^{ij} = \epsilon^{ijk} B_k$$

 $F^{12} = B_z, F^{23} = B_x, F^{31} = B_y$ $F^{21} = -B_z, F^{32} = -B_x, F^{13} = -B_y$ symbol

 $\epsilon^{ijk} = \begin{cases} +1 \text{ if (i,j,k) is even permutation of (1,2,3)} \\ -1 \text{ if (i,j,k) is odd permutation of (1,2,3)} \\ 0 \text{ or the set} \end{cases}$ otherwise $\epsilon^{123} = 1, \epsilon^{132} = -1, \epsilon^{312} = 1, \dots$



Antisymmetric Field Strength Tensor, F_{µv} $F_{\mu\nu} = -F_{\nu\mu}$

$$F_{\mu
u} = \eta_{\mulpha}\eta_{
ueta}F^{lphaeta}$$
 whe



Therefore,

 $\equiv F_{\mu\nu}F^{\mu\nu} = 2(\mathbf{B}\cdot\mathbf{B} - \mathbf{E}\cdot\mathbf{E})$ This is a scalar and is invariant under parity transformation.

ere $\eta_{\mu\alpha} = diag(-1, 1, 1, 1)$



Dual Field Strength T $\widetilde{F}^{\mu\nu} = -\widetilde{F}^{\nu\mu}$ $\widetilde{F}^{\mu\nu} = \frac{1}{2} \epsilon^{\mu\nu\alpha\beta} F_{\alpha\beta} \quad \text{where} \quad \epsilon_{L}$



• Therefore,

 $F\bar{F} \equiv F_{\mu\nu}\tilde{F}^{\mu\nu} = -4\mathbf{B}\cdot\mathbf{E}$

$$e^{\mu\nu\alpha\beta} = \begin{cases} +1 & \text{if } (\mu,\nu,\alpha,\beta) \text{ is even } p \\ +1 & \text{of } (0,1,2,3) \\ -1 & \text{if } (\mu,\nu,\alpha,\beta) \text{ is odd } p \\ -1 & \text{if } (\mu,\nu,\alpha,\beta) \text{ is odd } p \\ 0 & \text{of } (0,1,2,3) \\ 0 & \text{otherwise} \end{cases}$$

$$B_{y} \quad B_{z} \\ B_{z} \quad E_{y} \\ 0 & -E_{x} \\ D_{x} \quad 0 \end{pmatrix} \bullet \text{Equivalently,}$$

$$\tilde{F}^{0i} = B_{i} \\ \tilde{F}^{ij} = -\epsilon^{ijk}$$

This is a *pseudoscalar* and changes sign under parity transformation!







2.2 Action Principle for EM

In Heaviside units and c=1

• The answer is

 $I=-rac{1}{\Lambda}\int d^4x\;F^2+\int d^4x\;A_\mu j^\mu \quad d^4x=dtd^3{f x}$

with

Vector potential $F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$ where $A_{\mu} = (-\phi, \mathbf{A})$ Therefore, $\begin{cases} F_{i0} = \partial_i A_0 - \dot{A}_i = E_i \\ F_{ij} = \partial_i A_j - \partial_j A_i = \epsilon_{ijk} B_k \end{cases} \Rightarrow \mathbf{E} = -\nabla \phi - \dot{\mathbf{A}} \\ \mathbf{B} = \nabla \times \mathbf{A} \end{cases}$

$\partial_{\nu}F^{\mu\nu} = j^{\mu}, \quad \partial_{\nu}\tilde{F}^{\mu\nu} = 0$ What is the action that gives Maxwell's equations?





In Heaviside units and c=1

• The answer is

$$I = -\frac{1}{4} \int d^4x \ F^2 + \int$$

with

 $F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$ where $A_{\mu} = (-\phi, \mathbf{A})$

One set of Maxwell's equations is simply given by the definition of $F_{\mu\nu}$: $\partial_{\nu}\tilde{F}^{\mu\nu} = \partial_{\nu}\frac{1}{2}\epsilon^{\mu\nu\alpha\beta}(\partial_{\alpha}A_{\beta} - \partial_{\beta}A_{\alpha}) = \epsilon^{\mu\nu\alpha\beta}\partial_{\nu}\partial_{\alpha}A_{\beta} = 0$

$\partial_{\nu}F^{\mu\nu} = j^{\mu}, \quad \checkmark \partial_{\nu}\tilde{F}^{\mu\nu} = 0$ What is the action that gives Maxwell's equations?

 $d^4x A_{\mu}j^{\mu} d^4x = dtd^3\mathbf{x}$





In Heaviside units and c=1

$$I = -\frac{1}{4} \int d^4x \ F^2 + \int$$

stationary point. For a small change in $A_{\mu} \rightarrow A_{\mu} + \delta A_{\mu}$, the corresponding change in $I \rightarrow I + \delta I$ is also small.

$$\delta I = \int d^4 x \; F^{\mu
u} \partial_
u (\delta A_\mu) + \int \mathbf{Integration} \ \int d^4 x \; (-\partial_
u F^{\mu
u} + j^\mu) \delta .$$



 $\int d^4x A_{\mu} j^{\mu}$

• The idea: The equation of motion for A_{μ} is the path that gives a

 $\int d^4x \ (\delta A_{\mu}) j^{\mu} \qquad \stackrel{\text{Hint:}}{\stackrel{\delta}{\delta}(F^2)} = 2F^{\mu\nu} \delta F_{\mu\nu} \\ = -4F^{\mu\nu} \partial_{\nu} (\delta A_{\mu})$ $A_{\mu} = 0 \implies \partial_{\nu} F^{\mu\nu} = j^{\mu}$



Finding symmetries in the action It is like a treasure hunt!

$$I = -\frac{1}{4} \int d^4x \ F^2 + \int$$

- This action is invariant under spatial translation, rotation, and parity transformation.
- It is also invariant under the following "gauge transformation",

$$A_{\mu} \to A_{\mu} + \partial_{\mu} f$$

• Here, f is an arbitrary scalar function.

 $\int d^4x A_{\mu} j^{\mu}$

Integration by parts Hint: $\int d^4x \ (\partial_\mu f) j^\mu \stackrel{\bullet}{=} - \int d^4x \ f \partial_\mu j^\mu = 0$ due to the charge conservation: $\partial_{\mu}j^{\mu} = 0 \implies \dot{\rho} + \nabla \cdot \mathbf{j} = 0$ 15



Finding symmetries in the action It is like a treasure hunt!

$$I = -\frac{1}{4} \int d^4x \ F^2 + \int$$

- This action is invariant under spatial translation, rotation, and parity transformation.

$$\phi \rightarrow \phi - f$$

 It is also invariant under the following "gauge transformation", Integration by parts Hint: $\int d^4x \ (\partial_\mu f) j^\mu = - \int d^4x \ f \partial_\mu j^\mu = 0$ $\mathbf{A} \to \mathbf{A} + \nabla f$ • Here, f is an arbitrary scalar function. due to the charge conservation: $\partial_{\mu}j^{\mu} = 0 \implies \dot{\rho} + \nabla \cdot \mathbf{j} = 0$ 16

 $\int d^4x A_{\mu} j^{\mu}$



Problem Set 2 Playing with Maxwell

2. Show that FF is a total derivative and can be written as

$$F_{\mu\nu}\tilde{F}^{\mu\nu} = 2\partial_{\mu}(A_{\nu}\tilde{F}^{\mu\nu})$$

1. Derive Maxwell's equations from $\partial_{\nu}F^{\mu\nu} = j^{\mu}$, $\partial_{\nu}F^{\mu\nu} = 0$.

2.3 FF in the action

FF in the action?

$$I = -\frac{1}{4} \int d^4x \ F^2 + \int$$

- This action is sufficient to produce all of Maxwell's equations.
- Can we add $\int d^4x \; F ilde{F}$ to the action?
 - derivative (as shown in Problem Set 2).

 $\int d^4x A_{\mu} j^{\mu}$

• The answer is yes. However, this is only a surface term, since FF is a total

FF in the action **Chern-Simons term**

- - α: a dimensionless constant
 - θ : a dimensionless pseudoscalar field
- This is not a surface term! Integration by parts gives

$$I_{\rm CS} = \frac{1}{2} \alpha \int d^4 x \, ($$

Ni (1977); Turner, Widrow (1987); Carroll, Field, Jackiw (1990)

• Consider $I_{ m CS}=-rac{1}{4}lpha \int d^4x \; heta F ilde F$ with $F ilde F=2 \partial_\mu (A_ u ilde F^{\mu u})$

Why Chern-Simons Terms?

Panel Discussion of "Chern-Simons Terms" inciples of Theoretical Physics Occasion of Jim Simons' 87th Birthday

25. 2021 - CUNY Graduate Center

Organized by Dennis Sullivan

 $(\partial_{\mu}\theta)A_{\nu}\tilde{F}^{\mu\nu}$

Jim Simons in 2023

https://einstein-chair.github.io/simons2023/

- This is a special case of the so-called Chern-Simons term, $p_{\mu}A_{
u}F^{\mu
u}$





Carroll, Field, Jackiw (1990) **Consistency with gauge invariance** p_µ cannot be arbitrary

$$I_{\rm CS} = \frac{1}{2} \alpha \int d^4 x \ p_\mu A_\nu$$

- This action is invariant under the gauge transformation, $A_
 u o A_
 u + \partial_
 u f$ if $\partial_{\nu} p_{\mu} - \partial_{\mu} p_{\nu} = 0$ Hint: Use integration by parts and the identity $\partial_{\nu}F^{\mu\nu} = 0$
- For example:
 - p_{μ} is a constant vector and not dynamical.
 - p_{μ} is a gradient of a dynamical (pseudo)scalar field, such as $p_{\mu} = \partial_{\mu}\theta$.

 $F^{\mu
u}$







Carroll, Field, Jackiw (1990) **Consistency with gauge invariance** p_µ cannot be arbitrary

$$I_{\rm CS} = \frac{1}{2} \alpha \int d^4 x \ p_\mu A_\nu$$

- This action is invariant under the gauge transformation, $A_
 u o A_
 u + \partial_
 u f$ if $\partial_{\nu}p_{\mu} - \partial_{\mu}p_{\nu} = 0$ Hint: Use integration by parts and the identity $\partial_{\nu}F^{\mu\nu} = 0$
- For example: This implies the presence of a preferred direction in spacetime and violation of Lorentz invariance!
 - p_{μ} is a constant vector and not dynamical, or
 - p_{μ} is a gradient of a dynamical (pseudo)scalar field, such as $p_{\mu} = \partial_{\mu}\theta$.

 $F^{\mu
u}$







The main goals of this lecture series Let's find new physics!

- We will study the cosmological consequence of $I_{\rm CS} = -\frac{1}{4}\alpha \int d^4x \ \theta F\tilde{F}$
- Specifically, we ask if θ is
 - active during cosmic inflation,
 - responsible for dark matter, or
 - responsible for dark energy.



The main goals of this lecture series Let's find new physics!

- We will study the cosmological consequence of $I_{\rm CS} = -\frac{1}{4}\alpha \int d^4x \ \theta F \tilde{F}$
- - active during cosmic inflation,
 Cosmic microwave background,
 - responsible for dark matter, or Gravitational waves, and
 - responsible for dark energy.

• Specifically, we ask if θ is — • We will also study observational signatures in —

Large-scale structure of the Universe.



Is there a known example of this term in particle physics? Yes, a pion. Credit: HiggsTan

- A pion is a composite meson composed of a quark and an antiquark.
 - A neutral pion, π^0 , is composed of either u or dd, and is a pseudoscalar. (Chinowsky & Steinberger, 1954)
 - π^0 is coupled to photons via I_{CS} where
 - $\theta = \pi^0 / f_{\pi}$ with $f_{\pi} \sim 184$ MeV (pion decay constant)
 - $\alpha = 2\alpha_{EM}N_c/(3\pi)$ with $N_c = 3$ (the number of quark colors) and $\alpha_{EM} \sim 1/137$ (EM fine structure constant)
- π^{0} decays into 2 photons via this term, which has been observed. So, what we are going to study in this lecture is not completely crazy!



Correction to Maxwell's equations In Heaviside units and c=1

- We now derive the correction to Max $I = -\frac{1}{4}\int d^4x \ \left(F^2 + a^4\right) d^4x \ \left(F^2 + a^4$
- Finding the path that gives a stational $\delta I = \int d^4 x \, \left(F^{\mu\nu} + \alpha\theta\tilde{F}\right)$ $= \int d^4 x \, \left[-\partial_\nu (F^{\mu\nu} + \varphi)\right] d^4 x \, \left[-\partial_\mu (F^{\mu\nu} + \varphi)\right] d^4 x \, \left[-$

axwell's equations from

$$\begin{aligned}
d^{4}x = dt \\
d^{4}x A_{\mu}j^{\mu} \\
& \text{Hint: } \delta(F\tilde{F}) = \epsilon^{\mu\nu\alpha\beta}(\delta F_{\mu}) \\
& \text{Hint: } \delta(F\tilde{F}) = \epsilon^{\mu\nu\alpha\beta}(\delta F_{\mu}) \\
& = -4(\partial_{\nu}\delta A_{\mu}) \\
& = -4(\partial_{\mu}\delta A_{\mu}) \\
& = -4(\partial_{$$





Correction to Maxwell's equations In Heaviside units and c=1







2.4 Parity Violation in EM Waves

Warm-up: The wave equation for A^µ Maxwell's equations in vacuum

Maxwel

I's equations in vacuum
$$\partial_{\nu} F^{\mu\nu} = 0$$
 gives

$$-\Box A^{\mu} + \eta^{\mu\alpha} \partial_{\alpha} (\partial_{\nu} A^{\nu}) = 0$$

$$\Box = \eta^{\alpha\beta} \partial_{\alpha} \partial_{\beta} = -\frac{\partial^2}{\partial t^2} + \nabla^2 \qquad \eta^{\alpha\beta} = \text{diag}(-1, 1)$$

$$A^{\mu} = \eta^{\mu\alpha} A_{\alpha} = (\phi, \mathbf{A})$$
"Lorenz gauge condition"

where

equations in vacuum
$$\partial_{\nu} F^{\mu\nu} = 0$$
 gives

$$\Box A^{\mu} + \eta^{\mu\alpha} \partial_{\alpha} (\partial_{\nu} A^{\nu}) = 0$$

$$\Box = \eta^{\alpha\beta} \partial_{\alpha} \partial_{\beta} = -\frac{\partial^2}{\partial t^2} + \nabla^2 \qquad \eta^{\alpha\beta} = \text{diag}(-1)$$

$$u^{\mu} = \eta^{\mu\alpha} A_{\alpha} = (\phi, \mathbf{A})$$

s equations in vacuum
$$\partial_{\nu} F^{\mu\nu} = 0$$
 gives
 $-\Box A^{\mu} + \eta^{\mu\alpha} \partial_{\alpha} (\partial_{\nu} A^{\nu}) = 0$
 $\Box = \eta^{\alpha\beta} \partial_{\alpha} \partial_{\beta} = -\frac{\partial^2}{\partial t^2} + \nabla^2 \qquad \eta^{\alpha\beta} = \text{diag}(-1)$
 $A^{\mu} = \eta^{\mu\alpha} A_{\alpha} = (\phi, \mathbf{A})$

by choosing $\Box f = -\partial_{\nu}A^{\nu}$ in A_{μ}

• Now, using invariance under the gauge transformation, we can set $\partial_{\nu}A^{\nu}=0$

$$ightarrow A_{\mu} + \partial_{\mu} f$$
 . Then . . .





Warm-up: The wave equation for A^µ The use case of the gauge invariance

Maxwell

's equations in vacuum
$$\partial_{\nu} F^{\mu\nu} = 0$$
 and $\partial_{\nu} A^{\nu} = 0$ gives

$$\Box A^{\mu} = 0 \quad \longrightarrow \quad \text{The equation for a wave} \\ \text{traveling at the speed of light!} \\ \Box = \eta^{\alpha\beta} \partial_{\alpha} \partial_{\beta} = -\frac{\partial^2}{\partial t^2} + \nabla^2 \\ A^{\mu} = \eta^{\mu\alpha} A_{\alpha} = (\phi, \mathbf{A}) \text{ with } \dot{\phi} + \nabla \cdot \mathbf{A} = 0$$

where

equations in vacuum
$$\partial_{\nu} F^{\mu\nu} = 0$$
 and $\partial_{\nu} A^{\nu} = 0$ gives
 $A^{\mu} = 0 \quad \longrightarrow \quad \text{The equation for a wave}$
traveling at the speed of light!
 $\Box = \eta^{\alpha\beta}\partial_{\alpha}\partial_{\beta} = -\frac{\partial^2}{\partial t^2} + \nabla^2$
 $A^{\mu} = \eta^{\mu\alpha}A_{\alpha} = (\phi, \mathbf{A}) \text{ with } \dot{\phi} + \nabla \cdot \mathbf{A} = 0$

s equations in vacuum
$$\partial_{\nu} F^{\mu\nu} = 0$$
 and $\partial_{\nu} A^{\nu} = 0$ gives

$$\Box A^{\mu} = 0 \quad \clubsuit \quad \text{The equation for a wave} \\ \text{traveling at the speed of light!} \\ \Box = \eta^{\alpha\beta} \partial_{\alpha} \partial_{\beta} = -\frac{\partial^2}{\partial t^2} + \nabla^2 \\ A^{\mu} = \eta^{\mu\alpha} A_{\alpha} = (\phi, \mathbf{A}) \text{ with } \dot{\phi} + \nabla \cdot \mathbf{A} = 0$$

• The number of degrees of freedom for A^{μ} is 3 due to the Lorenz gauge condition.



Physical degrees of freedom of EM waves 3? 2?

- We know that photons must have only 2 helicity states, λ=±1 (two circular polarization states).
 - Shouldn't the number of physical degrees of freedom be 2, instead of 3? The answer is yes.
- The Lorenz gauge does not fully specify A^µ. We can still add

$$A_{\mu}
ightarrow A_{\mu} + \partial_{\mu} f_2$$
 which satisfies $\Box f_2 = 0$

Choosing f₂ will fully specify A^µ. This leaves 2 degrees of freedom.

EM waves must be transverse

• It is common to choose f and f₂ such that



- We have 2 conditions. That leaves 2 degrees of freedom for EM waves.
- This choice is consistent with the Lorenz gauge condition $\phi +
 abla \cdot \mathbf{A} = 0$
- $\nabla \cdot \mathbf{A} = 0$ requires that the EM wave be *transverse*, i.e., the change in **A** is perpendicular to the direction of propagation of the EM wave.
- We will use this condition throughout the lecture.

The arrows show directions of the electric field vector **E**.





Correction to the EM wave equation With the Chern-Simons term

 $\partial_{\nu}F^{\mu\nu} + \alpha(\partial_{\nu}\theta)\tilde{F}^{\mu\nu} = 0$

• With $A^0 = \phi = 0$ in the Lorenz gauge, we find

 $-\Box A^i + \alpha(\partial_\nu \theta) \tilde{F}^{i\nu} = 0$

 $\ddot{\mathbf{A}} - \nabla^2 \mathbf{A} + \alpha \left| -\dot{\theta} \right|$



$$\left[\nabla \times \mathbf{A} \right] + \left(\nabla \theta \right) \times \dot{\mathbf{A}} = 0$$

Correction to the EM wave equation! <u>Note</u>: **A** is a vector and θ is a pseudoscalar.

Helicity Basis Going to Fourier space

- Fourier transform of A(t,x) is A(t,x)
 - The EM wave propagates in the direction of **k**. The change in A_k is perpendicular to **k**.

"Coulomb gauge"
$$abla \cdot \mathbf{A}(t,\mathbf{x}) = 0 o \mathbf{k} \cdot \mathbf{A}_{\mathbf{k}}(t) = 0$$

• Choose **k** to be on the $z(=x^3)$ axis. The helicity states, $\lambda = \pm 1$, are given for each Fourier mode by

$$A_{\pm} = \frac{A_{\mathbf{k}}^1 \mp i A_{\mathbf{k}}^2}{\sqrt{2}}$$

$$\mathbf{x} = (2\pi)^{-3/2} \int d^3 \mathbf{k} \, \mathbf{A}_{\mathbf{k}}(t) e^{i\mathbf{k}\cdot\mathbf{x}}$$

X³

$$(A^1, A^2, 0)$$

Helicity Basis Transformation property under rotation • To show that A_{\pm} represents the helicity states, rotate the spatial coordinates around the z axis in the right-handed system by an angle arphi . • The helicity states, $\lambda = \pm 1$, transform as 1', **A**²', **0**) **Right-handed** anded 35

$$A_{\lambda} \rightarrow A'_{\lambda} = e^{i\lambda\varphi}A_{\lambda}$$

Helicity
A_+: Right-handed state
A_: Left-handed state







Correction to the EM wave equation In the helicity basis

$$\ddot{\mathbf{A}} - \nabla^2 \mathbf{A} + \alpha \left[-\dot{\theta} (\nabla \times \mathbf{A}) + (\nabla \theta) \times \dot{\mathbf{A}} \right] = 0$$
Correction to the EM wave equation!

time, $\theta(t, \mathbf{x}) \rightarrow \theta(t)$. Then in Fourier space

$$\ddot{\mathbf{A}}_{\mathbf{k}} + k^{2}\mathbf{A}_{\mathbf{k}} - i\alpha\dot{\bar{\theta}}(\mathbf{k} \times \dot{\bar{\theta}})$$
$$\dot{\bar{A}}_{\pm} + \left(k^{2}\mp k\alpha\dot{\bar{\theta}}\right)$$

<u>Note</u>: **A** is a vector and θ is a pseudoscalar.

• To simplify the analysis, assume that θ is homogeneous and depends only on

$$\mathbf{A}) = \mathbf{0}$$

 $A_{+} = 0$

The equation of motion depends on handedness!



Recap: Day 2

- There are 2 scalars composed of $F_{\mu\nu}$ and $F_{\mu\nu}$.
 - F^2 is a scalar, whereas FF is a pseudoscalar.
 - The action $\int dt d^3 \mathbf{x} F^2$ gives Maxwell's equations, whereas $\int dt d^3 \mathbf{x} FF$ is a surface term.
- A Chern-Simons interaction between a pseudoscalar field θ and photons, $\int dt d^3 \mathbf{x} \, \theta FF$, modifies Maxwell's equations.
 - The equation of motion for Line waves are This is the signature of violation of parity symmetry! $\ddot{A}_{\pm} + \left(k^2 \mp k \alpha \dot{\bar{\theta}}\right) A_{\pm} = 0$
 - The equation of motion for EM waves depends explicitly on helicity states. What is the cosmological implication of this term? Let's find out!

